

# 1 Insights on Atmospheric Oxidation Processes by Performing Factor Analyses on

## 2 Sub-ranges of Mass Spectra

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13

14 Abstract

With the recent developments in mass spectrometry, combined with the strengths of factor analysis techniques, our understanding of atmospheric oxidation chemistry has improved significantly. The typical approach for using techniques like positive matrix factorization (PMF) is to input all measured data for the factorization in order to separate contributions from different sources and/or processes to the total measured signal. However, while this is a valid approach for assigning the total signal to factors, we have identified several cases where useful information can be lost if solely using this approach. For example, gaseous molecules emitted from the same source can show different temporal behaviors due differing loss terms, like condensation at different rates due to different molecular masses. This conflicts with one of PMF's basic assumptions of constant factor profiles. In addition, some ranges of a mass spectrum may contain useful information, despite contributing only minimal fraction to the total signal, in which case they are unlikely to have a significant impact on the factorization result. Finally, certain mass ranges may contain molecules formed via pathways not available to molecules in other mass ranges, e.g. dimeric species versus monomeric species. In this study, we attempted to address these challenges by dividing mass spectra into sub-ranges and applying the newly developed binPMF method to these ranges separately. We utilized a dataset from a chemical ionization atmospheric pressure interface time-of-flight (CI-APi-TOF) mass spectrometer as an example. We compare the results from these three different ranges, each corresponding to molecules of different volatilities, with binPMF results from the combined range. Separate analysis showed clear benefits in dividing factors for molecules of different volatilities more accurately, in

34 resolving different chemical processes from different ranges, and in giving a chance for high-  
35 molecular-weight molecules with low signal intensities to be used to distinguish dimeric species with  
36 different formation pathways. As two major insights from our study, we identified daytime dimer  
37 formation (diurnal peak around noon) which may contribute to new particle formation in Hyytiälä, as  
38 well as dimers from NO<sub>3</sub> oxidation process. We recommend PMF users to try running their analyses  
39 on selected sub-ranges in order to further explore their datasets.

40

## 41 **1 Introduction**

42 Huge amounts of volatile organic compounds (VOC) are emitted to the atmosphere every year  
43 (Guenther et al., 1995; Lamarque et al., 2010), which play a significant role in atmospheric chemistry  
44 and affect the oxidative ability of the atmosphere. The oxidation products of VOC can contribute to  
45 the formation and growth of secondary organic aerosols (Kulmala et al., 2013; Ehn et al., 2014; Kirkby  
46 et al., 2016; Troestl et al., 2016), affecting air quality, human health, and climate radiative forcing  
47 (Pope III et al., 2009; Stocker et al., 2013; Zhang et al., 2016; Shiraiwa et al., 2017). Thanks to the  
48 advancement in mass spectrometric applications, like the aerosol mass spectrometer (AMS)  
49 (Canagaratna et al., 2007) and chemical ionization mass spectrometry (CIMS) (Bertram et al.,  
50 2011; Jokinen et al., 2012; Lee et al., 2014) our capability to detect these oxidized products, as well as  
51 our understanding of the complicated atmospheric oxidation pathways in which they take part, have  
52 been greatly enhanced.

53 Monoterpenes (C<sub>10</sub>H<sub>16</sub>), one common group of VOC emitted in forested areas, have been shown to  
54 be a large source of atmospheric secondary organic aerosol (SOA). The oxidation of monoterpenes  
55 produces a wealth of different oxidation products (Oxygenated VOC, OVOC), including highly  
56 oxygenated organic molecules (HOM) with molar yields in the range of a few percent, depending on  
57 the specific monoterpene and oxidant (Ehn et al., 2014; Bianchi et al., 2019). Bianchi et al. (2019)  
58 summarized that HOM can be either Extremely Low Volatility Organic Compounds (ELVOC), Low  
59 Volatility Organic Compounds (LVOC), or Semi-volatile Organic Compounds (SVOC)  
60 (classifications by Donahue et al. 2012), depending on their exact structures. For less oxygenated  
61 products, the majority are likely to fall into the SVOC or the Intermediate VOC (IVOC) range. The  
62 volatility of the OVOC will determine their dynamics, including their ability to contribute to the  
63 formation of SOA and new particles (Bianchi et al., 2019; Buchholz et al., 2019).

64 The recent developments of CIMS techniques has allowed researchers to observe unprecedented  
65 numbers of OVOC, in real-time (Riva et al., 2019). This ability to measure thousands of compounds  
66 is a great benefit, but also a large challenge for the data analyst. For this reason, factor analytical  
67 techniques have often been applied to reduce the complexity of the data by finding co-varying signals

68 that can be grouped into common factors (Huang et al., 1999). For aerosol and gas-phase mass  
69 spectrometry, positive matrix factorization, PMF (Paatero and Tapper, 1994; Zhang et al., 2011) has  
70 been the most utilized tool. The factors have then been attributed to sources (e.g. biomass burning  
71 organic aerosol) or processes (e.g. monoterpane ozonolysis) depending on the application and ability  
72 to identify spectral signatures (Yan et al., 2016; Zhang et al., 2017). In the vast majority of these PMF  
73 applications to mass spectra, the mass range of ions has been maximized in order to provide as much  
74 input as possible for the algorithm. This approach was certainly motivated in early application of  
75 PMF on e.g. offline filters, with chemical information of metals, water-soluble ions, and organic and  
76 elemental carbon (OC and EC), where the number of variables is counted in tens, and the number of  
77 samples in tens or hundreds (Zhang et al., 2017). However, with gas-phase CIMS, we often have up  
78 to a thousand variables, with hundreds or even thousands of samples, meaning that the amount of data  
79 itself is unlikely to be a limitation for PMF calculation. In this work, we aimed to explore potential  
80 benefits of dividing the spectra into sub-ranges before applying factorization analysis.

81 An inherent requirement of factorization approaches is that the factor profiles, in this case the relative  
82 abundances of ions in the mass spectra, of each factor stay nearly constant. Due to the complexity  
83 and number of atmospheric processes affecting the formation, transformation, and loss of VOC,  
84 OVOC and aerosol, this often does not hold, and is one of the main limitations of factorization  
85 approaches. Given the different volatilities of OVOC, it may even be expected that molecules from  
86 the same source may have very different loss time scales, which may affect the factor analysis. For  
87 compounds of low volatility, such as many HOM, the main atmospheric loss process is typically  
88 condensation onto aerosol particles, with chemical sink being negligible (Bianchi et al., 2019). If, on  
89 the other hand, a compound does not irreversibly condense, oxidation reactions can also affect its  
90 lifetime. Volatility issue has been studied and reported for AMS data, with different volatilities of  
91 various OA types (Huffman et al., 2009; Crippa et al., 2014; Paciga et al., 2016; Äijälä et al., 2017).  
92 Semi-volatile oxygenated organic aerosol (SV-OOA) and Low-volatility oxygenated organic aerosol  
93 (LV-OOA) can both be mainly produced from biogenic sources, but get separated based on different  
94 volatilities by PMF (El Haddad et al., 2013). Sekimoto et al. (2018) found that the two profiles  
95 resolved with VOC emitted from biomass burning had different estimated volatilities. As the  
96 volatility of a molecule is linked to its molecular mass (Peräkylä et al., 2020), it may be beneficial to  
97 apply PMF separately to mass ranges where one can expect the loss processes to be similar, thereby  
98 resulting in more constant factor profiles. In this way, distinct sources are hopefully separated by  
99 PMF, with minimized influence of differing volatilities from one source.

100 The number of PMF or other factorization studies utilizing CIMS data remains very limited.  
101 “Traditional” PMF analyses have so far, to our knowledge, only been applied to nitrate-based

102 chemical ionization atmospheric pressure interface time-of-flight (CI-APi-TOF) data (Yan et al.,  
103 2016;Massoli et al., 2018). One study has also utilized non-negative matrix factorization (NNMF) to  
104 look at diurnal trends of Iodide ToF-CIMS data (Lee et al., 2018). The lack of more studies utilizing  
105 PMF, or other factorization techniques, on CIMS data is likely partly due to the complexity of the  
106 data, with multiple overlapping ions hampering HR peak fitting (Zhang et al., 2019). In addition,  
107 variable factor profiles may hamper PMF's ability to correctly separate the factors. The two CI-APi-  
108 TOF studies utilized the nearly the entire measured spectrum (from around 200 Th to around 600 Th),  
109 either in unit mass resolution (UMR) or high resolution (HR) peak fitting data (Yan et al.,  
110 2016;Massoli et al., 2018). Massoli et al. (2018) estimated the volatility of the molecules they detected,  
111 finding that all the six extracted factors had notable contributions from IVOC, SVOC and (E)LVOC.  
112 These compound groups will have clearly different loss mechanisms, and thereby loss rates,  
113 suggesting that variation in factor profiles is inevitable, even if the source was identical for all  
114 molecules in the factor. We hypothesize that this effect further hampers the correct factorization, and  
115 further that this effect can be reduced by dividing the spectra into separate ranges, with each sub-  
116 range containing molecules with roughly similar loss mechanisms and rates.

117 As an additional motivation to separate different ranges from the mass spectrum, it is not only the  
118 loss mechanisms, but also the formation pathways that may differ. For example, atmospheric  
119 oxidation chemistry of organics is, to a large extent, the chemistry of peroxy radicals ( $\text{RO}_2$ ) (Orlando  
120 and Tyndall, 2012).These  $\text{RO}_2$  are initiated by VOC reacting with oxidants like ozone, or the hydroxyl  
121 ( $\text{OH}$ ) or nitrate ( $\text{NO}_3$ ) radicals, while their termination occurs mainly by bimolecular reactions with  
122  $\text{NO}$ ,  $\text{HO}_2$  and/or other  $\text{RO}_2$ . Some product molecules can be formed from any of the three termination  
123 pathways, while for example ROOR “dimers” can only be formed from  $\text{RO}_2+\text{RO}_2$  reactions (Berndt  
124 et al., 2018a;Berndt et al., 2018b). This also means that there can be several different pathways to  
125 form dimers from the same precursors VOC, by combining  $\text{RO}_2$  formed from the same or different  
126 oxidants. As an example of the latter, an ROOR dimer can contain one moiety produced from ozone  
127 oxidation and another moiety from  $\text{NO}_3$  oxidation (Yan et al., 2016). Thus, their concentration is  
128 dependent on both the precursor VOC concentration, and the concentrations of both oxidants. Such a  
129 molecule will not have a direct equivalent in any of the monomer products: even though monomers  
130 can form from  $\text{RO}_2 + \text{R}'\text{O}_2$  reactions, the products from  $\text{RO}_2$  are not dependent on the source of the  
131  $\text{R}'\text{O}_2$ . This may complicate the identification of certain dimer factors by PMF if the entire spectrum  
132 is analyzed at once, and therefore separation of the monomer and dimer products before the PMF  
133 analysis could be advantageous.

134 Recently, we proposed a new PMF approach, binPMF, to simplify the analysis of mass spectral data  
135 (Zhang et al., 2019). This method divides the mass spectrum into narrow bins, typically some tens of

136 bins per integer mass, depending on the mass resolving power of the instrument, before performing  
137 PMF analyses. In this way, binPMF does not require any time-consuming, and potentially subjective  
138 high resolution peak fitting, and can thus be utilized for data exploration at a very early stage of data  
139 analysis. Data preparation is nearly as simple as in the case of UMR analysis, yet it utilizes much  
140 more spectral information as it does not sum up signal over all ions at each integer mass. In addition  
141 to saving time and effort in data analysis, the results are less sensitive to mass calibration fluctuations.  
142 Finally, the binning also greatly increases the number of input variables, which has the advantage that  
143 factor analysis with smaller mass ranges becomes more feasible, as more meaningful variation is  
144 supplied to the algorithm.

145 We designed this study to explore the benefits of separate analysis of different mass ranges from mass  
146 spectra. We used a previously published ambient dataset measured by a CI-APi-TOF, and conducted  
147 binPMF analysis with three different mass ranges, roughly corresponding to different volatility ranges.  
148 We compared the results from the sub-range analyses with each other and with results from binPMF  
149 run on the combined ranges. We found that dimers generated during daytime and dimers initiated by  
150  $\text{NO}_3$  oxidation can be separated from our dataset by utilizing the sub-ranges, but not with the full  
151 range. We believe that this study will provide new perspectives for future studies analyzing gas-phase  
152 CIMS data.

153

## 154 **2 Methodology**

155 The focus of this work is on retrieving new information from mass spectra by applying new analytical  
156 approaches. Therefore, we chose a dataset that has been presented earlier, though without PMF  
157 analysis, by Zha et al. (2018), and was also used in the first study describing the binPMF method  
158 (Zhang et al., 2019). The measurements are described in more details below in section 2.1, while the  
159 data analysis techniques used in this work are presented in section 2.2.

160

### 161 **2.1 Measurements**

#### 162 **2.1.1 Ambient site**

163 The ambient measurements were conducted at the Station for Measuring Ecosystem–Atmosphere  
164 Relations (SMEAR) II in Finland (Hari and Kulmala, 2005) as part of the Influence of Biosphere–  
165 Atmosphere Interactions on the Reactive Nitrogen budget (IBAIRN) campaign (Zha et al., 2018).  
166 Located in the boreal environment in Hyytiälä, SMEAR II is surrounded with coniferous forest and  
167 has limited anthropogenic emission sources nearby. Diverse measurements of meteorology, aerosol  
168 and gas phase properties are continuously conducted at the station. Details about the meteorological

169 conditions and temporal variations of trace gases during IBAIRN campaign are presented by Zha et  
170 al. (2018) and Liebmann et al. (2018).

171

### 172 **2.1.2 Instrument and data**

173 Data were collected with a nitrate ( $\text{NO}_3^-$ )-based chemical ionization atmospheric pressure interface  
174 time-of-flight mass spectrometer (CI-APi-TOF, Jokinen et al., 2012) with about  $4000 \text{ Th Th}^{-1}$  mass  
175 resolving power, at ground level in September, 2016. In our study, the mass spectra were averaged to  
176 1 h time resolution from September 6<sup>th</sup> to 22<sup>nd</sup> for further analysis. We use the thomson (Th) as the  
177 unit for mass/charge, with  $1 \text{ Th} = 1 \text{ Da/e}$ , where  $e$  is the elementary charge. As all the data discussed  
178 in this work are based on negative ion mass spectrometry, we will use the absolute value of the  
179 mass/charge, although the charge of each ion will be negative. The masses discussed in this work  
180 includes the contribution from the nitrate ion, 62, unless specifically mentioned. Furthermore, as the  
181 technique is based on soft ionization with  $\text{NO}_3^-$  ions, any multiple charging effects are unlikely, and  
182 therefore the reported mass/charge values in thomson can be considered equivalent to the mass of the  
183 ion in Da.

184 The forest site of Hyytiälä is dominated by monoterpene emissions (Hakola et al., 2006). The main  
185 feature of previous CI-APi-TOF measurements in Hyytiälä (Ehn et al., 2014; Yan et al., 2016) has  
186 been a bimodal distributions of HOM, termed monomers and dimers, as they are formed of either one  
187 or two  $\text{RO}_2$  radicals, respectively. For the analysis in this study, we chose three mass/charge ( $m/z$ )  
188 ranges of 50 Th each (Figure 1), corresponding to regions between which we expect differences in  
189 formation or loss mechanisms. In addition to regions with HOM monomers and HOM dimers, one  
190 range was chosen at lower masses, in a region presumably mainly consisting of molecules that are  
191 less likely to condense onto aerosol particles (Peräkylä et al., 2020).

192

### 193 **2.2 Positive matrix factorization (PMF)**

194 After the model of PMF was developed (Paatero and Tapper, 1994), numerous applications have been  
195 conducted with different types of environmental data (Song et al., 2007; Ulbrich et al., 2009; Yan et  
196 al., 2016; Zhang et al., 2017). By reducing dimensionality of the measured dataset, PMF model greatly  
197 simplifies the data analysis process with no requirement for prior knowledge of sources or pathways  
198 as essential input. The main factors can be further interpreted with their unique/dominant markers  
199 (elements or masses).

200 The basic assumption for PMF modelling is mass balance, which assumes that ambient concentration  
201 of a chemical component is the sum of contributions from several sources or processes, as shown in  
202 equation (1).

$$\mathbf{X} = \mathbf{TS} \times \mathbf{MS} + \mathbf{R} \quad (1)$$

In equation (1), **X** stands for the time series of measured concentration of different variables (*m/z* in our case), **TS** represents the temporal variation of factor contributions, **MS** stands for factor profiles (mass spectral profiles), and **R** is the residual as the difference of the modelled and the observed data. The matrices **TS** and **MS** are iteratively calculated by a least-squares algorithm utilizing uncertainty estimates, to pursue minimized *Q* value as shown in equation (2), where  $S_{ij}$  is the estimated uncertainty, an essential input in PMF model.

$$Q = \sum \sum \left( \frac{R_{ij}}{S_{ii}} \right)^2 \quad (2)$$

PMF model was conducted by multi-linear engine (ME-2) (Paatero, 1999) interfaced with Source Finder (SoFi, v6.3) (Canonaco et al., 2013). Signal-to-noise ratio (SNR) was calculated as  $\text{SNR}_{ij} = \text{abs}(X_{ij}) / \text{abs}(S_{ij})$ . When the Signal-to-noise ratio (SNR) is below 1, the signal of  $X_{ij}$  will be down-weighted by replacing the corresponding uncertainty  $S_{ij}$  by  $S_{ij}/\text{SNR}_{ij}$  (Visser et al., 2015). Future studies should pay attention to the potential risk when utilizing this method since down-weighting low signals element-wise will create a positive bias to the data. Robust mode was operated in the PMF modelling, where outliers ( $\left| \frac{R_{ij}}{S_{ij}} \right| > 4$ ) were significantly down-weighted (Paatero, 1997).

219 2.3 binPMF

As a newly developed application of PMF for mass spectral data, binPMF has no requirement for chemical composition information, while still taking advantage of the HR mass spectra, saving effort and time (Zhang et al., 2019). To explore the benefits of analyzing separated mass ranges, we applied binPMF to the three separated ranges. The three ranges were also later combined for binPMF analysis as comparison with the previous results. The PMF model requires both data matrix and error matrix as input, and details of the preparation of data and error matrices are described below.

227 2.3.1 Data matrix

Different from normal UMR or HR peak fitting, in binPMF, the mass spectra are divided into small bins after baseline subtraction and mass axis calibration. Linear interpolation was first conducted to the mass spectra with a mass interval of 0.001 Th. Then the interpolated data was averaged into bins of 0.02 Th width. We selected three ranges for further analysis based on earlier studies (Ehn et al., 2014; Yan et al., 2016; Bianchi et al., 2019; Peräkylä et al., 2020).

- 233 - Range 1,  $m/z$  250 – 300 Th,  $51 \text{ unit masses} \times 25 \text{ bins per unit mass} = 1275 \text{ bins/variables}$ ,  
 234 consisting mainly of molecules with five to nine carbon atoms and four to nine oxygen atoms  
 235 in our dataset.
- 236 - Range 2,  $m/z$  300 – 350 Th,  $51 \times 25 = 1275$  bins, mainly corresponding to HOM monomer  
 237 products, featured with nine to ten C- and seven to ten O-atoms.
- 238 - Range 3,  $m/z$  510 – 560 Th,  $51 \times 30 = 1530$  bins, mainly corresponding to HOM dimer products,  
 239 with carbon numbers of sixteen to twenty and eleven to fifteen O-atoms.
- 240 To avoid unnecessary computation, only signal regions with meaningful signals in the mass spectra  
 241 were binned (Zhang et al., 2019). For a nominal mass  $N$ , the signal region included in further analyses  
 242 was between  $N-0.2$  Th and  $N+0.3$  Th for Range 1 and 2, and between  $N-0.2$  Th and  $N+0.4$  Th for  
 243 Range 3. The wider signal regions in Range 3 is due to wider peaks at higher masses. The data were  
 244 averaged into 1-h time resolution and in total we had 384 time points in the data matrix.

245

### 246 2.3.2 Error matrix

247 The error matrix represents the estimated uncertainty for each element of the data matrix and is crucial  
 248 for iterative calculation of the  $Q$  minimum. Equation (3) is used for error estimation (Polissar et al.,  
 249 1998),

$$250 \quad S_{ij} = \sigma_{ij} + \sigma_{\text{noise}} \quad (3)$$

251 where  $S_{ij}$  represents the uncertainty of  $m/z j$  at time  $i$ ,  $\sigma_{ij}$  stands for counting statistics uncertainty  
 252 and is estimated as follows,

$$253 \quad \sigma_{ij} = a \times \frac{\sqrt{I_{ij}}}{\sqrt{t_s}} \quad (4)$$

254 where  $I$  is the signal intensity term, in unit of counts per second (cps),  $t_s$  stands for length of averaging  
 255 in seconds, while  $a$  is an empirical coefficient to compensate for unaccounted uncertainties (Allan et  
 256 al., 2003; Yan et al., 2016) and is 1.28 in our study as previously estimated from laboratory  
 257 experiments (Yan et al., 2016). The  $\sigma_{\text{noise}}$  term was estimated as the median of the standard  
 258 deviations from signals in the bins in the region between nominal masses, where no physically  
 259 meaningful signals are expected.

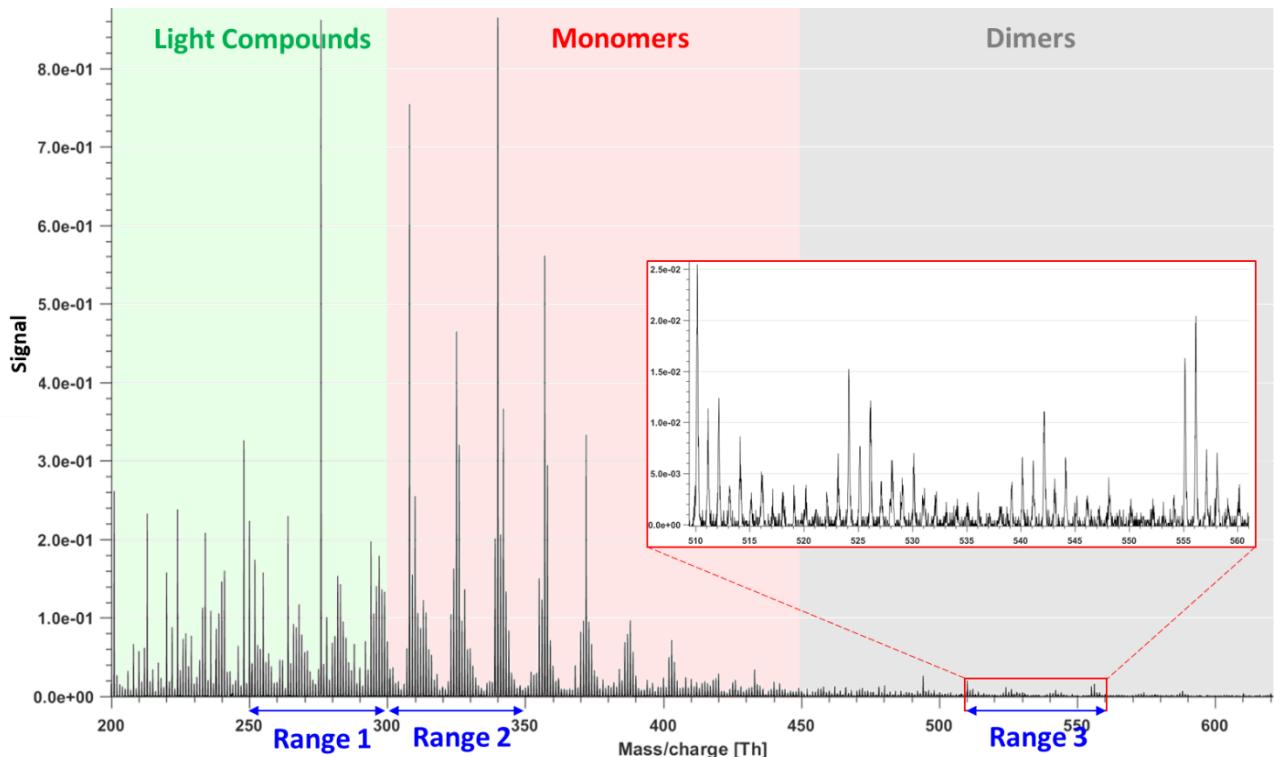
260

## 261 3 Results

### 262 3.1 General overview of the dataset/spectrum

263 During the campaign, in autumn, 2016, the weather was overall sunny and humid with average  
 264 temperature of 10.8 °C and relative humidity (RH) of 87% (Zha et al., 2019). The average

265 concentration of NO<sub>x</sub> and O<sub>3</sub> was 0.4 ppbv and 21 ppbv, respectively. The average total HOM  
266 concentration was ~ 10<sup>8</sup> molecules cm<sup>-3</sup>.



267  
268 Figure 1. Example of mass spectrum with 1-h time resolution measured from a boreal forest environment  
269 during the IBAIRN campaign (at 18:00, Finnish local time, UTC+2). The mass spectrum was divided into  
270 three parts and three sub-ranges were chosen from different parts for further analysis in our study. The  
271 nitrate ion (62 Th) is included in the mass.

272  
273 Figure 1 shows the 1 h averaged mass spectrum taken at 18:00 on September 12, as an example of  
274 the analyzed dataset. In addition to exploring the benefits of this type of sub-range analysis in relation  
275 to different formation or loss pathways, separating into sub-ranges may also aid factor identification  
276 for low-signal regions. As shown in Figure 1, there is a difference of 1-2 orders of magnitude in the  
277 signal intensity between Range 3 and Ranges 1-2. If all Ranges are run together, we would expect  
278 that the higher signals from Ranges 1 and 2 will drive the factorization. While if run separately,  
279 separating formation pathways of dimers in Range 3 will likely be easier. As dimers have been shown  
280 to be crucial for the formation of new aerosol particles from monoterpene oxidation (Kirkby et al.,  
281 2016; Troestl et al., 2016; Lehtipalo et al., 2018), this information may even be the most critical in  
282 some cases, despite the low contribution of these peaks to the total measured signal.  
283 binPMF was separately applied to Range 1, 2, 3, and a ‘Range combined’ which comprised all the  
284 three sub-ranges. All the PMF runs for the four ranges were conducted from two to ten factors and  
285 repeated three times for each factor number, to assure the consistency of the results. Factorization

286 results and evolution with increasing factor number are briefly described in the following sections,  
287 separately for each Range (sections 3.2 – 3.5). It is worth noting that the factor order in factor  
288 evolution does not necessarily correspond to that of the final results. The factor orders displayed in  
289 Figures 2-5 have been modified for further comparison between different ranges. More detailed  
290 discussion and comparison between the results are presented in Section 4.

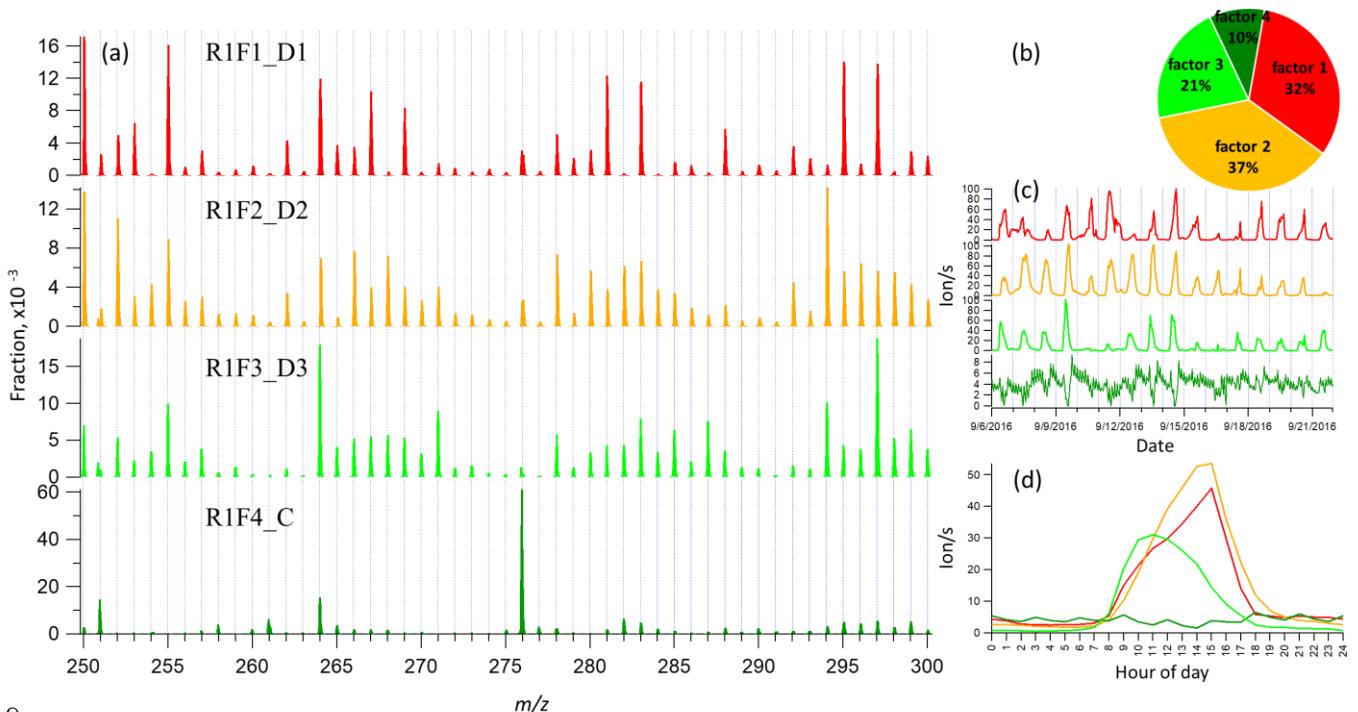
291

### 292 **3.2 binPMF on Range 1 (250 – 300 Th)**

293 As has become routine (Zhang et al., 2011;Craven et al., 2012), we first examined the mathematical  
294 parameters of our solutions. From two to ten factors,  $Q/Q_{exp}$  decreased from 2.8 to 0.7 (Fig S1 in  
295 supplementary information), and after three factors, the decreasing trend was gradually slowing down  
296 and approaching one, which is the ideal value for  $Q/Q_{exp}$  as a diagnostic parameter. The unexplained  
297 variation showed a decline from 18% to 8% from two to ten factors.

298 In the two-factor results, two daytime factors were separated, with peak time both at 14:00 - 15:00.  
299 One factor was characterized by large signals at 250 Th, 255 Th, 264 Th, 281 Th, 283 Th, 295 Th,  
300 297 Th. The other factor was characterized by large signals at 294 Th, 250 Th, 252 Th, 264 Th, 266  
301 Th, 268 Th, and 297 Th. In Hyytiälä, as reported in previous studies, odd masses observed by the  
302 nitrate CI-APi-TOF are generally linked to monoterpene-derived organonitrates during the day (Ehn  
303 et al., 2014;Yan et al., 2016). When the factor number increased to three, the two earlier daytime  
304 factors remained similar with the previous result, while a new factor appeared with a distinct sawtooth  
305 shape in the diurnal cycle. The main marker in the spectral profile was 276 Th, with a clear negative  
306 mass defect. When one more factor was added, the previous three factors remained similar as in the  
307 three-factor solution, and a new morning factor was resolved, with 264 Th and 297 Th dominant in  
308 the mass spectral profile, and a diurnal peak at 11:00.

309 As the factor number was increased, more daytime factors were separated, with similar spectral  
310 profiles to existing daytime factors and various peak times. No nighttime factors were found in the  
311 analysis even when the factor number reached ten. We chose the four-factor result for further  
312 discussion, and Figure 2 shows the result of Range 1, with spectral profile, time series, diurnal cycle  
313 and averaged factor contribution during the campaign. As shown in Figure 2d, Factors 1-3 are all  
314 daytime factors, while Factor 4 has no clear diurnal cycle, but a distinct sawtooth shape. Factor 4  
315 comes from a contamination of perfluorinated acids, from the inlet's automated zeroing every three  
316 hours during the measurements (Zhang et al., 2019). The zeroing periods have been removed from  
317 the dataset before binPMF analysis, but the contamination factor was still resolved. This factor is  
318 discussed in more detail in sections 4.1 and 4.4.



319

320 Figure 2 Four-factor result for Range 1, for (a) factor spectral profiles, (b) averaged factor contribution during  
 321 the campaign, (c) time series and (d) diurnal trend. Details on the factors' naming schemes are shown in Table  
 322 1.

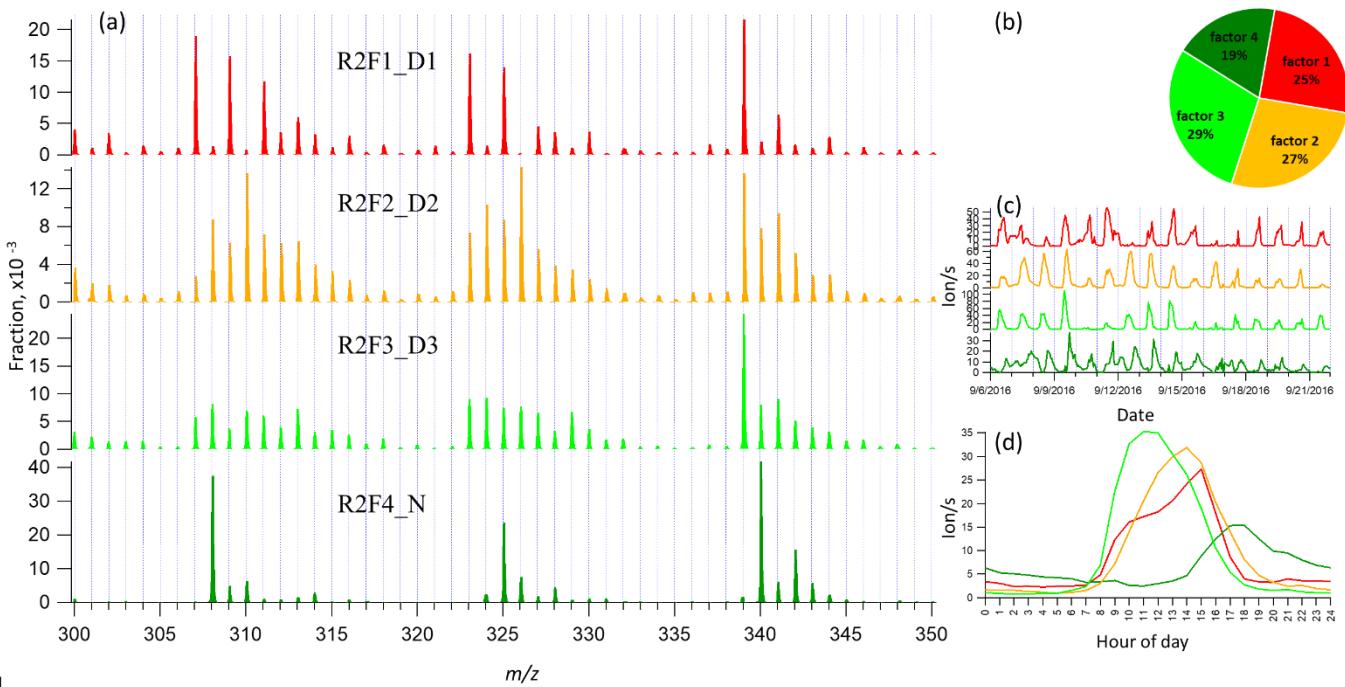
323

### 324 **3.3 binPMF on Range 2 (300-350 Th)**

325 This range covers the monoterpene HOM monomer range, and binPMF results have already been  
 326 discussed by Zhang et al. (2019) as a first example of the application of binPMF on ambient data.  
 327 Our input data here is slightly different. In the previous study, the 10 min automatic zeroing every  
 328 three hours was not removed before averaging to 1 hour time resolution while here, we have removed  
 329 this data. Overall, the results are similar as in our earlier study, and therefore the result are just briefly  
 330 summarized below for further comparison and discussion in Section 4. Similar to Range 1, both the  
 331  $Q/Q_{exp}$  (2.2 to 0.6) and unexplained variation (16% to 8%) declined with the increased factor number  
 332 from two to ten.

333 When the factor number was two, one daytime and one nighttime factor were separated, with diurnal  
 334 peak times at 14:00 and 17:00, respectively. The nighttime factor was characterized by masses at 340  
 335 Th, 308 Th and 325 Th (monoterpene ozonolysis HOM monomers (Ehn et al., 2014)) and remained  
 336 stable throughout the factor evolution from two to ten factors. With the addition of more factors, no  
 337 more nighttime factors got separated while the daytime factor was further separated and more daytime  
 338 factors appeared, peaking at various times in the morning (10:00 am), at noon or in the early afternoon  
 339 (around 14:00 pm and 15:00 pm). High contribution of 339 Th can be found in all the daytime factor

340 profiles. As the factor number reached six, a contamination factor appeared, characterized by large  
 341 signals at 339 Th and 324 Th, showing negative mass defects (Figure S2 in the Supplement). The  
 342 factor profile is nearly identical to the contamination factor determined in Zhang et al. (2019), where  
 343 the zeroing periods were not removed, causing larger signals for the contaminants. In our dataset,  
 344 where the zeroing periods were removed, no sawtooth pattern was discernible in the diurnal trend,  
 345 yet it could still be separated even though it only contributed 3% to Range 2. More about the  
 346 contamination factors from different ranges will be discussed in Section 4.4. Since the aim of this  
 347 study is mainly to explore the benefits of analyzing different ranges of the mass spectrum, we chose  
 348 to show the four-factor result below, to simplify the later discussion and comparison. Figure 3 shows  
 349 four-factor result of Range 2, with spectral profile, time series, diurnal cycle and averaged factor  
 350 contribution during the campaign.

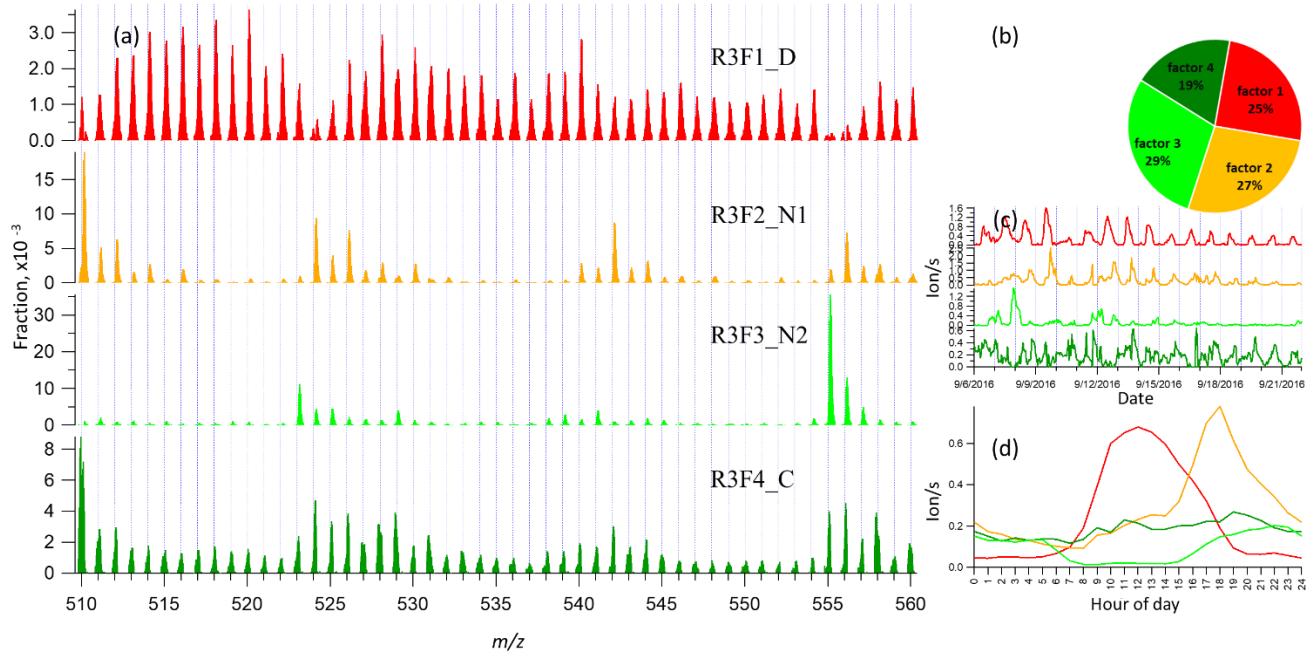


351  
 352 Figure 3 Four-factor result for Range 2, for (a) factor spectral profiles, (b) averaged factor contribution during  
 353 the campaign, (c) time series and (d) diurnal trend. Details on the factors' naming schemes are shown in Table  
 354 1.  
 355

### 356 3.4 binPMF on Range 3 (510-560 Th)

357 Range 3 represents mainly the monoterpene HOM dimers (Ehn et al., 2014). Similar to Range 1 and  
 358 2, both the  $Q/Q_{exp}$  (1.5 to 0.6) and unexplained variation (18% to 15%) showed decreasing trend with  
 359 the increased factor number (2-10). As can be seen from Figure 1, data in Range 3 had much lower  
 360 signals, compared to that of the Range 1 and 2, explaining the higher unexplained variation for Range  
 361 3.

362 In the two-factor result for Range 3, one daytime and one nighttime factor appeared, with diurnal  
 363 peak times at noon and 18:00, respectively. The nighttime factor was characterized by ions at 510 Th,  
 364 524 Th, 526 Th, 542 Th, and 555 Th, 556 Th, while the daytime factor showed no dominant marker  
 365 masses, yet with relatively high signals at 516 Th, 518 Th and 520 Th. As the number of factors  
 366 increased to three, one factor with almost flat diurnal trend was separated, with dominant masses of  
 367 510 Th, 529 Th, 558 Th. Most peaks in this factor had negative mass defects, and this factor was  
 368 again linked to a contamination factor. The four-factor result resolved another nighttime factor with  
 369 a dominant peak at 555 Th, and effectively zero contribution during daytime. As the factor number  
 370 was further increased, the new factors seemed like splits from previous factors with similar spectral  
 371 profiles. We therefore chose four-factor result also for Range 3 (results shown in Fig. 4) for further  
 372 discussion.

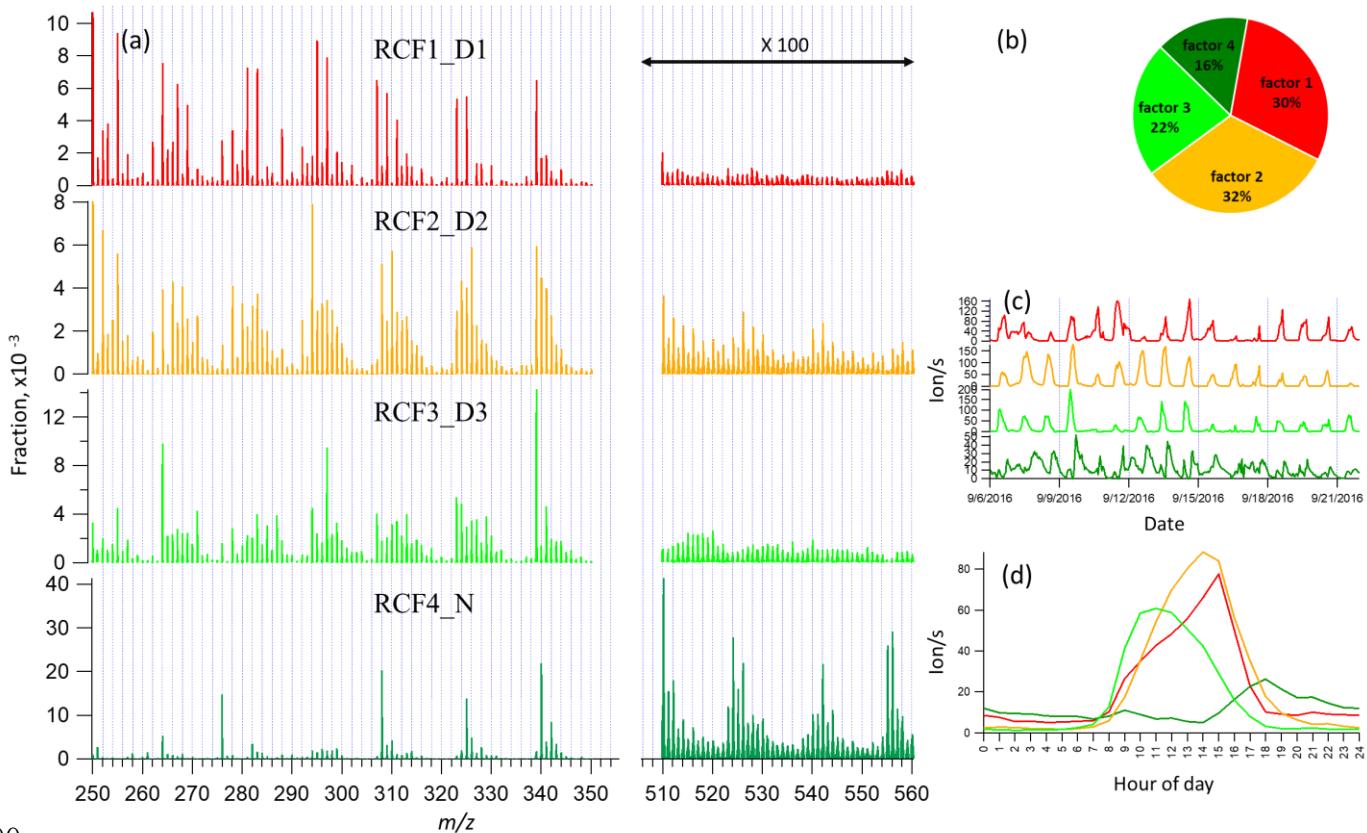


373  
 374 Figure 4 Four-factor result for Range 3, for (a) factor spectral profiles, (b) averaged factor contribution during  
 375 the campaign, (c) time series and (d) diurnal trend. Details on the factors' naming schemes are shown in Table  
 376 1.  
 377

### 3.5 binPMF on Range Combined (250-350 Th & 510-560 Th)

379 As comparison to the previous three ranges, we conducted the binPMF analysis on Range Combined,  
 380 which is the combination of the three ranges. The results of this range are fairly similar to those of  
 381 Ranges 1 and 2, as could be expected since the signal intensities in these ranges were much higher  
 382 than in Range 3. As the number of factors increased (2-10), both the  $Q/Q_{exp}$  (1.3 to 0.6) and  
 383 unexplained variation (16% to 8%) showed a decreasing trend.

384 In the two-factor result, one daytime factor and one nighttime factor were separated. In the nighttime  
 385 factor, most masses were found at even masses, and the fraction of masses in Range 3 was much  
 386 higher than that in daytime factor. In contrast, in the daytime factor, most masses were observed at  
 387 odd masses and the fraction of signal in Range 3 was much lower. During the day, photochemical  
 388 reactions as well as potential emissions increase the concentration of NO, which serves as peroxy  
 389 radical ( $\text{RO}_2$ ) terminator and often outcompetes  $\text{RO}_2$  cross reactions in which dimers can be formed  
 390 (Ehn et al., 2014). Thus, the production of dimers is suppressed during the day, yielding instead a  
 391 larger fraction of organic nitrates, as has been shown also previously (Yan et al., 2016).  
 392 With the increase of the number of factors, more daytime factors were resolved with different peak  
 393 times. When the factor number reached seven, a clear sawtooth-shape diurnal cycle occurred, i.e. the  
 394 contamination factor, caused by the zeroing. As more factors were added, no further nighttime factors  
 395 were separated, and only more daytime factors appeared. To simplify the discussion and inter-range  
 396 comparison, we also here chose the four-factor result for further analysis, as it already provided  
 397 enough information for our main goal in this study. Figure 5 shows the four-factor result of Range  
 398 Combined, with spectral profile, time series, diurnal cycle and averaged factor contribution during  
 399 the campaign. The signals in range of 510-560 Th were enlarged 100-fold to be visible.



400

401 Figure 5 Four-factor result for Range Combined, for (a) factor spectral profiles, (b) averaged factor  
402 contribution during the campaign, (c) time series and (d) diurnal trend. Details on the factors' naming  
403 schemes are shown in Table 1.

404

## 405 **4 Discussion**

406 In Section 3, results by binPMF analysis were shown for Ranges 1, 2, 3 and Combined. In this section,  
407 we discuss and compare the results from the different ranges. To simplify the inter-range comparison,  
408 we chose four-factor results for all the four ranges, with the abbreviations shown in Table 1. From  
409 Range 1, three daytime factors and a contaminations factor were separated. In Range 2, three daytime  
410 factors and one nighttime factor (abbreviated as R2F4\_N) were resolved. The R2F4\_N factor was  
411 characterized by signals at 308 Th ( $C_{10}H_{14}O_7\cdot NO_3^-$ ), 325 Th ( $C_{10}H_{15}O_8\cdot NO_3^-$ ), and 340 Th  
412 ( $C_{10}H_{14}O_9\cdot NO_3^-$ ), and can be confirmed as monoterpene ozonolysis products (Ehn et al., 2014; Yan et  
413 al., 2016). With the increase of factor number to six, the contamination factor got separated also in  
414 this mass range. In Range 3, one daytime factor, two nighttime factors and a contamination factor  
415 were separated. The first nighttime factor (R3F2\_N1) had large peaks at 510 Th ( $C_{20}H_{32}O_{11}\cdot NO_3^-$ )  
416 and 556 Th ( $C_{20}H_{30}O_{14}\cdot NO_3^-$ ), dimer products that have been identified during chamber studies of  
417 monoterpene ozonolysis (Ehn et al., 2014). The molecule observed at 510 Th has 32 H-atoms,  
418 suggesting that one of the RO<sub>2</sub> involved would have been initiated by OH, which is formed during  
419 the ozonolysis of alkenes such as monoterpenes at nighttime (Atkinson et al., 1992; Paulson and  
420 Orlando, 1996). The other nighttime factor (R3F3\_N2) was dominated by ions at 523 Th  
421 ( $C_{20}H_{31}O_8NO_3\cdot NO_3^-$ ) and 555 Th ( $C_{20}H_{31}O_{10}NO_3\cdot NO_3^-$ ), representing nighttime monoterpene  
422 oxidation involving NO<sub>3</sub>. As these dimers contain only one N-atom, and 31 H-atoms, we can assume  
423 that they are formed from reactions between an RO<sub>2</sub> formed from NO<sub>3</sub> oxidation and another RO<sub>2</sub>  
424 formed by ozone oxidation. These results match well with the profiles in a previous study by Yan et  
425 al. (2016). The results of Range Combined are very similar to Range 2, with one nighttime factor and  
426 three daytime factors. The contamination factor was separated with increase of factor number to seven.  
427

428 Table 1. Summary of PMF results for the different mass ranges

Range	Factor number	Factor name <sup>a</sup>	Dominant peaks	Peak time
1 (250 - 300 Th)	1	R1F1_D1	250, 255, 295, 297	15:00
	2	R1F2_D2	250, 252, 294	15:00
	3	R1F3_D3	264, 297	11:00
	4	R1F4_C	276	- <sup>b</sup>
2 (300 - 350 Th)	1	R2F1_D1	307, 309, 323, 325, 339,	15:00
	2	R2F2_D2	310, 326, 339,	14:00
	3	R2F3_D3	339	11:00
	4	R2F4_N	308, 325, 340	18:00

	1	R3F1_D	516, 518, 520, 528, 540	12:00
3 (510 – 560 Th)	2	R3F2_N1	510, 524, 542, 556	18:00
	3	R3F3_N2	523, 555	22:00
	4	R3F4_C	510, 558	_ <sup>b</sup>
	1	RCF1_D1	250, 255, 295, 339	15:00
Combined (1, 2, 3)	2	RCF2_D2	250, 252, 294, 339	14:00
	3	RCF3_D3	264, 297, 339	11:00
	4	RCF4_N	308, 340, 510, 524, 555, 556	18:00

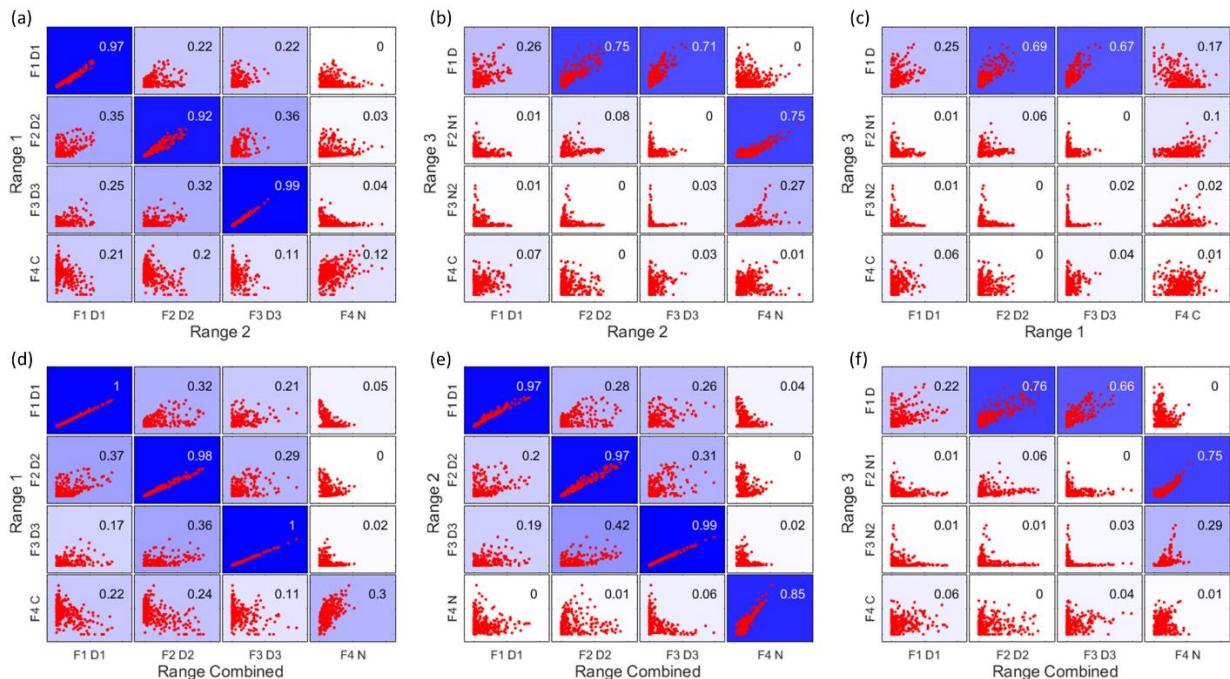
<sup>a</sup> Factor name is defined with range name, factor number and name. For example, RxFy represents Factor y in Range x. RC stands for Range Combined. For the factor name, D is short for daytime, N for Nighttime, C for contamination.

<sup>b</sup> The contamination factor in Range 1 shows sawtooth pattern; while in Range 3 shows no diurnal pattern.

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#### 433 4.1 Time series correlation

434 In Figure 6, the upper panels show the time series correlations among the first three ranges. As  
435 expected based on the results above, generally the daytime factors, and the two nighttime  
436 monoterpane ozonolysis factors (R2F4\_N and R3F2\_N1) correlated well, respectively. However, the  
437 contamination factors did now show strong correlation between different ranges, even though  
438 undoubtedly from the same source. More about the contamination factors will be discussed in Section  
439 4.1.4. The lower panels in Figure 6 displays the correlations between the first three ranges and the  
440 Range Combined, and clearly demonstrates that the results of Range Combined is mainly controlled  
441 by high signals from Range 1 and 2. More detailed aspects of the comparison between factors in  
442 different ranges is given in the following sections. The good agreements between factors from  
443 different subranges also help to verify the robustness of the solutions.



444  
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Figure 6 Time series correlations among Range 1, 2, 3 (upper panels a-c), and between the first three ranges and the Range Combined (lower panels d-f). The abbreviations for different factors are the same in Table 1, with F for factor, D for daytime, N for nighttime and C for contamination, e.g. F1D1 for Factor 1 daytime 1.

448 The coefficient of determination,  $R^2$ , is marked in each subplot by a number shown in the right upper corners  
449 and by the blue colors, with stronger blue indicating higher  $R^2$ .  
450

## 451 **4.2 Daytime processes**

### 452 **4.2.1 Factor comparison**

453 As mentioned above, with increasing number of factors, usually more daytime factors will be resolved,  
454 reflecting the complicated daytime photochemistry. The three daytime factors between Range 1 and  
455 2 agreed with each other quite well (Figure 6a). However, R1F1\_D1 and R2F1\_D1 did not show  
456 strong correlation with the only daytime factor in Range 3 (R3F1\_D), while the other two daytime  
457 factors in both Range 1 and 2, i.e. R1F2\_D2, R1F3\_D3, R2F2\_D2, R2F3\_D3, correlated well with  
458 R3F1\_D from Range 3.

459 The 1<sup>st</sup> daytime factors from Range 1 and 2, R1F1\_D1 and R2F1\_D1, were mainly characterized by  
460 odd masses 255 Th, 281 Th, 283 Th, 295 Th, 297 Th, 307 Th, 309 Th, 311 Th, 323 Th, 325 Th, 339  
461 Th. The factors are dominated by organonitrates. Organic nitrate formation during daytime is  
462 generally associated with the termination of RO<sub>2</sub> radicals by NO. This termination step is mutually  
463 exclusive with the termination of RO<sub>2</sub> with other RO<sub>2</sub>, which can lead to dimer formation. If the NO  
464 concentration is the limiting factor for the formation of these factors, the low correlations between  
465 the NO-terminated monomer factors, and the dimer factors, is to be expected. In contrast, if the other  
466 daytime factors mainly depend on oxidant and monoterpene concentrations, some correlation  
467 between those, and the daytime dimer factor, is to be expected, as shown in Figure 6b, c.

468 All the spectral profiles resolved from Range Combined binPMF analysis inevitably contained mass  
469 contribution from 510 – 560 Th, even the daytime factor from Range Combined (RCF1\_D1) which  
470 did not show clear correlation with R3F1\_D from Range 3 (Figure 6e).

471 The 2<sup>nd</sup> and 3<sup>rd</sup> daytime factors in Range 1 and 2, R1F2\_D2, R1F3\_D3, R2F2\_D2, R2F3\_D3, had  
472 high correlations with R3F1\_D in Range 3. Daytime factors in Range Combined (RCF2\_D2 and  
473 RCF3\_D3) also showed good correlation with R3F1\_D in Range 3. However, if we compare R3F1\_D  
474 and the mass range of 510 – 560 Th of the daytime factors in Range Combined, just with a quick look,  
475 we can readily see the difference. The daytime factor separated in Range 3 (R3F1\_D) has no obvious  
476 markers in the profile. With the increase of factor number (up to ten factors), no clearly new factors  
477 were separated in Range 3, but instead the previously separated factors were seen to split into several  
478 factors. However, the spectral pattern in R3F1\_D is different from that in the mass range of 510 –  
479 560 Th in RCF2\_D2. The factorization of Range Combined was mainly controlled by low masses  
480 due to their high signals. The signals at high masses were forced to be distributed according to the

481 time series determined by small masses. Ultimately, this will lead to failure in factor separation for  
 482 this low-signal range.

#### 483 **4.2.2 Daytime dimer formation**

484 Dimers are primarily produced during nighttime, due to NO suppressing RO<sub>2</sub> + RO<sub>2</sub> reactions in  
 485 daytime (Ehn et al., 2014; Yan et al., 2016). However, in this study, we found one clear daytime factor  
 486 in Range 3 (R3F1\_D, peak at local time 12:00, UTC+2) by sub-range analysis. With high loadings  
 487 from even masses including 516, 518, 520, 528, 540 Th, this only daytime factor in dimer range  
 488 correlated very well with two daytime factors in Ranges 1 and 2 (R1F2\_D2, R1F3\_D3, R2F2\_D2,  
 489 R2F3\_D3) (Figure 6b and c). Table 2 include the correlation matrix of all PMF and factors and  
 490 selected meteorological parameters. Strong correlation between R3F1\_D with solar radiation was  
 491 found, with R = 0.79 (Table 2). This may indicate involvement of OH oxidation in producing this  
 492 factor.

493 Table 2 Correlation between factors and meteorological parameters and gases

	R1F1_D1	R1F1_D2	R1F1_D3	R1F1_C	R2F1_D1	R2F2_D2	R2F3_D3	R2F4_N	R3F1_D	R3F2_N1	R3F3_N2	R3F4_C	RCF1_D1	RCF2_D2	RCF3_D3	RCF4_N
O <sub>3</sub>	0.51	0.59	0.35	-0.18	0.47	0.57	0.36	0.43	0.55	0.33	0.27	0.22	0.49	0.57	0.33	0.34
NO	0.13	-0.01	0.24	-0.03	0.18	-0.02	0.24	-0.22	0.13	-0.19	-0.17	0.03	0.13	0.00	0.26	-0.18
NOx	-0.05	-0.22	-0.10	0.09	-0.01	-0.23	-0.11	-0.13	-0.16	-0.21	-0.04	0.04	-0.04	-0.22	-0.09	-0.11
RH	-0.46	-0.80	-0.63	0.30	-0.43	-0.82	-0.64	-0.27	-0.78	-0.39	-0.07	-0.07	-0.43	-0.82	-0.60	-0.21
T	0.66	0.72	0.40	-0.24	0.65	0.66	0.41	0.39	0.65	0.30	0.14	0.19	0.66	0.68	0.38	0.24
UVB	0.52	0.63	0.82	-0.40	0.52	0.68	0.84	-0.30	0.79	-0.08	-0.27	0.08	0.49	0.68	0.83	-0.29

494  
 495 As previous studies have shown, dimers greatly facilitate new particle formation (NPF) (Kirkby et  
 496 al., 2016; Troestl et al., 2016; Lehtipalo et al., 2018), and this daytime dimer factor may represent a  
 497 source of dimers that would impact the initial stages of NPF in Hyytiälä. Mohr et al. (2017) reported  
 498 a clear diel pattern of dimers (sum of about 60 dimeric compounds of C<sub>16-20</sub>H<sub>13-33</sub>O<sub>6-9</sub>) during NPF  
 499 events in 2013 in Hyytiälä, with minimum at night and maximum after noon, and estimated these  
 500 dimers can contribute ~5% of the mass of sub-60 nm particles. The link between the dimers presented  
 501 in that paper and those reported here will require further studies, as will the proper quantification of  
 502 the dimer factor identified here.

503

#### 504 **4.3 Nighttime processes**

##### 505 **4.3.1 Factor comparison**

506 Since high-mass dimers are more likely to form at night due to photochemical production of NO in  
 507 daytime, which inhibits RO<sub>2</sub> + RO<sub>2</sub> reactions, Range 3 had the highest fraction of nighttime signals

508 of all the sub-ranges. While Range 3 produced two nighttime factors, Ranges 2 and Combined showed  
509 one, and Range 1 had no nighttime factor. The difference between the two results also indicates the  
510 advantage of analyzing monomers and dimers separately.

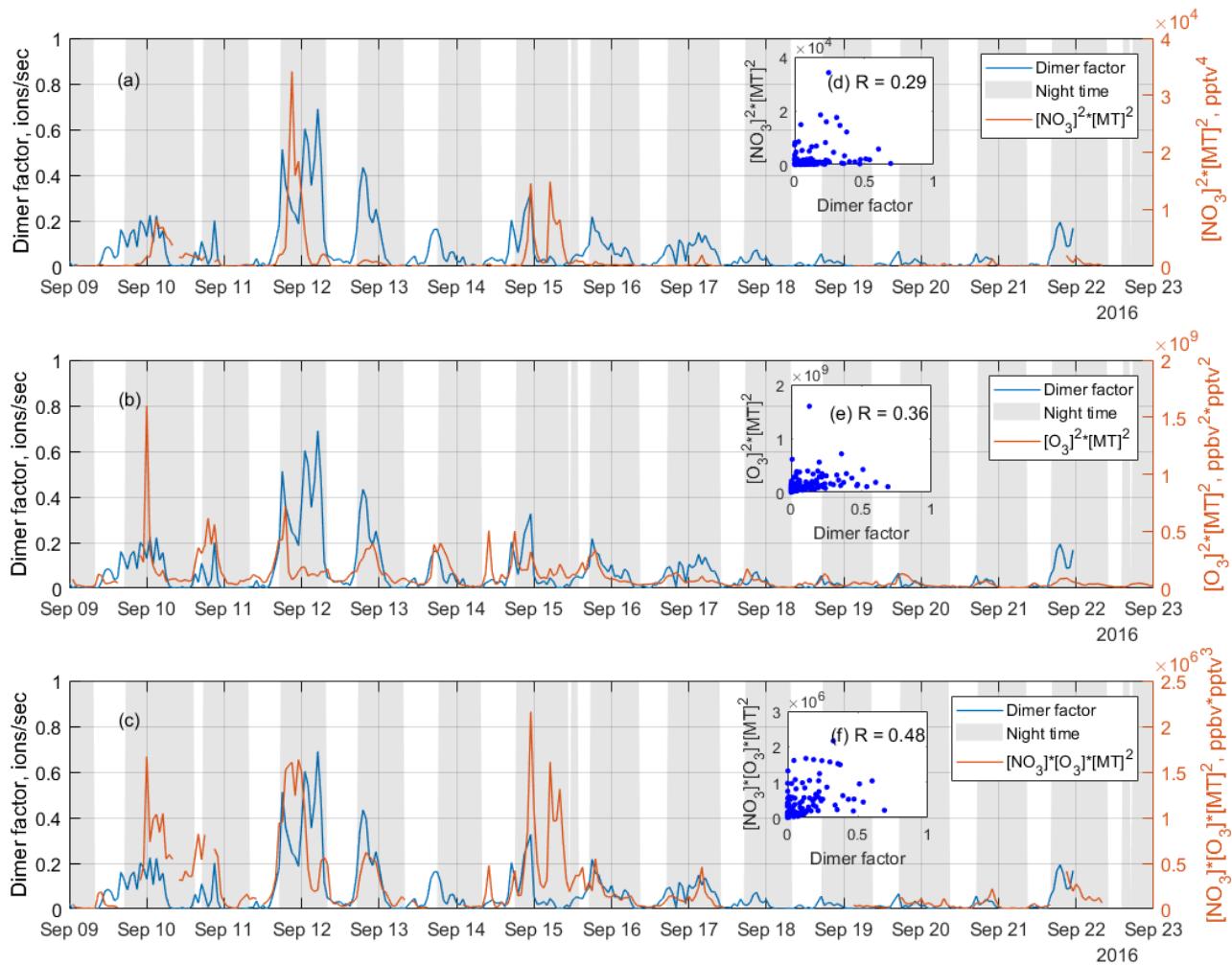
511 The two nighttime factors in Range 3 can be clearly identified as arising from ozonolysis (R3F2\_N1)  
512 and a mix of ozonolysis and NO<sub>3</sub> oxidation (R3F2\_N2) based on the mass spectral profiles, as  
513 described above. The organonitrate at 555 Th, C<sub>20</sub>H<sub>31</sub>O<sub>10</sub>NO<sub>3</sub>·NO<sub>3</sub><sup>-</sup>, is a typical marker for NO<sub>3</sub>  
514 radical initiated monoterpene chemistry (Yan et al., 2016). However, several interesting features  
515 become evident when comparing to the results of Range 2 and Combined. Firstly, only one nighttime  
516 factor (R2F4\_N, RCF4\_N) was separated in each of these ranges, and that shows clear resemblance  
517 with ozonolysis of monoterpenes as measured in numerous studies, e.g. (Ehn et al., 2012;Ehn et al.,  
518 2014). Secondly, the high correlation found in Figure 6b between the ozonolysis factors (i.e.,  
519 R2F4\_N, R3F2\_N1, RCF4\_N), further supports the assignment. However, factor R2F4\_N is the only  
520 nighttime factor in the monomer range, suggesting that NO<sub>3</sub> radical chemistry of monoterpenes in  
521 Hyytiälä does not form substantial amounts of HOM monomers. The only way for the CI-APi-TOF  
522 to detect products of monoterpene-NO<sub>3</sub> radical chemistry may thus be through the dimers, where one  
523 highly oxygenated RO<sub>2</sub> radical from ozonolysis reacts with a less oxygenated RO<sub>2</sub> radical from NO<sub>3</sub>  
524 oxidation.

525 In the results by Yan et al. (2016) the combined UMR-PMF of monomers and dimers did yield a  
526 considerable amount of compounds in the monomer range also for the NO<sub>3</sub> radical chemistry factor.  
527 There may be several reasons for this discrepancy. One major cause for differences between the spring  
528 dataset of Yan et al. (2016) and the autumn dataset presented here, is that nighttime concentrations  
529 of HOM was greatly reduced during our autumn campaign. The cause may have been fairly frequent  
530 fog formation during nights, and also the concentration of e.g. ozone decreased nearly to zero during  
531 several nights (Zha et al., 2018). It is also possible that the NO<sub>3</sub> radical-related factor by Yan et al.  
532 (2016) is probably a mixture of NO<sub>3</sub> and O<sub>3</sub> radical chemistry, while the monomer may thus be  
533 attributed to the O<sub>3</sub> part. Alternatively, the different conditions during the two measurement periods,  
534 as well as seasonal difference in monoterpene mixtures (Hakola et al., 2012), caused variations in the  
535 oxidation pathways.

### 537 **4.3.2 Dimers initiated by NO<sub>3</sub> radicals**

538 Previous studies show that NO<sub>3</sub> oxidation of α-pinene, the most abundant monoterpene in Hyytiälä  
539 (Hakola et al., 2012), produces fairly little SOA mass (yields <4 %), while β-pinene shows yields of  
540 up to 53 % (Bonn and Moorgat, 2002;Nah et al., 2016). The NO<sub>3</sub>+β-pinene reaction results in low  
541 volatile organic nitrate compounds with carboxylic acid, alcohol, and peroxide functional groups (Fry

542 et al., 2014; Boyd et al., 2015), while  $\text{NO}_3 + \alpha\text{-pinene}$  reaction will typically lose the nitrate functional  
 543 group and form oxidation products with high vapor pressures (Spittler et al., 2006; Perraud et al.,  
 544 2010). Most monoterpene-derived HOM, including monomers, are low-volatile (Peräkylä et al.,  
 545 2020) and thus a low SOA yield indicates a low HOM yield. Thus, while there are to our knowledge  
 546 no laboratory studies on HOM formation from  $\text{NO}_3$  oxidation of  $\alpha\text{-pinene}$ , a low yield can be expected  
 547 based on SOA studies.



548

549 Figure 7 Time series of the  $\text{NO}_3$  oxidation dimer factor (blue line), and the product of (a)  $[\text{NO}_3]^2 \times$   
 550  $[\text{monoterpene}]^2$ , (b)  $[\text{O}_3]^2 \times [\text{monoterpene}]^2$ , and (c)  $[\text{NO}_3] \times [\text{O}_3] \times [\text{monoterpene}]^2$ , where  $[\cdot]$  represents  
 551 concentration in unit of pptv for  $\text{NO}_3$  radicals and monoterpene, ppbv for  $\text{O}_3$ , while the scatter plots are shown  
 552 as inserts, (d), (e), (f), respectively. The scatter plots and correlation coefficients  $R$  are only calculated from  
 553 nighttime data, which is selected based on solar radiation, to eliminate the influence from daytime oxidation  
 554 processes.

555

556 As discussed above, a dimer factor (R3F2\_N2) was identified as being a crossover between  $\text{NO}_3$   
 557 radical initiated and  $\text{O}_3$  initiated  $\text{RO}_2$  radicals. Figure 7 shows the time series of this factor, as well as  
 558 the product of  $[\text{NO}_3]^2 \times [\text{monoterpene}]^2$ ,  $[\text{O}_3]^2 \times [\text{monoterpene}]^2$ , and  $[\text{NO}_3] \times [\text{O}_3] \times [\text{monoterpene}]^2$ .  
 559 These products are used to mimic the formation rates of the  $\text{RO}_2$  radicals reacting to form the dimers,

either from pure  $\text{NO}_3$  oxidation (Fig. 7a), pure  $\text{O}_3$  oxidation (7b), or the mixed reaction between  $\text{RO}_2$  from the two oxidants (7c). The  $\text{NO}_3$  concentration was estimated in Liebmann et al. (2018) for the same campaign. Monoterpenes were measured using a proton transfer reaction time of flight mass spectrometer (PTR-TOF-MS). More details on measurement of  $\text{NO}_3$  proxy and monoterpene can be found in Liebmann et al. (2018).

As shown in Figure 7, the time series of the dimer factor tracks those of  $[\text{NO}_3] \times [\text{monoterpene}]$  and  $[\text{O}_3] \times [\text{monoterpene}]$  reasonably well, but shows the highest correlation with the product of  $[\text{NO}_3] \times [\text{O}_3] \times [\text{monoterpene}]^2$ . This further supports this dimer formation as a mixed processes of ozonolysis and  $\text{NO}_3$  oxidation. The heterogeneity of the monoterpene emissions in the forest, and the fact that no dimer loss process is included, partly explain the relatively low correlation coefficients. The sampling inlets for PTR-TOF were about 170 m away from the  $\text{NO}_3$  reactivity measurement (Liebmann et al., 2018), which in turn was some tens of meters away from the HOM measurements. Thus, this analysis should be considered qualitative only.

The nitrate dimer factor (R3F2\_N2) was dominated by the organonitrate at 555 Th,  $\text{C}_{20}\text{H}_{31}\text{O}_{10}\text{NO}_3\cdot\text{NO}_3^-$ . However, unlike the pure ozonolysis dimer factor which had a corresponding monomer factor ( $R = 0.86$  between factor R2F4\_N and R3 F2\_N1), this  $\text{NO}_3$ -related dimer factor did not have an equivalent monomer factor. This suggests that the  $\text{NO}_3$  oxidation of the monoterpene mixture in Hytiälä does not by itself form much HOM, but in the presence of  $\text{RO}_2$  from ozonolysis, the  $\text{RO}_2$  from  $\text{NO}_3$  oxidation can take part in HOM dimer formation. This further implies that, different from previous knowledge based on single-oxidant experiments in chambers,  $\text{NO}_3$  oxidation may have a larger impact on SOA formation in the atmosphere where different oxidants exist concurrently. This highlights the need for future laboratory studies to consider systems with multiple oxidants during monoterpene oxidation experiments, to truly understand the role and contribution of different oxidants, and  $\text{NO}_3$  in particular.

#### 4.4 Fluorinated compounds

During the campaign, an automated instrument zeroing every three hours was conducted. While the zeroing successfully removed the low-volatile HOM and  $\text{H}_2\text{SO}_4$ , the process also introduced contaminants into the inlet lines, e.g. perfluorinated organic acids from Teflon tubing. Each zeroing process lasted for 10 min. In the data analysis, we removed all the 10-min zeroing periods, and averaged the data to 1-h time resolution, but contaminants were still identified in all ranges by binPMF. However, the correlation between contamination factors from different ranges is low (Figure 6c).

592 To future investigate the low factor correlations of the same source, three fluorinated compounds with  
593 different volatilities,  $(CF_2)_3CO_2HF \cdot NO_3^-$  (275.9748 Th),  $(CF_2)_5C_2O_4H^-$  (338.9721 Th), and  
594  $(CF_2)_6CO_2HF \cdot NO_3^-$  (425.9653 Th), were examined in fine time resolution, i.e. 1 min. The time series  
595 and 3-h cycle of the three fluorinated compounds were shown in Figure S3 and S4 in Supplement.  
596 The correlation coefficients dropped greatly before and after the zero period was removed, from 0.9  
597 to 0.3 for  $R^2$  between 276 Th and 339 Th, and 0.8 to 0.1 between 276 Th and 426 Th (Fig. S5a, b).  
598 Similar effect is also found with the 1 h averaged data (Fig. S5c, d). It is evident that the three  
599 fluorinated compounds were from the same source (zeroing process), but due to their different  
600 volatilities, they were lost at different rates. This, in turn, means that the spectral signature of this  
601 source will change as a function of time, at odds with one of the basic assumptions of PMF.  
602 The analysis of the fluorinated compounds in our system was here merely used as an example to show  
603 that volatility can impact source profiles over time. In Figure S5, it can be clearly seen that the profile  
604 of Range Combined is noisier than that of Range 3, probably due to the varied fractional contributions  
605 of contamination compounds to the profile. In ambient data, products from different sources can have  
606 undergone atmospheric processing, altering the product distribution. This analysis highlighted the  
607 importance of differences in the sink terms due to different volatilities of the products. This may be  
608 an important issue for gas phase mass spectrometry analysis, potentially underestimated by many  
609 PMF users, as it is likely only a minor issue for aerosol data, for which PMF has been applied much  
610 more routinely. If failing to achieve physically meaningful factors using PMF on gas phase mass  
611 spectra, our recommendation is to try applying PMF to sub-ranges of the spectrum, where IVOC,  
612 SVOC and (E)LVOC could be analyzed separately.

613

#### 614 **4.5 Atmospheric insights**

615 Based on the new data analysis technique binPMF applied in sub-ranges of mass spectra, we were  
616 able to separate two particularly intriguing atmospheric processes, the formation of daytime dimers  
617 as well as dimer formation involving  $NO_3$  radicals, which otherwise could not have been identified  
618 in our study.

619 With a diurnal peak around noon time, the daytime dimers identified in this study correlate very well  
620 with daytime factors in monomer range. Strong correlation between this factor and solar radiation  
621 indicate the potential role of OH oxidation in the formation of daytime dimers. By now, very few  
622 studies have reported the observations of daytime dimers. As dimers are shown to be able to take part  
623 in new particle formation (NPF) (Kirkby et al., 2016), this daytime dimer may contribute to the early  
624 stages of NPF in the boreal forest.

625 The second process identified in our study is the formation of dimers that are a crossover between  
626  $\text{NO}_3$  and  $\text{O}_3$  oxidation. Such dimers have been identified before (Yan et al., 2016). However, we were  
627 not able to identify corresponding HOM monomer compounds. This finding indicates that while  $\text{NO}_3$   
628 oxidation of the monoterpenes in Hyytiälä may not undergo autoxidation to form HOM by themselves,  
629 they can contribute to HOM dimers when the  $\text{NO}_3$ -derived  $\text{RO}_2$  react with highly oxygenated  $\text{RO}_2$   
630 from other oxidants. Multi-oxidant systems should be taken into consideration in future experimental  
631 studies on monoterpene oxidation processes.

632

## 633 **5 Conclusions**

634 The recent development in mass spectrometry, combined with factor analysis such as PMF, has  
635 greatly improved our understanding of complicated atmospheric processes and sources. However,  
636 one of PMF's basic assumptions is that factor profiles remain constant in time, yet for atmospheric  
637 gas-phase species, reactions and sinks may violate this assumption. In this study, we conducted  
638 separate binPMF analysis on three different sub-ranges to explore the potential benefits of such an  
639 approach for producing more physically meaningful factors.

640 With binPMF applied on sub-ranges, our study identified daytime dimers, presumably initiated by  
641  $\text{OH}/\text{O}_3$  with a diurnal peak at around noon, which may contribute to NPF in Hyytiälä. Also, based on  
642 the sub-range binPMF analysis, we successfully separated  $\text{NO}_3$ -related dimers which did not have a  
643 corresponding monomer factor. The  $\text{NO}_3$ -related factor was consistent with earlier observations (Yan  
644 et al., 2016), but would not have been identified from this dataset without utilizing the different sub-  
645 ranges. In future laboratory experiments, more complex oxidation systems may be useful in order to  
646 understand the role of  $\text{NO}_3$  oxidation in SOA formation. Apart from these two findings, we also find  
647 other benefits by applying binPMF on sub-ranges of the mass spectra.

648 First, volatility affects the PMF results. Different compounds emitted from the same source showed  
649 different temporal trends, likely due to differences in volatilities. This increased the difficulties for  
650 PMF to separate this source in the combined data set, and the resolved profile was less accurate than  
651 that of the sub-ranges. Future studies of gas-phase mass spectra should pay attention to this volatility  
652 effect on factor analysis.

653 Secondly, chemistry or sources contributing to the particular range can be better separated. Only the  
654 binPMF analysis on Range 3, where HOM dimers are typically observed, resolved two nighttime  
655 factors, characterized by monoterpene oxidation related to  $\text{NO}_3$  and  $\text{O}_3$  oxidation.

656 Thirdly, peaks with smaller signal intensities can be correctly assigned. The signal intensities between  
657 different parts of the mass spectrum may vary by orders of magnitude. In the combined case, the  
658 results were almost completely controlled by the higher signals from smaller masses. The separate

659 analysis on Range 3 allowed the low signals to provide important information. In addition, running  
660 binPMF on different separate mass ranges also allows us to compare the factors obtained from the  
661 different ranges and help to verify the results.

662

668 **Data availability.** The data used in this study are available from the first author upon request: please  
669 contact Yanjun Zhang ([yanjun.zhang@helsinki.fi](mailto:yanjun.zhang@helsinki.fi)).

677 **Author contribution.** ME and YZ designed the study. QZ and MR collected the data; data analysis  
678 and manuscript writing were done by YZ. All coauthors discussed the results and commented the  
679 manuscript.

680 **Competing interests.** The authors declare that they have no conflict of interest

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