

1 **Supercooled Liquid Water Cloud observed and analysed at the Top**
2 **of the Planetary Boundary Layer above Dome C, Antarctica**

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21 **Abstract**

22 A comprehensive analysis of the water budget over the Dome C (Concordia, Antarctica)
23 station has been performed during the austral summer 2018-2019 as part of the Year of Polar
24 Prediction (YOPP) international campaign. Thin (~100-m **deep**) supercooled liquid water
25 (SLW) clouds have been detected and analysed using remotely sensed observations at the
26 station (tropospheric depolarization LIDAR, microwave radiometer HAMSTRAD, net surface
27 radiation from Baseline Surface Radiation Network **BSRN**), radiosondes and using satellite
28 observations (CALIOP/CALIPSO) combined with a specific configuration of the Numerical
29 Weather Prediction model: ARPEGE-SH (**Action de Recherche Petite Echelle Grande Echelle**
30 **– Southern Hemisphere**). **The analysis shows that SLW clouds were present from November to**
31 **March, with the greatest frequency occurring in December and January when ~50% of the days**
32 **in summer time exhibited SLW clouds.** Two case studies are used to illustrate this phenomenon.
33 On 24 December 2018, the atmospheric planetary boundary layer (PBL) evolved following a
34 **typical** diurnal variation, **which** is to say with a warm and dry mixing layer at local noon thicker
35 than the cold and dry stable layer at local midnight. Our study showed that the SLW clouds
36 were observed at Dome C within the entrainment and the capping inversion zones at the top of
37 the PBL. ARPEGE-SH was not able to correctly estimate the ratio between liquid and solid
38 water inside the clouds with the **Liquid Water Path (LWP)** strongly underestimated **by a factor**
39 **1000 compared to observations.** The lack of simulated SLW in the model impacted the net
40 surface radiation that was 20-30 W m⁻² higher in the BSRN observations than in the ARPEGE-
41 SH calculations, mainly attributable to **the BSRN** longwave downward surface radiation being
42 50 W m⁻² greater than that of ARPEGE-SH. **The second case study takes place on** 20 December
43 2018, **when** a warm and wet episode impacted the PBL with no clear diurnal cycle of the PBL
44 top. SLW cloud appearance within the entrainment and capping inversion zones **coincided with**
45 **the warm and wet event.** The amount of liquid water measured by HAMSTRAD was ~20 times

46 greater in this perturbed PBL than in the typical PBL. Since ARPEGE-SH was not able to
47 accurately reproduce these SLW clouds, the discrepancy between the observed and calculated
48 net surface radiation was even greater than in the typical PBL case, reaching $+50 \text{ W m}^{-2}$, mainly
49 attributable to the downwelling longwave surface radiation from BSRN being 100 W m^{-2}
50 greater than that of ARPEGE-SH. The model was then run with a new partition function
51 favouring liquid water for temperatures below -20°C down to -40°C . In this test mode,
52 ARPEGE-SH has been able to generate SLW clouds with modelled LWP and net surface
53 radiation consistent with observations during the typical case whereas, during the perturbed
54 case, the modelled LWP was 10 times less than the observations and the modelled net surface
55 radiation remained lower than the observations by $\sim 50 \text{ W m}^{-2}$. Accurately modelling the
56 presence of SLW clouds appears crucial to correctly simulate the surface energy budget over
57 the Antarctic Plateau.

58

59 1. Introduction

60 Antarctic clouds play an important role in the climate system by influencing the Earth's
61 radiation balance, both directly at high southern latitudes and, indirectly, at the global level
62 through complex teleconnections (Lubin et al., 1998). In Antarctica, there are very few
63 observational stations and most of them are located on the coast, a fact that limits the type and
64 characteristics of clouds observed. Nevertheless, prior studies suggest that cloud properties vary
65 geographically, with a fractional cloud cover around the South Pole of about 50 to 60% in all
66 seasons, and a cloud cover of about 80 to 90% near the coast (Bromwich et al., 2012; Listowski
67 et al., 2019). Based on spaceborne observations, Adhikari et al. (2012) observed that low-level
68 cloud occurrence over the Antarctic Plateau is consistently between 20-50% with the highest
69 values occurring in winter and the lowest values consistently occurring over the Eastern
70 Antarctic Plateau. Furthermore, cloud parameters such as the hydrometeors size and the
71 microphysical structure are also very difficult to retrieve in Antarctica. Nevertheless, some in
72 situ aircraft measurements exist particularly over the Western Antarctic Peninsula (Grosvenor
73 et al., 2012; Lachlan-Cope et al., 2016) and nearby coastal areas (O'Shea et al., 2017) that
74 provide ice mass fraction, concentration and particle size relative to cloud temperature, cloud
75 type and formation mechanism which have provided new insights to polar cloud modelling.
76 These studies also highlighted sea-ice production of ice-condensation nuclei, which is
77 important in winter both coastally and at Dome C (Legrand et al., 2016). Additionally, Grazioli
78 et al. (2017) observed precipitating crystal characteristics at Dumont d'Urville using a
79 combination of ground-based radars, in situ cameras and precipitation sensors, and looked at
80 the role that the katabatic winds play in the formation, modification and sublimation of ice
81 crystals. Over the Antarctic Plateau, where the atmosphere is colder and drier than along the
82 coast, ice crystal clouds are mainly observed with crystal sizes ranging from 5 to 30 μm
83 (effective radius) in the core of the cloud; mixed-phase clouds are preferably observed near the

84 coast (Listowski et al., 2019) with larger ice crystals and water droplets (Lachlan-Cope, 2010;
85 Lachlan-Cope et al., 2016; Grosvenor et al., 2012; O’Shea et al., 2017; Grazioli et al., 2017).

86 The time and geographical distribution of tropospheric clouds over the Antarctic region
87 has been recently studied using the raDAR/liDAR-MASK (DARDAR) spaceborne products
88 (Listowski et al., 2019). The authors determined that clouds are mainly constituted of ice above
89 the continent. The presence of Supercooled Liquid Water (SLW, the water staying in liquid
90 phase below 0°C) clouds shows variations according to temperature and sea ice fraction,
91 decreasing sharply poleward, with an abundance two to three times less over the Eastern
92 Antarctic Plateau than over the Western Antarctic. The inability of mesoscale high-resolution
93 models and operational numerical weather prediction models to accurately calculate the net
94 surface radiation due to the presence of clouds (particularly of SLW clouds) in Antarctica
95 causes biases of several tens of watt per square meters (Listowski and Lachlan-Cope, 2017,
96 King et al., 2006, 2015; Bromwich et al., 2013) impacting the radiative budget of the Antarctic
97 and beyond (Lawson and Gettelman, 2014; Young et al. 2019). The year-long study of mixed-
98 phase clouds at South Pole with a micropulse LIDAR presented in Lawson and Gettelman
99 (2014) showed that SLW clouds occur more frequently than observed in earlier aircraft studies,
100 and are underestimated in models leading to biases in the surface radiation budget. In the present
101 study, we explore these biases further, moving the focus to the modelling and simultaneous
102 observations of low-level SLW clouds and surface radiation over the Eastern Antarctic Plateau,
103 specifically at Dome C.

104 With the support of the World Meteorological Organization (WMO) World Weather
105 Research Programme (WWRP), the Polar Prediction Project (PPP) international programme
106 has been dedicated to the development of improved weather and environmental prediction
107 services for the polar regions, on time scales from hours to seasons
108 (<https://www.polarprediction.net>). Within this project, the Year of Polar Prediction (YOPP),

109 from 2018 to 2019, aims at enabling a significant improvement in environmental prediction
110 capabilities for the polar regions and beyond, by coordinating a period of intensive observing,
111 modelling, verification, user-engagement and educational activities. The Water Budget over
112 Dome C (H₂O-DC) project (<https://apps3.awi.de/YPP/pdf/stream/52>) has been endorsed by
113 YOPP for studying the water budget by means of ground-based measurements of water (vapour,
114 solid and liquid) and clouds, by active (backscatter LIDAR) and passive (microwave
115 radiometer) remote sensing, and operational meteorological analyses. The Dome C (Concordia)
116 station is located in the Eastern Antarctic Plateau (75°06'S, 123°21'E, 3233 m above mean sea
117 level, amsl).

118 H₂O-DC concentrates on the Year of Polar Prediction Special Observing Period of
119 measurements in the Antarctic (SOP-SH), from 16 November 2018 to 15 February 2019.
120 During this time frame, several instruments have been employed.

121 1) The H₂O Antarctica Microwave Stratospheric and Tropospheric Radiometer
122 (HAMSTRAD, Ricaud et al., 2010a) to obtain vertical profiles of temperature and water
123 vapour, Integrated Water Content (IWC) or precipitable water, and Liquid Water Path (LWP),
124 with an adjustable time resolution fixed at 60 seconds during the YOPP campaign.

125 2) The tropospheric depolarization LIDAR (Tomasi et al., 2015) to obtain vertical profiles
126 of backscattering and depolarization ratio.

127 These two H₂O-DC data sets have been complemented in the present analysis by the 3
128 following observational datasets.

129 3) The Baseline Surface Radiation Network (BSRN) net surface radiances at the station.

130 4) The temperature profiles from radiosondes launched twice daily at the station during
131 YOPP.

132 5) The spaceborne observations (backscatter and polarization) from the
133 CALIOP/CALIPSO LIDAR in the vicinity of the station.

134 In addition, a specific **Antarctic** configuration of the global ARPEGE model from Météo-
135 France (Pailleux et al., 2015) is used to characterize the water budget above Dome C
136 considering the gas, liquid **and solid** phases to study the genesis of clouds (ice/liquid).

137 The aim of the present study is to combine all these observations and simulations in order
138 to 1) detect the presence of SLW clouds above Dome C, 2) analyse the formation and evolution
139 of such SLW clouds and 3) estimate the radiative impact of such clouds on the net surface
140 radiation. We concentrate the analyses on two case studies observed during the YOPP
141 campaign: one case when the Planetary Boundary Layer (PBL) exhibited a “typical” diurnal
142 cycle (24 December 2018) and a second case when the diurnal cycle of the PBL was perturbed
143 by a warm and wet episode (20 December 2018).

144 The data sets used in our study are presented in section 2. **The methodology employed is**
145 **explained in section 3.** The analyses of the SLW clouds during the typical and the perturbed
146 PBL periods are detailed in sections **4 and 5**, respectively. **The observed and modelled impact**
147 **of SLW clouds on the surface net radiation is described in section 6.** Section 7 includes a
148 **discussion of the results and the conclusion synthesizes the study in section 8.**

149

150 **2. Datasets**

151 **2.1. The HAMSTRAD Radiometer**

152 HAMSTRAD is a microwave radiometer **that profiles** water vapour (H₂O), liquid water and
153 tropospheric temperature above Dome C. Measuring at **both** 60 GHz (oxygen molecule line
154 (O₂) to deduce the temperature) and 183 GHz (H₂O line), this unique, state-of-the-art
155 radiometer was installed on site for the first time in January 2009 (Ricaud et al., 2010a **and b**).

156 The measurements of the HAMSTRAD radiometer allow the retrieval of the vertical profiles
157 of H₂O and temperature from the ground to **10-km** altitude with vertical resolutions of 30 to 50
158 m in the **PBL**, 100 m in the free troposphere and 500 m in the upper troposphere-lower

159 stratosphere. The time resolution is adjustable and fixed at 60 seconds during the YOPP
160 campaign. Note that an automated internal calibration is performed every 12 atmospheric
161 observations and lasts about 4 minutes. Consequently, the atmospheric time sampling is 60
162 seconds for a sequence of 12 atmospheric measurements and a new atmospheric sequence is
163 performed after 4 minutes. The temporal resolution on the instrument allows for detection and
164 analysis of atmospheric processes such as the diurnal evolution of the PBL (Ricaud et al., 2012)
165 and the presence of clouds and diamond dust (Ricaud et al., 2017). In addition, two other
166 parameters can be estimated.

167 1) The Integrated Water Vapour (IWV) or precipitable water (kg m^{-2}) obtained by
168 integrating the absolute humidity profile from the surface to 10 km altitude.

169 2) The Liquid Water Path (g m^{-2}) that gives the amount of liquid water integrated along the
170 vertical.

171 IWV has been validated against radiosondes at Dome C between 2010 and 2014 showing a
172 5-10% wet bias of HAMSTRAD compared to the sondes (Ricaud et al., 2015) that were
173 uncorrected for sensor heating or time lag effect that may produce a 4% dry bias (Miloshevich
174 et al., 2006). The $1\text{-}\sigma$ RMS error on the 7-min integration time IWV is 0.05 kg m^{-2} or $\sim 5\%$
175 (Ricaud et al., 2013).

176 The HAMSTRAD-observed LWP has only been presented when the instrument was
177 installed at the Pic du Midi station (2877 amsl, France) during the calibration/validation period
178 in 2008 prior to its set up in Antarctica in 2009 (Ricaud et al., 2010a). Because the instrument
179 has been designed and developed for measuring water vapour in very dry and cold environments
180 such as those encountered at the Dome C station all year long, the radiometer functionality is
181 better adapted for the Dome C site than for the Pic du Midi site. It has not been possible to
182 validate LWP observations at the Pic du Midi station. The H₂O-DC project has thus provided a

183 unique opportunity to perform such a qualitative validation against LIDAR observations of
184 SLW.

185

186 **2.2. The tropospheric depolarization LIDAR**

187 A tropospheric depolarization LIDAR (532 nm) has been operating at Dome C since 2008
188 (see http://lidarmax.altervista.org/englidar/_Antarctic%20LIDAR.php). The LIDAR provides
189 5-min tropospheric profiles of aerosols and clouds continuously, from 20 to 7000 m above
190 ground level (agl), with a resolution of 7.5 m. LIDAR depolarization (Mishchenko et al., 2000)
191 is a robust indicator of non-spherical shape for randomly oriented cloud particles. A
192 depolarization ratio below 10% is characteristic of SLW clouds, while higher values are
193 produced by ice particles. The possible ambiguity between SLW clouds and oriented ice plates
194 is avoided at Dome C by operating the LIDAR 4° off-zenith (Hogan and Illingworth, 2003).
195 The LIDAR observations at Dome C have already been used to study the radiative properties
196 of water vapour and clouds in the far infrared (Palchetti et al., 2015). As a support to LIDAR
197 data interpretation, time-lapse webcam videos of local sky conditions are also collected.

198

199 **2.3. The BSRN Network**

200 The BSRN sensors at Dome C are mounted at the Astroconcordia/Albedo-Rack sites, with
201 upward and downward looking, heated and ventilated standard Kipp&Zonen CM22
202 pyranometers and CG4 pyrgeometers providing measurements of hemispheric downward and
203 upward broadband shortwave (SW, 0.3–3 μm) and longwave (LW, 4–50 μm) fluxes at the
204 surface, respectively. These data are used to retrieve values of net surface radiation (defined as
205 the difference between the downward and upward fluxes). All these measurements follow the
206 rules of acquisition, quality check and quality control of the BSRN as described in Driemel et
207 al. (2018).

208

209 **2.4. Radiosondes**

210 Vertical temperature and humidity profiles have been measured on a daily basis at Dome C
211 since 2005, employing RS92 Vaisala radiosondes. The radiosonde data were taken using the
212 standard Vaisala evaluation routines without any correction of sensor heating or time lag effect.
213 The sondes are known to have a cold bias of 1.2 K from the ground to about 4 km altitude
214 (Tomasi et al., 2011 and 2012) and a dry bias of 4% on IWV (Miloshevish et al., 2006), mainly
215 between 630 and 470 hPa, with a correction factor for humidity varying within 1.10–1.15 for
216 daytime (Miloshevish et al., 2009). During YOPP and the two case studies, launches were
217 performed twice per day at 00:00 and 12:00 UTC.

218

219 **2.5. CALIOP on board CALIPSO**

220 Orbiting at 705-km altitude, the CALIPSO (Cloud Aerosol Lidar and Infrared Pathfinder
221 Satellite Observations) mini-satellite has been observing clouds and aerosols since 2006 to
222 better understand the role of clouds and aerosols in climate. To accomplish this mission, the
223 CALIPSO satellite is equipped with a LIDAR, a camera and an infrared imager (Winker et al.,
224 2009). CALIOP (Cloud-Aerosol Lidar with Orthogonal Polarization) is a dual-wavelength (532
225 and 1064 nm) backscatter LIDAR. It provides high-resolution vertical profiles of clouds and
226 aerosols along the orbit track (Young et al., 2009). We have used version V3.40 data retrieved
227 from <https://www-calipso.larc.nasa.gov/>.

228

229 **2.6. The ARPEGE-SH Model**

230 A special Antarctic configuration of the operational global model ARPEGE was used for
231 the YOPP SOP-SH period (16/11/2018–15/02/2019). This configuration named ARPEGE-SH
232 is based on the operational global model used for Numerical Weather Prediction (NWP)

233 ARPEGE (Pailleux et al., 2015), but with its highest horizontal resolution centred over Dome
234 C instead of over France, as set up in ARPEGE. A 4D variational (4DVar) assimilation was
235 performed every 6 h. The meteorological analyses were given by the ARPEGE-SH system
236 together with the 24-hour forecasts at the node the closest to the location of Dome C. Two
237 analyses at 00:00 and 12:00 UTC were represented in the present study together with hourly
238 forecasts initialized by the two analyses from 01:00 to 11:00 and from 13:00 to 24:00 UTC,
239 respectively. The horizontal resolution during the SOP-SH period was 7.5 km at Dome C. The
240 vertical resolution during the SOP-SH period was constituted by 105 vertical levels, the first
241 one being set at 10 m, with 12 levels below 1 km and 35 levels below 3 km. Several ARPEGE-
242 SH output parameters were selected for analysis: cloud fraction, ice, water vapour and liquid-
243 water mixing ratio, temperature, Total Column Ice (TCI, ice integrated along the vertical),
244 LWP, IWV, and net surface radiation. For each of the model vertical level, the value of the
245 cloud fraction ranges between 0 and 1 and is defined as the fraction of the cloud within the
246 model horizontal grid box. The total cloud fraction at each level is a combination between the
247 resolved cloud, the cloud from the shallow convection and the cloud from the deep convection.
248 The resolved cloud is based on a pdf function with critical relative humidity profile. The shallow
249 convection cloud (below 4000 m) is based on the cloud water/ice tendencies computed by the
250 shallow mass flux scheme with a maximum value at 0.3. For the deep convection, the cloudiness
251 is computed with the vertical divergence of the precipitation flux. The diurnal variation of the
252 top of the PBL is calculated by ARPEGE-SH as the level where the turbulence kinetic energy
253 becomes lower than $0.01 \text{ m}^2 \text{ s}^{-2}$.

254

255 **2.7. The NCEP temperature fields**

256 In order to assess the synoptic state of the atmosphere during the two case studies above
257 Dome C against the climatological state of the atmosphere in summer over Antarctica, we have

258 used the temperature fields at 600 hPa from the National Centers for Environmental Prediction
259 (NCEP) from 2009 to 2019 (Kanamitsu et al., 2002). These are NCEP-Department of Energy
260 (NCEP/DOE) Atmospheric Model Intercomparison Project (AMIP-II) Reanalysis (Reanalysis-
261 2) 6-hourly air temperature at 2.5°x2.5° horizontal resolution over the globe.

262

263 **2.8. The HYSPLIT back-trajectories**

264 In order to assess the origin of airmasses associated to the two case studies, ten-day back-
265 trajectories originated from the Dome C station at 500 and 1000 m above ground level have
266 been calculated on 20 and 24 December 2018 at 12:00 UTC from the Hybrid Single-Particle
267 Lagrangian Integrated Trajectory model (HYSPLIT) model (Stein et al., 2015; Rolph et al.,
268 2017) (<https://www.ready.noaa.gov/HYSPLIT.php>).

269

270 **3. Methodology**

271 In this article, we present two case studies from the SOP-SH that illustrate the occurrence
272 of low-level SLW clouds at Dome C. Both cases occurred in December 2018, within 5 days of
273 each other, which allows direct comparison between the cases without concerns for seasonal
274 variations in radiation.

275 The first case study presented was on 24 December 2018 and was representative of a
276 climatological summer atmosphere in contrast to the second case study (20 December 2018)
277 when the atmosphere was very different from a climatological summer atmosphere. We have
278 considered in Figure 1 the temperature fields from the NCEP at 600 hPa to highlight the state
279 of the atmosphere above Antarctica with a focus over the Dome C station at different periods:
280 a) decadal average over December-January from 2009 to 2019, b) YOPP average over
281 December 2018-January 2019, c) daily average over 24 December 2018, d) 20 December 2018
282 at 00:00 UTC, e) 20 December 2018 at 12:00 UTC, and f) 21 December 2018 at 00:00 UTC.

283 The climatological summer temperature field at 600 hPa has been calculated by averaging the
284 December and January data from 2009 to 2019 and the mean synoptic state of the YOPP
285 campaign during the summer 2018-2019 has been calculated by averaging data from early
286 December 2018 to end of January 2019. The synoptic state of the first case study was selected
287 on 24 December 2018 averaged from 00:00 to 24:00 UTC and for the second case study on 20
288 December 2018 at 00:00 UTC and 12:00 UTC, and on 21 December 2018 at 00:00 UTC. Firstly,
289 the summer atmosphere during YOPP was very consistent with the decadal climatological state
290 of the atmosphere both over Antarctica and the Dome C station (temperature less than 245 K).
291 Secondly, the synoptic state of the atmosphere on 24 December 2018 (1st case study), although
292 warmer (> 258 K) over some parts of the Antarctic Plateau ($60^{\circ}\text{E}-90^{\circ}\text{E}$) is, over Dome C,
293 consistent with the YOPP summer synoptic state and the climatological summer temperatures
294 of ~ 246 K. Thirdly, on 20 December 2018 (2nd case study), on tongue of warm air (254-260 K)
295 originated from the oceanic coast in the sector $0-30^{\circ}\text{W}$ (00:00 UTC) reaches Dome C 24 hours
296 later with temperatures increasing from 252 to 256 K, about 10 K greater than on 24 December
297 2018. Ten-day back trajectories calculated from HYSPLIT (see Figure Supp1) initiated at
298 Dome at 500 and 1000 m above ground level remain over the Antarctic Plateau on 24 December
299 2018 (1st case study) whereas are originated to the oceanic coast in the sector $0-30^{\circ}\text{W}$ on 20
300 December 2018 (2nd case study). This is consistent with previous studies (Ricaud et al., 2017)
301 showing that inland-originated air masses bring cold and dry air to Dome C whilst ocean-
302 originated air masses bring warm and wet air to Dome C.

303 In the following, we will label the 1st case study on 24 December 2018 as typical case and
304 the 2nd case study as perturbed case. We will show that, in the typical case, the SLW cloud
305 occurred over a 24-hour period that was characterized by a typical summertime, diurnal PBL
306 cycle, where the mixed-layer develops over the course of the day, reaches a quite stable height
307 and then collapses to the surface toward the end of the day, around 12 UTC (Ricaud et al.,

308 2012). The first case provides insight into the impact of SLW clouds on the local radiative
309 fluxes. The perturbed case provides a contrasting situation where the diurnal cycle of the PBL
310 was perturbed by the sudden arrival of very moist and warm air of oceanic origin (see Ricaud
311 et al., 2017). We analyse how this episode affected the presence and evolution of SLW clouds
312 and their influence on the surface energy budget. Note that, in the remaining of the article, the
313 data will be presented according to their height above ground level (agl) unless explicitly shown
314 as above mean sea level (amsl).

315

316 **4. Typical diurnal cycle of the PBL**

317 The first case study occurred on 24 December 2018 during a typical diurnal PBL cycle.
318 All the results are presented in Universal Time Coordinated (UTC) with local time (LT) being
319 eight hours ahead of UTC (LT = UTC + 8 hr). As described in Ricaud et al. (2012), the typical
320 summer boundary layer at Dome C is very similar to that described by Stull (1988). Although
321 sunlight is present throughout the day, the variation in magnitude is enough to allow a stable
322 boundary layer from 18:00 to 06:00 LST, similar to a stable nocturnal boundary layer. There
323 is then a transition from a stable boundary layer to a mixed layer around 06:00 LST with the
324 increase in the solar irradiation, which reaches a maximum around solar noon. Then around
325 18:00 LST, the stable boundary layer starts to form again, with a quasi-mixed layer about it.
326 The height of the summertime boundary layer at Dome C typically ranges between 100 and 400
327 m. The presence of SLW clouds at the top of the PBL together with the diurnal evolution of the
328 PBL will be discussed in more detail in the section 7.2.

329

330 **4.1. Clouds**

331 The presence of clouds is highlighted by the LIDAR backscatter and depolarization profiles
332 shown in Figures 2a and b, respectively. High values of LIDAR backscatter ($\beta > 100 \beta_{\text{mol}}$, with

333 β_{mol} the molecular backscatter) indicate that clouds and/or precipitation are present
334 intermittently throughout the day with some significant differences. First, vertical “stripes” of high
335 backscatter values are visible from 10 to 400 m height before 10:00 UTC and after 19:00 UTC,
336 associated with high values of depolarization ratio ($> 20\%$), characteristic of precipitating ice
337 crystals. Second, high values of β associated with very low depolarization ratio ($< 5\%$) occur
338 within a thin layer of approximately 100-m depth around 500 m from 08:00 to 22:00 UTC, with
339 some breaks around 11:00 and 19:00-21:00 UTC. From the LIDAR observations, this
340 combination of high backscatter and low depolarization ratio signifies the presence of a SLW
341 cloud (Figure 2c).

342 The NWP model ARPEGE-SH calculates cloud fraction, ice water and liquid water mixing
343 ratios (kg kg^{-1}) for 24 December 2018 (Figures 3a, b and c, respectively). We note that the
344 outputs from ARPEGE-SH at 00:00 and 12:00 UTC are the analyses and, for the remaining
345 time, the outputs are the hourly forecasts. ARPEGE-SH predicts the presence of clouds (cloud
346 fraction > 0.95) for most of the day except around 11:00 and 23:00 UTC (Fig. 3a). Before 12:00
347 UTC, the cloud is mainly confined between 300 and 600-800 m whilst, after 12:00 UTC, it
348 spreads from the surface to 800 m. There are also high-level clouds at 2000-3000 m height but
349 with a cloud fraction between 0.50 and 0.70. The majority of the clouds produced by ARPEGE-
350 SH are mainly composed of ice crystals (Fig. 3b) with some traces of droplets (Fig. 3c) due to
351 the model’s partitioning between ice and liquid where all condensed water is ice below -20°C .
352 The liquid water clouds derived from the LIDAR observations are superposed over the SLW
353 clouds calculated by ARPEGE-SH. The modelled values of liquid water ($\sim 4 \cdot 10^{-6} \text{ g m}^{-3}$) are very
354 low, far lower than the values of 0.1 g m^{-3} observed for coastal polar stratus clouds (see e.g.
355 O’Shea et al., 2017; Lachlan-Cope et al., 2016; Young et al., 2016). It is evident that ARPEGE-
356 SH fails in estimating: 1) the vertical distribution of liquid water (a thin layer is observed around
357 500 m whereas the modelled cloud layer extends from the surface to 800 m); 2) its temporal

358 evolution (presence of SLW cloud almost all day long in ARPEGE-SH compared to SLW
359 clouds from 08:00 to 22:00 UTC in the observations); and 3) the liquid vs. ice mixing ratio, the
360 former being in the model several orders of magnitudes lower than the latter, in contrast to the
361 observations.

362 The presence of clouds above the station can also be inferred from vertically-integrated
363 variables such as: 1) TCI calculated by ARPEGE-SH, 2) LWP from HAMSTRAD and
364 ARPEGE-SH, and 3) IWV from HAMSTRAD and ARPEGE-SH (Figures 4a, b and c,
365 respectively). The ARPEGE-SH TCI on 24 December 2018 (Fig. 4a) oscillates between 10 and
366 30 g m^{-2} except around 12:00 UTC when a clear minimum occurs ($\sim 3 \text{ g m}^{-2}$), underscoring the
367 fact that ARPEGE-SH obtains ice clouds for the entire day, except at 12:00 UTC. The
368 HAMSTRAD LWP shows an obvious increase from ~ 1.0 to $\sim 2.0\text{-}3.0 \text{ g m}^{-2}$ when the presence
369 of SLW cloud is indicated by LIDAR observations (Fig. 4b). The ARPEGE-SH LWP is, on
370 average, 10^3 times lower than that observed by HAMSTRAD, highlighting the fact that
371 ARPEGE-SH misrepresents features of the SLW clouds over Dome C. The $1\text{-}\sigma$ RMS error on
372 the 1-min integration time for the HAMSTRAD LWP can be estimated to be $\sim 15\%$. Based on
373 the comparisons between the HAMSTRAD LWP and the LIDAR observations of SLW clouds
374 during the YOPP campaign, we can estimate that the LWP bias is about 1.0 g m^{-2} . We cannot
375 rule out that these biases might also be related in part to differences in the observation
376 wavelengths employed (submicrons for the LIDAR and microwaves for HAMSTRAD) that
377 could favour large particles (HAMSTRAD) against small particles (LIDAR). Biases might also
378 be due to the observing geometry that differs between the LIDAR (close to zenith viewing) and
379 HAMSTRAD (atmospheric scans at 10 angles from zenith to $\sim 3^\circ$ elevation). HAMSTRAD and
380 ARPEGE-SH IWV (Fig. 4c) vary from $0.65\text{-}1.05 \text{ kg m}^{-2}$ throughout the day on 24 December
381 2018, with an agreement to within 0.1 kg m^{-2} (i.e. $\sim 10\text{-}15\%$), which is consistent with previous
382 studies (Ricaud et al., 2017).

383 Observation of clouds from space-borne sensors has two main advantages: 1) it
384 complements the ground-based cloud observations at Dome C (namely ice/liquid water), and
385 2) it provides an estimate of the vertical and horizontal extents of the detected cloudy layers.
386 Note that the CALIPSO spaceborne LIDAR operates at the same wavelength as the backscatter
387 LIDAR at Dome C, with the same method for discriminating ice from liquid water.
388 Consequently, the two LIDARs should give consistent information for the detected cloud phase.
389 However, the presence of an optically thick cloud may extinguish the CALIOP signal
390 underneath as was already presented in Ricaud et al. (2017) when studying episodes of thick
391 (5-km deep) clouds and diamond dust (ice crystals in suspension close to the surface). The main
392 difficulty with this approach is related to the temporal and spatial sampling of the spaceborne
393 instrument, namely finding a satellite overpass coincident both in time and location with the
394 cloud observed at Dome C. This, unfortunately, decreases the number of overpasses that is
395 scientifically exploitable. Nevertheless, on 24 December 2018, 2 orbits of CALIOP/CALIPSO
396 passed close to Dome C at times when SLW clouds were observed by ground-based
397 instruments. We show the vertical feature mask and ice/water phase from the pass closest to the
398 station (~220 km), from 15:50 to 16:03 UTC (Figures 5a and b, respectively). Firstly, we note
399 the presence of a cloud a few hundred meters deep near the surface in the vicinity of Dome C
400 (Fig. 5a; note that the CALIOP/CALIPSO altitude is above sea level and Dome C is at an
401 altitude of 3233 m amsl). Secondly, this cloud is composed of SLW (Fig. 5b), confirming the
402 analysis based on the observations from the LIDAR and the HAMSTRAD radiometer.
403 Furthermore, we can state that this SLW cloud is not a local phenomenon but has a horizontal
404 extent of ~450 km along the orbit track. Considering the CALIOP total and perpendicular
405 attenuated backscatter data at 532 nm on 24 December 2018 at 16:00 and 14:00 UTC (Figures
406 Supp2 and Supp3, respectively), we note that: 1) the SLW cloud is located between 3.7 and 3.8
407 km amsl, that is to say a height from ~450 to ~550 m agl, and 2) since the CALIOP signal is

408 able to reach the surface underneath the SLW cloud, ice is not detected by the space-borne
409 instrument. This is consistent with the observations performed at Dome C. The other orbit from
410 14:11 to 14:25 UTC (Figure Supp4) is slightly more distant than the one shown in Figure 5
411 (~360 km), but it exhibits a similar SLW cloud located between ~450 and ~550 m agl, over an
412 even greater horizontal extent of ~700 km along the orbit track.

413

414 **4.2. Vertical profiles of temperature and water vapour**

415 On 24 December 2018, temperatures from both HAMSTRAD and ARPEGE-SH ranged
416 from 240 to 250 K (-33 to -23°C) from the surface to 1-km agl, compatible with the presence
417 of SLW clouds. The diurnal variations of temperature and water vapour anomalies calculated
418 by ARPEGE-SH and measured by HAMSTRAD are shown in Figure 6. For each height, the
419 daily-averaged value has been subtracted. This has the advantages of highlighting areas of
420 maximum and minimum changes along the vertical, and reduces biases when comparing the
421 two data sets. Absolute anomalies (K) are presented for temperatures whilst relative anomalies
422 (%) are shown for water vapour.

423 The diurnal variation of the ARPEGE-SH temperature (Fig. 6a) from the surface to 1 km
424 shows a warm atmosphere before 12:00 UTC and a fast cooling one afterward. HAMSTRAD
425 shows a similar cooling (Fig. 6b), but the transition is not so abrupt and occurs later, around
426 15:00 UTC. The diurnal amplitude is greater in ARPEGE-SH (~5 K) than in HAMSTRAD (~3
427 K). The diurnal variation of the water vapour in ARPEGE-SH (Fig. 6c) from the surface to 1
428 km shows a wet atmosphere before 12:00 UTC and a drier atmosphere after, again with an
429 abrupt transition. From HAMSTRAD, the diurnal variation of the water vapour (Fig. 6d) from
430 the surface to 1 km is more complex, alternating wet and dry phases, which is particularly
431 obvious at 500-m altitude: wet (00:00-03:00 UTC), dry (03:00-08:00 UTC), wet (08:00-09:00
432 UTC), dry (09:00-12:00 UTC), wet (12:00-22:00 UTC) and dry (22:00-24:00 UTC). The time

433 evolution of the SLW cloud (Fig. 2c) and the diurnal variation of the top of the PBL as
434 calculated by ARPEGE-SH are superposed on all the panels of Figure 6. We note that the SLW
435 cloud appeared just below the ARPEGE-SH-estimated PBL top, around 08:00 UTC, and
436 persisted around the same altitude after 12:00 UTC even though the top of the PBL had
437 dramatically decreased down to the surface. In addition, the SLW cloud persisted after 12:00
438 UTC in a layer that is cooler than earlier in the day, but slightly warmer than the air above and
439 below it. However, the model shows that this layer is drier while the observations suggest it is
440 wetter.

441

442 4.3. Potential Temperature Gradient

443 We now consider the mechanisms that allow the SLW cloud to persist in a thin layer (about
444 100-m deep) around 500-600 m altitude. Even if the PBL gets thinner after 12:00 UTC, a
445 residual mixed layer remains above (see e.g. Figure 1.7 of Stull, 2012; Figure 12 top of Ricaud
446 et al., 2012 and definition of a residual layer from the American Meteorological Society at
447 http://glossary.ametsoc.org/wiki/Residual_layer). This layer, where turbulence is sporadic or
448 even absent, lies above the surface-connected stable layer, and can be viewed as a fossil of the
449 mixed layer developed during the previous mixing period. The transition from the boundary
450 layer to the free atmosphere is characterized by a local maximum of the potential temperature
451 (θ) vertical gradient ($\partial\theta/\partial z$).

452 Figure 7 shows $\partial\theta/\partial z$ computed from ARPEGE-SH, with the evolution of the PBL top and
453 the SLW cloud superposed. Black areas correspond to neutral conditions ($\partial\theta/\partial z \sim 0$), whereas
454 the coloured ones relate to stable stratification according to the colour scale in the Figure. The
455 SLW cloud, once appeared at the top of the PBL around 08:00 UTC, persists after 12:00 UTC
456 in a layer around 500-600 m located just below the local maximum of $\partial\theta/\partial z$, even after the PBL
457 collapses down to the surface.

458 Figure 7 shows $\partial\theta/\partial z$ field and the evolution of the mixed layer top, both computed from
459 ARPEGE-SH output – the latter defined according to whether the turbulent kinetic energy
460 exceeds a defined threshold – and the observed SLW cloud superposed. Black areas correspond
461 to neutral conditions ($\partial\theta/\partial z \sim 0$), whereas the coloured ones relate to stable stratification
462 according to the colour scale in the Figure. The SLW cloud, once appeared at the top of the
463 PBL around 08:00 UTC, persists after 12:00 UTC in a layer around 500-600 m coinciding with
464 the top of the residual mixed layer (see above for the definition) even after the ARPEGE-
465 defined mixed layer top collapses down to the surface.

466 Figures 8a, b and c show the vertical profiles of θ (K) and $\partial\theta/\partial z$ (K km⁻¹) as calculated
467 from temperature measured by the radiosondes and analysed by ARPEGE-SH at Dome C on
468 24 December 2018 at 00:00 and 12:00 UTC and on 25 December 2018 at 00:00 UTC,
469 respectively. The presence and the depth of the SLW cloud detected from LIDAR observations
470 are highlighted in the Figure. The atmosphere as analysed by ARPEGE-SH is about 3-5 K
471 warmer than the observations. From 100 m upward, the maximum of $\partial\theta/\partial z$ is measured at 400,
472 550 and 600 m on 24 December 2018 at 00:00 and 12:00 UTC and on 25 December 2018 at
473 00:00 UTC, respectively with an amplitude of 10, 12 and 40 K km⁻¹, respectively. ARPEGE-
474 SH cannot reproduce the fine vertical structure of $\partial\theta/\partial z$. For example, the simulated maxima
475 of $\partial\theta/\partial z$ (Fig. 8) are slightly higher (600, 700 and 600 m for the same dates, respectively) and
476 less intense than those of radiosondes (8, 8 and 18 K km⁻¹, respectively).

477

478 **5. Perturbed diurnal cycle of the PBL**

479 On the second case study, 20 December 2018, the diurnal cycle of the PBL was perturbed
480 by the sudden arrival of very moist, warm air of oceanic origin. During this warming period,
481 the boundary layer remains mixed and does not form a stable boundary layer even when the
482 solar forcing decreases. This will be discussed in detail in the section 7.2.

483

484 **5.1. Clouds**

485 As in section 3.1, the high LIDAR backscatter ($\beta > 100 \beta_{\text{mol}}$) and low depolarization
486 (<5%) showed the presence of SLW clouds (Figures 9a, b and c, respectively). Before 13:00
487 UTC, there is no trace of clouds above Dome C, while from 13:00 to 23:00 UTC SLW clouds
488 are detected between 200 and 600 m. On all panels, we superimposed the PBL top calculated
489 by the ARPEGE-SH model. We note that the PBL top does not drop to the surface after 12:00
490 UTC as typically occurs, like on 24 December 2018, but rather remains between 100 and 200
491 m. Consistent with the conclusions derived from the observations of 24 December 2018, the
492 SLW cloud, once present, stays just above the height of the PBL top.

493 The cloud fraction, ice water and liquid water mixing ratios (kg kg^{-1}) calculated by
494 ARPEGE-SH on 20 December 2018 are shown in Figures 10a, b and c, respectively. Contrary
495 to the observations, the model simulates mixed-phase clouds (maximum cloud fraction of
496 ~30%), mainly composed of ice, prior to 12:00 UTC; from 00:00 to 06:00 UTC, the clouds are
497 forecasted below the PBL top. After 12:00 UTC, clouds appear 1-2 hours later in the model
498 than in the observations, at 14:00-15:00 UTC, just below the PBL top (maximum cloud fraction
499 of ~100%). The modelled cloud is mainly composed of ice with some traces of SLW above the
500 PBL around 15:00-16:00 UTC. In all occurrences, the liquid water amounts produced by the
501 model are extremely small, nearly non-existent. We note the presence of high altitude cirrus
502 (ice) clouds calculated by ARPEGE-SH after 12:00 UTC around 3-4 km height, while not
503 observed likely because the LIDAR light is attenuated by the SLW layer. As on 24 December
504 2018, the model fails to reproduce the presence of the SLW layer observed by the LIDAR near
505 the PBL top.

506 The diurnal evolutions of the TCI calculated by ARPEGE-SH, the LWP from
507 HAMSTRAD and ARPEGE-SH, and the IWV from HAMSTRAD and ARPEGE-SH on 20

508 December 2018 are presented in Figures 11a, b and c, respectively, with the presence of SLW
509 clouds derived from the LIDAR observations superimposed on Fig. 11b. Ice clouds are
510 calculated by ARPEGE-SH mainly around 15:00-16:00 UTC, with TCI values comparable to
511 those on 24 December 2018. SLW clouds are deduced from HAMSTRAD LWP between 13:00
512 and 23:00 UTC which coincides well with the SLW clouds observed by the LIDAR. The
513 maximum LWP values observed during this episode are much higher ($\sim 50 \text{ g m}^{-2}$) than on 24
514 December 2018 ($\sim 2\text{-}3 \text{ g m}^{-2}$). Again, the ARPEGE-SH LWP is negligible ($\sim 10^3$ times less than
515 observations). In parallel with the rapid increase of LWP, the observed IWV also jumps from
516 ~ 0.5 to $\sim 2.3 \text{ kg m}^{-2}$ within one hour after 13:00 UTC. ARPEGE-SH also calculates an increase
517 of IWV but lagged by one hour and much less intense ($\sim 1.3 \text{ kg m}^{-2}$). Additionally, the model
518 produces a systematically dryer atmosphere compared to HAMSTRAD by about 0.5 kg m^{-2}
519 after 16:00 UTC, although before the cloudy period that starts at 12:00 UTC, ARPEGE-SH and
520 HAMSTRAD IWV are consistent to within $\pm 0.2 \text{ kg m}^{-2}$.

521 On 20 December 2018, after 13:00 UTC when SLW clouds have been detected at Dome
522 C, both CALIPSO overpasses are far away from Dome C and, for the closest overpass at 13:17
523 UTC (closest distance to Dome C is 500 km), a very thick ice cloud at about 3 km agl prevents
524 the LIDAR radiation from reaching the surface (Figure Supp5). Unfortunately, no meaningful
525 information can be ascertained from the spaceborne observations on that day relevant to SLW
526 clouds in the vicinity of Dome C.

527

528 **5.2. Vertical profiles of temperature and water vapour**

529 The diurnal variations of the temperature and water vapour anomalies on 20 December
530 2018 as calculated by ARPEGE-SH and measured by HAMSTRAD are shown in Figure 12. In
531 ARPEGE-SH, a sharp transition between a warm and a cool atmosphere is evident at 12:00
532 UTC below the top of the PBL. In HAMSTRAD, from 00:00 to 06:00 UTC, the atmosphere

533 starts warming and then from 06:00 to 13:00 UTC, cools gradually to a minimum. After 13:00
534 UTC, HAMSTRAD temperatures reveal a warming starting from the surface and progressively
535 thickening until reaching the top of the PBL by the end of the day. Above the PBL, the
536 HAMSTRAD-observed and ARPEGE-SH-calculated temporal evolution of temperature and
537 water vapour are in an overall agreement. In the PBL, the model simulates a moistening around
538 05:00 UTC, but the most striking event is a sudden drying at 12:00 UTC. In HAMSTRAD,
539 there is a continuous drying from 00:00 UTC, followed by an obvious transition at 13:00 UTC,
540 opposite to that of ARPEGE-SH at 12:00 UTC. The warm and wet atmosphere observed after
541 13:00 UTC develops a mixed layer, consequently the PBL top no longer collapses to a stable
542 layer, in contrast to what was observed on 24 December. Furthermore, the SLW clouds present
543 in the entrainment zone steadily remain at the PBL top until the end of the day.

544

545 **5.3. Potential Temperature Gradient**

546 **Figure 13** shows $\partial\theta/\partial z$ (K km^{-1}) from ARPEGE-SH, with the evolution of the PBL top and
547 the SLW cloud superimposed. In these perturbed conditions, the SLW clouds are present a few
548 tens of meters above the top of the PBL after 12:00 UTC. The PBL top is located in a layer
549 coinciding with the local maximum of $\partial\theta/\partial z$, around 100-300 m, and does not dramatically
550 decrease to the surface for the rest of the day.

551 **Figures 14a, b and c** show the vertical profiles of θ (K) and $\partial\theta/\partial z$ (K km^{-1}) as calculated
552 from temperature measured by the radiosondes and analysed by ARPEGE-SH at Dome C on
553 20 December 2018 at 00:00 and 12:00 UTC and on 21 December 2018 at 00:00 UTC,
554 respectively. **The presence and the depth of the SLW cloud detected from LIDAR observations**
555 **are highlighted in the Figure.** The ARPEGE-SH profiles are about 0-5 K warmer than the
556 observations. From 50 m upward, the maximum of $\partial\theta/\partial z$ is measured at 75, 150 and 375 m on
557 20 December 2018 at 00:00 and 12:00 UTC and on 21 December 2018 at 00:00 UTC,

558 respectively, with a corresponding amplitude of 75, 40 and 55 K km⁻¹. The location of the
559 observed maximum in the potential temperature gradient is consistent with the ARPEGE-SH
560 calculations on 20 December 2018 prior to the warm and wet episode: at 00:00 UTC (Fig. 14a),
561 the calculated $\partial\theta/\partial z$ is maximum at 75 m and reaches 100 K km⁻¹. However, at 12:00 UTC (Fig.
562 14b) the modelled $\partial\theta/\partial z$ peaks at 200 m (slightly higher than observed) with a value of 50 K
563 km⁻¹. On the following day at 00:00 UTC (Fig. 14c), $\partial\theta/\partial z$ calculated by ARPEGE-SH shows
564 two maxima at 100 and 450 m with an amplitude of 45 and 25 K km⁻¹, respectively, while the
565 observations demonstrate a single maximum just below 400 m.

566

567 **6. Impact of SLW Clouds on Net Surface Radiation**

568 The presence of clouds over Dome C has a strong impact on the net surface radiation as
569 demonstrated by Ricaud et al. (2017). This can be seen clearly in the time-series of upwelling
570 and downwelling longwave and shortwave fluxes observed by BSRN for the two case studies.

571

572 **6.1 Typical PBL Case – 24 December 2018**

573 Figure 15 (top) shows the time evolution of the net surface radiation as measured by the
574 BSRN instruments and as calculated by ARPEGE-SH on 24 December 2018, superimposed
575 with SLW cloud height. We also show the time evolution of the difference between surface
576 radiation (W m⁻²) observed by BSRN and calculated by ARPEGE-SH on 24 December 2018,
577 in longwave downward (LW↓), longwave upward (LW↑), shortwave downward (SW↓) and
578 shortwave upward (SW↑) components, superimposed with LWP. We highlight 4 periods with
579 images taken from the webcam installed on the shelter hosting the LIDAR and HAMSTRAD:
580 a) at 00:25 UTC (cirrus clouds, no SLW cloud), b) at 03:56 UTC (cirrus clouds, no SLW cloud),
581 c) at 09:46 UTC (SLW cloud) and d) at 17:20 UTC (SLW cloud). The net surface radiation
582 shows maxima between 00:00 and 05:00 UTC (08:00-13:00 LT) and minima between 11:00

583 and 13:00 UTC (19:00-21:00 LT) in the ARPEGE-SH and BSRN time series. When SLW
584 clouds are present in the observations (08:00-10:00, 12:00-19:00 and around 21:00 UTC),
585 whilst absent in ARPEGE-SH, the measured net surface radiation is systematically greater than
586 the simulated one by 20-30 W m^{-2} . In the presence of SLW clouds after 12:00 UTC, this
587 difference is mainly attributable to $\text{LW}\downarrow$ component, BSRN values being 50 W m^{-2} greater than
588 those of ARPEGE-SH. Thus, SLW clouds tend to radiate more LW radiation toward the ground
589 (like greenhouse gases) than more transparent clouds, like cirrus, do. There are differences from
590 -30 to +60 W m^{-2} between observed and calculated $\text{SW}\downarrow$ and $\text{SW}\uparrow$ components but this
591 difference falls within $\pm 10 \text{ W m}^{-2}$ for the net SW surface radiation ($\text{SW}\downarrow - \text{SW}\uparrow$). The reflective
592 impact of SLW layers can also be seen after 12:00 UTC: unlike observed SLW clouds,
593 ARPEGE-SH simulates ice clouds, and therefore too high $\text{SW}\downarrow$ values. The difference between
594 observed and simulated values of this parameter thus increases, as can be seen on the Figure.
595 But because of the high values in surface albedo, a compensating effect occurs on the surface
596 reflected SW fluxes, and the resulting impact on net radiation is quite weak (the time series of
597 the observed – simulated difference in incoming and reflected SW flux follow each other quite
598 well). The major impact on net radiation is therefore related to the longwave fluxes.

599

600 **6.2 Perturbed PBL Case – 20 December 2018**

601 **Figure 16 (top)** shows the net surface radiation as measured by the BSRN photometric
602 instruments and as calculated by ARPEGE-SH for 20 December 2018, superimposed with the
603 SLW clouds. We also show the time evolution of difference in surface radiation (W m^{-2})
604 observed by BSRN and calculated by ARPEGE-SH on 20 December 2018 for $\text{LW}\downarrow$, $\text{LW}\uparrow$,
605 $\text{SW}\downarrow$ and $\text{SW}\uparrow$ components, superimposed with LWP. We highlight 4 periods with snapshots
606 taken from the webcam: 1) 07:15 UTC (clear sky), 2) 12:35 UTC (clear sky), 3) 13:30 UTC
607 (SLW cloud) and 4) 21:00 UTC (SLW cloud). Before 13:00 UTC, there are no clouds above

608 Dome C whilst after 13:00 UTC clouds are present. The diurnal evolution of the modelled and
609 observed net surface radiation shows a maximum of $\sim+50 \text{ W m}^{-2}$ in ARPEGE-SH and $\sim+85 \text{ W}$
610 m^{-2} in BSRN over the period 00:00-04:00 UTC, and a minimum of about -50 W m^{-2} around
611 12:00-13:00 UTC on both time series. Nevertheless, when SLW clouds are observed at 13:00
612 UTC, the observed net surface radiation jumps to $+10 \text{ W m}^{-2}$, a feature not reproduced in the
613 model. The difference between the BSRN-observed and ARPEGE-SH-modelled net surface
614 radiation is larger than $+30 \text{ W m}^{-2}$ when SLW clouds are present, reaching $+60 \text{ W m}^{-2}$ when the
615 LWP measured by HAMSTRAD is at its maximum (50 g m^{-2} at 13:00 UTC). This is twice the
616 difference observed in the non-perturbed PBL episode detailed in section 3.4. This underlines
617 again the strong impact SLW clouds may have on the radiation budget over Antarctica. In the
618 presence of SLW clouds after 13:00 UTC, the difference in net surface radiation is mainly
619 attributable to $\text{LW}\downarrow$ component, BSRN values being 100 W m^{-2} greater than those of ARPEGE-
620 SH. The $\text{SW}\downarrow$ and $\text{SW}\uparrow$ also decrease due to the high reflectivity of the SLW layer seen at
621 12:00 UTC and again at 15:00 UTC. Note that there are differences from -100 to $+60 \text{ W m}^{-2}$
622 between observed and calculated $\text{SW}\downarrow$ and $\text{SW}\uparrow$ components but this difference falls below 20
623 W m^{-2} for the net SW surface radiation ($\text{SW}\downarrow - \text{SW}\uparrow$). Both SW components decrease after
624 17:00 UTC. Some of this may be due to: 1) increasing LWP, and 2) the presence of precipitating
625 ice crystals and/or blowing snow (characterized by red spots on Figure 9b) that are increasing
626 optical depth and decreasing transmission/visibility (webcam images in Figure 16d) although
627 surface wind was rather weak ($3\text{-}10 \text{ m s}^{-1}$, not shown).

628

629 **7. Discussions**

630 **7.1. SLW Clouds vs Mixed-Phase Clouds**

631 In order to evaluate whether the observed cloud is constituted of liquid and/or mixed phase
632 water, we have considered the raw signals recorded by the LIDAR. For the two dates under

633 consideration (Figures Supp6 and Supp7 relative to 24 and 20 December 2018, respectively),
634 we have represented (top) the P signal as the signal received with the same polarization as the
635 laser (unpolarized component). Any suspended object can contribute to P signal. We have also
636 represented the S (cross-polarized) LIDAR signal (bottom) that is only produced by non-
637 spherical (obviously frozen at Dome C) particles and, to a smaller extent, by multiple scattering
638 in water clouds.

639 First of all, an elevated P signal above ~ 400 m on 24 December 2018 ($P \geq 0.1$ mV) and
640 above ~ 200 m on 20 December 2018 ($P \geq 0.3$ mV) is associated with a cloud as shown in
641 sections 4.1 and 5.1. Inside these clouds, the S signal is always very low: $S \sim 0.003$ mV on 24
642 December 2018 and ~ 0.01 mV on 20 December 2018. Consequently, the S signal is very weak
643 and corresponds to a maximum of $\sim 3\%$ of the corresponding P signal. Some S signal is
644 nevertheless present in the cloud and could be given by multiple scattering inside the truly liquid
645 water cloud and/or the effective presence of ice particles.

646 When considering the LIDAR depolarization diurnal evolutions presented in Figures 2b
647 and 9b associated to the two dates, ice particles could have been disappeared in the low
648 depolarization ratio S/P of the SLW layer because the P signal inside the SLW cloud is very
649 high compared to the S signal. But when considering the P and S signals distinctively (Figs.
650 Supp6 and Supp7), the S signal remains very weak in the SLW cloud compared to the P signal
651 whatever the date considered. Consequently, even if the presence of some ice particles scattered
652 within the SLW layers cannot be excluded from the S signal plot, the very low depolarization
653 of the layers leads to classify them as a liquid cloud.

654 The important point is that the optical properties of the layer, relevant for the radiative
655 budget in the shortwave, such as optical extinction, optical depth, asymmetry factors, etc. are
656 bound to the P signal, being e.g. optical extinction in the visible proportional to the lidar P
657 signal. Thus, the shortwave radiative characteristics of the cloud are driven by the P signal, and

658 thus by liquid water. The layer is thus a truly SLW layer, being that its ice component, even if
659 present, is irrelevant from a radiative point of view.

660

661 **7.2. SLW Clouds and PBL**

662 During the YOPP SOP-SH, SLW clouds were observed in the LIDAR data for 15 days in
663 December (49% of days) and 13 days in January (47%), which is a similar rate of occurrence
664 to other years (53% in December 2016 and 2018; 51% in January 2018 and 2019) (Figure 17).
665 A day is flagged with a SLW cloud occurrence when a SLW cloud has been detected in the
666 LIDAR observations for a period longer than 1 hour. The clouds observed during the SOP-SH
667 are typically located at the top of the PBL (100 to 400 m height) and are 50-100 m thick.

668 The presence of SLW clouds in the atmosphere is strongly dependent on the temperature
669 field. From Fig. 2.33 of Pruppacher and Klett (2012), the percentage of clouds containing no
670 ice becomes non-negligible at temperatures greater than -35°C , although SLW clouds have been
671 observed at lower temperatures over Russia (-36°C) and the Rocky Mountains in the USA ($-$
672 40.7°C). Recent laboratory measurements show that liquid water can exist down to -42.55°C
673 (Goy et al., 2018).

674 Considering that the SLW clouds at Dome C are so thin, they resemble stratocumulus, as
675 can be observed at middle latitudes. The diurnal cycle of the SLW cloud also evokes that of
676 oceanic stratocumulus, with a trend to fragmentation and/or dissipation during the “day” (local
677 noon) because of solar absorption and to a solid deck state during the “night” (local midnight)
678 because of reversed buoyancy due to cloud top longwave cooling. We use here the “night” and
679 “day” terms for convenience, though solar radiation remains positive 24-hr long at this period
680 of the year. During the SOP-SH, SLW clouds were observed in the LIDAR data for
681 approximately 48% of days (Fig. 17) but it is not yet evident whether they were formed during
682 the “day” (local noon) when the mixed layer becomes thick enough to reach the condensation

683 level, and vertically broadened during the “night”, or created during the “night” (local midnight)
684 and then dissipated during the coming “day”. Complementary observations would be needed,
685 in particular turbulence profiles from the surface to above the top of boundary-layer clouds, to
686 determine what is the coupling/decoupling diurnal cycle of these clouds.

687 The diurnal evolution of the top of the PBL is consistent with previous studies carried out
688 at Dome C (e.g. Argentini et al., 2005; King et al., 2006; Ricaud et al., 2012; Casasanta et al.,
689 2014), with a top higher when there is a relatively warm mixed layer than in colder stable
690 conditions.

691 The collocation of the positive potential temperature gradient with the height of the SLW
692 clouds is consistent with the schematic representation of the diurnal variation of the PBL
693 illustrated by Stull (2012) and adapted by Ricaud et al. (2012) for the Eastern Antarctic Plateau.
694 Figure 18 is a modified version of Figure 12 from Ricaud et al. (2012) to take into account the
695 impact of the clouds on the PBL structure. Starting with the simplest, cloud-free case, we have
696 during the convective (mixing) period a mixed layer at the top of which is located the
697 “entrainment zone”, so-named because air parcels coming from the above free troposphere are
698 entrained into the mixed layer below under the effect of overshooting thermals and
699 compensating descending currents. When clouds form at the top of the PBL (boundary-layer
700 clouds), we consider that the PBL locally (i.e. where clouds are present) extends to the top of
701 these clouds. The PBL is clearly separated from the above stable free troposphere by the so-
702 called “capping inversion”. The cloud layers as well as the capping inversion zone are thin, of
703 the order of 100 m. When the stable layer forms close to the surface, the SLW cloud may persist
704 over the residual mixed layer, as may persist the capping inversion zone which can also be
705 qualified as “residual”. The stable layer is then progressively eroded, when the incoming
706 available energy becomes large enough to ensure turbulent mixing from the surface. The new
707 mixing layer thus grows through the previous stable layer and residual mixed layer, up to it

708 reaches the residual capping inversion. The stratification of the different layers is characterized
709 by the simplified potential temperature profiles in Figure 18. Considering both the potential
710 temperature gradients and the vertical extent of the SLW cloud, these layers are quite thin, less
711 than 100-m deep.

712

713 **7.3 SLW Clouds in ARPEGE-SH**

714 In comparison with observations, ARPEGE-SH consistently underestimates LWP by
715 several orders of magnitude. This is due in part to the partitioning into liquid and ice phases in
716 the model which is a simple function of temperature such that, below -20°C , all cloud particles
717 are iced. The inability of ARPEGE-SH to reproduce the observed liquid water content of the
718 cloud leads to an underestimate of the simulated downwelling longwave radiation relative to
719 observations, and an overestimate of both upwelling and downwelling shortwave flux. This
720 effect is particularly notable in the perturbed PBL case study where the high moisture content
721 leads to an enhanced longwave effect. As the SLW cloud horizontal extent in the first case
722 study is about 280 km and persists over more than 12 hours (section 3.1), the discrepancy in the
723 net surface radiation between observation and NWP model may have a strong impact on the
724 calculation of the radiation budget over Antarctica. Lawson and Gettelman (2014) showed that
725 better representation of liquid water in modelled mixed phase clouds in Global Climate Models
726 led to an increase of $7.4 \text{ W}\cdot\text{m}^{-2}$ in the cloud radiative effect over Antarctica.

727 Furthermore, even when considering analyses of ARPEGE-SH at 00:00, 06:00, 12:00 and
728 18:00 UTC and associated forecasts (not shown), neither IVW nor LWP are significantly
729 modified, and SLW remains underestimated. The 4Dvar analysis is not able to correct the dry
730 bias especially during the case of 20 December 2018 probably because it is influenced by a
731 large-scale advection. The underestimation of the SLW in ARPEGE-SH can be explained by
732 the fact that: 1) the underestimation of liquid water is mainly a physical problem in the model

733 related to the ice/liquid partition function vs temperature (see below) and 2), since the cloud
734 water is not a model control variable in the 4DVar scheme, it cannot be analysed.

735 We have thus tried to modify the ice partition function (ice/liquid water vs temperature)
736 used in the ARPEGE-SH operational model (Figure Supp8). We noticed that, for temperatures
737 below -20°C , water was present only in the solid form in the model. A test has been performed
738 for 20 and 24 December 2018 with ARPEGE-SH by considering a new ice partition function
739 allowing the presence of liquid water for temperature between -20°C and -40°C (Figure Supp8).
740 The analyses were done at 00:00 UTC and the forecasts from 01:00 to 24:00 UTC. This run
741 was labelled as ARPEGE-SH-TEST.

742 For 24 December 2018, and consistently with Fig. 3, we have drawn on Fig. Supp9 the
743 diurnal evolutions of different variables calculated by ARPEGE-SH-TEST: a) the Cloud
744 Fraction, b) the Ice Water mixing ratio and c) the Liquid Water mixing ratio. Similarly, and
745 consistently with Fig. 4, Figure Supp11 presents: a) the ARPEGE-SH-TEST TCI, b) the LWP
746 measured by HAMSTRAD and calculated by ARPEGE-SH-TEST and c) the IWV measured
747 by HAMSTRAD and calculated by ARPEGE-SH-TEST. Eventually, and consistently with Fig.
748 9, Figure Supp13 presents the net surface radiation observed by BSRN and calculated by
749 ARPEGE-SH-TEST, and the difference between surface radiation of longwave downward,
750 longwave upward, shortwave downward and shortwave upward components observed by
751 BSRN and calculated by ARPEGE-SH-TEST. In the same manner, for the case of 20 December
752 2018, Figs. Supp10, Supp12 and Supp14 echo Figs. 11, 12 and 16, respectively.

753 On 24 December 2018 (typical case), the new partition function significantly improves the
754 modelled SLW, with liquid water content about 1000 times greater in ARPEGE-SH-TEST than
755 in ARPEGE-SH, and LWP varying from ~ 0 to $\sim 3 \text{ g m}^{-2}$ consistently with HAMSTRAD to
756 within $\pm 0.5 \text{ g m}^{-2}$. The impact on the net surface radiation is obvious with an excellent
757 agreement between ARPEGE-SH-TEST and BSRN to within $\pm 20 \text{ W m}^{-2}$. Unfortunately, on

758 20 December 2018 (perturbed case), even if the impact on SWL clouds is important (liquid
759 water content multiplied by a factor 100), LWP is still a factor 10 less in ARPEGE-SH-TEST
760 than in HAMSTRAD. ARPEGE-SH-TEST still fails to reproduce the large increase in liquid
761 water and IWV at 13:00 UTC since the local maximum is calculated 2 hours later. The impact
762 on the net surface radiation is weak with ARPEGE-SH-TEST underestimating the net surface
763 radiation by 50 W m^{-2} compared to observations, mainly attributable to the downwelling
764 longwave surface radiation from BSRN being 100 W m^{-2} greater than that of ARPEGE-SH-
765 TEST.

766 Finally, the bias on the net surface radiation and the underestimation of IWV and LWP of
767 the model compared to the observations is strongly reduced when using a new ice partition
768 function in ARPEGE-SH-TEST. This suggests that LWP has more impact than IWV on LW↓
769 due to the small quantities of specific humidity at Dome C.

770

771 **8. Conclusions**

772 A comprehensive water budget study has been performed during the Year of Polar Programs
773 SOP-SH at Dome C (Concordia, Antarctica) from mid-November 2018 to mid-February 2019.
774 Supercooled liquid water (SLW) clouds were observed and analysed by means of remote-
775 sensing ground-based instrumentation (tropospheric depolarization LIDAR, HAMSTRAD
776 microwave radiometer, BSRN net surface radiation), radiosondes, spaceborne sensor
777 (CALIOP/CALIPSO depolarization LIDAR) and the NWP ARPEGE-SH. The analysis shows
778 that SLW clouds were present from November to March, with the greatest frequency occurring
779 in December and January since ~50% of the days in summer time exhibited SLW clouds. The
780 clouds observed during the SOP-SH are typically located at the top of the boundary layer (100
781 to 400 m height) and are 50-100 m thick.

782 The analyses focused on two periods showing 1) a typical diurnal cycle of the PBL on 24
783 December 2018 (warm and dry, local mixing layer followed by a thinner cold and dry, local
784 stable layer which develops when the surface has cooled down) and 2) a perturbed diurnal cycle
785 of the PBL on 20 December 2018 (a warm and wet episode prevented from a clear diurnal cycle
786 of the PBL top). In both cases thin (~100-m thick) SLW clouds have been observed by ground-
787 based and spaceborne LIDARs developing within the entrainment and the capping inversion
788 zones at the top of the PBL. Spaceborne lidar observations revealed horizontal extensions of
789 these clouds as large as 280 and 550 km for the 24 and 20 December cases, respectively.
790 ARPEGE-SH was not able to correctly estimate the ratio between liquid and solid water inside
791 the cloudy layers, with SLW always strongly underestimated by a factor 1000 in the studied
792 cases, mainly because the liquid/ice partition function used in the model favours ice at
793 temperatures less than -20°C . Consequently, the net surface radiation was affected by the
794 presence of SLW clouds during these two episodes. The net surface radiation observed by
795 BSRN was 20-30 W m^{-2} higher than that modelled in ARPEGE-SH on 24 December 2018
796 (typical diurnal cycle of the PBL), this difference reaching $+50 \text{ W m}^{-2}$ on 20 December 2018
797 (perturbed diurnal cycle of the PBL), consistent with the total observed liquid water being 20
798 times greater in the perturbed PBL diurnal cycle than in the typical PBL diurnal cycle. The
799 difference in the net surface radiation is mainly attributable to longwave downward surface
800 radiation, BSRN values being 50 and 100 W m^{-2} greater than those of ARPEGE-SH in the
801 typical and perturbed cases, respectively.

802 The ice/liquid partition function used in the ARPEGE-SH NWP has been modified to favour
803 liquid water at temperatures below -20°C down to -40°C . For the two study cases, the model
804 run with this new partition function has been able to generate SLW clouds. During the typical
805 case, modelled LWP was consistent with observations and, consequently, the net surface
806 radiation calculated by the model agreed with measurements to within $\pm 20 \text{ W m}^{-2}$. During the

807 perturbed case, modelled LWP was a factor 10 less than observations and, consequently, the
808 model underestimated the net surface radiation by $\sim 50 \text{ W m}^{-2}$ compared to observations.

809 Time coincident ground-based remote-sensed measurements of water (vapour, liquid and
810 solid), temperature and net surface radiation are available at Dome C since 2015. Consequently,
811 a comprehensive statistical analysis of the presence of SLW clouds will be performed in the
812 near future. Coupled with modelling studies (NWP ARPEGE-SH, mesoscale models), an
813 estimation of the radiative impact of these clouds on the local climate will then be performed.

814

815 **Data availability**

816 HAMSTRAD data are available at <http://www.cnrm.meteo.fr/spip.php?article961&lang=en>
817 (last access: 28 August 2019). The CALIOP images are accessible at <http://www-calipso.larc.nasa.gov/>
818 (last access: 28 August 2019). The tropospheric depolarization LIDAR
819 data are reachable at http://lidarmax.altervista.org/englidar/_Antarctic%20LIDAR.php (last
820 access: 28 August 2019). Radiosondes are available at <http://www.climantartide.it> (last access:
821 28 August 2019). BSRN data can be obtained from the ftp server (<https://bsrn.awi.de/data/data-retrieval-via-ftp/>)
822 (last access: 28 August 2019). The ARPEGE data and corresponding
823 technical information are available from the YOPP Data Portal and from the ftp server ([ftp.umr-](ftp.umr-cnrm.fr)
824 [cnrm.fr](ftp.umr-cnrm.fr) with user: yopp and password: Arpege) (last access: 28 August 2019). The NCEP data
825 are available at <https://www.esrl.noaa.gov/psd/> and the back-trajectory calculations can be
826 performed at <https://www.ready.noaa.gov/HYSPLIT.php>.

827

828 **Author contributions**

829 PR, MDG, AL, and PG provided the observational data while EB, NA and VG developed
830 the model code and performed the simulations. PD, JLA and DV contributed to the data
831 interpretation. All the co-authors participated in the data analysis. PR prepared the manuscript

832 with contributions from all co-authors. DV, EB, NA, MDG and PD also contributed
833 significantly to the revision of the manuscript supervised by PR.

834

835 **Competing interests**

836 The authors declare that they have no conflict of interest.

837

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851 <https://www.esrl.noaa.gov/psd/> and the NOAA Air Resources Laboratory to have accessed the
852 HYSPLIT model through <https://www.ready.noaa.gov/HYSPLIT.php>. We would like to thank
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854

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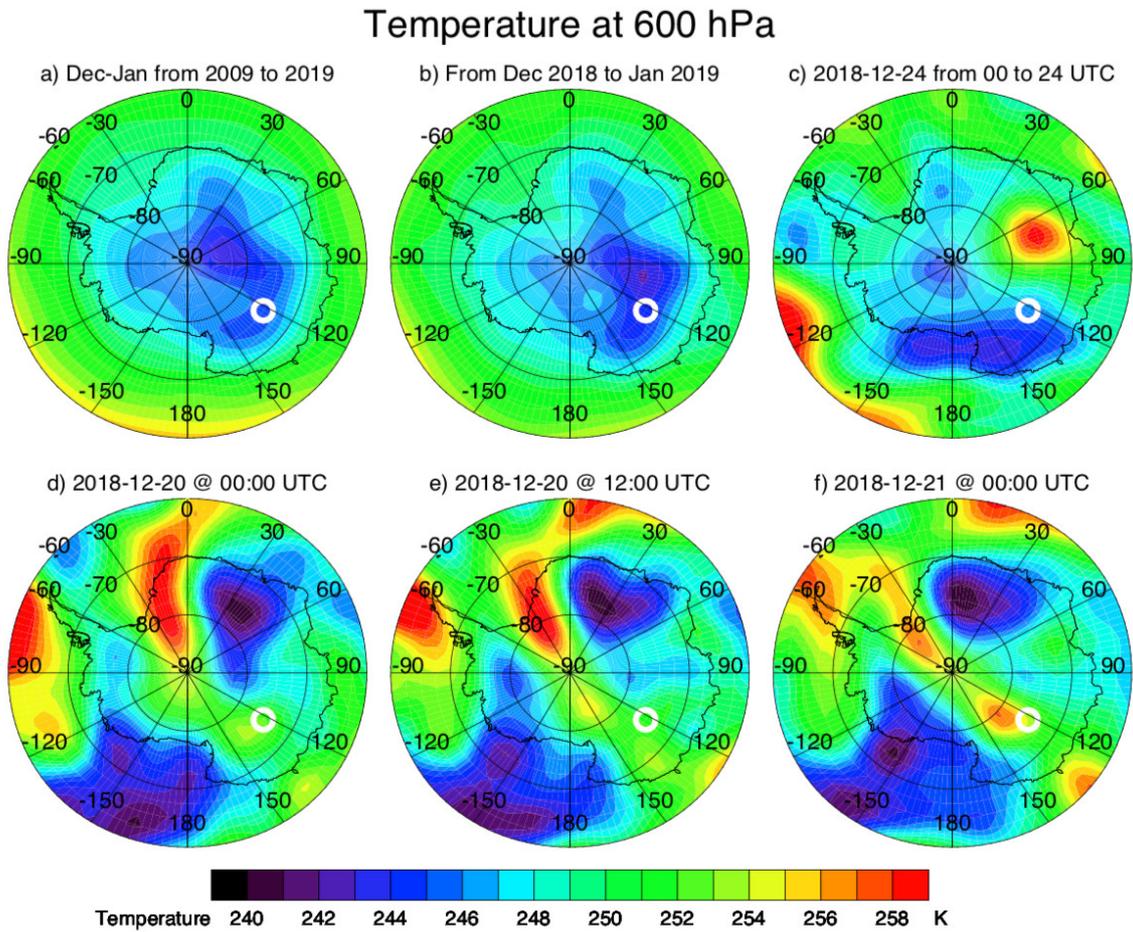
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Figures

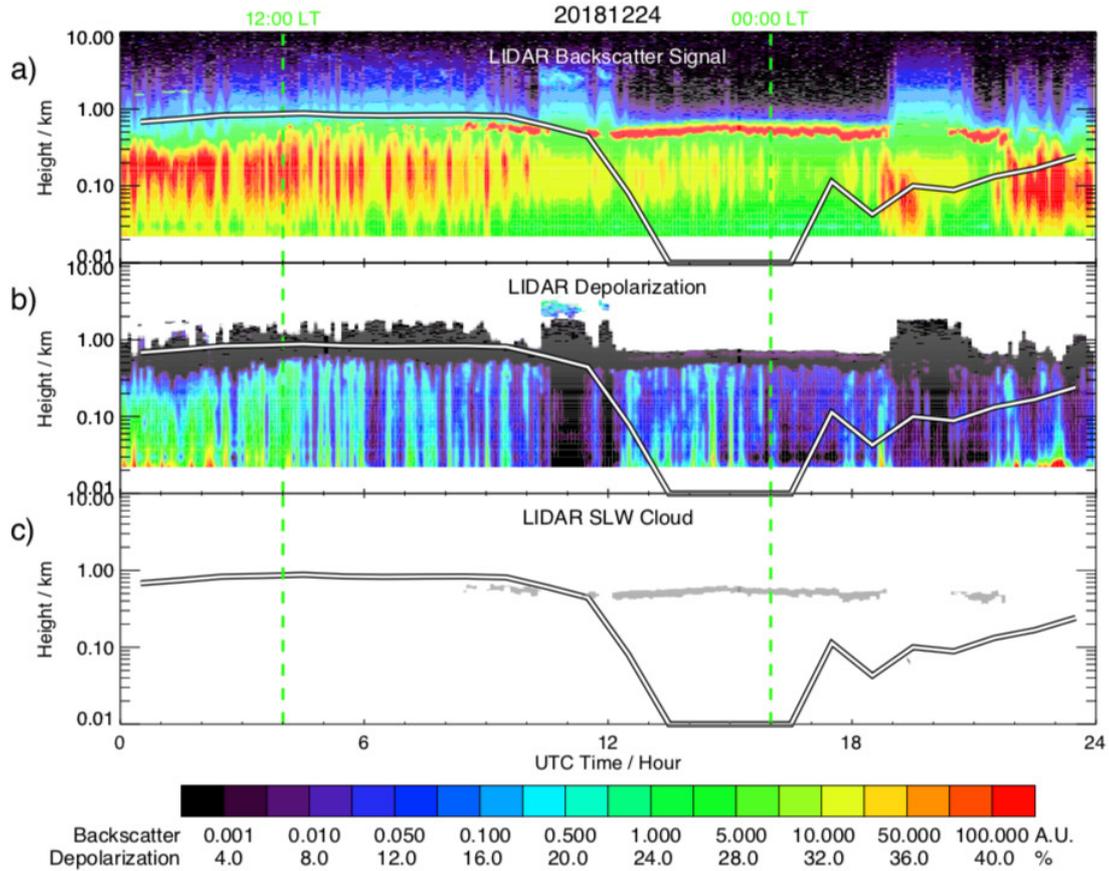
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991

992 **Figure 1:** Temperature fields from NCEP at 600 hPa: a) decadal average over December-
993 January from 2009 to 2019, b) YOPP average over December 2018-January 2019, c) daily
994 average over 24 December 2018, d) 20 December 2018 at 00:00 UTC, e) 20 December 2018 at
995 12:00 UTC, and f) 21 December 2018 at 00:00 UTC. The white circle represents the position
996 of the Dome C station.

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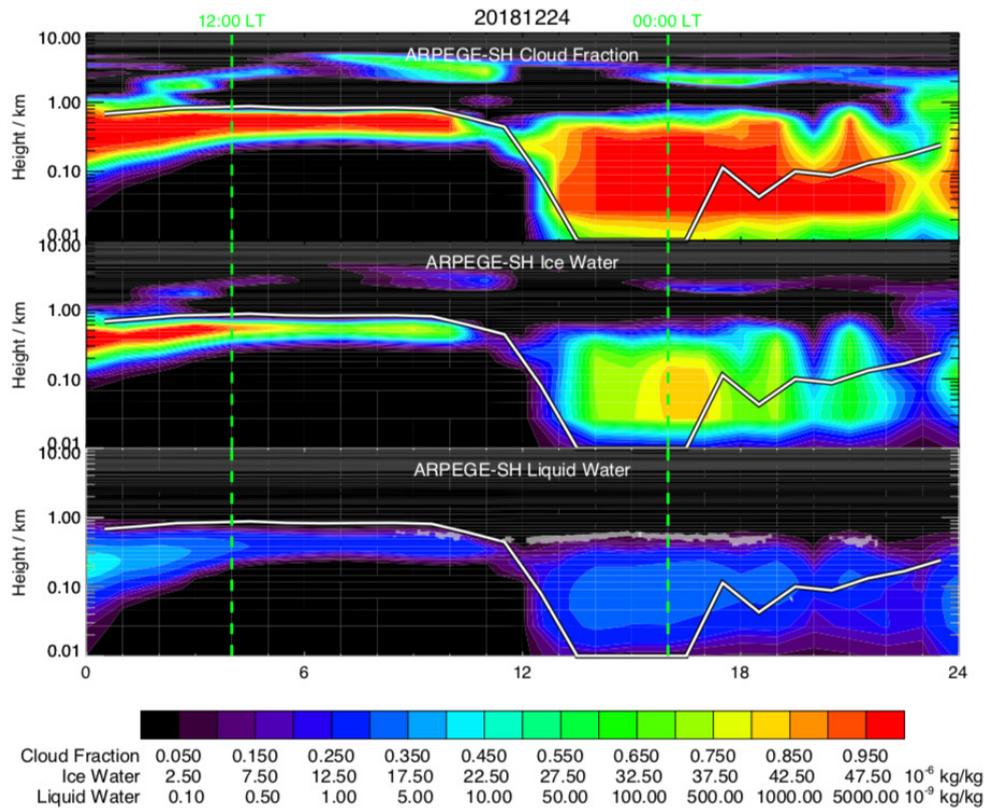


999

1000 **Figure 2:** Diurnal variation on 24 December 2018 (UTC Time) along the vertical of: a) the
 1001 backscatter signal (Arbitrary Unit, A.U.), b) the depolarization ratio (%) measured by the
 1002 aerosol LIDAR, and c) the Supercooled Liquid Water (SLW) cloud height (grey) deduced from
 1003 the aerosol LIDAR ($\beta_c > 100 \beta_{mol}$, depolarization $< 5\%$). Superimposed to all the Figures is the
 1004 top of the Planetary Boundary Layer calculated by the ARPEGE-SH model (black-white thick
 1005 line). Two vertical green dashed lines indicate 12:00 and 00:00 LT.

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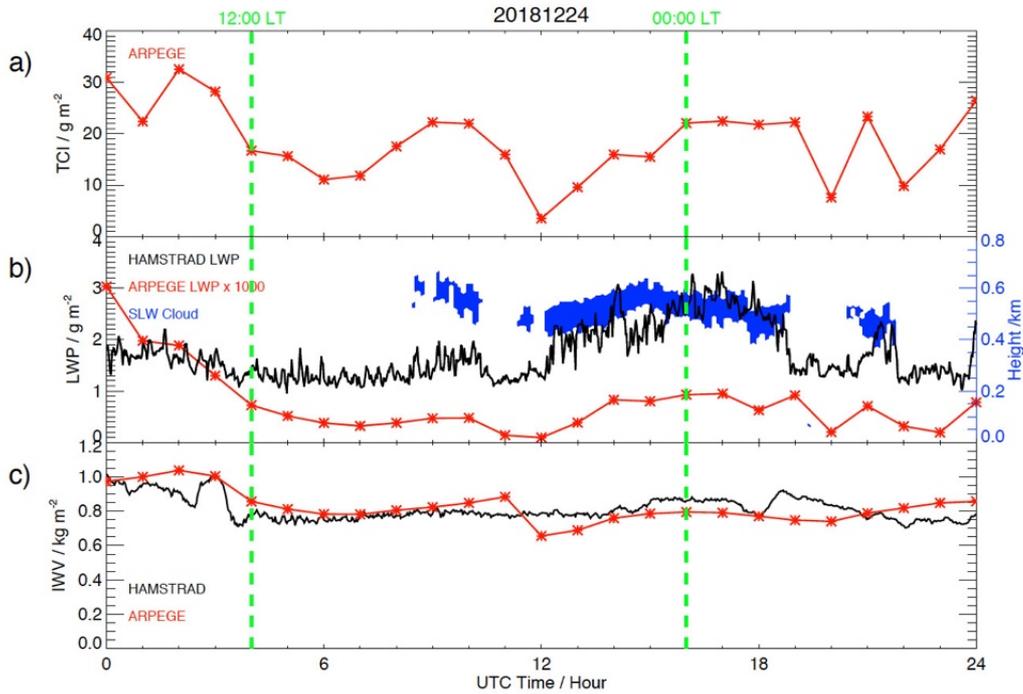


1008

1009 **Figure 3:** Time-height cross section on 24 December 2018 (UTC Time) of: a) the Cloud
1010 Fraction (0-1), b) the Ice Water mixing ratio (10^{-6} kg kg $^{-1}$) and c) the Liquid Water mixing ratio
1011 (10^{-9} kg kg $^{-1}$) calculated by the ARPEGE-SH model. Superimposed to all the panels is the top
1012 of the Planetary Boundary Layer calculated by the ARPEGE-SH model (black-white thick line).
1013 Superimposed in panel c is the SLW cloud (grey area) deduced from the LIDAR observations
1014 (see Fig. 1c). Two vertical green dashed lines indicate 12:00 and 00:00 LT.

1015

1016

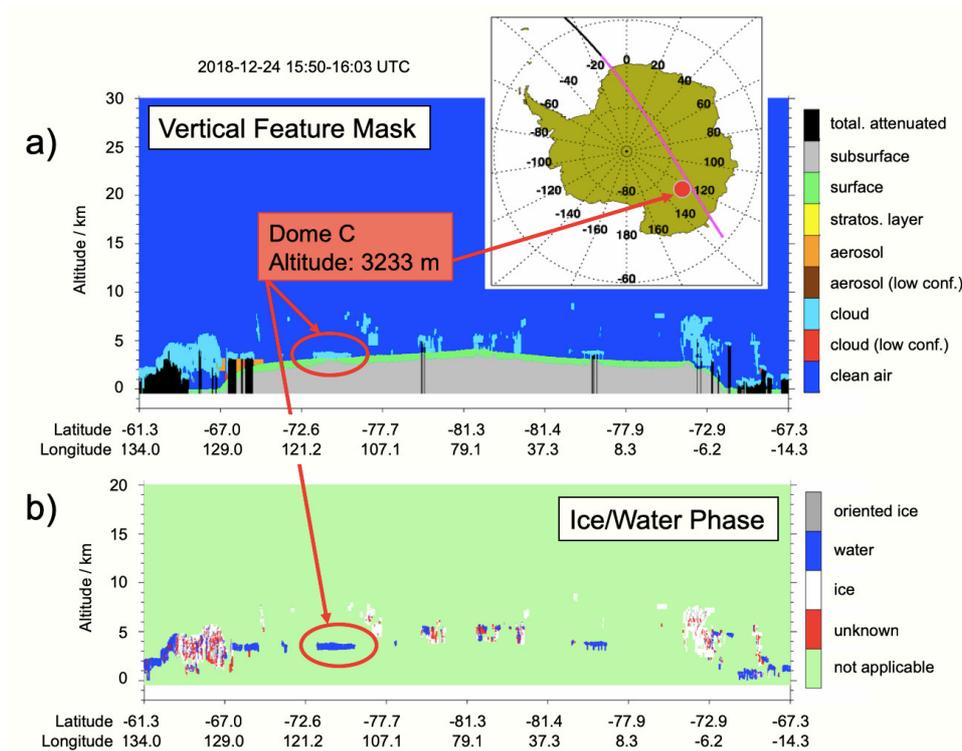


1017

1018 **Figure 4:** Diurnal variation on 24 December 2018 (UTC Time) of: a) the Total Column of Ice
 1019 (TCI) (g m^{-2}) calculated by ARPEGE-SH (red crossed line), b) the Liquid Water Path (LWP)
 1020 measured by HAMSTRAD (g m^{-2} , black solid line) and calculated by ARPEGE-SH ($\times 1000 \text{ g}$
 1021 m^{-2} , red crossed line) and c) the Integrated Water Vapour (IWV, kg m^{-2}) measured by
 1022 HAMSTRAD (black solid line) and calculated by ARPEGE-SH (red crossed line).
 1023 Superimposed to panel b) is the SLW cloud thickness (blue area) deduced from the LIDAR
 1024 observations (see Fig. 1c) (blue y-axis on the right of the Figure). Note LWP from ARPEGE-
 1025 SH has been multiplied by a factor 1000. Two vertical green dashed lines indicate 12:00 and
 1026 00:00 LT.

1027

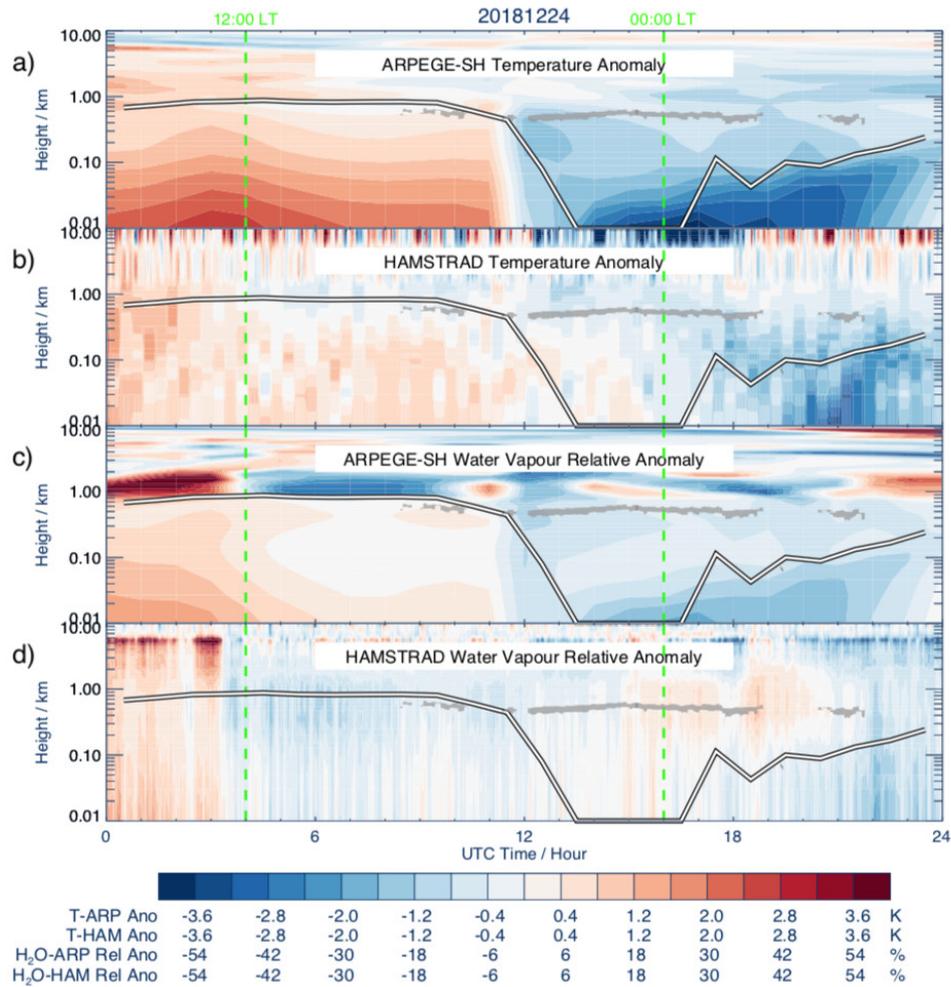
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1030 **Figure 5:** CALIOP/CALIPSO airborne LIDAR observations **version V3.40** along one orbit
 1031 on 24 December 2018 (15:50-**16:03** UTC) in the vicinity of Dome C (75°S, 123°E): a) the
 1032 Vertical Feature Mask highlighting a cloud (light blue) near the surface (red circle) and b) the
 1033 Ice/Water Phase Mask highlighting a SLW (dark blue) cloud near the surface (red circle). The
 1034 ground-track of the sensor (pink) has been embedded at the top of the Figure, with the location
 1035 of Dome C marked (red filled circle). Note that the altitude is relative to the sea surface, with
 1036 the height of surface of Dome C at an elevation of 3233 m amsl. **Figure adapted from the**
 1037 **original image available at [https://www-](https://www-calipso.larc.nasa.gov/products/lidar/browse_images/std_v34x_showdate.php?browse_date=2018-12-24)**
 1038 **[calipso.larc.nasa.gov/products/lidar/browse_images/std_v34x_showdate.php?browse_date=20](https://www-calipso.larc.nasa.gov/products/lidar/browse_images/std_v34x_showdate.php?browse_date=2018-12-24)**
 1039 **[18-12-24](https://www-calipso.larc.nasa.gov/products/lidar/browse_images/std_v34x_showdate.php?browse_date=2018-12-24)**.

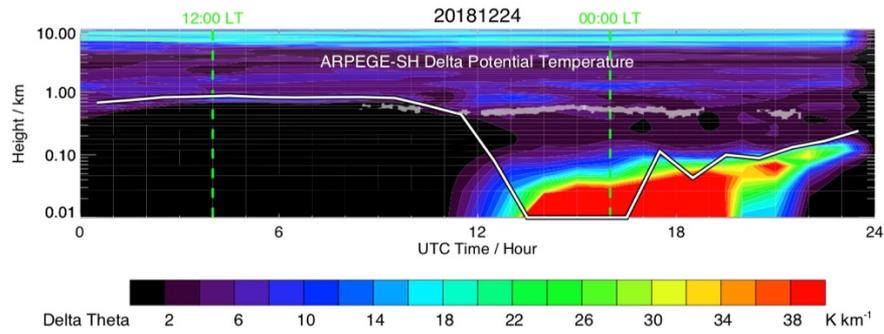
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1041
 1042 **Figure 6:** Time-height cross section on 24 December 2018 (UTC Time) of a) the temperature
 1043 anomaly (K) calculated by ARPEGE-SH and b) observed by HAMSTRAD, c) the water vapour
 1044 relative anomaly (%) calculated by ARPEGE-SH and d) observed by HAMSTRAD.
 1045 Superimposed to all the Figures are the SLW cloud altitude (grey area) deduced from the
 1046 LIDAR observations (see Fig. 1c) and the top of the Planetary Boundary Layer calculated by
 1047 the ARPEGE-SH model (black thick line). Two vertical green dashed lines indicate 12:00
 1048 and 00:00 LT.

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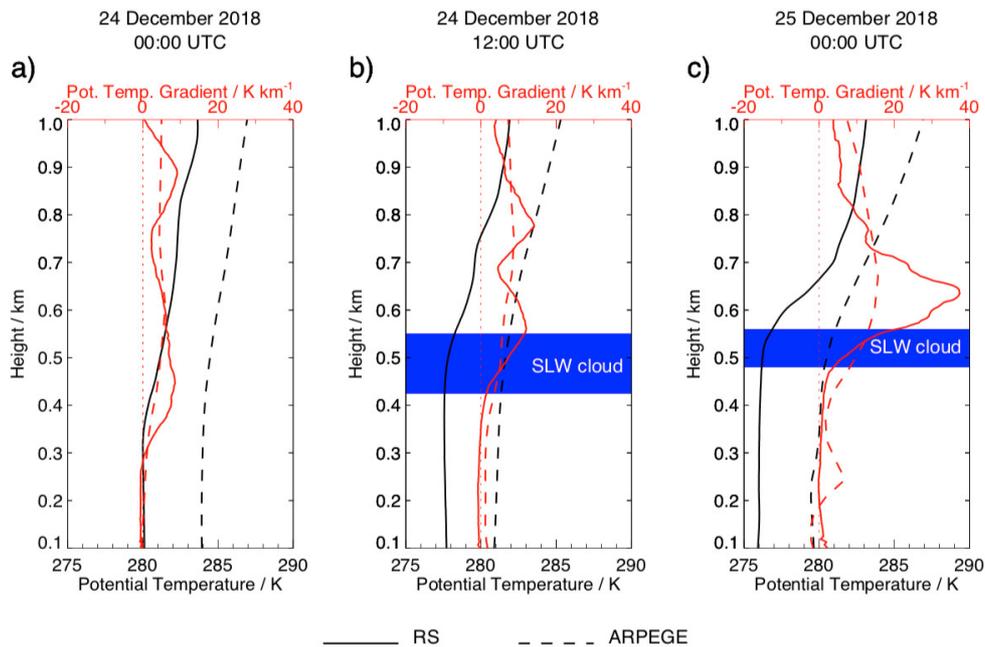


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1052 **Figure 7:** Time-height cross section of $\partial\theta/\partial z$ (K km^{-1}) calculated from ARPEGE-SH
 1053 temperature on 24 December 2018 (UTC Time). Superimposed are the SLW cloud altitude
 1054 (grey area) deduced from the LIDAR observations (see Fig. 1) and the top of the Planetary
 1055 Boundary Layer calculated by the ARPEGE-SH model (black-white thick line). Two vertical
 1056 green dashed lines indicate 12:00 and 00:00 LT.

1057

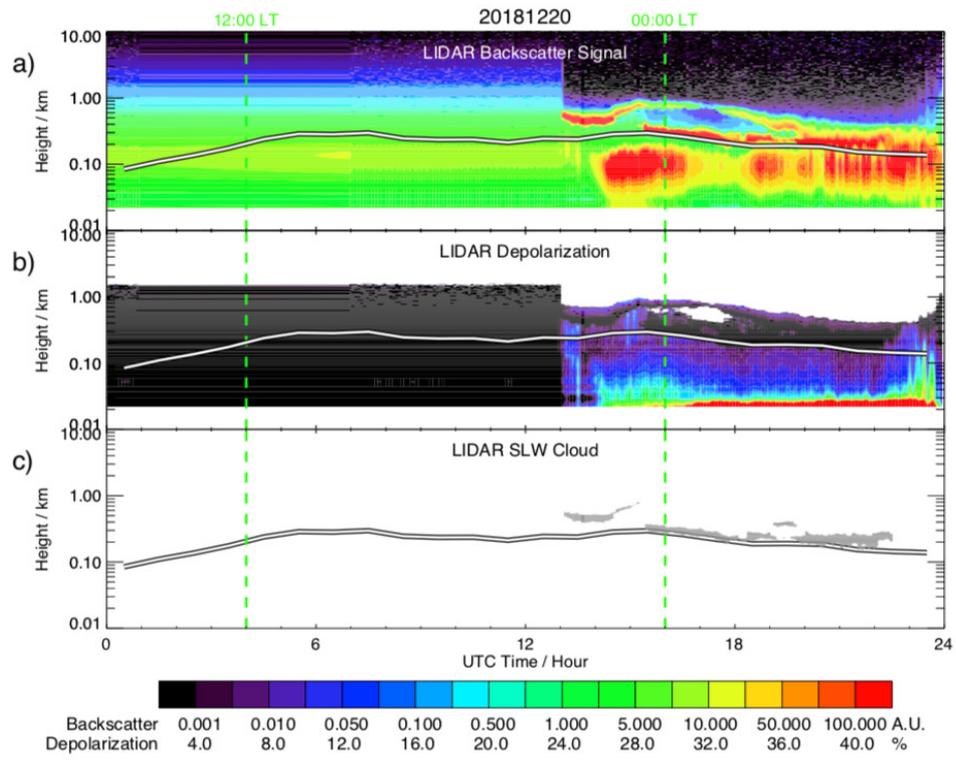
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1061 **Figure 8:** Vertical profiles of potential temperature θ (black) and the gradient in potential
 1062 temperature $\partial\theta/\partial z$ (red) as calculated from temperature measured by the radiosondes (solid line)
 1063 and analysed by ARPEGE-SH (dashed line) at Dome C on 24 December 2018 at a) 00:00 and
 1064 b) 12:00 UTC, and c) on 25 December 2018 at 00:00 UTC. The presence and the depth of the
 1065 SLW cloud detected from LIDAR observations are indicated by a blue area.

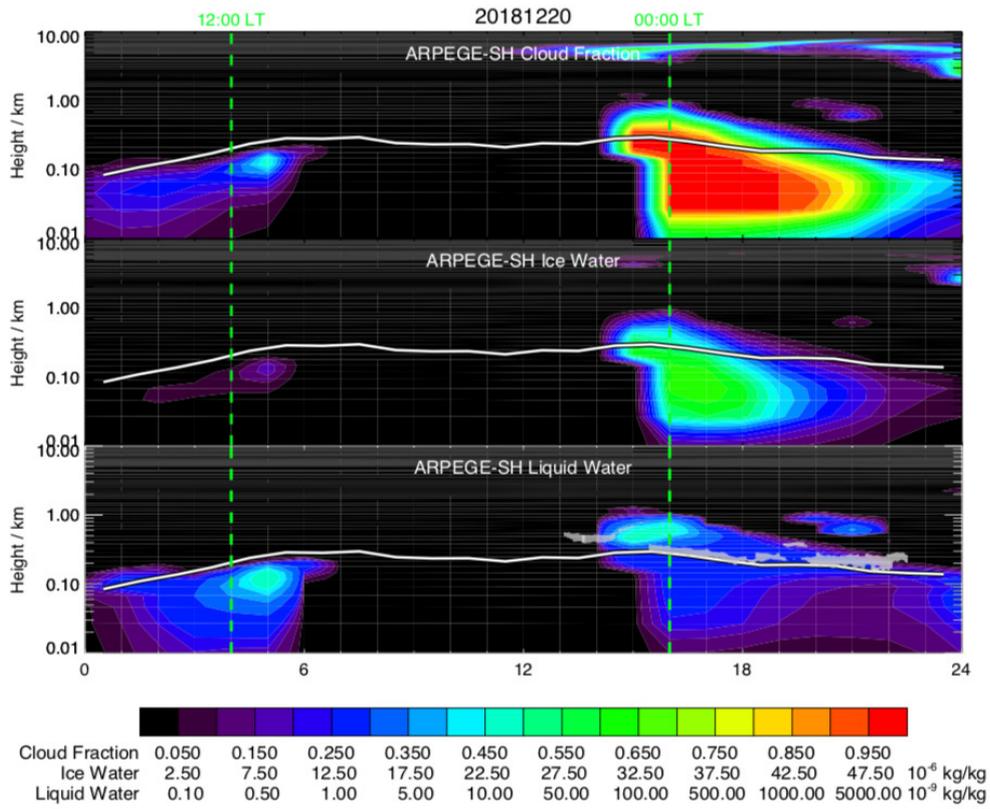
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1068 **Figure 9:** Same as **Figure 2** but for 20 December 2018.

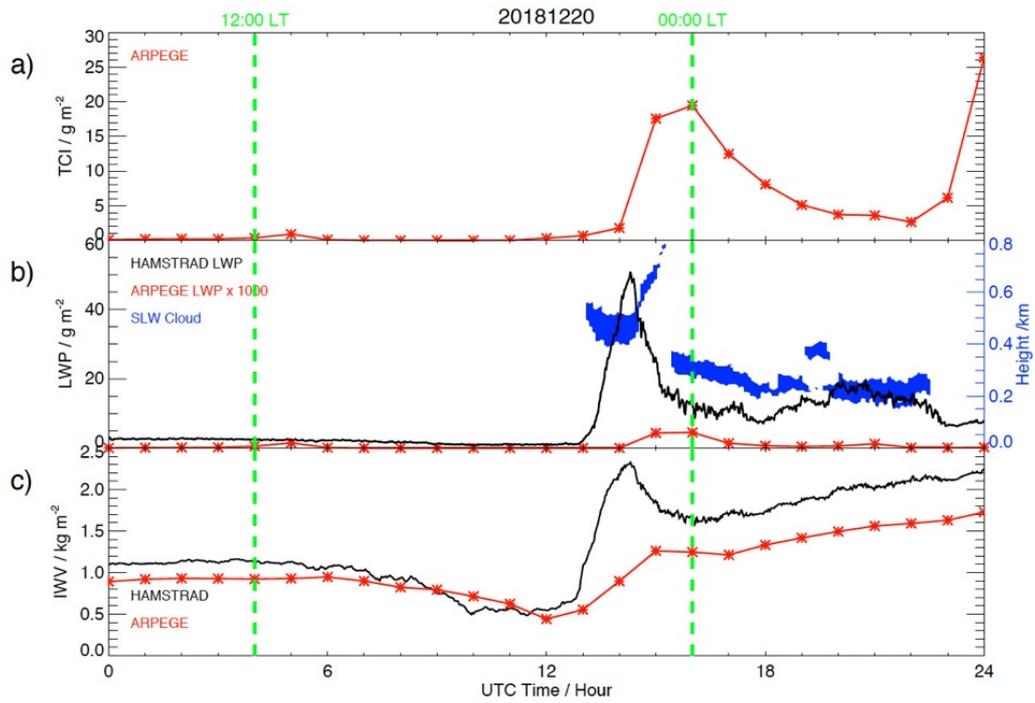
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1071 **Figure 10:** Same as **Figure 3** but for 20 December 2018.

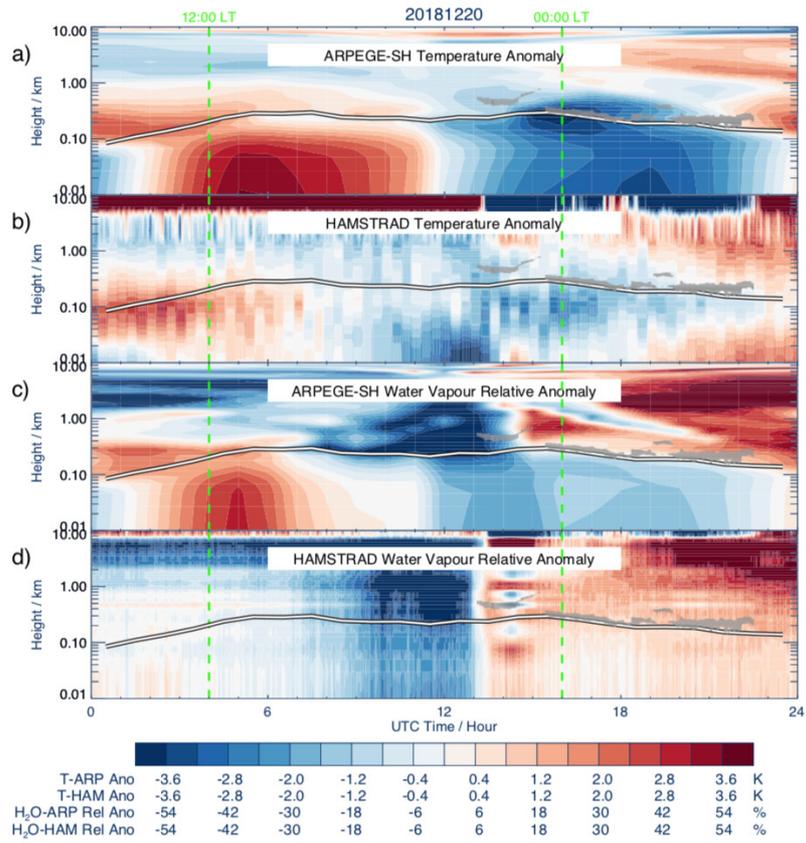
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1074 **Figure 11:** Same as **Figure 4** but for 20 December 2018.

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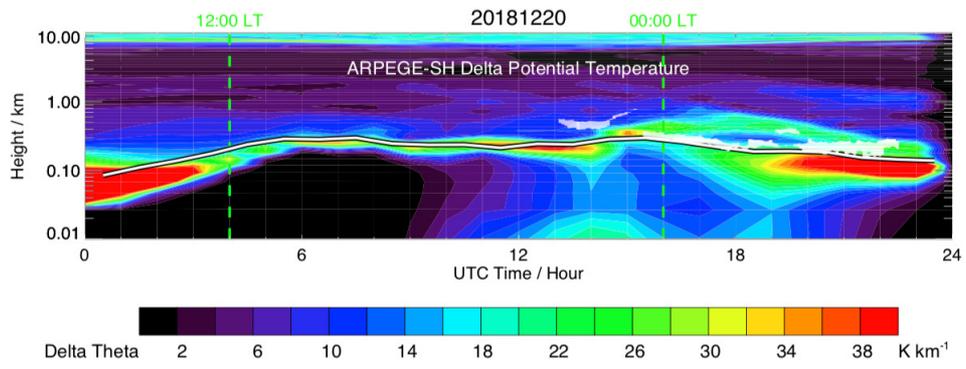
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1077 **Figure 12:** Same as **Figure 6** but for 20 December 2018.

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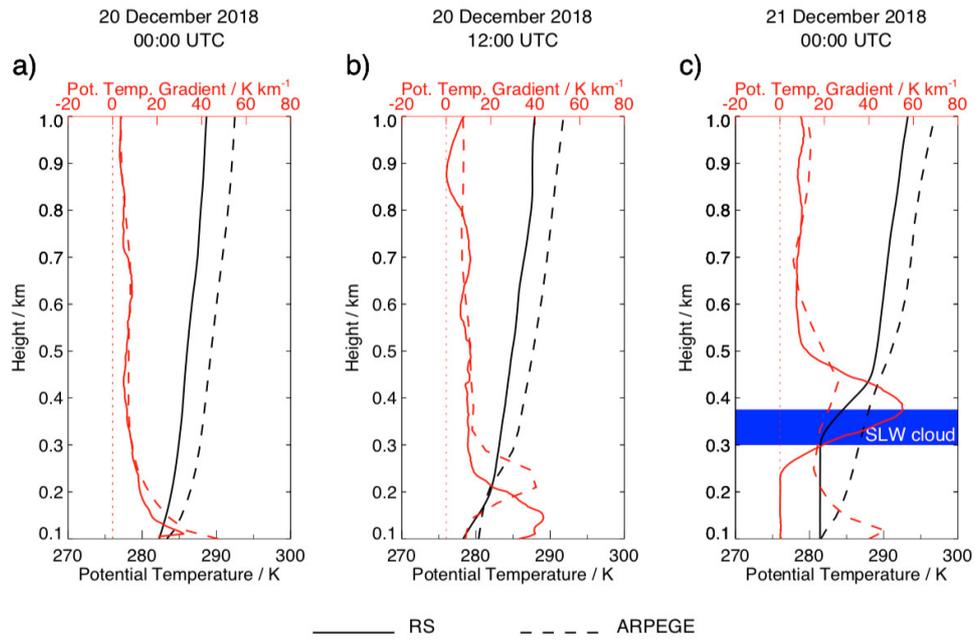


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1082 **Figure 13:** Same as **Figure 7** but for 20 December 2018.

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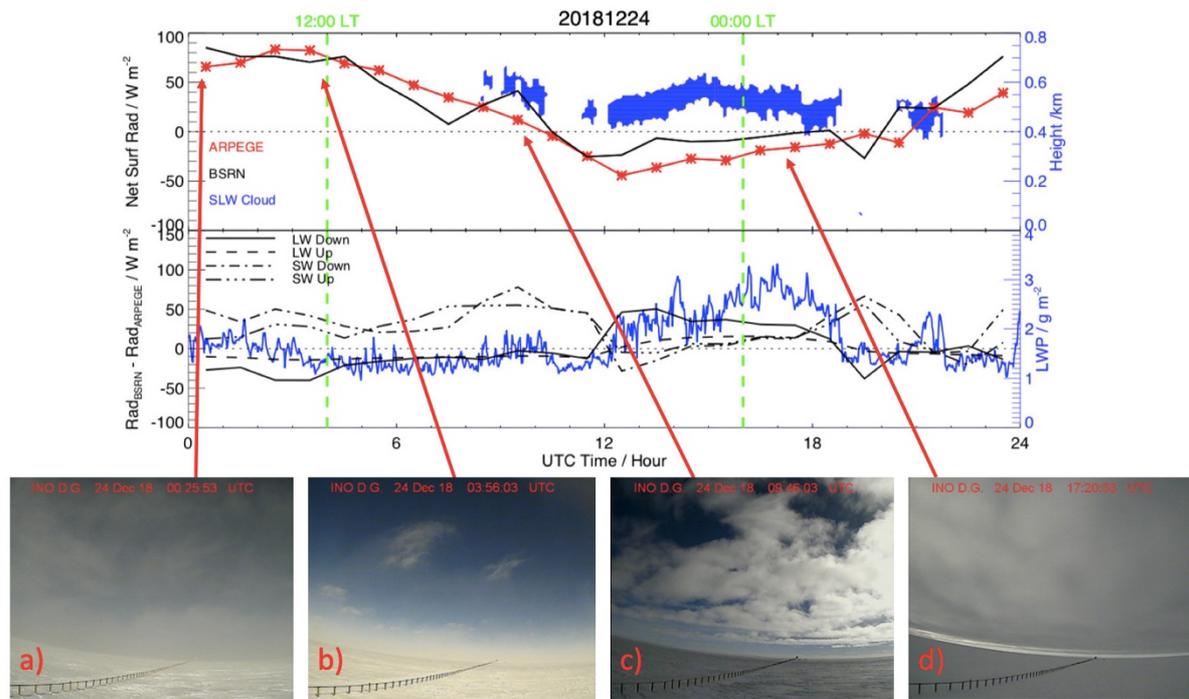


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1086 **Figure 14:** Same as **Figure 8** but on 20 December 2018 at a) 00:00 and b) 12:00 UTC, and c)

1087 on 21 December 2018 at 00:00 UTC.

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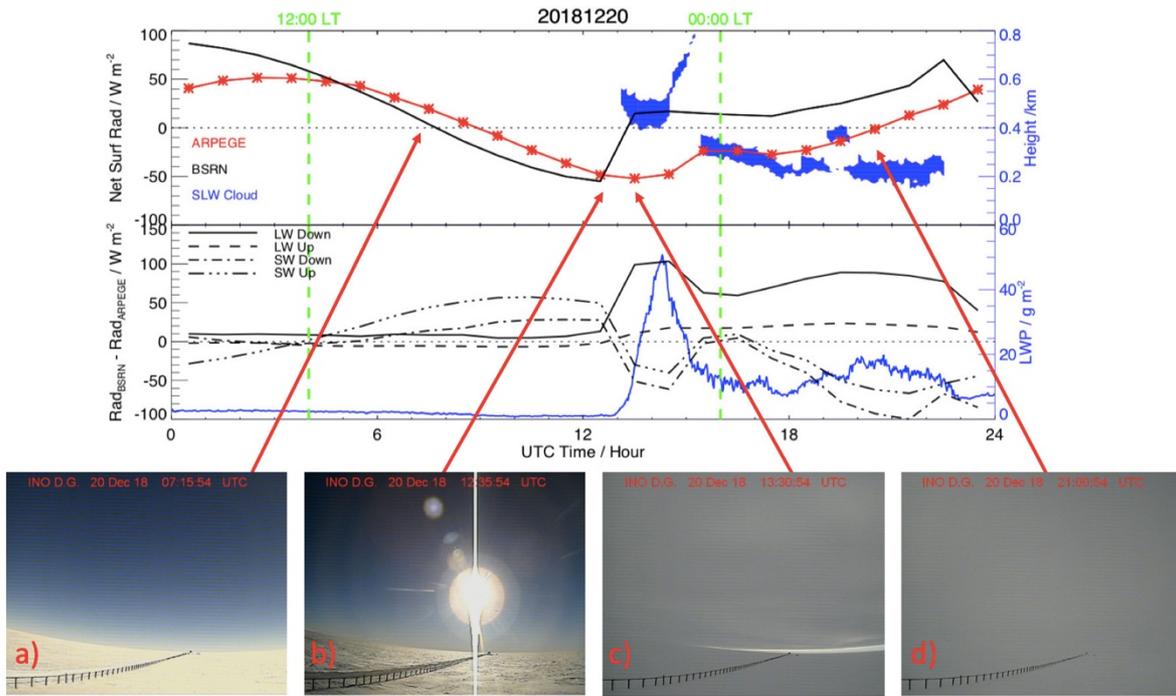


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1090 **Figure 15:** (Top) Diurnal variation of the net surface radiation ($W m^{-2}$) observed by BSRN
 1091 (black solid line) and calculated by ARPEGE-SH (red crossed line) on 24 December 2018 in
 1092 UTC Time. Superimposed is the SLW cloud height (blue) deduced from the LIDAR. (Middle)
 1093 Diurnal variation of the difference between surface radiation ($W m^{-2}$) observed by BSRN and
 1094 calculated by ARPEGE-SH on 24 December 2018 for longwave downward (black solid),
 1095 longwave upward (black dashed), shortwave downward (black dashed dotted) and shortwave
 1096 upward (black dashed triple dotted) components. Superimposed is LWP (blue) measured by
 1097 HAMSTRAD. (Bottom) Four webcam images showing the cloud coverage at: a) 00:25 UTC
 1098 and b) 03:56 UTC (cirrus clouds, no SLW cloud), c) 09:46 UTC (SLW cloud) and d) 17:20
 1099 UTC (SLW cloud). Two vertical green dashed lines indicate 12:00 and 00:00 LT.

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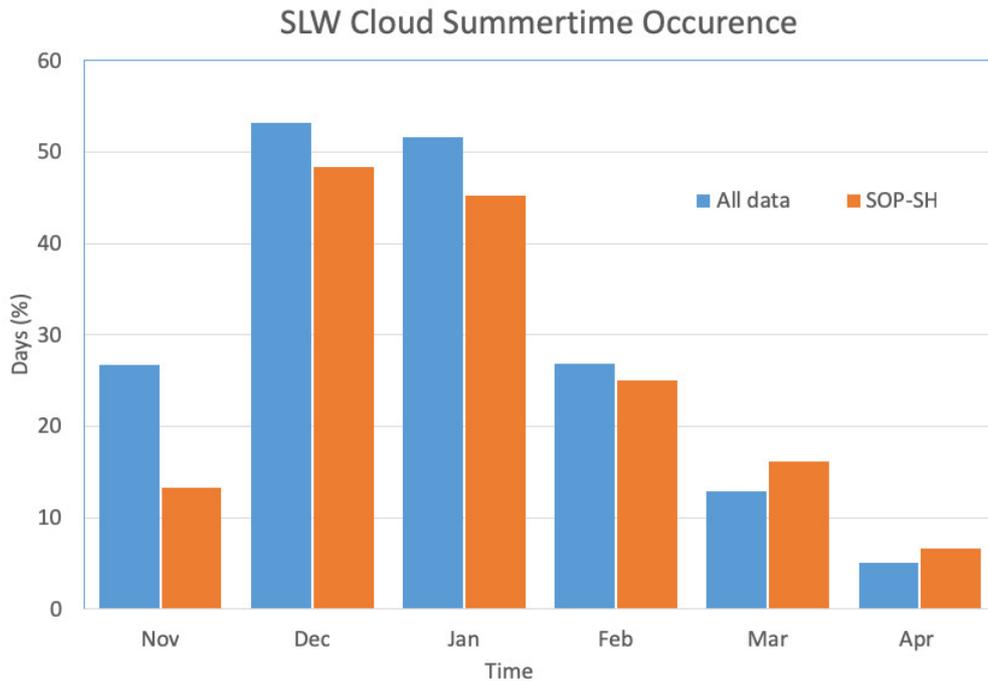
1102

1103 **Figure 16:** Same as **Figure 15** but for 20 December 2018 whilst the 4 webcam images were

1104 selected at: a) 07:15 and b) 12:35 UTC (clear sky), c) 13:30 and d) 21:00 UTC (SLW cloud).

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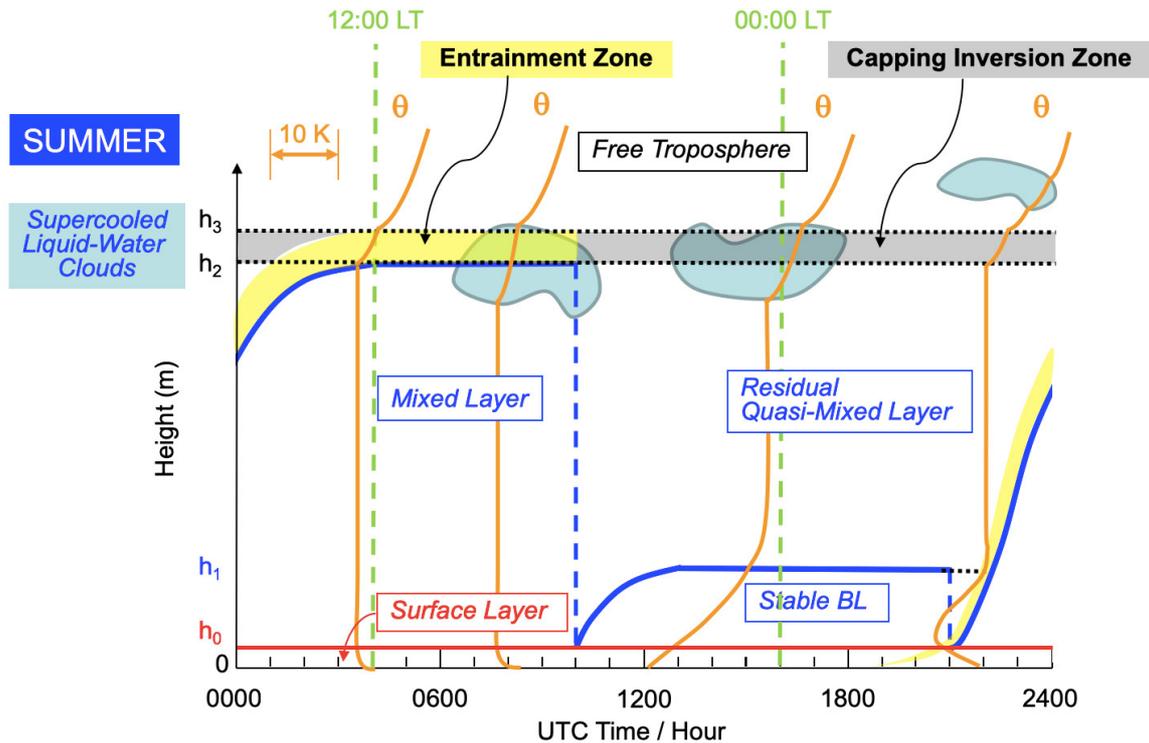
1108 **Figure 17:** Percentage of days per month that SLW clouds were detected within the LIDAR
 1109 data for more than 1 hour per day over the following summer periods: all data (blue) represent
 1110 the following months November (2016-2018), December (2016-2018), January (2018-2019),
 1111 February (2018-2019) and March (2018-2019) whilst SOP-SH (orange) represent the YOPP
 1112 campaign (November 2018 to April 2019). A day is flagged with a SLW cloud occurrence when
 1113 a SLW cloud has been detected in the LIDAR observations for a period longer than 1 hour.

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1119 **Figure 18:** Figure modified and updated from Fig. 12 of Ricaud et al. (2012) showing the
 1120 diurnal evolution (UTC Time) of the different layers in the Planetary Boundary Layer (PBL)
 1121 with h_0 the top of the surface layer, h_3 the daily overall top of the PBL, and h_1 the top of the
 1122 intermediate stable layer within the PBL. The orange lines symbolize the vertical profiles of
 1123 potential temperature θ , and the light blue areas the SLW clouds. The layer between h_2 and h_3
 1124 is named “capping inversion zone”. The yellow area represents the “entrainment zone” at the
 1125 top of the (cloudy or cloud-free) mixed layer. When the mixed layer is fully developed, the
 1126 entrainment zone coincides with the capping inversion zone. Note that $LT = UTC + 8 h$,
 1127 midnight and noon in the local time reference being indicated by the green dashed lines.

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