



1 Statistical Analysis of Contrail to Cirrus Evolution during the Contrail and Cirrus Experiments
2 (CONCERT)

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15

16 **Abstract:**

17 Air traffic affects the cloudiness, and thus the climate, by emitting exhaust gases and particles. The
18 study of the evolution of contrail properties is very challenging due to the complex interplay of vortex
19 dynamics and atmospheric environment (e.g. temperature, supersaturation). Despite substantial
20 progress in recent years, the optical, microphysical, and macrophysical properties of contrails and
21 ambient cirrus during contrail formation and subsequent ageing are still subject to large uncertainties
22 due to instrumental and observational limitations and the large number of variables influencing the
23 contrail life cycle. In this study, various contrail cases corresponding to different aircraft types and
24 atmospheric conditions are investigated using a statistical method based on the in situ optical
25 measurements performed during the CONCERT campaigns 2008 and 2011. These two aircraft
26 campaigns encompass more than 17 aircraft contrail cases. A Principal Component Analysis (PCA)
27 of the angular scattering coefficients measured by the Polar Nephelometer has been implemented in
28 order to classify the sampled ice cloud measurements in 6 clusters representative of different
29 development stages of the contrails (primary wake, young contrail, contrail-cirrus and natural cirrus).
30 Based on the information derived from air traffic control, extinction coefficients, asymmetry
31 coefficients, nitrogen oxide concentrations, relative humidity with respect to ice (RHI) and particle
32 size distributions are analyzed for each cluster to provide a characterization of the evolution of ice-
33 cloud properties during the contrail to cirrus evolution. The PCA demonstrates that contrail optical
34 properties are well suited to identify and discriminate the different contrail growth stages and to
35 provide an independent method for the characterization of the evolution of contrail properties.

36

37 **1 Introduction**

38 Aircraft exhaust plumes have a significant impact on climate and tropospheric chemistry (Lee et al.,
39 2010; IPCC, 1999). The Intergovernmental Panel for Climate Change IPCC special report on aviation
40 (1999) estimates that NO_x emissions from subsonic aircraft increase ozone concentrations by up to
41 6% at cruise level. Short and long lived pollution species have different impact on atmospheric
42 chemical composition depending on the flight level (Frömming et al, 2012). Emissions of water
43 vapor, black carbon (BC) / soot particles, sulfate (SO₄) aerosols and nitrogen oxides (NO_x) contribute



44 to the modification of the chemical composition of the upper troposphere on shorter timescales (Lee
45 et al., 2010, Gettelman and Chen, 2013; Liou et al., 2013). The long term climate impact is mainly
46 driven by CO₂ emissions. Modelling studies have shown that the direct radiative forcing from aviation
47 is expected to represent 3-4% (50-60 mW m⁻²) of the anthropogenic forcing (Lee et al., 2010; De
48 Leon et al., 2012) and could reach 87 mW m⁻² in 2025 (Chen and Gettelman, 2016). Aircraft induced
49 cloudiness has also an important impact on climate, although the quantitative assessment of the
50 radiative forcing remains a major source of uncertainties (Lee et al., 2010).

51 Contrail formation is mainly controlled by the thermodynamic properties of the ambient air and by
52 the aircraft emissions. The conditions for contrail formation can be determined by the Schmidt-
53 Appleman Criterion (SAC) (Schumann, 1996). Contrail chemical composition can have a significant
54 impact on the contrail formation (Kärcher et al., 2009). Indeed, the emission index (e.g. soot emission
55 index in kg-fuel⁻¹) is directly linked to the contrail microphysical properties, as the total number
56 densities and ice crystal diameters. Several studies in the past have been dedicated to the evolution of
57 concentrations of nitrogen oxide (NO) and sulfur dioxide (SO₂). Rapidly oxidized by OH (post-
58 combustor reaction; Jurkat et al., 2011), NO and SO₂ are transformed into a series of species including
59 nitrous acid (HONO), nitric acid (HNO₃), and sulfuric acid (H₂SO₄). Part of the nitric acid, e.g.,
60 formed in the young plume is taken up by contrail ice particles (Kärcher and Voigt, 2006; Voigt et
61 al., 2006; Schäuble et al., 2009). HONO can be a temporary reservoir of OH by photolysis reactions,
62 and H₂SO₄ a precursor for radiatively active sulfate particles also contributing to soot particle coating
63 (Jurkat et al., 2011). OH-induced reactions, formation and reaction schemes of greenhouse gases
64 (H₂O, O₃, CH₄), as well as emission of black carbon and sulfate aerosols have significant impact on
65 climate (Frömming et al., 2012; Gettelmann and Chen, 2013). Two different processes of contrail
66 formation have been studied: combustion condensation trails and aerodynamic condensation trails.
67 Different studies (Gierens and Dilger, 2013; Jansen and Heymsfield, 2015) have illustrated
68 characteristics of aerodynamically controlled contrail formation associated to warmer temperatures
69 (observations at temperatures above -38°C). For contrails primarily initiated by the combustion
70 processes, the mixing of hot and humid exhaust gases with cooler and dryer ambient air increases the
71 local relative humidity in the exhaust plume leading to the formation of contrails when the saturation
72 with respect to liquid water is reached. Thus, soot and sulfate aerosols emitted by the aircraft (Moore
73 et al., 2017) may act as condensation nuclei to form liquid droplets. Homogeneous ice nucleation of
74 the liquid droplets can occur when the exhaust cools down through mixing with the ambient
75 temperature, while preserving ice saturation. Small ice crystals are then formed in the jet phase within
76 some tenths of a second (Kärcher and Yu, 2009). The further life-cycle of contrails depends on the
77 interaction with the wake vortices behind aircraft and the ambient atmosphere (Irvine et al., 2012;
78 Graf et al., 2012; Duda et al., 2013; Carleton et al., 2013; Schumann and Heymsfield, 2017). The ice
79 crystals in the young contrails are captured within two counter-rotating wake vortices in the
80 downwash behind the aircraft induced by the aircraft lift, which induce adiabatic compression,
81 heating, and partial sublimation of the ice crystals within the primary wake (Lewellen and Lewellen,
82 2001; Sussmann and Gierens, 2001, Unterstrasser et al., 2008, Unterstrasser et al., 2016; Kärcher and
83 Voigt, 2017). This primary wake may disappear if ambient air is subsaturated with respect to ice. In
84 the case of supersaturation, the secondary wake becomes visible, thereby detraining ice particles from
85 the primary wake at a higher level (Sussmann and Gierens, 1999, Kaufmann et al., 2014). The initially
86 almost spherical ice crystals become increasingly aspherical and grow by uptake of water vapor as
87 long as saturation with respect to ice is prevailing. For ambient humidity above ice saturation,
88 contrails can persist after the vortex breakdown, spread and evolve into contrail cirrus (Schumann
89 and Heymsfield, 2017). The associated cloud cover (larger than for linear contrails alone) of contrail
90 cirrus then shows increasing impact on the radiative forcing (Burkhardt and Kärcher, 2011;
91 Schumann et al., 2015).



92 The assessment of the contrail radiative forcing requires, in particular, an accurate estimation of the
93 cloud cover, the visible optical depth, the single scattering characteristics, the ice crystal effective
94 size and habit (Yang et al., 2010; Spangenberg et al., 2013). Satellite observations provide a
95 comprehensive dataset to study statistically the contrail to cirrus evolution. The combined contrail
96 tracking algorithms on the Spinning Enhanced Visible and Infrared Imager (SEVIRI) on board the
97 Meteosat Second Generation (MSG) satellites with properties inferred by the Moderate Imaging
98 Spectroradiometer (MODIS) on board the Terra satellite was used by Vazquez-Navarro et al., (2015)
99 to characterize the properties of 2300 contrails. Properties included lifetime (mean values of 1h), the
100 width (8 km), the length (130 km), the optical thickness (0.34), the altitude (11.7 km) and the radiative
101 forcing (-26 W m^{-2} for shortwave forcing over land) of these contrails. However, detailed *in situ*
102 optical and microphysical measurements are still needed to evaluate satellite products and to develop
103 more appropriate retrieval algorithm. In particular, distinguishing contrails from natural cirrus
104 remains extremely challenging from satellite observations. Although the optical and microphysical
105 properties of young contrails (linear contrails) differ from natural cirrus properties, the contrail
106 properties are highly time dependent and persistent contrail cirrus can be embedded in thin cirrus
107 clouds. Recent *in situ* measurements (Voigt et al., 2017) show that the microphysical properties of
108 contrail cirrus can still be distinguished from natural cirrus at contrail cirrus ages up to several hours.

109 Most of the studies (Jessberger et al., 2013; Lewellen et al., 2012; Schumann et al., 2013) separate
110 the contrail analysis between the two wakes. Indeed, the primary and secondary wake properties
111 depend strongly on atmospheric conditions and aircraft type (emission index, vortex, flight level,
112 ambient humidity, temperature,...). In the primary wake, contrail ice crystals are quasi-spherical with
113 values of the effective diameter (D_{eff}) typically lower than $4 \mu\text{m}$ (Schumann et al., 2011; Gayet et
114 al., 2012; Järvinen et al., 2016; Schumann et al., 2017b). Also the total number concentration of ice
115 particles is typically larger than 1000 cm^{-3} a few seconds after contrail formation (Baumgardner and
116 Gandrud, 1998; Petzold et al., 1997) and subsequently decreases by dilution to concentrations below
117 200 cm^{-3} within less than a minute after contrail generation (Poellot et al., 1999; Schröder et al., 2000;
118 Gayet et al., 2012). Gayet et al. (2012) reported mean values of ice water content of 3 mg m^{-3} and
119 maximum extinction coefficients close to 7 km^{-1} . The recent overview on contrail studies presented
120 in Schumann et al. (2017b) reports several microphysical properties at different stages, for different
121 atmospheric conditions as well as comparisons with the Contrail Cirrus Prediction (CoCIP) model
122 simulations. Their study highlights a large variability (which increases with contrail age) of contrail
123 properties. Comparing primary and secondary wakes, several studies reported findings on the
124 secondary wake and its evolution into contrail cirrus. Detrained from the primary wake and submitted
125 to saturated ambient air with respect to ice, ice crystals grow rapidly, while crystal concentration
126 decreases. Within the first minutes after formation, measurements exhibit aspherical ice crystals
127 characterized by effective sizes up to $6 \mu\text{m}$, IWC ranging between 2.5 and 10 mg m^{-3} , extinction
128 between 2 and 3 km^{-1} , with crystal concentrations typically lower than 100 cm^{-3} (Goodman et al.,
129 1998; Voigt et al., 2010; Kübbeler et al., 2011; Gayet et al., 2012; Jeßberger et al., 2013; Schumann
130 et al., 2013; Poellot et al., 1999; Febvre et al., 2009; Kaufmann et al., 2014). Aged contrails can persist
131 and evolve into contrail cirrus if the ambient air is saturated with respect to ice, however those studies
132 are limited by the lack of unambiguous identification (Schumann et al., 2017a). Also after a few
133 minutes, difficulties appear for the pilot to track the contrail by visual navigation, which is due to
134 contrail and contrail cirrus spreading in the free troposphere. Observations of the ice crystal shape
135 and growth over several tens of minutes and up to an hour illustrate that crystal effective size can
136 easily reach $20 \mu\text{m}$ and beyond with number concentrations ranging from 1 to 5 cm^{-3} (Lawson et al.,
137 1998; Schäuble et al., 2009), extinction less than 0.5 km^{-1} (Febvre et al., 2009), and IWC up to 10
138 mg m^{-3} (Schröder et al., 2000; De Leon et al., 2012). At this stage, within a sustained ice-
139 supersaturated environment, contrail microphysical properties may still differ from those of natural
140 cirrus (Voigt et al., 2017) with concentrations of ice crystals larger than $100 \mu\text{m}$ in the order of 0.1



141 cm^{-3} . These crystals typically show bullet rosette type habits (Heymsfield et al., 1998; Heymsfield et
142 al., 2010). Optical depth values can reach a value of 2.3 (Atlas and Wang, 2010), corresponding to an
143 extinction of 0.023 km^{-1} . Nevertheless, the transition from contrails to cirrus highly depends on the
144 ambient saturation conditions and modelling studies with typical atmospheric conditions suggest time
145 evolution of optical and microphysical properties from contrail to contrail cirrus clouds (Burkhardt
146 and Kärcher, 2011; Unterstrasser et al., 2016 ; Schumann et al., 2015).

147 In this study, we report on a method presenting a powerful alternative for classifying cloud events
148 into young contrail, contrail-cirrus and natural cirrus. The method is applied to aircraft data of the
149 CONCERT (Contrail and Cirrus Experiment) campaigns (Voigt et al., 2010, 2011, 2014). The
150 methodology consists of implementing a principal component analysis (PCA) of the angular light
151 scattering data from a Polar Nephelometer. The PCA results of the different type of contrails (different
152 clusters) are then utilized with corresponding optical, microphysical, and chemical properties in order
153 to validate hypothesis on contrail phase definitions (young contrails to cirrus contrails). This paper
154 starts with an illustration of the properties of contrails and cirrus clouds observed during two specific
155 CONCERT flights (19 November 2008 and 16 September 2011) encompassing a series of different
156 contrail evolution phases. These two flights containing a variety of contrail-cirrus information can be
157 regarded as an analytical framework producing results which then can be compared to contrail-cirrus
158 properties of other flights.

159 **2 CONCERT projects and data processing**

160 2.1 CONCERT campaigns

161 CONCERT-1 and CONCERT-2 campaigns took place in October/November 2008 and
162 August/September 2011, respectively. These two campaigns with the DLR Falcon 20 E research
163 aircraft were based in Oberpfaffenhofen, Germany, and sampled contrails and cirrus at mid-latitudes
164 in the Northern Hemisphere. The overall objective has been to reduce uncertainties on the
165 microphysical, chemical, and radiative properties of contrails behind aircraft of different types and to
166 improve the evaluation of contrail's impact on climate. Besides the primary objectives focusing on
167 contrails, few CONCERT flights were dedicated to emissions of Etna and Stromboli volcanos (Voigt
168 et al., 2014; Shcherbakov et al., 2016). Also a few stratospheric intrusions were observed during the
169 flight missions. In total, 23 flights were recorded during the two measurement campaigns, wherein
170 12 flights were entirely focused on aircraft contrail chasing. Overall, more than 17 different aircraft
171 exhausts plumes have been probed. Particularly, the CONCERT-2 campaign mainly focused on
172 observing persistent contrails, and hence on the evolution of contrails into contrail cirrus.

173 During both CONCERT campaigns, the DLR research aircraft Falcon was equipped with a set of
174 instruments to measure the optical and microphysical properties of cloud particles and also the trace
175 gas composition in the UTLS (Upper Troposphere / Lower Stratosphere) region. Voigt et al. (2010)
176 provide a detailed description of the aircraft instrumentation. We briefly introduce the instruments
177 used in this study.

178 2.2 Aircraft instrumentation

179 The microphysical and optical particle properties of contrails and cirrus presented in this study were
180 mainly derived from the PMS Forward Scattering Spectrometer Probe 300 (FSSP-300), the Polar
181 Nephelometer (PN), and the PMS 2D-C hydrometeor imaging probe. The combination of these
182 independent techniques characterizes cloud particles within a range of diameters varying from 0.5
183 μm to 2 mm.



184 The PN (Gayet et al., 1997) measures the angular scattering coefficients (non-normalized scattering
185 phase function) of an ensemble of water droplets or ice crystals or a mixture of those particles ranging
186 from a few micrometers to approximately 1 mm in diameter. These particles intersect a collimated
187 laser beam, at a wavelength of 804 nm, near the focal point of a parabolic mirror. The light scattered
188 at angles from 3.49° to 172.5° is reflected onto a circular array of 56 near-uniformly positioned
189 photodiodes. In this study, reliable measurements were performed at 30 scattering angles ranging
190 from ±15° to ±162°. The measurements allow to distinguish particle phase (water droplets or ice
191 crystals) and to derive single scattering properties such as the extinction coefficient and the
192 asymmetry coefficient with uncertainties of 25% and 4%, respectively (Gayet et al., 2002; Jourdan et
193 al., 2010).

194 Particle size distributions and corresponding microphysical and optical integrated properties (IWC,
195 Deff, N, and extinction) were derived from FSSP-300 measurements (Baumgardner et al., 1992). This
196 instrument measures the intensity of forward scattered light from cloud particles passing through the
197 laser beam, with cloud particles in the diameter range 0.35-20 µm. In the forward angular region
198 (from 4° to 12°), scattering is mainly described by the particle diffraction pattern and therefore
199 depends on the refractive index, the shape, and the size of the particles. The method of data processing
200 and size calibration used during the CONCERT campaigns have been presented in Gayet et al. (2012)
201 (Appendix A). We briefly recall that the asymmetry parameter derived from the PN was used to
202 discriminate nearly spherical particles ($g \geq 0.85$) from non-spherical ones ($g < 0.85$) at 804 nm. For
203 spherical ice particles, Mie calculations were used to derive the size bin limits and the corresponding
204 extinction efficiency. Results were adjusted to the calibrated probe response. Additionally, in order
205 to minimize Mie ambiguities related to the FSSP-300 size response, 31 channels were rebinned to 13
206 channels with a diameter ranging from 0.5 µm to 18 µm (upper channels 30 and 31 were excluded
207 from the data analysis). For non-spherical particles, the size of the contrail particles is expressed in
208 terms of an equivalent surface or area diameter, i.e. the diameter of a sphere that has the same area
209 than the projected area of the measured non spherical particle image (Mishchenko et al., 1997;
210 Schumann et al., 2011). The particles were assumed to be rotationally symmetric ice ellipsoids with
211 an aspect ratio of 0.5. Accordingly, and contrary to the method used for spherical particles, 15 size
212 bins ranging from 0.5 µm to 18 µm were defined based on T-Matrix calculations following Borrmann
213 et al., (2000).

214 The bi-dimensional optical array spectrometer probe (2DC) provides information on the crystal size
215 and shape within a nominal size range from 25 µm to 800 µm by recording cloud particles shadow
216 images with a 25 µm resolution. The method of data processing used in this study is described in
217 detail in Gayet et al. (1996) and Febvre et al. (2009). Reconstruction of truncated particles has been
218 considered for the PSD calculations and the sampling surfaces have been derived according to
219 Heymsfield and Parrish (1978). In order to improve the statistical significance of low particle
220 concentrations, a 5-s running mean was applied. The bulk parameters were calculated assuming the
221 surface-equivalent diameter relationships of Heymsfield (1972) and Locatelli and Hobbs, (1974). As
222 the sensitivity of the probe to small particles decreases with airspeed (Lawson et al., 2006), particles
223 smaller than 100 µm may not be detectable at the Falcon airspeed of typically 180 m s⁻¹. This may
224 result in larger uncertainties of up to 100% in the derived microphysical parameters such as the IWC
225 (Gayet et al., 2002 and 2004).

226 Depending on the spherical or non-spherical shape of ice crystals, ice water content IWC, extinction
227 coefficient Ext, and effective diameter D_{eff} were calculated independently according to Garret et al.
228 (2003) and Gayet et al. (2012). For spherical ice crystals ($g_{PN} \geq 0.85$), optical and microphysical
229 properties are calculated from the following equations:



$$Ext = \frac{\pi}{4} \sum_i \beta_{ext}^i N_i D_i^2 \quad (1)$$

$$IWC_{spherical} = \frac{\pi}{6} \rho_{ice} \sum_i N_i D_i^3 \quad (2)$$

230 where β_{ext}^i is the extinction efficiency (depending on spherical or aspherical particle characterization),
 231 D_i the mean diameter in channel i , N_i the number concentration, and ρ_{ice} the bulk ice density (0.917
 232 g cm^{-3}).

233

234 For non-spherical ice crystals ($g_{PN} < 0.85$ and for particle diameters larger than $70 \mu\text{m}$), an equivalent
 235 diameter method is used (Gayet et al., 2004). For an ice crystal with an area A , the particle equivalent
 236 diameter D_{equ} and equivalent mass x_{equ} are defined as :

237

$$A \leq 0.049 \text{ mm}^2 \quad D_{equ} = 0.82 A^{0.48} \quad (3)$$

$$A > 0.049 \text{ mm}^2 \quad D_{equ} = 0.56 A^{0.32} \quad (4)$$

$$x_{equ} = \frac{\pi}{6} \rho_{water} D_{equ}^3 \quad (5)$$

$$IWC_{non-spherical} = \rho_{ice} \sum_i N_i x_{equ} \rho_{water} \quad (6)$$

238 With ρ_{water} the bulk water density (1 g cm^{-3}).

239 In young tropospheric aircraft plumes, the main chemical component to be measured is NO_y , mainly
 240 composed of NO and NO_2 . During CONCERT campaigns trace gas measurements of NO and NO_y
 241 mixing ratio were performed using the chemiluminescence technique (Schlager et al., 1997) with a
 242 time resolution of 1 s. The accuracy (and precision) of the NO and NO_y measurements are estimated
 243 with 7% (and 10%) and 10% (and 15%), respectively (Ziereis et al., 2000).

244 Chemical ionization mass spectrometry combined with a SF_5^- ion source was used to detect the
 245 concentration of HNO_3 , SO_2 , and HONO in the exhaust plumes and the UTLS (Jurkat et al., 2010,
 246 Jurkat et al., 2011). Mass spectra were sampled with an ion trap mass spectrometer with resolution of
 247 < 0.3 atomic mass units and averaged over five spectra resulting in an overall time resolution of 1.6
 248 s. Detection limits for HONO and SO_2 for 1.6 s time resolution were 72 and 67 pmol mol^{-1} and for
 249 HNO_3 36 pmol mol^{-1} for 32 s time resolution. During CONCERT 2011 a quadrupole mass
 250 spectrometer was employed on the Falcon (Voigt et al., 2014) with detection limits of 15 pmol mol^{-1}
 251 for HNO_3 and 8 pmol mol^{-1} for SO_2 at 20 s time resolution (Jurkat et al., 2016).

252 The detection of water vapor and relative humidity with respect to ice (RHI) is important to
 253 characterize contrail ice crystals. Water vapor has been measured with the chemical ionization mass
 254 spectrometer AIMS-H₂O during CONCERT-2 (Kaufmann et al., 2014; 2016). In addition,
 255 hygrometers using the Lyman- α technique (Zöger et al., 1999; Meyer et al., 2015), and frost point
 256 hygrometers (Heller et al., 2017) were implemented on the Falcon during CONCERT-1 and 2.

257 3 Results

258

259

3.1 Overview of the cloud properties sampled during the reference cases



260 The purpose of this section is to give an overview of the contrail optical properties and more
261 interestingly to evaluate the ability of the Polar Nephelometer measurements to identify contrails.
262 Two flights, performed on 16 September 2011 during CONCERT-2 (flight 16b) and on 19 November
263 2008 during CONCERT-1 (flight 19b), respectively, were selected for their variety of contrails and
264 cirrus sampled during these two flights. The two flights are considered as a benchmark to illustrate
265 the potential of the PCA methodology described in Sect. 3.2.

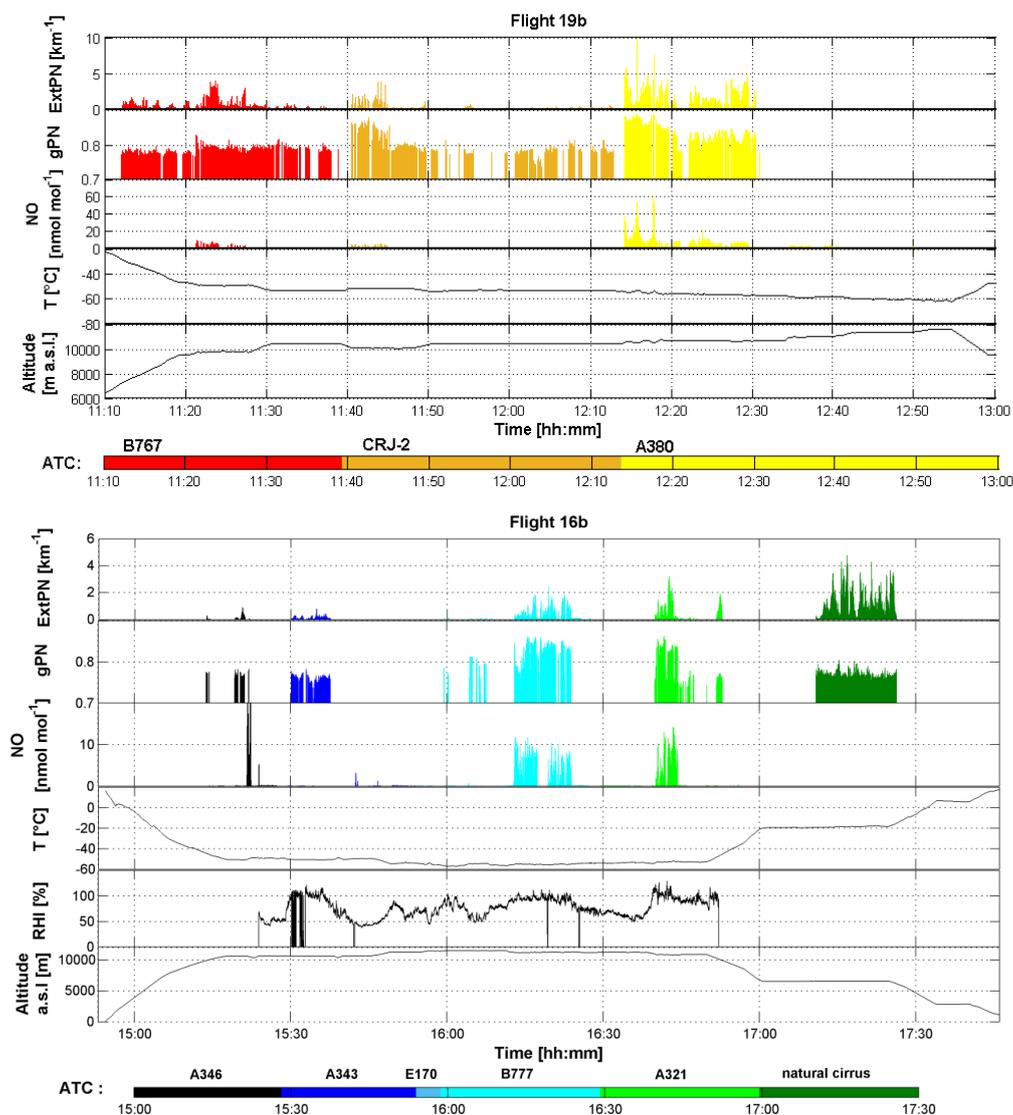


Figure 1 : Time series at 1 s resolution for flights a) 19b and b) 16b. From top to bottom: extinction coefficient (in km^{-1}) and asymmetry parameter measured by the Polar Nephelometer at 804 nm, concentration of nitric oxide (in nmol mol^{-1}) measured by chemiluminescence technique, temperature (in $^{\circ}\text{C}$), relative humidity with respect to ice (in %), and altitude a.s.l. (in m). Temporal series are colored according to time, and grey indicators illustrate contrail information from Air Traffic Control (ATC) information as provided by a dedicated flight controller.



266 Figure 1 displays the time series of the extinction coefficient (ExtPN) at 804 nm, asymmetry
267 parameter (gPN) at 804 nm, relative humidity (RHI), and the nitric oxide (NO) concentration for
268 flights 19b and 16b. RHI measured with the AIMS mass spectrometer is shown for flight 16b. For
269 RHI measurements on flight 19b we refer to Kübbeler et al. (2011); Gayet et al. (2012); Jessberger et
270 al. (2013); Schumann et al. (2013). For both flights, Air Traffic Control (ATC) provided information
271 on the aircraft characteristics (aircraft type, engine type, fuel flow, weight, engine power setting,...)
272 responsible for the formation of encountered contrails. In a first classification approach, the time
273 series are color coded according to the ATC information of consecutive sampling of different
274 contrails.

275 The PN Extinction coefficient coupled with the asymmetry parameter seems to be a reasonable proxy
276 to detect contrails and cirrus clouds (see amongst other references, Voigt et al., 2010). ExtPN values,
277 by definition, depend on the cloud particle concentration and size. Values typically beyond 0.1 km^{-1}
278 correspond in general to cloud events well correlated to supersaturation with respect to ice conditions
279 ($\text{RHI} > 100\%$). Figure 1 proves that relatively high values of extinction can be found in flights 19b
280 and 16b which are linked to the presence of contrails or cirrus clouds. Moreover, the temporal
281 distributions of these values are coherent with ATC information for both flights. For instance, most
282 of the contrails induced by commercial aircraft exhaust plumes translate into significant extinction
283 coefficient values. The ExtPN values are between 0.2 km^{-1} and 10 km^{-1} for contrails induced by A346,
284 A340, and A380 commercial aircraft. Cirrus clouds are detected with more variable extinction values
285 mostly larger than 0.5 km^{-1} . Most of the aircraft induced contrails are detected by the PN with the
286 exception of the ones stemming from the E170 airplane. At 15:50 during flight 16b, ATC identified
287 the E170 position close to the Falcon flight trajectory, however the ExtPN and the NO mixing ratio
288 remained very low. Hence, the E170 contrail was not probed by Falcon. In the following we assume
289 that only periods with ExtPN values above 0.1 km^{-1} are considered a reasonably reliable signature for
290 contrails sampled during the flight campaigns.

291 The absolute values of the asymmetry parameter gPN provide additional information of the cloud
292 particle shape. Indeed, gPN is a good indicator of the degree of sphericity of ice crystals (Gayet et al.,
293 2012). Ice clouds with gPN values higher or equal to 0.85 are typically composed of spherical ice
294 crystals, whereas lower values are indicative of aspherical ice particles. In a supersaturated
295 environment of contrails, crystals grow by water vapor deposition and become increasingly aspherical
296 with time. This is why spherical ice crystals prevail in very young contrails with an asymmetry
297 coefficient around 0.85 with RHI above 100%. Subsequently, gPN is decreasing when water vapor
298 diffusion is generating more and more aspherical crystal shapes at ice supersaturation. This can be
299 observed for A321 chasing during flight 16b with gPN decreasing to 0.75 whilst RHI remains around
300 100%, whereas for B777 chasing, no gPN decrease is observed at $\text{RHI} < 80\%$. Also natural cirrus
301 clouds are mainly composed of non-spherical ice crystals, possibly with hexagonal shapes. These
302 clouds can be easily discriminated from young contrails as they exhibit a much lower asymmetry
303 parameter typically below 0.75 (see amongst others Jourdan et al., 2003b, Febvre et al., 2009).
304 However, no accurate ambient RHI data can be retrieved for measurements in natural cirrus due to
305 instrumental calibration problems. A good example of the evolution of gPN is the CRJ-2 contrail
306 observed between 11:40 and 11:45 during flight 19b. The sequence illustrates the potential of the
307 gPN measurement to characterize the evolution of contrail properties, with decreasing crystal
308 sphericity documented by the decreasing asymmetry parameter from 0.88 to 0.79 (uncertainties
309 around 0.04) after only 5 min and down to 0.77 after 20 min. A more stable variation of gPN values
310 (around 0.78 ± 0.02) is then observed until 12:10 after 30 min of contrail ageing associated with
311 crystal growth by water vapor diffusion. A similar decrease in gPN has been noted by Gayet et al.
312 (2012) in the ageing contrail from an A380 aircraft and is visible for the B767 and the A321 contrails.



313 NO concentration measurements can also be used to discriminate natural cirrus clouds from ice clouds
314 influenced by aircraft traffic. At the typical altitude of 10 km, NO environmental concentrations are
315 close to background values. In contrast, NO concentrations in young contrails may reach several tens
316 of nmol mol^{-1} (Voigt et al., 2010). Figure 1 shows a good correlation between the expected
317 localization of young contrails and NO concentrations. The dilution effect into the upper troposphere
318 causes an important decay of chemical concentrations. For instance, the first few seconds of the A380
319 chasing during flight 19b are characterized by a high NO concentration (up to 40 nmol mol^{-1}) followed
320 by a fast decrease to 10 nmol mol^{-1} in the next 15 min, and less than 5 nmol mol^{-1} beyond 15 min.
321 NO concentrations finally decrease to background levels within hours (e.g. Voigt et al., 2017). This
322 decrease of the NO concentration is in accordance with the decrease of the extinction coefficient
323 (from 10 to 0.2 km^{-1}) and asymmetry parameter (from 0.88 to 0.77). Thus NO was mainly used as
324 additional contrail indicator. However, during some aircraft chasing events, NO concentrations were
325 near background levels, while mass spectrometric measurements (not shown here) indicate elevated
326 concentrations of HONO, HNO_3 , and SO_2 representative for contrail chemical species.

327 The above case studies of flights 19b and 16b clearly show that the optical properties of contrail type
328 ice clouds (supported by the ATC information) in conjunction with specific trace gas concentration
329 measurements can be used to discriminate contrails from natural ice cloud events. A first order
330 analysis of these parameters can be used to roughly distinguish young contrails (mostly quasi-
331 spherical ice crystals) from aged contrails (mostly aspherical ice crystals) and natural cirrus
332 (background NO concentrations). This analysis is mainly qualitative and based solely on a few
333 integrated parameters (Fig. 1). A more robust statistical method should be used to accurately separate
334 the different contrail phases and also natural cirrus. In the following section, relationships between
335 contrail and ice cloud properties and their scattering properties are investigated more extensively to
336 assess whether the information content of the PN scattering measurements is sufficient to document
337 changes in the contrail microphysical properties.

338 **3.2 Statistical Method**

339 In this section, we present a methodology based on the statistical analysis of the optical signature of
340 ice clouds and in particular contrails. The goal is to classify the contrail properties according to the
341 aircraft origin and evolution stage. The main objective of the Principal Component Analysis (PCA)
342 is data reduction in order to allow a better physical interpretation of the light scattering patterns
343 derived from the Polar Nephelometer measurements (Legendre and Legendre, 1998; Jourdan et al.,
344 2003). In this study, optical properties of ice crystals in the evolving contrail environment are
345 analyzed to evaluate contrail evolution. This statistical analysis was already successfully applied to
346 discriminate mixed phase clouds (Jourdan et al., 2010), liquid clouds, and ice clouds, (Jourdan et al.,
347 2003) as well as to characterize porous aerosol in degassing plumes (Shcherbakov et al., 2016) using
348 light-scattering properties measured by the Polar Nephelometer.

349 **3.2.1 Reference definition**

350 The PCA is first applied to the PN angular scattering coefficients measurements performed during
351 flights 16b and 19b which are here considered as our reference datasets. Initially, a correlation matrix
352 is calculated to characterize the link between each scattering angle. The PCA is designed to generate
353 a new limited set of uncorrelated parameters, called principal components C_{ij} representative of the
354 original data set variability.

355 A first implementation of the PCA is performed to detect unreliable data or out of order photodiodes.
356 For instance, seven photodiodes presented a low signal to noise ratio and were excluded from the



357 dataset. Flight sequences characterized by $\text{ExtPN} < 0.1$ were also removed. Finally, flight sequences
 358 dedicated to aircraft chasing and ice cloud sampling were considered to perform a second PCA.

359 Then, the analysis is performed on the remaining angular scattering coefficients (4669 Angular
 360 Scattering Coefficients (ASC) representing PN measurements of flights 16b and 19b) now restricted
 361 to 25 angles θ ranging from 15° to 155° . The new set of variables or coordinates, C_{ij} , can be expressed
 362 with the scalar product of the vector of reduced angular scattering coefficients $\overline{\sigma}_j(\theta)$ for the j^{th}
 363 measurements, expressed in log scale, and the l^{th} eigenvector $\xi_l(\theta)$ (i.e. principal component) of the
 364 total data set correlation matrix (Jourdan et al., 2010).

$$C_{ij} = (\overline{\ln \sigma_j} - \langle \overline{\ln \sigma} \rangle)^T \cdot \overline{\xi}_l \quad (4)$$

365 where $\langle \overline{\ln \sigma} \rangle$ represents the average ASC of the dataset.

366 The first three eigenvectors $\overline{\xi}_l(\theta)$ of the correlation matrix are displayed in Fig. 2 along with their
 367 normalized eigenvalues λ_l , representing more than 99% of the variability of the PN angular scattering
 368 coefficients (ASC).

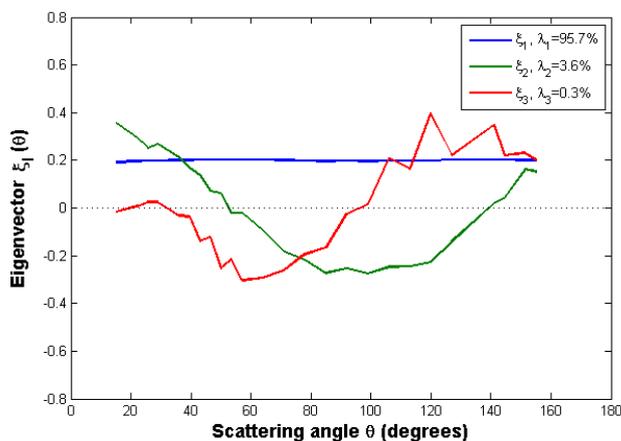


Figure 2 : First three eigenvectors for the flights 16b and 19b.

369 The first eigenvector $\xi_1(\theta)$ is approximately constant versus scattering angle and represents 95.7%
 370 of the total variance. It means that this principal component is representative of changes of the
 371 magnitude of phase functions without any changes in their global shape. This behavior means that
 372 95.7% of the ASC variations are linked to changes of the cloud particle extinction. Results show a
 373 good correlation ($r^2 = 0.98$) between the first eigenvector and the extinction derived from the PN
 374 measurements (ExtPN).

375 The second eigenvector $\xi_2(\theta)$ reverses sign twice at scattering angles equal to 50° and 140° with an
 376 extremum around 90° . Accordingly, 3.6% of the angular scattering variability corresponds to a
 377 redistribution of scattered energy from the angular region (50° - 140°) to scattering angles lower than
 378 50° and higher than 140° . Light-scattering modeling studies demonstrate that the scattering behavior
 379 in the angular region between 60° and 140° is sensitive to the particle shape and thermodynamic
 380 phase (Jourdan et al., 2010). A strong linear correlation ($r^2=0.97$) between the second eigenvector and
 381 the asymmetry coefficient (gPN) at 804 nm is found.



382 The third eigenvector represents only 0.3% of the total variance. However, this eigenvector provides
 383 additional information in scattering regions which are not well described by the first two principal
 384 components. It has opposite signs in the angular region (30°-90°) and (90°-155°) with maximum
 385 extremal values at 60° and 120°. The shape of the third eigenvector describes the forward/backward
 386 hemisphere partitioning of the scattering. Baran et al. (2012), Xie et al. (2006), and Xie et al. (2009)
 387 showed that the scatter pattern for angles between 120° and 160°, corresponding to ice bow-like
 388 effects, is sensitive to quasi-spherical particles. Moreover, these backscattering angles ($\theta > 120^\circ$) and
 389 scattering angles around 22° and 46° (corresponding to halo features) can also be linked to the particle
 390 habits and surface roughness (Xie et al., 2009, Jourdan et al., 2010).

391 Based on these three first principal components, each phase function (or ASC) measured by the PN

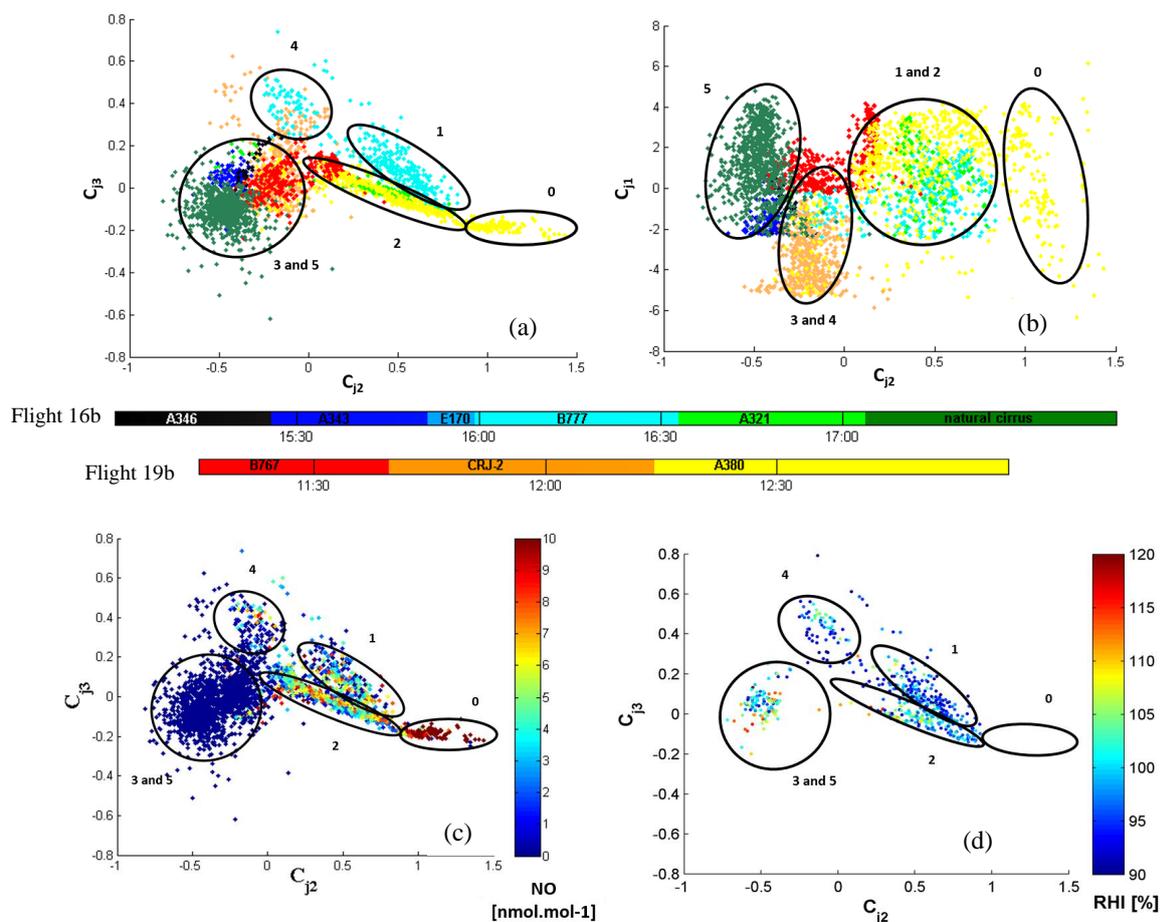


Figure 3: Expansion coefficient diagram for flights 16b and 19b (excluding 19b for RHI information): third versus second principal component for a), c) and d), and first versus second principal component for b). Data points are color coded according to ATC information for a) and b), by NO concentration for c), and by RHI values for d). The 6 typical scattering regimes (0-5) are indicated and numbered accordingly.

392 can be expressed with a good accuracy as a linear combination of these components (Jourdan et al.,
 393 2010). The PN data are projected into a new space defined by the three principal components (3D-
 394 space) instead of the 25-dimensional space of ASC. The scatterplots of the C_{j3} and C_{j1} expansion



395 coefficients versus the C_{j2} coefficient are represented on Fig. 3a and b respectively. Fig. 3a illustrates
396 the features of the ASC measurements in one of the most comprehensive way. Each point corresponds
397 to a measured phase function documented over 25 angles. The variability of C_{j2} coefficients is
398 significant with values ranging from -1 to 1.5. The angular variation of the second principal
399 component indicates that large values of C_{j2} ($C_{j2} > 0.75$) correspond to ASC with low side scattering
400 (60° - 130°) and higher forward scattering (15° - 40°) and somehow higher backscattering (145° - 155°).
401 This behavior is connected to an increase of the asymmetry parameter with an increase of C_{j2} values.
402 Thus, the fraction of spherical particles increases with increasing C_{j2} . In the region defined by
403 negative values of C_{j2} the density of points is relatively high. These cloud events exhibit optical
404 properties characterized by a large side scattering and low asymmetry parameter. Therefore, specific
405 cloud sequences sharing similar scattering properties can be identified based on this second principal
406 component. Young contrails characterized by quasi-spherical ice crystals have high positive values
407 of C_{j2} while cirrus clouds and contrail cirrus exhibit high negative values.

408 In the space of the third principal component high positive values of C_{j3} imply that less energy is
409 scattering in the forward hemisphere and thus more energy is scattered in the backward hemisphere.
410 The variability of the expansion coefficients is less pronounced as ASC are distributed between -0.4
411 and 0.6. Most of the measured ASC do not significantly differ from the average ASC in the angular
412 ranges (30° - 90°) and (90° - 155°). However, some specific clusters linked to scattering behavior can
413 be identified for values of C_{j3} greater than 0.1 and lower than -0.1. These threshold values also depend
414 of the position of the ASC on the second principal component. Finally, the first principal component
415 is directly linked to the extinction coefficient. High values of C_{j1} are representative of optically dense
416 cloud sequences.

417 Based on the time series displayed in Fig. 1, data points corresponding to contrails are also color
418 coded according to their aircraft origin illustrated on Fig. 3a and b. From these information and based
419 on the first three principal components, 6 clusters (see numbered ellipsoids in Fig. 3) representative
420 of particular scattering behavior can be roughly identified. Figure 3a suggests an increase of C_{j2} and
421 a decrease of C_{j1} with increasing aircraft size. Figure 3c shows an increase of C_{j2} for increasing NO
422 mixing ratio. Some contrails or ice cloud events are clearly delimited by a single area in the C_{j3} versus
423 C_{j2} and also C_{j2} versus C_{j1} diagrams. For instance cirrus clouds are gathered in cluster 5. Most of the
424 contrails induced by the B767, A340 and CRJ2 aircraft are associated to cluster 3 or 5. It means that
425 these cloud events share similar optical properties characterized by a low asymmetry parameter, high
426 side scattering behavior, and supersaturated ambient conditions with respect to ice. More
427 interestingly, Fig. 3 shows that some contrail events are smeared out over several areas or clusters.
428 Contrails relative to the A380 aircraft are dispatched in cluster 0 and 2 while the ones corresponding
429 to the B777 are spread out between clusters 1 and 4. This clearly indicates that the contrails are
430 evolving in space and/or time along the Falcon flight track while chasing the respective contrails.
431 This evolution can also be seen in the in-situ measurements of NO concentration color coded on Fig.
432 3c. Cloud regions influenced by air traffic can be discriminated from clouds formed by natural
433 processes based on the NO concentration values. While clusters 3 and 5 are characterized by very
434 low NO concentrations (close to zero) above background, clusters 0, 1, 2, and 4 correspond to higher
435 concentrations representative of a significant aircraft exhaust influence. For instance, a clear trend
436 shows that an increase of NO concentration translates into higher values of C_{j2} . Hence, contrails
437 characterized by a low side scattering due to the presence of spherical ice crystals correspond to high
438 NO concentration. This behavior can be a signature of young contrail properties. Elder or aged
439 contrails composed of a higher fraction of non-spherical crystals or growing more aspherically are
440 expected to exhibit an enhanced side scattering and a lower asymmetry parameter associated to lower
441 NO concentrations.



442 The contrail and cirrus classification based on ASC measurements appears to be consistent with the
443 independent trace gas measurements. Each cluster represented on Fig. 3 can be linked to a distinct
444 cloud event. Therefore, the combination of flights 16b and 19b can provide a relevant test-bed
445 database to discriminate contrail properties. Young contrails (spherical ice crystals) are associated to
446 clusters 0, 1 or 2, whereas aged contrails (aspherical ice crystals and higher RHI values) with more
447 pristine ice are categorized in clusters 3 and 4, and finally natural cirrus (low NO concentrations) are
448 found in cluster 5. A less precise analysis (using onboard camera) reveals that cluster 0 corresponds
449 essentially to the primary wake created below the secondary wake behind an aircraft. These different
450 clusters are defined arbitrarily according to ATC information and according to their optical
451 differences through the three first principal components. In the following, the clusters will be
452 referenced according to the contrail evolution stage

- 453 - Cluster 0 : Primary Wake (PW)
- 454 - Cluster 1 : Young Contrail 1 (YC1)
- 455 - Cluster 2 : Young Contrail 2 (YC2)
- 456 - Cluster 3 : Aged Contrail Clean (ACC)
- 457 - Cluster 4 : Aged Contrail (AC)
- 458 - Cluster 5 : Cirrus Cloud (CC).

459 Thus, the next step is to validate these cluster definitions according to the different tracers. One has
460 to keep in mind that some points are still arbitrarily attributed to a particular cluster without strong
461 physical justification.

462 3.2.2 Application to other CONCERT flights

463 In this section we investigate the possibility to complement the previous analysis with additional
464 cloud optical measurements performed during other CONCERT flights to increase the robustness of
465 the method.

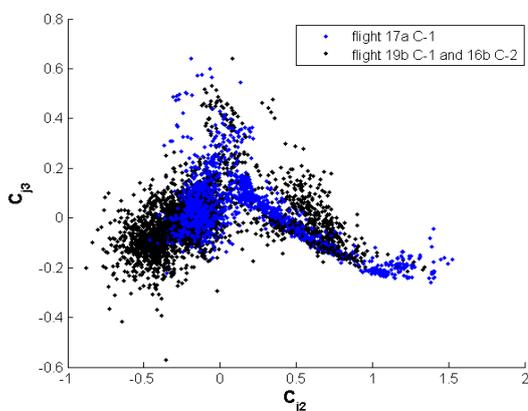


Figure 4: Example of data projection in the C_2/C_3 space where data from flight 17a (blue data points) are superimposed on the data from the benchmark flights 19b and 16b (black data points).

466 The principal components obtained on the basis of the measurements performed during flight 16b
467 and 19b are considered as our reference axes. Now, the ASC measured during other flights can be
468 projected on this space of principal components. The coordinates of these flights are calculated from



469 Eq. (4). An example of this data projection is illustrated in Fig. 4 where flight 17a is represented in
 470 the C_{j2}/C_{j3} space. Every PCA data point can be attributed to one cluster previously defined by the
 471 PCA implemented with flights 16b and 19b data (black points). In other words, every ASC measured
 472 during another flight can be merged (projected) into the expansion coefficient diagram relative to the
 473 base measurements performed during flights 16b and 19b. Data points sharing similar optical
 474 properties will still be close to each other. According to the cluster definition, Fig. 4 shows that
 475 different contrail phases have been experienced during flight 17a. Data points are mostly grouped
 476 into cluster ACC, but are also present in clusters AC, YC2, and PW. Finally, cloud data encountered
 477 during this flight are mainly categorized as young and aged contrails.

478 We follow this methodology to project and classify each additional “contrail” flight performed
 479 during both CONCERT campaigns. In order to attribute each measurement (or point) to one cluster
 480 and to enhance the statistical significance of our clustering analysis we compute the Mahalanobis
 481 distance (De Maesschalck et al., 2000). Clusters are defined by their means (or centers), standard
 482 deviations (or widths), and cross-correlations (or tilts). The Mahalanobis distance is given by the
 483 equation:

$$D_M(x)_i = \sqrt{(x - \mu_i)^T S_i^{-1} (x - \mu_i)} \quad (5)$$

484 with D_M the Mahalanobis distance between point χ and the i^{th} cluster center, μ_i the N-dimensional
 485 mean of this cluster and S_i its covariance matrix. Each data point can be associated to a specific cluster
 486 corresponding to the shorter Mahalanobis distance, and the ellipsoids’ eccentricity and width can be
 487 adjusted.

Day / Aircraft		Cluster	Cluster						Number of points	Age (s)
			PW	ACC	AC	YC1	YC2	CC		
			1st wake	aged contrail		young contrail		Cirrus		
17a C-1	TOTAL							1435		
	A340-311							359	61 - 144	
17b C-1	TOTAL							2715		
	B737-500							310	77 - 151	
	A340-642							100	82 - 139	
	NC							189	-	
19a C-1	TOTAL							2152		
	A319-111							628	94 - 129	
	A340-311							175	63 - 90	
19b C-1	TOTAL							1647		
	B767-300							319	77 - 107	
	CRJ-2							151	80 - 95	
	A380-841							677	109 - 240	
20 C-1	TOTAL							1434		
	B737-300							64	90 - 290	
16b C-2	TOTAL							1511		
	A340-600							128	100 - 132	
	B777							378	120 - 160	
	A321							135	70 - 95	
17 C-2	TOTAL							2904		
	NC1							498	-	
	NC2							233	-	
24 C-2	TOTAL							1380		
	B777							371	112 - 178	

Table 1: Classification relative to the six clusters on the C_{j2}/C_{j3} representation of the PCA of all data points for each flight of the two CONCERT campaigns (C-1 in November 2008 and C-2 in September 2011). The length of the bars represent the relative contribution of data points of individual contrails (blue bars) and also entire flights (black bars) to the 6 individual clusters.

488 The classification relative to the six clusters shown on the C_{j3} vs C_{j2} and the C_{j1} and C_{j2} expansion
 489 diagrams is summarized in Table 1. A total of 8 flights (6 additional flights) representing 4426 ASC
 490 measurements was processed. The lengths of the bars in Table 1 represent the relative contributions
 491 of data points to the different clusters: a) black bar merge cloud data points (with extinction coefficient



492 higher than 0.1 km^{-1}) for entire flights and b) blue bars present individual aircraft contrails within
493 specific flights. An important fraction (at least more than 30%) of data points is detected in clusters
494 ACC and CC for each flight during the two campaigns. This indicates that these data points are
495 sampled in aged contrail and sometimes natural cirrus. For flights more clearly performed in well
496 visible contrails outside natural cirrus (earlier development stage and/or intensified persistent elder
497 contrails), significant fractions of data points are associated to clusters PW, YC1, and YC2 (young
498 contrails) for both CONCERT-1 and CONCERT-2 campaigns. However, these flights are also
499 characterized by a significant contribution of data points to cluster ACC (aged contrails clean) and to
500 a minor extent in cluster AC (aged contrails, mostly corresponding to measurements performed
501 during two different B777 contrail chasing events).

502 These results are in reasonable agreement with previous conclusions (this subsection) about cluster
503 definitions and associated contrail / ice cloud characteristics. Very young contrails have been chased
504 during CONCERT-1 (flights 19a, 19b and 20) and during CONCERT-2 (flights 16b, 17 and 24).
505 Another interesting result is related to flight 17 during CONCERT-2 (flight 17 C-2). No contrail
506 information has been communicated from ATC, however the Falcon has been flying apparently in
507 visible contrails, probably too old for ATC recognition. Data points can mainly be attributed to cluster
508 CC, and to a minor extent to cluster ACC and cluster YC2. This observation suggests that significantly
509 aged contrails have been sampled, resembling strikingly natural cirrus clouds. Indeed, crystal
510 formation and growth processes in contrails and natural cirrus suggest that very old contrails more
511 and more resemble natural cirrus properties. From Table 1 it is obvious that an important amount of
512 data points had been sampled in natural cirrus during this flight. All these natural cirrus data points
513 appear in the black bars but only to a minor extent in the blue bars limited to ATC communicated
514 contrail sequences.

515 ATC information on contrail ages has been collected during each chasing. Some chasings have been
516 performed less than 100 s after contrail formation. This is the case for the A340 contrail during flight
517 19a and for the CRJ-2 contrail during flight 19b of CONCERT-1 and for the A321 contrail during
518 flight 16b of CONCERT-2. One can notice that the contrail ages are well correlated to chosen cluster
519 definitions, revealing that contrail data relative to the A340 are included in cluster YC1 and YC2
520 (young contrails) for more than 53% of the data points, and nearly 65% for the CRJ-2 and 88% for
521 the A321. According to our cluster classification, only 12% of the data points gathered during these
522 three flights correspond to aged contrail (cluster ACC and AC) categories in contrast to other
523 CONCERT-1 and CONCERT-2 flights (with more than 30% of data points associated to ACC and
524 AC). Even though it is still difficult to associate contrail ages to measurement points, the “contrail
525 age” ranges are in agreement with the cluster definitions.

526 **4 Evolution of contrail properties**

527 **4.1 Optical and chemical cluster properties**

528 As demonstrated in the previous part, cloud events can be separated according to their light-scattering
529 properties. Six clusters were defined based on two flights with significant number of data points for
530 each cluster. In this section we present mean optical, chemical, and microphysical properties for each
531 of the six clusters. These mean properties have been calculated over all data points associated to the
532 6 individual clusters (all flights, both CONCERT campaigns). Figure 5a, 5c, and 5d illustrate the
533 normalized frequency distributions of the asymmetry parameter (gPN), the extinction coefficient
534 (ExtPN), and NO concentrations for the six clusters, respectively. Figure 5b represents mean
535 normalized scattering phase functions, also for the 6 clusters.



536 The asymmetry parameter gPN statistics for the six clusters shown in Fig. 5a provide the most
537 relevant information on cloud characteristics and the related context of contrail evolution/age. In
538 agreement with findings in Gayet et al. (2012), aged contrails (cluster ACC and AC) and natural
539 cirrus (cluster CC) correspond to gPN values between 0.72 and 0.80, younger contrails (cluster YC1
540 and YC2) to gPN between 0.80 and 0.86, and primary wake measurements (cluster PW) to gPN above
541 0.86. This result can be explained with the time evolution of ice crystal shapes after exhaust from
542 quasi-spherical ice particle to, e.g. column, needle, bullet, and bullet-rosette type crystals. In the
543 primary wake, the pressure increases associated in the descending vortex. This leads to adiabatic
544 heating and subsequent sublimation processes of the ice crystals (Lewellen and Lewellen, 2001;
545 Unterstrasser et al., 2016) and explains a return to spherical shapes and high values of the asymmetry
546 coefficients.

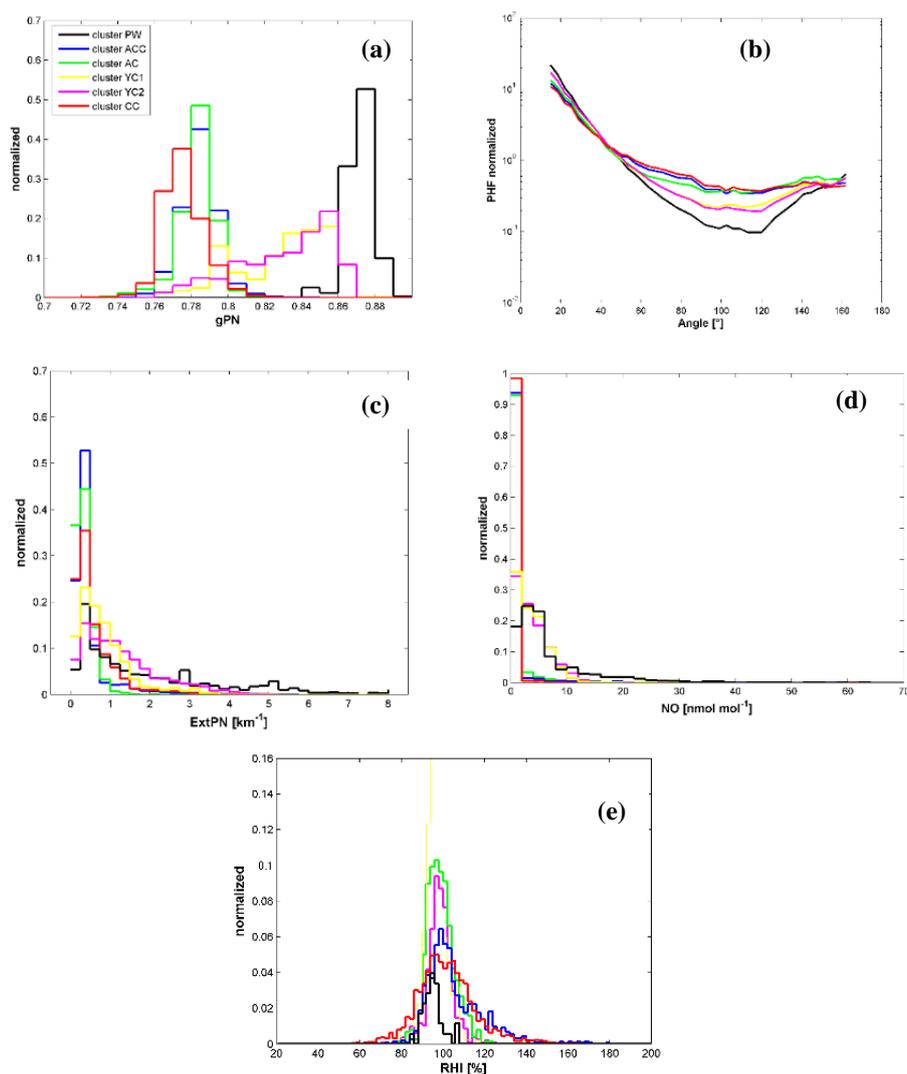


Figure 5: Normalized histograms of a) asymmetry coefficient, b) phase function, c) extinction retrieved by Polar Nephelometer, d) NO concentration for all flights, and e) RHI conditions for CONCERT-2 flights.



547 The normalized phase functions are presented in Fig. 5b. Primary wake phase functions (cluster PW)
548 are clearly different from young contrail phase functions (cluster YC1 and YC2), which are
549 themselves different from aged contrails (cluster ACC and AC) and natural cirrus (cluster CC) phase
550 functions. The main difference between the averaged phase functions is observed for the side
551 scattering region (50° - 140°) which is related to changes of ice particles shapes and to the proportion
552 of spherical ice crystals within the contrails. This behavior is expected and also in agreement with
553 position of cluster PW, YC2 and YC1 on the expansion coefficient diagram (Fig. 2). The decrease of
554 the C_{j2} coefficient is associated to a side scattering enhancement. Therefore, very young contrails
555 composed of a majority of spherical ice crystals are characterized by phase functions with a
556 substantial scattering at forward angles associated with much lower scattering at sideward angles. As
557 the contrails evolve these features smooth out leading to phase functions with a featureless behavior
558 and a more flat appearance at side scattering angles. Finally, the averaged normalized phase functions
559 of old contrails and natural cirrus are resembling each other. This also explains that they are difficult
560 to discriminate with the PCA.

561 The extinction coefficient statistics are presented in Fig. 5c. All the aged contrails (cluster ACC and
562 AC) exhibit extinction coefficients lower than 2 km^{-1} . The same statement applies for 80% of the
563 sampled natural cirrus (cluster CC). For younger contrails (cluster YC1 and YC2) extinction
564 coefficients can reach 5 km^{-1} . Largest extinction coefficients are achieved in primary wake
565 measurements sorted into cluster PW with extinction coefficients reaching values up to 8 km^{-1} . Still,
566 the main fraction (more than 50% of data points) of young contrail data yields extinction coefficients
567 between 0 and 1 km^{-1} .

568 Concentrations of chemical species also allow characterizing contrail/cirrus cloud data. The
569 concentration depends strongly on the type of the pursued aircraft. Figure 5d shows mean
570 concentrations of nitrogen oxide NO data points attributed to the six individual clusters. Young
571 contrail NO concentrations (cluster PW, YC1 and YC2) can reach values up to 10 nmol mol^{-1} and up
572 to 60 nmol mol^{-1} in the primary wake. In contrast, in aged contrails and in natural cirrus (cluster ACC,
573 AC and CC) NO concentrations do not exceed 10 nmol mol^{-1} (which is true for more than 97.5%,
574 99.6%, and 99.7% of data points for clusters ACC, AC, and CC, respectively). Indeed, after exhaust,
575 concentrations of nitrogen oxide NO and sulfur dioxide SO_2 created by combustion reactions decrease
576 rapidly due to the dispersion in the upper troposphere and reactions with other molecules.

577 Finally, saturation conditions with respect to ice are presented in Fig. 5e for all clusters and for
578 CONCERT-2 flights only. The predominant measured ambient relative humidity in all clusters is
579 around 95%. Cluster ACC and CC (blue and red lines) show higher RHI values (more than 120%)
580 than other clusters. Thus, this can explain the formation of natural cirrus and persistent contrails for
581 these ambient conditions. Contrary to all other clusters, no supersaturation is observed for cluster PW
582 (in black), defined as primary wake measurements. This result is in agreement with the definition of
583 the primary wake, which is still in the non-persistent phase of the contrail.

584 The above results highlight that the principal component analysis, based on the ASC measurements
585 described in Sect. 3, allows to discriminate different types of contrails. Specific optical and chemical
586 properties can thus be characterized for each contrail phase and can be related to their evolution. An
587 interesting aspect is that the PCA analysis facilitates to connect clusters of optical properties to
588 microphysical characteristics of the contrails within specific clusters.

589 **4.2 Microphysical cluster properties**

590 Microphysical properties are assessed using the combination of FSSP-300 and 2DC measurements.
591 They have been analyzed for each cluster for hydrometeor diameters between $0.5 \mu\text{m}$ and $800 \mu\text{m}$.



592 Figure 6 shows mean volume particle size distributions (PSD) for all six clusters measured from the
593 FSSP-300 and 2DC with respective limited instrumental size ranges. A linear interpolation in
594 logarithmic space has been applied for each PSD in the size range from 17 μm to 70 μm which is not
595 accurately documented by the two instruments. These results include all flights of the study (8 flights
596 from CONCERT-1 and 2). It is important to note that more than 500 data points are included in each
597 cluster with a maximum of 6300 data points for cluster ACC.

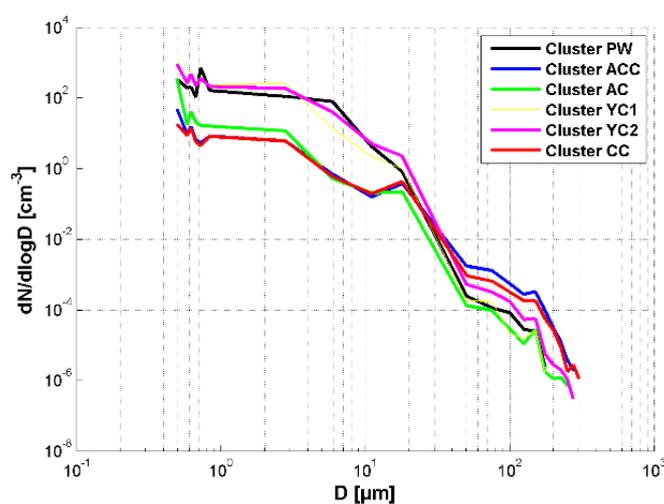


Figure 6: Number particle size distributions for each cluster including all data points of all flights. FSSP-300 measurements from 0.5 to 17 μm and 2DC measurements from 70 μm to 800 μm . The data are linearly interpolated in logarithm space between 17 μm and 70 μm .

598 Figure 6 shows that the mean number PSDs for each cluster are consistent with the cluster definition.
599 Indeed, two categories of PSD can be observed. Within the FSSP-300 size range, PSD relative to old
600 contrails (cluster ACC and AC) and cirrus (cluster CC) exhibit more than one order of magnitude
601 lower number concentrations of small ice crystal compared to young contrails (cluster YC1 and YC2)
602 and primary wake (cluster PW). Within the two groups, no significant differences can be noticed due
603 to uncertainties of the FSSP-300 number concentration measurements of 30% for concentrations
604 around 5cm^{-3} and 75% for concentrations around 0.5cm^{-3} (Gayet et al., 2002). The differences
605 between these two groups can be explained by the production of small ice crystals in fresh exhaust
606 plumes followed by rapid dilution during subsequent minutes after the exhaust. Within the 2DC
607 range, the PSDs are also in agreement with the cluster definitions. A higher concentration of large ice
608 crystals with diameters around 100 μm and beyond are expected for natural cirrus and significantly
609 well-developed contrails. This is particularly well illustrated on the PSD in the 2D-C size range where
610 a higher concentration of large ice crystal is observed for clusters ACC and CC compared to the
611 younger contrails. However, these PSD do not allow discriminating young contrails in primary wake
612 (cluster PW) from contrails in the secondary wake (cluster YC1 and YC2).



Extinction (km^{-1})		Mean	std	Mediane	prctile 25	prctile 75
cluster	PW	3,904	4,386	1,972	0,528	6,125
	ACC	0,088	0,282	0,028	0,009	0,057
	AC	0,079	0,186	0,024	0,001	0,071
	YC1	2,037	2,363	1,307	0,440	2,667
	YC2	2,163	2,816	1,227	0,316	2,634
	CC	0,113	0,237	0,065	0,030	0,127

IWC (mg m^{-3})		Mean	std	Mediane	prctile 25	prctile 75
cluster	PW	15,46	21,56	6,26	0,87	22,79
	ACC	8,74	37,77	0,35	0,01	3,33
	AC	1,65	19,76	0,02	0,00	0,22
	YC1	5,29	11,45	1,46	0,17	5,78
	YC2	7,85	23,40	1,36	0,14	7,77
	CC	28,69	145,73	0,96	0,05	2,87

NTOTAL (cm^{-3})		Mean	std	Mediane	prctile 25	prctile 75
cluster	PW	125,94	98,37	109,54	54,24	166,57
	ACC	5,57	17,86	1,29	0,76	2,41
	AC	17,89	33,48	2,92	1,21	32,90
	YC1	155,80	159,42	106,65	38,43	207,70
	YC2	164,20	173,17	103,81	42,26	211,85
	CC	6,06	10,12	3,75	2,17	6,81

Table 2: Optical and microphysical properties for each cluster according interpolated particle size distributions from FSSP-300 and 2DC measurements.

613 Table 2 presents ice water content (IWC, in mg m^{-3}) and total number of ice crystals (NTOTAL, in
 614 particles cm^{-3}) derived from the measured PSD for each cluster. The extinction coefficient (in km^{-1})
 615 obtained from the PN measurements is also displayed. Despite the large uncertainties associated to
 616 both instruments and the interpolation method (for ice crystals with diameters ranging from 17 μm
 617 and 70 μm), these results again prove that each cluster can be connected to specific contrail phases.
 618 The microphysical and optical properties of cluster PW are in agreement with the cloud properties
 619 excepted in the primary wakes. The extinction coefficient has a mean value of 3.9 km^{-1} , IWC is close
 620 to 15.5 mg m^{-3} , and the number concentration yields a typical value of 125 particles cm^{-3} . Young
 621 (clusters YC1 and YC2) and aged contrails (clusters ACC and AC) exhibit distinctive differences in
 622 their optical and microphysical properties. Higher extinction coefficients and ice number
 623 concentration, 2 km^{-1} and 160 cm^{-3} , respectively, characterize young contrails compared to aged
 624 contrails with 0.08 km^{-1} and 10 particles cm^{-3} , respectively. Cluster CC corresponds to natural cirrus
 625 clouds where significant atmospheric spreading and ice growth occurred. Thus, within this cluster the
 626 extinction coefficients (mean values of 0.1 km^{-1}) as well as the number concentration of ice crystals
 627 (around 6 particles cm^{-3}) are very low. The IWC is higher with a mean value of 28.7 mg m^{-3} due to
 628 ice crystal growth in supersaturated conditions.

629 However, it is difficult to discriminate young contrail cases (YC1 and YC2) based on their
 630 microphysical properties. Clusters ACC and AC microphysical properties are also similar but ACC
 631 IWC and number concentrations are closer to the ones of the cirrus case indicating a more evolved
 632 stage of the observed ACC contrail cluster.

633 Conclusions

634 In this study, a new form of statistical analysis of contrail to cirrus evolution is presented based on
 635 two intensive contrail measurement campaigns, CONCERT-1 and CONCERT-2. The data are used



636 to study optical and microphysical properties of contrails during their evolution from young contrails
637 to contrail-cirrus clouds, and ambient natural cirrus clouds. The combination of optical,
638 microphysical, and chemical airborne measurement data was used to present an extended view of
639 cloud properties, and to merge those results with ATC flight information about sampled contrails.

640 A Principal Component Analysis (PCA) methodology has been applied to the measured Polar
641 Nephelometer scattering phase function data in order to facilitate a distinction between cloud
642 properties in different contrail phases. The PCA results were derived first for two reference flights
643 that sampled contrails and cirrus in various development stages, including the primary wake, the
644 young secondary wake, old contrails (few minutes after formation) and natural cirrus. For these
645 flights, the PCA clearly demonstrates the potential to separate different groups of clouds, justifying
646 the use of these two flights as a benchmark. The scattering phase functions measured during other
647 CONCERT flights were then projected into the space of principal components obtained from the two
648 reference flights. Mahalanobis distances were used to measure the separation between the additional
649 data points and the data in the predefined clusters. From the entire data set, the cloud properties in the
650 various contrail development stages can be separated and analyzed separately. The analysis
651 demonstrates that the clearest separation between clusters is related to particle shape, which is
652 significantly controlling the scattering phase function and the derived asymmetry parameter gPN. The
653 asymmetry parameter clearly separates young contrails (gPN of 0.72 to 0.80) from contrail/cirrus
654 with gPN ranging from 0.80 to 0.88. Since it is still difficult to evaluate the exact age of each
655 measurement, young and aged contrails are classified also by their optical and chemical properties.
656 The measured NO concentrations are useful to distinguish natural cirrus from old contrails. Despite
657 the important gap between the two instruments used to measure particle size distributions, particle
658 size spectra and related mean values of the ice particle number concentration, extinction and ice water
659 content have been determined for each cluster. The various clusters clearly show different size
660 distributions. In good agreement with previous findings on optical and chemical properties, we find
661 that young contrails have more than a factor of ten higher number concentrations of small ice crystals
662 (with diameters lower than 20 μm) than aged contrails and natural cirrus. On the other hand, aged
663 contrails and natural cirrus contain larger ice crystals, with diameters larger than 75 μm . The optical
664 and microphysical properties of the aged contrail cirrus are often similar to those found in the ambient
665 “natural” cirrus clouds. The results show that the PCA method allows to identify and discriminate
666 different contrail growth stages and to provide an independent method for the characterization of the
667 evolution of contrail properties.

668 Accurate modeling of cirrus or contrails’ single scattering properties is a primary condition for the
669 interpretation of remote sensing measurements. Therefore, measurements of the optical
670 characteristics of ice crystals in natural conditions are still needed for validation of numerical
671 techniques and for the determination of free parameters in light scattering models. In this context, the
672 results from the PCA could be used to develop representative parameterizations of the scattering and
673 geometrical properties of the ice crystals’ shapes and sizes observed in the visible wavelength range
674 that then have to be extrapolated into the near infrared.

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