Comments to the Author:

The revised manuscript has improved significantly. Most of the comments by the reviewers have been adequately addressed. However, the summary of previous measurements over the Eastern Pacific ocean is incomplete. There is several recent experimental studies from ship and aircraft, and modeling studies that integrate organic and inorganic iodine sources from the ocean surface, and into the stratosphere. These recent literature should be added, and discussed in context of the results presented here before the manuscript is ready for publication.

We thank the editor for his suggestions and comments, and hope that we have improved the manuscript accordingly. All corrections are marked in red in the manuscript below. We have added more data comparisons with focus on the tropical East Pacific especially in the atmospheric discussion sections. As we concentrate mainly on the MABL in our manuscript, we have chosen to mainly compare our data with MABL measurements as well.

Reviewer#1 has suggested to depicted "abiotic sources of HOI and HOBr in Fig 5". The Editor agrees with this suggestion. While the formation of VSLs is the primary focus of this paper, HOI is arguably the largest source of iodine from the ocean (Sherwen et al. 2016).

Fig. 5 is missing direct emissions of HOI, and should be revised to reflect the state-of-the-science, rather than the scope of this paper. This will broaden the appeal of the Figure & paper.

We added the inorganic iodine cycling via iodide and ozone schematically to the figure, and hope that it reflects better current knowledge.

Reviewer#2 points out that the authors do not explicitly include tropospheric iodine chemistry in their model. The authors circumvent this by sensitivity studies that vary the lifetime of iodine, which is fine, but no reference is given to support the lifetime. Please provide references in support of the iodine lifetime.

We have added a reference (Sherwen et al., 2016) to the manuscript.

With respect to the discussion of iodine injections to the stratosphere: the contribution of CH3I was recently compared with inorganic iodine injections (Saiz Lopez et al. 2015).
Discussion is missing about how the CH3I data presented by the authors compare with these estimates, and in context of the relevance of inorganic iodine sources to the stratosphere.

We have added a short discussion of such a comparison. Note however, that this is not straightforward and somewhat speculative since neither Saiz-Lopez et al. (2015) nor Ordóñez et al. (2012) show their CH$_3$I emissions or boundary layer concentrations. Our paper on the other hand only provides boundary layer estimates and does not include any simulations of stratospheric contributions.

The statement "that for the Peruvian upwelling region... higher iodocarbon sources lead to larger IO abundances" is partially supported by Fig. 8. There is equal evidence to the contrary in that Figure. The statement that "a considerable part of the atmospheric IO variations [can be explained] with the variability of the oceanic organic sources" does not seem well supported. Which variability in IO are the authors referring to in Fig. 8b? A correlation plot should be added in support of this statement, or the language be modified. In the present form there is equal evidence in Fig. 8b to support that the contribution of combined iodocarbon fluxes to IO can be as low as few %. The manuscript language (abstract, discussion, conclusion) should be adjusted to reflect all data presented in this paper. At present it seems somewhat biased to conditions when the contributions of organic iodine fluxes were significant.

We have changed abstract, discussion and conclusion as requested by the editor to make clearer that during the second part of the cruise the contribution of the iodocarbon sources to atmospheric IO is very small. Additionally, we toned down the sentence referring to the variations of IO vs the variability of organic sources.
Biogenic halocarbons from the Peruvian upwelling region as tropospheric halogen source

Helmke Hepach\textsuperscript{1,6}, Birgit Quack\textsuperscript{1}, Susann Tegtmeier\textsuperscript{1}, Anja Engel\textsuperscript{1}, Astrid Bracher\textsuperscript{2}, Steffen Fuhlbrügge\textsuperscript{1}, Luisa Galgani\textsuperscript{1,7}, Elliot L. Atlas\textsuperscript{3}, Johannes Lampel\textsuperscript{4}, Udo Frieß\textsuperscript{4}, and Kirstin Krüger\textsuperscript{5}

Correspondence to: H. Hepach (hhepach@geomar.de)

[1] GEOMAR Helmholtz Centre for Ocean Research Kiel, Germany
[2] Alfred Wegener Institute (AWI), Helmholtz Centre for Polar and Marine Research Bremerhaven, Germany and Institute of Environmental Physics, University of Bremen, Germany
[3] Rosenstiel School of Marine and Atmospheric Science (RSMAS), University of Miami, USA
[4] Institute of Environmental Physics, University of Heidelberg, Germany
[5] Department of Geosciences, University of Oslo, Norway
[6] now at Environment Department, University of York, United Kingdom
[7] now at Department of Biotechnology, Chemistry and Pharmacy, University of Siena, Italy
Abstract

Halocarbons are produced naturally in the oceans by biological and chemical processes. They are emitted from surface seawater into the atmosphere, where they take part in numerous chemical processes such as ozone destruction and the oxidation of mercury and dimethyl sulfide. Here we present oceanic and atmospheric halocarbon data for the Peruvian upwelling zone obtained during the M91 cruise onboard the research vessel Meteor in December 2012. Surface waters during the cruise were characterized by moderate concentrations of bromoform (CHBr₃) and dibromomethane (CH₂Br₂) correlating with diatom biomass derived from marker pigment concentrations, which suggests this phytoplankton group as likely source. Concentrations measured for the iodinated compounds methyl iodide (CH₃I) of up to 35.4 pmol L⁻¹, chloroiodomethane (CH₂ClI) of up to 58.1 pmol L⁻¹ and diiodomethane (CH₂I₂) of up to 32.4 pmol L⁻¹ in water samples were much higher than previously reported for the tropical Atlantic upwelling systems. Iodocarbons also correlated with the diatom biomass and even more significantly with dissolved organic matter (DOM) components measured in the surface water. Our results suggest a biological source of these compounds as significant driving factor for the observed large iodocarbon concentrations. Elevated atmospheric mixing ratios of CH₃I (up to 3.2 ppt), CH₂ClI (up to 2.5 ppt) and CH₂I₂ (3.3 ppt) above the upwelling were correlated with seawater concentrations and high sea-to-air fluxes.

During the first part of the cruise, the enhanced iodocarbon production in the Peruvian upwelling contributed significantly to tropospheric iodine levels while this contribution was considerably smaller during the second part.

1 Introduction

Brominated and iodinated short-lived halocarbons from the oceans contribute to tropospheric and stratospheric chemistry (von Glasow et al., 2004; Saiz-Lopez et al., 2012b; Carpenter and Reimann, 2014). They are significant carriers of iodine and bromine into the marine atmospheric boundary layer (Salawitch, 2006; Jones et al., 2010; Yokouchi et al., 2011; Saiz-Lopez et al., 2012b), where they and their degradation products may also be involved in aerosol and ultra-fine particle formation (O'Dowd et al., 2002; Burkholder et al., 2004). Furthermore, they play an important role for ozone chemistry and other processes such as the oxidation of several atmospheric constituents in the troposphere (Saiz-Lopez et al., 2012a). Numerous modelling studies over the last years have shown that brominated short-lived compounds and their degradation products can be entrained into the stratosphere and enhance the halogen-driven ozone destruction (Carpenter and Reimann, 2014; Hossaini et al., 2015).
Recently, it was suggested that also oceanic iodine in organic or inorganic form can contribute to the stratospheric halogen loading, however only in small amounts due to its strong degradation (Tegtmeier et al., 2013; Saiz-Lopez et al., 2015).

While different source and sink processes determine the distribution of halocarbons in the oceanic surface water, the underlying mechanisms are largely unresolved. Biological activity plays a role for the production of bromoform (CHBr₃), dibromomethane (CH₂Br₂), methyl iodide (CH₃I), chloroiodomethane (CH₂I₂) and diiodomethane (CH₂I₂) (Gschwend et al., 1985; Tokarczyk and Moore, 1994; Moore et al., 1996), while CH₃I also originates from photochemical reactions with dissolved organic matter (DOM) (Moore and Zafiriou, 1994; Bell et al., 2002; Shi et al., 2014).

Biologically mediated halogenation of DOM (Lin and Manley, 2012; Liu et al., 2015) and bromination of compounds such as β-diketones via enzymes like bromoperoxidase (BPO) within or outside the algal cells (Theiler et al., 1978) are among the potentially important production processes of CHBr₃ and CH₂Br₂. Air-sea gas exchange into the atmosphere is the most important sink for both compounds in the surface ocean (Quack and Wallace, 2003; Hepach et al., 2015).

The biological formation of CH₃I has been investigated during laboratory and field studies (Scarratt and Moore, 1998; Amachi et al., 2001; Fuse et al., 2003; Smythe-Wright et al., 2006; Brownell et al., 2010; Hughes et al., 2011) identifying phyto- and bacterioplankton as producers, revealing large variability in biological production rates. However, biogeochemical modelling studies suggest that photochemistry may be more important for global CH₃I production (Stemmler et al., 2014). Due to their much shorter lifetime in surface water and the atmosphere, fewer studies investigated production processes of CH₂I₂ and CH₂I₂. CH₂I₂ has been suggested to be produced both by phytoplankton (Moore et al., 1996) and bacteria (Fuse et al., 2003; Amachi, 2008). The main source for CH₂I₂ is likely its production during the photolysis of CH₂I with a yield of 35 % based on a laboratory study (Jones and Carpenter, 2005). CH₂I₂ has also been detected in phytoplankton cultures (Tokarczyk and Moore, 1994) where it may originate from direct production or also from CH₂I₂ conversion. The main sink for both CH₂I₂ and CH₂I₂ is photolytical breakdown in the surface ocean resulting in lifetimes of less than 10 min (CH₂I₂) and 9 h (CH₂I₂), respectively, in the tropical ocean (Jones and Carpenter, 2005; Martino et al., 2006). Other sinks for these three iodocarbons are air-sea gas exchange and chloride substitution. The latter may play an important role for CH₃I in low latitudes at low wind speeds (Zafiriou, 1975; Jones and Carpenter, 2007).
Oceanic measurements of natural halocarbons are sparse (Ziska et al., 2013), but reveal that especially tropical and subtropical upwelling systems are potentially important source regions (Quack et al., 2007a; Raimund et al., 2011). Previously observed high tropospheric iodine monoxide (IO) levels in the tropical East Pacific have been related to short-lived iodinated compounds in surface waters (Schönhardt et al., 2008; Dix et al., 2013). However, iodocarbon fluxes have not been considered high enough to explain observed IO concentrations (Jones et al., 2010; Mahajan et al., 2010; Grossmann et al., 2013; Lawler et al., 2014), and recent global modelling studies suggested abiotic sources contributing on average about 75% to the IO budget (Prados-Roman et al., 2015). Such abiotic sources could be emissions of hypoiodous acid (HOI) and molecular iodine (I$_2$) as recently confirmed by a laboratory study (Carpenter et al., 2013).

This paper characterizes the Peruvian upwelling region between 5.0° S, 82.0° W and 16.2° S, 76.8° W with regard to the two brominated compounds CHBr$_3$ and CH$_2$Br$_2$ and the iodinated compounds CH$_3$I, CH$_2$ClI and CH$_2$I$_2$ in water and atmosphere. CH$_2$ClI and CH$_2$I$_2$ were measured for the first time in this region. Possible oceanic sources based on the analysis of phytoplankton species composition and different DOM components were evaluated and identified. Sea-to-air fluxes of these halogenated compounds were derived and their contribution to the tropospheric iodine loading above the tropical East Pacific were estimated by combining halocarbon, IO measurements and model calculations.

2 Methods

The M91 cruise of the RV Meteor from December 1 to 26, 2012 investigated the surface ocean and atmosphere of the Peruvian upwelling region (Bange, 2013). From the northernmost location of the cruise at 5.0° S and 82.0° W, the ship moved to the southernmost position at 16.2° S and 76.8° W with several transects perpendicular to the coast, alternating between open ocean and coastal upwelling (Fig. 1). All underway measurements were taken from a continuously operating pump in the ship’s hydrographic shaft from a depth of 6.8 m. Sea surface temperature (SST) and sea surface salinity (SSS) were measured continuously with a SeaCAT thermosalinograph from Seabird Electronics (SBE).

Deep samples were taken from 4 to 10 depths between 1 and 2000 m from 12 L Niskin bottles attached to a 24-bottle-rosette sampler equipped with a CTD and an oxygen sensor from SBE. Halocarbon samples were collected at 24 of the total 98 casts. The uppermost sample from the depth profiles (between 1 and 10 m) was included in the surface water measurements.
2.1 Analysis of halocarbon samples

Halocarbon samples were taken every three hours throughout the whole day from sea surface water and air. Water was sampled bubble free in 300 mL amber glass bottles. Surface water samples were analyzed on board with a purge and trap system attached to a GC-MS (combined gas chromatography and mass spectrometry) described in more detail in Hepach et al. (2014). The depth profile samples were analyzed with a similar setup: a purge and trap system was attached to a GC equipped with an ECD (electron capture detector). The precision of the measurements was within 10 % for all five halocarbons determined from duplicates and both systems were calibrated using the same liquid standards in methanol. The purge efficiency for all compounds in both set ups was larger than 98 % with a sample volume of 50 mL, a purging temperature of 70 °C, and a purge stream of 30 mL min⁻¹. Halocarbon measurements in seawater started only on December 9 due to instrument issues. Atmospheric halocarbon samples were taken in pre-cleaned stainless steel canisters on the monkey deck at a height of 20 m above sea level using a metal bellows pump starting on December 1, and were analyzed at the Rosenstiel School of Marine and Atmospheric Science (RSMAS) as described in Schauffler et al. (1998). For further details of atmospheric measurements see Fuhlbrügge et al. (2015a). Quantification was achieved using the NOAA standard SX3573 from GEOMAR.

2.2 Biological parameters

Phytoplankton composition was derived from pigment concentrations. Samples were taken in parallel with the halocarbon samples in the sea surface and up to six samples in depths between 3 and 200 m. Water was filtered through GF/F filters, which were stored at -80 °C until analysis after shock-freezing in liquid nitrogen. Pigments as described in Taylor et al. (2011) were analyzed using a HPLC technique according to Barlow et al. (1997). We used the diagnostic pigment analysis by Vidussi et al. (2001), subsequently refined by Uitz et al. (2006) by introducing pigment specific weight coefficients, to determine the chlorophyll a (Chl a) concentration of seven groups of phytoplankton which are assumed to comprise the entire phytoplankton community in ocean waters. Identified phytoplankton groups include diatoms, chlorophytes, dinoflagellates, haptophytes, cyanobacteria, cryptophytes and chrysophytes. Total chlorophyll a (TChl a) concentrations were calculated from the sum of the pigment concentrations of monovinyl Chl a, divinyl Chl a and chlorophyllide a.

Samples for the identification of DOM components were taken at 37 stations from a rubber boat from subsurface water at approximately 20 cm. Very-well mixed layers at these
measurement locations reach down to between 6 and 25 m, and DOM turn-over times for the respective compounds has been reported to be several days to months (Engel et al., 2011; Hansell, 2013). All samples were processed onboard and analyzed back in the home laboratory. Samples were analyzed for dissolved and total organic carbon (DOC and TOC), total dissolved nitrogen (TDN) and total nitrogen (TN) by a high-temperature catalytic oxidation method using a TOC analyzer (TOC-V CSH) from Shimadzu. Total, dissolved and particulate high molecular weight (HMW, >1kDa) combined carbohydrates (TCCHO, DCCHO and PCCHO), as well as total, dissolved and particulate combined HMW uronic acids (TURA, DURA and PURA), i.e. galacturonic acid and glucuronic acid were analyzed by High Performance Anion Exchange Chromatography coupled with Pulsed Amperometric Detection (HPAEC-PAD) after Engel and Händel (2011). For a more detailed description of both the sampling method and analysis see Engel and Galgani (2016).

2.3 Correlation analysis

Correlation analyses between all halocarbons, biological proxies and ambient parameters were carried out using Matlab® for all collocated surface and depth samples. All datasets were tested for normal distribution using the Lilliefors-test. Since most of the data were not distributed normally, Spearman’s rank correlation (hereinafter called $r_s$) was used. All correlations with a significance level of smaller than 5 % ($p < 0.05$) were regarded as significant.

2.4 Calculation of sea-to-air fluxes

Sea-to-air fluxes $F$ of halocarbons were calculated according to equation 1 with $k_w$ as the gas exchange coefficient parameterized according to Nightingale et al. (2000), $c_w$ the water concentrations from the halocarbon underway measurements, $c_{atm}$ from the simultaneous atmospheric measurements and $H$ as the Henry’s law constant to derive the equilibrium concentration.

$$F = k_w \cdot (c_w - \frac{c_{atm}}{H})$$

The gas exchange coefficient usually applied to derive carbon dioxide fluxes was adjusted for halocarbons using Schmidt number corrections as calculated in Quack and Wallace (2003), and Henry’s law coefficients as reported for each of the compounds by Moore et al. (1995) were applied. Wind speed and air pressure were averaged to 10 min intervals for the calculation of the instantaneous fluxes.
2.5 MAX-DOAS Measurements of IO

Multi-AXis Differential Optical Absorption Spectroscopy (MAX-DOAS) (Hönninger, 2002; Platt and Stutz, 2008) observations were conducted continuously at daytime from November 30 to December 25 2012 in order to quantify tropospheric abundances of IO, BrO, HCHO, Glyoxal, NO₂ and HONO along the cruise track, and for aerosol profiles also of O₄. The MAX-DOAS instrument and the measurement procedure are described in Grossmann et al. (2013) and Lampel et al. (2015). The primary quantity derived from MAX-DOAS measurements is the differential slant column density (dSCD), which represents the difference in path-integrated concentrations between two measurements in off-axis and zenith direction. From the MAX-DOAS observations of O₄ dSCD aerosol extinction profiles to estimate the quality of visibility were inferred using an optimal estimation approach described in Frieß et al. (2006) and Yilmaz (2012) after applying a correction factor of 1.25 to the O₄ dSCDs (Clémer et al., 2010). IO was analyzed in the spectral range from 418 – 438 nm following the settings in Lampel et al. (2015). IO was found up to 6 times above the detection limit (twice the measurement error).

2.6 FLEXPART simulations of tropospheric iodine

The atmospheric transport of the iodocarbons from the oceanic surface into the Marine Atmospheric Boundary Layer (MABL) was simulated with the Lagrangian particle dispersion model FLEXPART (Stohl et al., 2005) which has been used extensively in studies of long-range and mesoscale transport (Stohl and Trickl, 1999). FLEXPART is an off-line model driven by external meteorological fields. It includes parameterizations for moist convection, turbulence in the boundary layer, dry deposition, scavenging, and the simulation of chemical decay. We simulate trajectories of a multitude of air parcels describing transport and chemical decay of the emitted oceanic iodocarbons. For each data point of the observed sea-to-air flux, 100000 air parcels were released over the duration of the M91 cruise from a 0.1° x 0.1° grid box at the ocean surface centered at the measurement location. We used FLEXPART version 9.2 and the runs are driven by the ECMWF reanalysis product ERA-Interim (Dee et al., 2011) given at a horizontal resolution of 1° x 1° on 60 model levels. Transport, dispersion and convection of the air parcels are calculated from the 6-hourly fields of horizontal and vertical wind, temperature, specific humidity, convective and large scale precipitation and others. The chemical decay of the iodocarbons was prescribed by their atmospheric lifetime which was set to 4 days, 9 hours and 10 min for CH₃I, CH₂ClI, and CH₂I₂, respectively, according to current estimates (Jones and Carpenter, 2005; Martino et al., 2006; Carpenter and Reimann,
After degradation of the iodocarbons, the released iodine was simulated as inorganic iodine (I$_y$) tracer with a prescribed lifetime in the marine boundary layer with two days (personal communication R. von Glasow). Thus we did not include detailed tropospheric iodine chemistry, explicit removal of HOI, HI, IONO$_2$, and I$_x$O$_y$ through scavenging or heterogeneous recycling of HOI, IONO$_2$, and INO$_2$ on aerosols. In order to estimate the uncertainties arising from this simplification, we conducted two additional simulations, one with a very short lifetime of one day and one with a longer lifetime of three days. Following model simulations of halogen chemistry for air masses from different oceanic regions in Sommariva and von Glasow (2012), IO corresponds to 20% of the I$_y$ budget in the marine boundary layer on a daytime average. The IO to I$_y$ ratio shows moderate changes with daytime resulting in highest IO proportion at sunrise (~ 30%) and lowest IO proportion around noon (~ 15%) (see Fig. S6 in Sommariva and von Glasow (2012)). The ratio shows only very small variations for different air mass origins and thus the chemical conditions such as ozone and nitrogen species concentrations. Additionally, the ratio does not change much with altitude within the marine boundary layer. Based on the above estimates from Sommariva and von Glasow (2012), we used the IO to I$_y$ ratio as a function of daytime to estimate IO from I$_y$ every 3 hours. Daily averages of the IO abundance were compared to the MAX-DOAS IO measurements on board described in the previous section.

3 The tropical East Pacific – general description and state during M91

The tropical East Pacific is characterized by one of the strongest and most productive all-year-prevailing eastern boundary upwelling systems of the world (Bakun and Weeks, 2008). Temperatures drop to less than 16 °C when cold water from the Humboldt current is transported to the surface due to Ekman transport caused by strong equatorward winds (Tomczak and Godfrey, 2005), which is also connected to an upward transport of nutrients (Chavez et al., 2008). As a consequence of the enhanced nutrient supply and the high solar insolation, phytoplankton blooms, indicated by high Chl $a$ values, can be observed at the surface especially in the boreal winter months (Echevin et al., 2008). A strong oxygen minimum zone (OMZ) is formed due to enhanced primary production, sinking particles and weak circulation (Karstensen et al., 2008).

Low SSTs of mean (min – max) 19.4 (15.0 – 22.4) °C and high TChl $a$ values of on average 1.80 (0.06 – 12.65) µg L$^{-1}$ (Table 1, Fig. 1) were measured during our cruise. Diatoms were dominating the TChl $a$ concentration in the surface water with a mean of 1.66 (0.00 – 10.47) µg Chl $a$ L$^{-1}$, followed by haptophytes (mean: 0.25 µg Chl $a$ L$^{-1}$), chlorophytes (mean: 0.19
µg Chl a L$^{-1}$), cyanobacteria (mean: 0.09 µg Chl a L$^{-1}$), dinoflagellates (mean: 0.08 µg Chl a L$^{-1}$), cryptophytes (mean: 0.03 µg Chl a L$^{-1}$), and finally chrysophytes (mean: 0.03 µg Chl a L$^{-1}$). Diatoms were observed at all stations with concentrations above 0.5 µg Chl a L$^{-1}$ contributing more than 50 % of the algal biomass. They were significantly correlated with TChl a (Table 2) and with cryptophytes, which were elevated in very similar regions. Abundance of these phytoplankton groups was strongly anticorrelated with SST and SSS, indicating a close conjunction with the colder and more saline upwelling waters. Nutrients (nitrate, nitrite, ammonium and phosphate) were also measured during the cruise (see Czeschel et al. (2015) for further information). A weak anticorrelation of phytoplankton with the ratio of dissolved inorganic nitrogen and phosphate (sum of nitrate, nitrite and ammonium divided by phosphate, DIN:DIP) (Table 2) indicated that diatoms and cryptophytes were more abundant in aged upwelling, where nutrients were already slightly depleted or used up. The TChl a maximum was generally found in the surface ocean except for four stations with overall low TChl a (< 0.5 µg L$^{-1}$) where a subsurface maximum around 30 and 50 m was identified.

All regions with SSTs below the mean of 19.4 °C are considered as upwelling in the following sections for identifying different significant regions for halocarbon production. Based on this criterion, four upwelling regions (I – IV) close to the coast were classified (Fig. 1). The most intense upwelling (lowest SSTs, high nutrient concentrations) appeared in the northernmost region of the cruise track, region I, while higher TChl a and lower nutrients indicate a fully developed bloom in the southern part of the cruise (upwelling regions III and IV). Upwelling region II was characterized by a lower DIN:DIP ratio in contrast to region I. SSS with a mean of 34.95 (34.10 and 35.50) is lowest in upwelling region IV, which is likely influenced by local river input such as the rivers Pisco, Cañete and Matagente, and may explain the observed low salinities due to enhanced fresh water input in boreal winter (Bruland et al., 2005).

4 Halocarbons in the surface water and depth profiles during M91

4.1 Halocarbon distribution in surface water

Measurements of halocarbons in the tropical East Pacific are very sparse and no data were available for the Peruvian upwelling system before our campaign. Sea surface concentrations of CHBr$_3$ with a mean of 6.6 (0.2 – 21.5) and of CH$_2$Br$_2$ of 4.3 (0.2 – 12.7) pmol L$^{-1}$ were measured during M91 (Table 1, Fig. 2). These values are low in comparison to 44.7 pmol L$^{-1}$
CHBr₃ in tropical upwelling systems in the Atlantic, while our measurements of CH₂Br₂ compare better to these upwelling systems, from which maximum concentrations of 9.4 pmol L⁻¹ were reported (Quack et al., 2007a; Carpenter et al., 2009; Hepach et al., 2014; Hepach et al., 2015). CHBr₃ (0.2–20.7 pmol L⁻¹) and CH₂Br₂ (0.7–6.5 pmol L⁻¹) concentrations in the tropical East Pacific open ocean and Chilean coastal waters during a cruise from Punta Arenas, Chile to Seattle, USA in April 2010 (Liu et al., 2013) compare well to our data. The most elevated concentrations measured at the Chilean coast agree with our elevated northern coastal data. Some measurements also exist for the tropical West Pacific with on average 0.5 to 3 times our CHBr₃ and 0.2 to 1 times our CH₂Br₂ with the high average originating from a campaign close to the coast with macroalgal and anthropogenic sources (Krüger and Quack, 2013; Fuhlbrügge et al., 2015b). CHBr₃ and CH₂Br₂ have been proposed to have similar sources (Moore et al., 1996; Quack et al., 2007b). However, during our cruise, the correlation between the two compounds was comparatively weak (rₛ = 0.56), consistent with the findings of Liu et al. (2013), who ascribed the weaker correlation of these two compounds to formation in a common ecosystem rather than to the exact same biological sources. Maxima of CH₂Br₂ were observed in both upwelling regions III and IV, while CHBr₃ was highest in the most southerly upwelling IV (Fig. 2).

While we found the Peruvian upwelling and the adjacent waters to be only a moderate source region for bromocarbons, iodocarbons were observed in high concentration of 10.9 (0.4–58.1) for CH₂ClI, 9.8 (1.1–35.4) for CH₃I and 7.7 (0.2–32.4) pmol L⁻¹ for CH₂I₂ (Table 1, Fig. 3a). These concentrations identify the Peruvian upwelling as a significant source region of iodocarbons, especially considering the very short lifetimes of CH₂I₂ (10 min) and CH₂ClI (9 h) in tropical surface water (Jones and Carpenter, 2005). Hot spots were upwelling regions III and even more the less fresh upwelling of region IV (Fig. 2 and 3).

The occurrence of CH₃I in the tropical oceans (up to 36.5 pmol L⁻¹) has previously been attributed to a predominantly photochemical source (Richer and Wallace, 2004; Jones et al., 2010), explaining its global hot spots in the subtropical gyres and close to the tropical western boundaries of the continents (Ziska et al., 2013; Stemmler et al., 2014). Previous campaigns in the East Pacific obtained concentrations of up to 21.7 and up to 8.8 pmol L⁻¹ (Butler et al., 2007), and around 1.0 to 1.5 pmol L⁻¹ (Mahajan et al., 2012), but not directly in the upwelling.

No oceanic observations of CH₂ClI and CH₂I₂ have been published so far for the tropical East Pacific. Concentrations of CH₂ClI of up to 24.5 pmol L⁻¹ were measured in the tropical and subtropical Atlantic ocean (Abrahamsson et al., 2004; Chuck et al., 2005; Jones et al., 2010).
and up to 17.1 pmol L$^{-1}$ for CH$_2$I$_2$ (Jones et al., 2010; Hepach et al., 2015), which is lower but in the range of our measurements from the Peruvian upwelling.

Correlations between the compounds indicate similar sources for all measured halocarbons, except for CH$_3$Br$_2$, with upwelling region IV as hot spot area (Fig. 2). The strongest correlation was found for CHI with CH$_2$ClI ($r_5 = 0.83$). CH$_2$I$_2$ and CH$_2$ClI are often found to correlate very well with each other (Tokarczyk and Moore, 1994; Moore et al., 1996; Archer et al., 2007), usually attributed to the formation of CH$_2$ClI during photolysis of CH$_2$I$_2$. In comparison, the weaker correlation between CH$_2$ClI and CH$_3$I$_2$ ($r_5 = 0.59$) during our cruise may be the result of additional sources for CH$_2$ClI (see also section 5).

4.2 Halocarbon distribution in depth profiles

Depth profiles of halocarbons reveal maxima at the surface and around the Chl $a$ maximum, usually attributed to biological production of these compounds. CHBr$_3$ and CH$_2$Br$_2$ profiles (Fig. 4a – b) showed distinct maxima in the deeper Chl $a$ maximum during a large part of the cruise, while some profiles were characterized by elevated concentrations in the surface usually associated with upwelling water. Both kinds of profiles are consistent with previous studies finding maxima in the deeper water column in the open ocean and surface maxima in upwelling regions (Yamamoto et al., 2001; Quack et al., 2004; Hepach et al., 2015). During the northern part of M91 (upwelling III), CH$_2$Br$_2$ in the water column was more elevated than CHBr$_3$ (Fig. 4a), while during the remaining part of the cruise, CHBr$_3$ was usually higher (Fig. 4b).

Though most of the stations were characterized by subsurface maxima of iodocarbons, which were mostly located between 10 and 50 m (see example in Fig. 4, upper panel), surface maxima were often observed in upwelling region IV (see example in Fig. 4, lower panel), the region with highest iodocarbon concentrations. Profiles with surface maxima were generally characterized by much higher concentrations of these compounds, although we cannot completely preclude subsurface maxima at these locations owing to the sampling interval. CH$_2$I$_2$ was hardly detected in deeper water in the northern part of our measurements (Fig. 4, upper panel). Surface maxima in depth profiles of CH$_3$I and CH$_2$ClI were connected to surface maxima of several phytoplankton species, mainly diatoms ($r_5 = 0.57$ and 0.62). Direct and indirect biological and photochemical formation account as possible sources for these maxima. CH$_2$I$_2$ was usually strongly depleted in the surface in contrast to the deeper layers due to its rapid photolysis, which may also have been a source for surface CH$_2$ClI. Subsurface maxima occurred both below and within the mixed layer (see the example in Fig. 4d indicated
by the temperature-, salinity- and density profiles). Maxima in the mixed layer probably appear because of very fast production (Hepach et al., 2015), while maxima below the mixed layer are supported by accumulation due to reduced mixing.

All five halocarbons were strongly depleted in waters below 50 m. These deeper layers were also characterized by very low oxygen values, known as strong OMZ below the biologically active layers (Karstensen et al., 2008). A possible reason for the strong depletion of the halocarbons is their bacterial mediated reductive dehalogenation occurring under anaerobic conditions (Bouwer et al., 1981; Tanhua et al., 1996).

5  Relationship of surface halocarbons to environmental parameters

Physical and chemical parameters as well as biological proxies such as TChl a and phytoplankton group composition were investigated using correlation analysis in order to examine marine sources of halocarbons.

5.1  Potential bromocarbon sources

Bromocarbons were weakly, but significantly anticorrelated with SSS and SST ($r_s$ between -0.29 and -0.57), indicating sources in the upwelled water (Table 2). They showed a positive correlation with diatoms ($r_s = 0.58$ for both compounds), the dominant phytoplankton group in the region. Diatoms have already been found to be involved in bromocarbon production in several laboratory and field studies (Tokarczyk and Moore, 1994; Moore et al., 1996; Quack et al., 2007b; Hughes et al., 2013). Thus, these findings are in agreement with current assumptions that this group may contribute directly or indirectly to bromocarbon production. During M91, CH$_2$Br$_2$ was more abundant in cooler, nutrient-rich water than CHBr$_3$, leading to a stronger correlation with TChl a and SST, indicating an additional source associated with fresh upwelling. No significant correlations were found for bromocarbons with polysaccharidic DOM (Table 3), implying that DOM components analyzed during the cruise were not involved in bromocarbon production, at least not in the upper water column. Bromocarbon production from DOM has also been suggested to be slow (Liu et al. 2015), which could shift larger bromocarbon concentrations to later times after our cruise.

5.2  Iodinated compounds and phytoplankton

In general, the iodocarbons correlated stronger with biological parameters than the bromocarbons. Diatoms were found to correlate very strongly with all three iodocarbons ($r_s = 0.73$ with CH$_3$I, $r_s = 0.79$ with CH$_2$ClI and $r_s = 0.72$). Weak but significant anticorrelations
with DIN:DIP and SST suggest that iodocarbons were associated with cool and slightly DIN depleted water. The occurrence of large amounts of iodocarbons seemed to be associated with an established diatom bloom. The production of CH$_3$I, CH$_2$ClI and CH$_2$I$_2$ by a number of diatom species has been observed in several studies before (Moore et al., 1996; Manley and de la Cuesta, 1997), consistent with our findings. The very high correlation of cryptophytes with iodocarbons was likely based on the co-occurrence of these species with diatoms (Table 2 and description in section 3).

5.3 Iodinated compounds and DOM

Correlations of the three iodinated compounds to polysaccharidic DOM components in subsurface water revealed a strong relationship of the iodocarbon abundance with polysaccharides and in particular uronic acids (Table 3). CH$_3$I and CH$_2$ClI showed strong correlations with particulate uronic acids (both $r_s = 0.84$), total uronic acids ($r_s = 0.83$ and 0.88) and dissolved polysaccharides ($r_s = 0.82$ and 0.90). The correlations of CH$_2$I$_2$ with polysaccharides were less strong, but significant ($r_s = 0.68$ with particulate, $r_s = 0.66$ with total and $r_s = 0.55$ with dissolved). The above listed DOM components were also significantly correlated to diatoms ($r_s = 0.68$ with polysaccharides and $r_s = 0.75$ with uronic acids), which were a potential source for the accumulated organic matter in the subsurface. The exact composition of surface water DOM is determined by ecosystem compositions. Polysaccharides with uronic acids as an important constituent have for example been shown to contribute largely to the DOM pool in a diatom rich region (Engel et al., 2012).

Hill and Manley (2009) tested several diatom species for their production of halocarbons in a laboratory study, and suggested that a major formation pathway for polyhalogenated compounds may actually not be from direct algal production, but rather indirectly through their release of hypoiodous (HOI) and hypobromous acid (HOBr), which then react with the present DOM (Lin and Manley, 2012; Liu et al., 2015). The formation of HOI and HOBr within the algae is enzymatic with possible chloroperoxidase (CPO), BPO and iodoperoxidase (IPO) involvement. While CPO and BPO may produce both HOBr and HOI, IPO only leads to HOI. Moore et al. (1996) suggested that the occurrence of BPO and IPO in the phytoplankton cells may be highly species dependent. This leads to the assumption that diatoms abundant in the Peruvian upwelling contained more IPO than BPO, which could explain the higher abundance of iodocarbons relative to bromocarbons during M91. The formation of CH$_3$I through DOM may be different than the production of CH$_2$ClI and CH$_2$I$_2$. While CH$_2$I$_2$ is suggested to be formed via haloform-type reactions (Carpenter et al.,
2005), CH$_3$I is produced using a methyl-radical source (White, 1982). The relationship of CH$_3$I with DOM can be the result of both photochemical and biological production pathways: DOM, which was observed in high concentrations in the biologically productive waters, can act as the methyl-radical source during photochemical production of CH$_3$I (Bell et al., 2002).

A second possible biological pathway of methyl iodide production takes place via bacteria and micro algae, which can utilize methyl transferases in their cells. HOI plays a significant role in this production pathway by providing the iodine to the methyl group (Yokouchi et al., 2014).

The Peruvian upwelling was a strong source for the iodocarbons CH$_3$I, CH$_2$ClI and CH$_2$I$_2$, and a weaker source for the bromocarbons CHBr$_3$ and CH$_2$Br$_2$. We propose a formation mechanism for this region as described in Fig. 5 based on measurements of short-lived halocarbons and biological parameters during M91. Diatoms, which can contain the necessary enzymes for halocarbon formation, were identified as important source based on their strong correlations with the bromo- and iodocarbons and with polysaccharidic DOM. The very good correlations of iodocarbons with polysaccharides and uronic acids are a hint that these DOM components may have been important substrates for iodocarbon production potentially produced from the present diatoms. The higher iodocarbon concentrations can likely be explained by phytoplankton species containing more IPO than BPO, leading to a stronger production of iodocarbons. Additionally, the particular type of DOM may also have regulated the production of specific halocarbons (Liu et al., 2015), in this case CH$_3$I, CH$_2$ClI and CH$_2$I$_2$.

One interesting feature of our analysis is the fact that CH$_2$I$_2$ when compared to the other two iodocarbons showed weaker correlations with the polysaccharides possibly due to its shorter surface water lifetime. Moreover, CH$_2$I$_2$ and CH$_2$ClI showed weaker correlations in the Peruvian upwelling than during other cruises in the tropical Atlantic, namely MSM18/3 (Hepach et al., 2015) and DRIVE (Hepach et al., 2014). Combining the two arguments of a short CH$_2$I$_2$ lifetime and only a weak correlation between CH$_2$I$_2$ and CH$_2$ClI, this may indicate an additional source for CH$_2$ClI similar to CH$_3$I, explaining why CH$_3$I and CH$_2$ClI correlate much better with each other than with CH$_2$I$_2$.

6 From the ocean to the atmosphere

6.1 Sea-to-air fluxes of iodocarbons

Due to high oceanic iodocarbon concentrations measured in sea surface water of the Peruvian upwelling and despite the moderate prevailing wind speeds of 6.17 (0.42 – 15.47) m s$^{-1}$, high
iodocarbon sea-to-air fluxes were calculated in contrast to the rather low bromocarbon emissions during our cruise (Fuhlbrügge et al., 2015a). The highest average fluxes of the three iodocarbons of 954 (21 – 4686) were calculated for CH$_3$I, followed by 834 (-24 – 5652) for CH$_2$ClI, and finally 504 (-126 – 2546) pmol m$^{-2}$ h$^{-1}$ for CH$_2$I$_2$ (Table 1). These were on average 4 to 7 times higher than CHBr$_3$ and 2 to 4 times higher than the CH$_2$Br$_2$ sea-to-air fluxes during the cruise.

Our estimated fluxes of CH$_3$I are in the range of emissions calculated for the tropical and subtropical Atlantic of 625 to 2154 pmol m$^{-2}$ h$^{-1}$ (Chuck et al., 2005; Jones et al., 2010). Moore and Groszko (1999), who performed a study between 40$^{\circ}$ N and 40$^{\circ}$ S close to our investigation region but not covering the Peruvian upwelling, calculated on average 666 pmol m$^{-2}$ h$^{-1}$, which is 0.7 times our flux. Sea-to-air fluxes of CH$_2$ClI from the same studies were reported to range on average between 250 and 1138 pmol m$^{-2}$ h$^{-1}$ with the largest fluxes originating from the Mauritanian upwelling region. These are 0.3 to 1.4 times the fluxes we calculated, showing that the Peruvian upwelling region is at the top end of oceanic CH$_2$ClI emissions. We are only aware of two studies focusing on emissions of CH$_2$I$_2$ from the Atlantic tropical ocean (Jones et al., 2010; Hepach et al., 2015) which are on average 0.2, respectively 1.4 times the fluxes from the tropical East Pacific. The larger sea-to-air fluxes reported in Hepach et al. (2015) from the equatorial Atlantic cold tongue are mainly a result of much lower atmospheric mixing ratios there, increasing the concentration gradient, and additionally higher wind speeds, increasing the exchange coefficient $k_w$.

Summarizing this section, the large production of iodocarbons in the Peruvian upwelling led to enhanced emissions of these compounds to the troposphere despite very low wind speeds. An additional factor influencing halocarbon emissions is the low height and insolation of the MABL, where halocarbons accumulate above the air-sea interface. The large sea-to-air fluxes and low wind speeds should result in high tropospheric iodocarbons, which was indeed observed and is discussed in the following section.

### 6.2 Atmospheric iodocarbons

Atmospheric mixing ratios of the three iodocarbons were elevated during M91 with up to 3.2 ppt for CH$_3$I, up to 2.5 ppt for CH$_2$ClI and up to 3.3 ppt for CH$_2$I$_2$ (Table 1), likely a result of the strong production and emissions of these compounds.

CH$_3$I data were generally elevated in comparison to other eastern Pacific measurements of up to 1.1 ppt to 2.1 ppt CH$_3$I (Butler et al., 2007; Mahajan, et al., 2012), but lower than in the tropical central and East Atlantic around the equator, characterized by higher atmospheric
CH$_3$I of over 5 ppt (Ziska et al., 2013). Gómez Martín et al. (2013) report 1.7 times higher maximum mixing ratios of CH$_3$I of 5.4 ppt from the Galapagos islands, possibly rather influenced by very local sources around the measuring site, e.g. from seaweed or macroalgae. Both CH$_2$ClI and CH$_2$I$_2$ were also elevated in comparison to previous oceanic measurements, where e.g. 0.01 to 0.99 ppt CH$_2$ClI were measured for remote locations in the Atlantic and Pacific (Chuck et al., 2005; Varner et al., 2008) and only up to 0.07 ppt were reported for CH$_2$I$_2$ at a remote site in the Pacific (Yokouchi et al., 2011). Coastal areas with high macroalgal abundance were characterized by high CH$_2$ClI of up to 3.4 ppt (Varner et al., 2008) and up to 3.1 ppt CH$_2$I$_2$ (Carpenter et al., 1999; Peters et al., 2005), while 19.8 ppt (Peters et al., 2005) were measured at Mace Head, Ireland and Lilia, Brittany in the North Atlantic.

The different atmospheric lifetimes of the three iodocarbons, ranging between 4 d (CH$_3$I), 9 h (CH$_2$ClI) and 10 min (CH$_2$I$_2$) (Carpenter and Reimann, 2014), partly explain the observed differences in their distributions. Although atmospheric CH$_3$I was generally elevated in regions of high oceanic CH$_3$I (Fig. 3) in upwelling regions III and IV, the atmospheric and oceanic data did not show a significant correlation. The CH$_3$I lifetime of several days allows atmospheric CH$_3$I to mix within the MABL, possibly masking a correlation between local source regions and elevated mixing ratios.

The two shorter-lived iodinated compounds CH$_2$ClI and CH$_2$I$_2$ generally showed a stronger influence of local marine sources. Both species correlate significantly with their oceanic concentrations with $r_s = 0.60$ (CH$_2$ClI) and $r_s = 0.64$ (CH$_2$I$_2$). Oceanic CH$_2$ClI and CH$_2$I$_2$ were emitted into the boundary layer where they could accumulate during night (see comparison with global radiation in Fig. 3b), and were rapidly degraded during day time via photolysis, which is their main sink in the troposphere (Carpenter and Reimann, 2014). This accumulation is especially apparent when the data is separated into “day” and “night” according to global radiation measurements assuming that no radiation implies complete darkness. During the day, oceanic CH$_2$ClI and CH$_2$I$_2$ correlate weakly with $r = 0.5$ for CH$_2$ClI and no significant correlation was found for CH$_2$I$_2$ ($r = 0.4$), also owing to the fact that hardly any CH$_2$I$_2$ was measured during daylight. However, oceanic CH$_2$ClI and CH$_2$I$_2$ correlate significantly with their atmospheric counterparts with $r = 0.68$ and 0.92 in the night hours (Fig. 7). Moderate average wind speeds in the upwelling regions (Fig. 6) and stable atmospheric boundary layer conditions (Fuhlbrügge et al., 2015a) supported the accumulation of these compounds.
The Peruvian upwelling was in general characterized by elevated atmospheric iodocarbons as a result of their large sea-to-air fluxes caused by strong biological production. The upwelling could sustain elevated atmospheric levels of e.g. CH$_3$I, trapped iodocarbons and their degradation products in a stable MABL, and may have therefore contributed significantly to the tropospheric inorganic iodine budget, which is discussed in the following.

6.3 Contributions to tropospheric iodine

After their emission from the ocean and their chemical degradation in the marine boundary layer, iodocarbons contribute to the atmospheric inorganic iodine budget, $I_y$. The importance of this contribution compared to abiotic sources is currently under debate and analyzed for various oceanic environments (Mahajan et al., 2010; Grossmann et al., 2013; Prados-Roman et al., 2015). So far, no correlations of IO with organic iodine precursor species have been observed (Grossmann et al., 2013) and correlations between IO and Chl $a$ were often found to be negative (Mahajan et al., 2012; Gómez Martín et al., 2013). Chemical modelling studies, undertaken to explain the contributions of organic and inorganic oceanic iodine sources, simulated that only a small fraction of the atmospheric IO stems from the organic precursors with estimates of about 25 % on a global average (Prados-Roman et al., 2015). Both arguments, the missing correlations and the small contributions, indicate that the organic source gas emissions play a minor role for the atmospheric iodine budget. Given the special conditions of the Peruvian upwelling with cold nutrient rich waters, the strong iodocarbon sources, and a stable MABL and trade inversion, it is of interest to analyze the local contributions to the atmosphere in this region and to compare with estimates from other oceanic environments.

We focus our analysis on the section of the cruise where MAX-DOAS measurements of IO and simultaneous iodocarbon measurements in the surface water and atmosphere were made (roughly south of 10° S). Tropospheric VCDs of IO in the range of 2.5 – 6.0 ×10$^{12}$ molec cm$^{-2}$ were inferred from the MAX-DOAS measurements. Similar VCDs of IO were reported by Schönhardt et al. (2008) based on remote satellite measurements from SCIAMACHY, but around one order of magnitude higher VCDs were measured by Mahajan et al. (2012), which was however considerably further from the coast located than our measurements.

Volume mixing ratios of IO along the cruise track (Fig. 8a) derived from the MAX-DOAS measurements show a pronounced variability and maxima close to upwelling regions II and IV. Daytime averaged IO volume mixing ratios are displayed in Fig. 8b and range between 0.8 (on December 12 and December 22) and 1.5 ppt (on December 26). Overall the daytime
IO abundance in the MABL above the Peruvian upwelling was relatively high compared to measurements from the nearby Galapagos islands (~ 0.4 ppt) (Gómez Martín et al., 2013), and measurements closer to our investigation region with around between 0.2 and 0.6 ppt in the MABL (Dix et al., 2013; Wang et al., 2015), as well as from other tropical oceans such as the Malaspina 2010 circumnavigation (0.4 – 1 ppt) (Prados-Roman et al., 2015). Other measurement campaigns such as the Cape Verde measurements (Read et al., 2008) or the TransBrom Sonne in the West Pacific (Grossmann et al., 2013) found similar IO mixing ratios with values above 1 ppt, but significantly lower IO VCDs in the latter case. One of the first measurements of IO above the Peruvian upwelling was conducted by Volkamer et al. (2010) who report up to 3.5 ppt in the region. These larger mixing ratios in comparison to our study are possibly due to even more enhanced biological production as the upwelling is much more pronounced in October than in late December.

The M91 cruise track crisscrossed the waters between the coast and 200 km offshore multiple times, providing a comprehensive set of measurements over a confined area (see Fig. 8a) and allowing us to analyze the relation between IO and organic precursors. Assuming constant emissions over the cruise period we can link the oceanic sources with atmospheric IO observations at locations reached after hours to days of atmospheric transport. Therefore, we released FLEXPART trajectories from all sea surface measurement locations continuously over the whole measurement time period from December 8 to December 26 loaded with the oceanic iodine as prescribed by the observed iodocarbon emissions. Based on the simulations of transport and chemical decay described in Section 2.5, we derived organic and inorganic iodine mixing ratios individually for each air parcel. Mixing with air parcels impacted by other source regions was not taken into account. FLEXPART-based IO originating from organic precursors was derived as mean values over all air parcels in the MABL coinciding with the MAX-DOAS measurement locations within an area of 5 km x 5 km. Simplifying assumptions of a prescribed inorganic iodine lifetime (2 days) and IO to I$_y$ ratio (0.15 to 0.3) were made to derive the IO mixing ratios. Uncertainties were estimated based on additional runs with varying atmospheric lifetime of inorganic iodine (1 – 3 days).

For the first part of the cruise from December 8 to December 18, FLEXPART-derived IO mixing ratio estimates at the MAX-DOAS measurement locations (red line in Fig. 8b and c) explain between 40 and 70 % (55 % on average) of the measured IO assuming a lifetime for inorganic iodine of 2 days. As a consequence, about 0.5 ppt of IO is expected to originate from other, likely inorganic, iodine sources. For the scenario of a shorter I$_y$ lifetime (one day),
we find that the organic sources explain about 30% of the IO and for a relatively long lifetime (three days), 80% can be explained. In general, the air masses were transported along the coast in northwest direction and organic sources contribute to the IO budget along this transport path. Most of the IO results from CH₂ClI (Fig. 8c) which was transported some hours northwestwards (lifetime of 9 hours) before contributing to the atmospheric inorganic iodine budget. For the second part of the cruise from December 19 to December 26, the amount of IO estimated from organic precursors was much smaller and often close to zero. This very small organic contribution was caused by two facts. First, the instantaneous sources during the last part of the cruise were much smaller (Fig. 6b – d) and second, further southward situated sources, which also influence the iodine abundance in the cruise track region were not analyzed and thus not included in the simulations. Because of missing information on the source strength and distribution southwards of the cruise track, a proper comparison is only possible for the first part of the cruise before December 19. However, given that the MAX-DOAS measurements of IO remained relatively high during the second part of the cruise, it is likely that additional significant organic iodine sources existed further southwards. This assumption is also supported by the fact that atmospheric CH₃I mixing ratios remained relatively high during the second part of the cruise (50% compared to the earlier part, see Fig. 3b) while the water concentrations were close to zero. Consequently, a source region of CH₃I must have existed further southwards contributing to the observed mixing ratios of CH₃I and IO after some hours to days of atmospheric transport.

While the contribution of organic iodine to IO during the first part of the cruise is considerably higher than found in other regions, the amount of inorganic iodine precursors of 0.5 ppt necessary to explain total IO is very similar to the one derived in other studies (Prados-Roman et al., 2015). The higher organic contribution was consistent with the fact that there was an overall higher IO abundance compared to most other campaigns. Instantaneous IO and organic source gas emissions during M91 were not directly correlated. However, taking the transport within the first hours and days into account enables us to explain a considerable part of the atmospheric IO variations with the variability of the oceanic organic sources (Fig. 8b). Overall, we conclude that for the first part of the cruise (December 8 to December 18) the Peruvian upwelling region with special conditions in the ocean and atmosphere, with higher iodocarbon sources lead to larger IO abundances, while the absolute inorganic contribution is similar to other regions. For the second part of the cruise (December 19 to December 26), the organic contribution is considerably smaller,
however cannot be fully estimated due to missing information on the sources southwards of the cruise track.

7 Conclusions

The Peruvian upwelling at the west coast of South America was characterized for halocarbons for the first time during the M91 cruise. We measured moderate concentrations of the bromocarbons CHBr$_3$ and CH$_2$Br$_2$, while we observed exceptionally high concentrations of the iodocarbons CH$_3$I, CH$_2$I and CH$_2$I$_2$ in the surface seawater. CHBr$_3$ and CH$_2$Br$_2$ were significantly correlated with TChl $a$ and diatoms, suggesting biological formation of these compounds. Higher correlations of diatoms were found with the three iodocarbons, and even stronger correlations of the iodocarbons with the DOM components polysaccharides and uronic acids were observed. The polyhalogenated compounds CH$_2$I and CH$_2$I$_2$ were potentially formed via these DOM components with the likely involvement of diatoms. CH$_3$I may have been formed via photochemistry from the large pool of observed DOM and/or biologically via methyl transferases in micro algae and bacteria. The production of iodocarbons from DOM via the proposed mechanisms seems to have exceeded the bromocarbon production in the region in contrast to several previous studies in tropical Atlantic upwelling regions (Hepach et al., 2014, 2015).

Depth profiles showed subsurface maxima, common in the open ocean, and very pronounced and elevated surface maxima in regions of highest underway iodocarbon concentrations. The surface water was always depleted in CH$_2$I$_2$ with respect to the underlying water column due to its very rapid photolysis. The OMZ at depth was strongly depleted in all five measured halocarbons, suggesting an effective sink in the oxygen depleted waters.

The high oceanic iodocarbon concentrations and elevated emissions also led to elevated atmospheric mixing ratios in the marine boundary layer. Atmospheric CH$_2$I and CH$_2$I$_2$ showed clear diurnal cycles, accumulating during night and decreasing rapidly during day time. Despite previous suggestions that the tropospheric iodine loading is mainly a product from direct emission of HOI and I$_2$, we calculated important contributions of iodocarbons to the observed IO levels. Using FLEXPART, we estimated for the first part of the cruise a contribution of combined iodocarbon fluxes to IO of 30 to 80 % assuming an inorganic iodine lifetime between 1 and 3 days. This contribution of organoiodine is much higher than previously assumed (Prados-Roman et al., 2015), suggesting that iodocarbons therefore may contribute significantly to tropospheric iodine levels in regions of strong iodocarbon production mediated by phytoplankton (diatoms) and bacteria. For the second part of the
cruise, the organic input to atmospheric IO levels is considerably smaller. Note however that unknown sources southwards of the cruise track are not taken into account and might enhance the estimated contribution. Atmospheric CH$_3$I concentrations during the cruise were with 1.5 ppt on average larger than in most other oceanic regions. However, deep convective and subsequent slow transport is not expected to bring more than a few percent of the boundary layer CH$_3$I concentrations up to the cold point (Tegtmeier et al., 2013). Therefore the direct contribution of CH$_3$I to stratospheric iodine is, although not negligible, expected to be smaller than the 0.2 – 0.7 ppt estimated for inorganic iodine injections (Saiz Lopez et al. 2015).

Our observations reveal several uncertainties which need to be addressed in the future to better constrain the halocarbon budget and understand its role in a changing climate. Further studies of upwelling regions need to be performed in different seasons and years, since these regions are impacted by synoptic and climatic conditions, which are expected to have an impact on the strength of halocarbon emissions. A regular monitoring and better knowledge of halocarbon sources and emissions is severely needed, since these have numerous implications for atmospheric processes such as ozone chemistry and aerosol formation, which have been investigated in several atmospheric modeling studies. These studies have mostly applied Chl $a$ as proxy for halocarbon emissions. However, the potential involvement of DOM in the production of both iodo- and bromocarbons and the often weak correlation to Chl $a$ in the field raises the question whether Chl $a$ is a suitable parameter to estimate halocarbon concentrations. Laboratory studies are therefore crucial to help identifying more adequate parameters for predicting halocarbons in the ocean. Associated with the involvement of DOM in iodocarbon production is the occurrence of the relevant DOM components in large concentrations in the sea surface microlayer (SML) (Engel and Galgani, 2016). The SML has been shown to cover a wide range of oceanic regions (Wurl et al., 2011), which could represent a significant additional source to atmospheric iodocarbons. The potential of the SML to produce CH$_2$I$_2$ has been previously suggested by Martino et al. (2009), who proposed that HOI converted from iodide in the SML may react with the present DOM, which may apply to CH$_2$ClI and CH$_3$I as well. Furthermore, the SML is in direct contact to the atmosphere, and the direct exposure to light may enhance halocarbon emissions (see also Fig. 5). The influence of these halocarbon emissions on the tropospheric halogen loading is still very much under debate, and our results underline the importance to constrain the actual contribution of these compounds to tropospheric halogen chemistry.
Acknowledgements

We thank the chief scientist of the cruise M91 Hermann Bange, as well as the captain and the crew of the RV Meteor for their support. We would like to acknowledge Sonja Wiegmann for pigment analysis, Kerstin Nachtigall for nutrient measurements, and Stefan Raimund and Sebastian Flöter for helping with halocarbon measurements. This work was part of the German research projects SOPRAN II (grant no. FKZ 03F0611A) and III (grant no. FKZ 03F0662A) funded by the Bundesministerium für Bildung und Forschung (BMBF). Astrid Bracher’s contribution was funded by the Total Foundation project „Phytoscope”.

References


Figures

Figure 1. Ambient parameters during the M91 cruise: SST (dark red) and sea surface salinity (SSS) (black) in a) with the dashed line as the mean SST. Nutrients (purple is nitrate – NO$_3^-$, yellow is phosphate – PO$_4^{3-}$, black is silicate – SiO$_2$, light cyan is ammonium – NH$_4^+$) with the N to P ratio (dark blue dashed line) are shown in b). Total chlorophyll a (TChl a) is shown in the map in c). The light blue shaded areas stand for the regions where SST is below the mean, indicating upwelling of cold water.
Figure 2. Halocarbon surface water measurements are shown in a) and b) for the bromocarbons (note the colorbar in the upper panel) with a – CHBr_3 and b – CH_2Br_2. Iodocarbons can be found in c) – e) (note the colorbar in the lower panel) with c – CH_3I, d – CH_2ClI and e – CH_2I_2.
Figure 3. Surface water measurements of iodocarbons are presented in a) with CH$_3$I in red, CH$_2$ClI in blue and CH$_2$I$_2$ in black on the left side along with SST (dark red) on the right side. Additionally, atmospheric mixing ratios of CH$_3$I (red), CH$_2$ClI (blue) and CH$_2$I$_2$ (grey) on the left side together with global radiation (black, taken from pyranometer measurements referring to total short-wave radiation) on the right side are depicted in b). Note that all times are in UTC.

Figure 4. A cruise map including all CTD stations and SST is shown in a), while selected depth profiles of iodocarbons and bromocarbons can be seen in a) – b), together with ambient parameters such as potential density (cyan), oxygen (dark blue), salinity (black) and
temperature (dark red in d) – e), as well as phytoplankton groups (cryptophytes and diatoms), total chlorophyll \( a \) and fluorescence in f) – g). \( \text{CH}_2\text{I}_2 \) was undetectable at the first station (see also consistence with surface data).

Figure 5. Proposed mechanisms for formation of iodocarbons – release of HOI with the help of iodoperoxidases (IPO), followed by reaction with DOM (dissolved organic matter (DOM)) via iodine binding to DOM (DOM-I) to form \( \text{CH}_2\text{ClI} \) and \( \text{CH}_2\text{I}_2 \). \( \text{CH}_3\text{I} \) forms via photochemistry and/or biological formation via methyltransferases. The box indicates potential formation of halocarbons from DOM in the sea surface microlayer SML additionally to inorganic iodine sources there, implicated by the light blue box. For more information on the inorganic iodine chemistry in the SML, please refer to Carpenter et al. (2013).
Figure 6. The concentration gradient of CH$_3$I, CH$_2$ClI and CH$_2$I$_2$ along with wind speed (grey) is shown in a), while the sea-to-air flux is depicted in b) for CH$_3$I, c) for CH$_2$ClI and d) for CH$_2$I$_2$. Note the color bar on the right.

Figure 7. Atmospheric iodocarbons vs oceanic iodocarbons for CH$_3$I in a), CH$_2$ClI in b) and CH$_2$I$_2$ in c) with the filled circles representing data obtained during the day and open circles night time data, including least square lines (solid for day, dashed for night).
Figure 8. MAX-DOAS measurements of IO during the M91 campaign along the cruise track are shown a). Daytime averaged IO values from MAX-DOAS and coincident FLEXPART values are provided in b). The vertical bars correspond to uncertainties associated with the inorganic iodine lifetime in the MABL (1 – 3 days) and the daytime fraction of IO to I$_y$ (0.15 to 0.3). Contributions of the three oceanic iodocarbon sources to the modelled FLEXPART IO are also given in c).
### Tables

Table 1. Environmental parameters, as well as halocarbons in water, air and sea-to-air fluxes during the cruise. Means of sea surface temperature (SST), sea surface salinity (SSS) and wind speed are for 10-min-averages.

<table>
<thead>
<tr>
<th>Parameter</th>
<th>Unit</th>
<th>Mean</th>
<th>(min - max)</th>
</tr>
</thead>
<tbody>
<tr>
<td>SST</td>
<td>°C</td>
<td>19.4</td>
<td>(15.0 - 22.4)</td>
</tr>
<tr>
<td>SSS</td>
<td></td>
<td>34.95</td>
<td>(34.10 - 35.50)</td>
</tr>
<tr>
<td>TChl a</td>
<td>µg L⁻¹</td>
<td>1.80</td>
<td>(0.06 - 12.65)</td>
</tr>
<tr>
<td>Wind speed</td>
<td>m s⁻¹</td>
<td>6.17</td>
<td>(0.42 - 15.47)</td>
</tr>
<tr>
<td>CHBr₃</td>
<td>Water</td>
<td>6.6</td>
<td>(0.2 - 21.5)</td>
</tr>
<tr>
<td>CH₂Br₂</td>
<td>Air</td>
<td>2.9</td>
<td>(1.5 - 5.9)</td>
</tr>
<tr>
<td>CH₃I</td>
<td>Water</td>
<td>9.8</td>
<td>(1.1 - 35.4)</td>
</tr>
<tr>
<td>CH₃I</td>
<td>Air</td>
<td>1.5</td>
<td>(0.6 - 3.2)</td>
</tr>
<tr>
<td>CH₃I</td>
<td>Sea-to-air flux</td>
<td>130</td>
<td>(-550 - 2201)</td>
</tr>
<tr>
<td>CH₃I</td>
<td>Sea-to-air flux</td>
<td>273</td>
<td>(-128 - 1321)</td>
</tr>
<tr>
<td>Substance</td>
<td>Medium</td>
<td>Unit</td>
<td>Value</td>
</tr>
<tr>
<td>-----------</td>
<td>--------</td>
<td>-------------</td>
<td>---------</td>
</tr>
<tr>
<td></td>
<td></td>
<td>Sea-to-air flux</td>
<td>954</td>
</tr>
<tr>
<td>CH₂ClI</td>
<td>Water</td>
<td>pmol L⁻¹</td>
<td>10.9</td>
</tr>
<tr>
<td></td>
<td>Air</td>
<td>ppt</td>
<td>0.4</td>
</tr>
<tr>
<td>CH₂I₂</td>
<td>Water</td>
<td>pmol L⁻¹</td>
<td>7.7</td>
</tr>
<tr>
<td></td>
<td>Air</td>
<td>ppt</td>
<td>0.2</td>
</tr>
<tr>
<td></td>
<td></td>
<td>Sea-to-air flux</td>
<td>834</td>
</tr>
<tr>
<td></td>
<td></td>
<td>pmol m⁻² h⁻¹</td>
<td></td>
</tr>
<tr>
<td></td>
<td></td>
<td>Sea-to-air flux</td>
<td>504</td>
</tr>
<tr>
<td></td>
<td></td>
<td>pmol m⁻² h⁻¹</td>
<td></td>
</tr>
</tbody>
</table>
Table 2. Spearman’s rank correlation coefficients of correlations of all halocarbon surface data with several ambient parameters, as well as biological proxies. Bold numbers indicate correlations that are significant with $p < 0.05$ with a sample number of 107 for all environmental data and 46 for all phytoplankton and nutrient data considering all collocated surface data.

<table>
<thead>
<tr>
<th></th>
<th>CHBr$_3$</th>
<th>CH$_2$Br$_2$</th>
<th>CH$_3$I</th>
<th>CH$_2$ClI</th>
<th>CH$_3$I$_2$</th>
<th>SST</th>
<th>SSS</th>
<th>Global radiation</th>
<th>Diatoms</th>
<th>Crypto-phytes</th>
<th>Dino-flagellates</th>
<th>TChl $a$</th>
</tr>
</thead>
<tbody>
<tr>
<td>DIN:DIP</td>
<td>-0.26</td>
<td>-0.34</td>
<td><strong>-0.38</strong></td>
<td>-0.30</td>
<td>-0.32</td>
<td>0.00</td>
<td>0.41</td>
<td>0.13</td>
<td>-0.18</td>
<td>-0.17</td>
<td>-0.10</td>
<td>-0.08</td>
</tr>
<tr>
<td>TChl $a$</td>
<td><strong>0.48</strong></td>
<td><strong>0.56</strong></td>
<td>0.73</td>
<td>0.74</td>
<td>0.70</td>
<td><strong>-0.82</strong></td>
<td><strong>-0.77</strong></td>
<td><strong>-0.20</strong></td>
<td><strong>0.93</strong></td>
<td>0.85</td>
<td><strong>0.33</strong></td>
<td></td>
</tr>
<tr>
<td>Dino-flagellates</td>
<td>0.15</td>
<td>0.28</td>
<td>0.15</td>
<td>0.17</td>
<td>0.21</td>
<td>-0.23</td>
<td>-0.22</td>
<td>-0.01</td>
<td><strong>0.38</strong></td>
<td>0.26</td>
<td></td>
<td></td>
</tr>
<tr>
<td>Crypto-phytes</td>
<td><strong>0.38</strong></td>
<td><strong>0.54</strong></td>
<td>0.61</td>
<td>0.61</td>
<td>0.64</td>
<td><strong>-0.74</strong></td>
<td><strong>-0.79</strong></td>
<td>-0.18</td>
<td>0.73</td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>Diatoms</td>
<td><strong>0.58</strong></td>
<td><strong>0.58</strong></td>
<td>0.73</td>
<td>0.79</td>
<td>0.72</td>
<td><strong>-0.76</strong></td>
<td><strong>-0.72</strong></td>
<td>-0.18</td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>Global radiation</td>
<td>0.14</td>
<td>-0.10</td>
<td>-0.22</td>
<td>-0.03</td>
<td>-0.08</td>
<td><strong>0.20</strong></td>
<td>0.12</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>SSS</td>
<td><strong>-0.44</strong></td>
<td><strong>-0.48</strong></td>
<td><strong>-0.75</strong></td>
<td><strong>-0.69</strong></td>
<td><strong>-0.45</strong></td>
<td>0.68</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>SST</td>
<td><strong>-0.29</strong></td>
<td><strong>-0.57</strong></td>
<td><strong>-0.52</strong></td>
<td><strong>-0.62</strong></td>
<td><strong>-0.58</strong></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>CH$_3$I$_2$</td>
<td><strong>0.60</strong></td>
<td>0.43</td>
<td><strong>0.66</strong></td>
<td>0.59</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>Compound</td>
<td>0.64</td>
<td>0.70</td>
<td>0.83</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>-----------</td>
<td>------</td>
<td>------</td>
<td>------</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>CH₂ClI</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>CH₃I</td>
<td>0.66</td>
<td>0.46</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>CH₂Br₂</td>
<td>0.56</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
</tbody>
</table>
Table 3. Correlations of halocarbons with combined high molecular weight (HMW) carbohydrates (CCHO) and uronic acids (URA) from subsurface samples (T – total, d – dissolved, P – particulate) with a sample number of 29 for each variable.

<table>
<thead>
<tr>
<th></th>
<th>CHBr₃</th>
<th>CH₂Br₂</th>
<th>CH₃I</th>
<th>CH₂Cl₂</th>
<th>CH₂I₂</th>
</tr>
</thead>
<tbody>
<tr>
<td>TCCHO</td>
<td>0.15</td>
<td>0.28</td>
<td><strong>0.78</strong></td>
<td><strong>0.82</strong></td>
<td><strong>0.66</strong></td>
</tr>
<tr>
<td>dCCHO</td>
<td>0.39</td>
<td>0.48</td>
<td><strong>0.82</strong></td>
<td><strong>0.90</strong></td>
<td><strong>0.55</strong></td>
</tr>
<tr>
<td>PCCHO</td>
<td>-0.06</td>
<td>-0.10</td>
<td><strong>0.61</strong></td>
<td><strong>0.64</strong></td>
<td><strong>0.68</strong></td>
</tr>
<tr>
<td>TURA</td>
<td>0.31</td>
<td>0.34</td>
<td><strong>0.83</strong></td>
<td><strong>0.88</strong></td>
<td><strong>0.52</strong></td>
</tr>
<tr>
<td>dURA</td>
<td>-0.18</td>
<td>0.42</td>
<td><strong>0.48</strong></td>
<td><strong>0.79</strong></td>
<td><strong>0.50</strong></td>
</tr>
<tr>
<td>PURA</td>
<td>0.37</td>
<td>0.22</td>
<td><strong>0.84</strong></td>
<td><strong>0.84</strong></td>
<td><strong>0.54</strong></td>
</tr>
</tbody>
</table>