

1 **Ozone variability in the troposphere and the stratosphere from the first six years of IASI**  
2 **observations (2008-2013)**

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12

13 **Abstract**

14 In this paper, we assess how daily ozone (O<sub>3</sub>) measurements from the Infrared Atmospheric  
15 Sounding Interferometer (IASI) on the MetOp-A platform can contribute to the analyses of the  
16 processes driving O<sub>3</sub> variability in the troposphere and the stratosphere and, in the future, to the  
17 monitoring of long-term trends. The temporal evolution of O<sub>3</sub> during the first 6 years of IASI  
18 (2008-2013) operation is investigated with multivariate regressions separately in four different  
19 layers (ground-300 hPa, 300-150 hPa, 150-25 hPa, 25-3 hPa), by adjusting to the daily time  
20 series averaged in 20° zonal bands, seasonal and linear trend terms along with important  
21 geophysical drivers of O<sub>3</sub> variation (e.g. solar flux, quasi biennial oscillations). The regression  
22 model is shown to perform generally very well with a strong dominance of the annual harmonic  
23 terms and significant contributions from O<sub>3</sub> drivers, in particular in the equatorial region where  
24 the QBO and the solar flux contribution dominate. More particularly, despite the short period of  
25 IASI dataset available to now, two noticeable statistically significant apparent trends are inferred  
26 from the daily IASI measurements: a positive trend in the upper stratosphere (e.g. 1.74±0.77  
27 DU/yr between 30°S-50°S) which is consistent with other studies suggesting a turnaround for  
28 stratospheric O<sub>3</sub> recovery, and a negative trend in the troposphere at the mid-and high northern

29 latitudes (e.g.  $-0.26 \pm 0.11$  DU/yr between 30°N-50°N), especially during summer and probably  
30 linked to the impact of decreasing ozone precursor emissions. The impact of the high temporal  
31 sampling of IASI on the uncertainty in the determination of O<sub>3</sub> trend has been further explored  
32 by performing multivariate regressions on IASI monthly averages and on ground-based FTIR  
33 measurements.

34

## 35 **1 Introduction**

36 Global climate change is one of the most important environmental problems of today and  
37 monitoring the behavior of the atmospheric constituents (radiatively active gases and those  
38 involved in their chemical production) is key to understand the present climate and apprehend  
39 future climate changes. Long-term measurements of these gases are necessary to study the  
40 evolution of their abundance, changing sources and sinks in the atmosphere.

41

42 As a reactive trace gas present simultaneously in the troposphere and in the stratosphere, O<sub>3</sub>  
43 plays a significant role in atmospheric radiative forcing, atmospheric chemistry and air quality.  
44 In the stratosphere, O<sub>3</sub> is sensitive to changes in (photo-)chemical and dynamical processes and,  
45 as a result, undergoes large variations on seasonal and annual time scales. Measurements of O<sub>3</sub>  
46 total column have indicated a downward trend in stratospheric ozone over the period from 1980s  
47 to the late 1990s relative to the pre-1980 values, which is due to the growth of the reactive  
48 bromine and chlorine species following anthropogenic emissions during that period (WMO,  
49 2003). In response to the 1987 Montreal Protocol and its amendments, with a reduction of the  
50 Ozone-Depleting Substances (ODS; Newchurch et al., 2003), a recovery of stratospheric ozone  
51 concentrations to the pre-1980 values is expected (Hofmann, 1996). While earlier works have  
52 debated a probable turnaround for the ozone hole recovery (e.g. Hadjinicolaou et al., 2005;  
53 Reinsel et al., 2002; Stolarski and Frith, 2006), WMO already indicated in 2007 that the total  
54 ozone in the 2002-2005 period was no longer decreasing, reflecting such a turnaround. Since  
55 then several studies have shown successful identification of ozone recovery over Antarctica and  
56 over northern latitudes (e.g. Mäder et al., 2010; Salby et al., 2011; WMO, 2011; Kuttippurath et  
57 al., 2013; Knibbe et al., 2014; Shepherd et al., 2014). Nevertheless, the most recent papers as  
58 well as the WMO 2014 ozone assessment have warned, because of possible underestimation of

59 the true uncertainties in the ozone trends attributed to decreasing Effective Equivalent  
60 Stratospheric Chlorine (EESC), against overly optimistic conclusions with regard to a possible  
61 increase in Antarctic stratospheric ozone (Kramarova et al., 2014 ; WMO, 2014; Knibbe et al.,  
62 2014; de Laat et al., 2015; Kuttippurath et al., 2015; Varai et al., 2015). The causes of the  
63 observed stratospheric O<sub>3</sub> changes are hard to isolate and remain uncertain precisely considering  
64 the contribution of dynamical variability to the apparent trend and the limitations of current  
65 chemistry-climate models to reproduce the observations. The assessment of ozone trends in the  
66 troposphere is even more challenging due to the influence of many simultaneous processes (e.g.  
67 emission of precursors, long-range transport, stratosphere-troposphere exchanges –STE–), which  
68 are all strongly variable temporally and spatially (e.g. Logan et al., 2012; Hess and Zbinden,  
69 2013; Neu et al., 2014). Overall, there are still today large differences in the value of the O<sub>3</sub>  
70 trends determined from independent studies and datasets (mostly from ground-based and satellite  
71 observations) in both the stratosphere and the troposphere (e.g. Oltmans et al., 1998; 2006;  
72 Randel and Wu, 2007; Gardiner et al., 2008; Vigouroux et al., 2008; Jiang et al., 2008; Kyrölä et  
73 al., 2010; Vigouroux et al., 2014). In order to improve on this and because O<sub>3</sub> has been  
74 recognized as a Global Climate Observing System (GCOS) Essential Climate Variables (ECVs),  
75 the scientific community has underlined the need of acquiring high quality global, long-term and  
76 homogenized ozone profile records from satellites (Randel and Wu, 2007; Jones et al., 2009;  
77 WMO, 2007; 2011; 2014). This specifically has resulted in the ESA Ozone Climate Change  
78 Initiative (O<sub>3</sub>-CCI; <http://www.esa-ozone-cci.org/>).

79  
80 The Infrared Atmospheric Sounding Interferometer (IASI) onboard the polar orbiting MetOp,  
81 with its unprecedented spatiotemporal sampling of the globe, its high radiometric stability and  
82 the long duration of its program (3 successive instruments to cover 15 years) provides in  
83 principle an excellent means to contribute to the analyses of the O<sub>3</sub> variability and trends. This is  
84 further strengthened by the possibility of using IASI measurements to discriminate O<sub>3</sub>  
85 distributions and variability in the troposphere and the stratosphere, as shown in earlier studies  
86 (Boynard et al., 2009; Wespes et al., 2009; Dufour et al, 2010 ; Barret et al., 2011; Scannell et  
87 al., 2012; Wespes et al., 2012; Safieddine et al., 2013). Here, we use the first 6 years (2008-  
88 2013) of the new O<sub>3</sub> dataset provided by IASI on MetOp-A to perform a first analysis of the O<sub>3</sub>

89 time development in the stratosphere and in the troposphere. This is achieved globally by using  
90 zonal averages in 20° latitude bands and a multivariate linear regression model which accounts  
91 for various natural cycles affecting O<sub>3</sub>. We also explore in this paper to which extent the  
92 exceptional temporal sampling of IASI can counterbalance the short period of data available for  
93 assessing trends in partial columns.

94

95 In section 2, we give a short description of IASI and of the O<sub>3</sub> retrieved columns used here.  
96 Section 3 details the multivariate regression model used for fitting the time series. In Section 4,  
97 we evaluate how the ozone natural variability is captured by IASI and we present the time  
98 evolution of the retrieved O<sub>3</sub> profiles and of four partial columns (Upper Stratosphere –UST-;  
99 Middle-Low Stratosphere –MLST-; Upper Troposphere Lower Stratosphere –UTLS-; Middle-  
100 Low Troposphere –MLT-) using 20-degree latitudinal averages on a daily basis. The apparent  
101 dynamical and chemical processes in each latitude band and vertical layer are then analyzed on  
102 the basis of the multiple regression results using a series of common geophysical variables. The  
103 “standard” contributors in the fitted time series, as well as a linear trend term, are analyzed in the  
104 specified altitude layers. Finally, the trends inferred from IASI are compared against those from  
105 FTIR for six stations in the northern hemisphere.

106

## 107 **2 IASI measurements and retrieval method**

108 IASI measures the thermal infrared emission of the Earth-atmosphere between 645 and 2760  
109 cm<sup>-1</sup> with a field of view of 2×2 circular pixels on the ground, each of 12 km diameter at nadir.  
110 The IASI measurements are taken every 50 km along the track of the satellite at nadir, but also  
111 across-track over a swath width of 2200 km. IASI provides a global coverage twice a day with  
112 overpass times at 9:30 and 21:30 mean local solar time. The instrument is also characterized by a  
113 high spectral resolution which allows the retrieval of numerous gas-phase species (e.g. Clerbaux  
114 et al., 2009; Clarisse et al., 2012).

115

116 Ozone profiles are retrieved with the Fast Optimal Retrievals on Layers for IASI (FORLI)  
117 software developed at ULB/LATMOS. FORLI relies on a fast radiative transfer and on a  
118 retrieval methodology based on the Optimal Estimation Method (Rodgers, 2000). In the version

119 used in this study (FORLI-O<sub>3</sub> v20100815), the O<sub>3</sub> profile is retrieved for individual IASI  
120 measurement on a uniform 1 km vertical grid on 40 layers from surface up to 40 km. The a priori  
121 information (a priori profile and a priori covariance matrix) is built from the  
122 Logan/Labow/McPeters climatology (McPeters et al., 2007) and only one single O<sub>3</sub> a priori  
123 profile and variance-covariance matrix are used. The retrieval parameters and performances are  
124 detailed in Hurtmans et al. (2012). The FORLI-O<sub>3</sub> profiles and/or total and partial columns have  
125 undergone validation using available ground-based, aircraft, O<sub>3</sub> sonde and other satellite  
126 observations (Anton et al., 2011; Dufour et al., 2012; Gazeaux et al., 2012; Parrington et al.,  
127 2012; Pommier et al., 2012; Scannell et al., 2012; Oetjen et al., 2014). Generally, the results  
128 show good agreements between FORLI-O<sub>3</sub> and independent measurements with a low bias  
129 (<10%) in the total column and in the vertical profile, except in UTLS where a positive bias of  
130 10-15% is reported (Dufour et al., 2012; Gazeaux et al., 2012; Oetjen et al., 2014).

131  
132 In this study, only daytime O<sub>3</sub> IASI observations from good spectral fits (RMS of the spectral  
133 residual lower than  $3.5 \times 10^{-8} \text{ W/cm}^2 \cdot \text{sr} \cdot \text{cm}^{-1}$ ) have been analyzed. Daytime IASI observations  
134 (determined with a solar zenith angle to the sun < 80°) are characterized by a better vertical  
135 sensitivity to the troposphere associated with a higher surface temperature and a higher thermal  
136 contrast (Clerbaux et al., 2009; Boynard et al., 2009). Furthermore, cloud contaminated scenes  
137 with cloud cover < 13% (Hurtmans et al., 2012) were removed using cloud information from the  
138 Eumetcast operational processing (August et al., 2012).

139  
140 An example of typical FORLI-O<sub>3</sub> averaging kernel functions for one mid-latitude observation in  
141 July (45°N/66°E) is represented on Fig.1. The layers have been defined as: ground-300hPa  
142 (MLT), 300-150hPa (UTLS), 150-25 hPa (MLST) and above 25 hPa (UST), so that they are  
143 characterized by a DOFS (Degrees Of Freedom for Signal) close to 1 with a maximum  
144 sensitivity approximatively in the middle of the layers, except for the 300-150 hPa layer which  
145 has a reduced sensitivity. Taken globally, the DOFS for the entire profile ranges from ~2.5 in  
146 cold polar regions to ~4.5 in hot tropical regions, depending mostly on surface temperature, with  
147 a maximum sensitivity in the upper troposphere and in the lower stratosphere (Hurtmans et al.,  
148 2012). In the MLT, the maximum of sensitivity is around 4–8 km altitude for almost all

149 situations (Wespes et al., 2012). The sharp decrease of sensitivity down to the surface is inherent  
150 to nadir thermal IR sounding in cases of low surface temperature or low thermal contrast and  
151 indicates that the retrieved information principally comes from the *a priori* in the lowest layer.  
152 Figure 2 presents July 2010 global maps of averaged FORLI-O<sub>3</sub> partial columns for two partial  
153 layers (MLT and MLST), and of the associated DOFS and *a priori* contribution (calculated as  
154  $X_a - \mathbf{A}(X_a)$ , where  $X_a$  is the *a priori* profile and  $\mathbf{A}$ , the averaging kernel matrix, following the  
155 formalism of Rodgers (2000)). The two layers exhibit different sensitivity patterns: in the MLT,  
156 the DOFS typically range from 0.4 in the cold polar regions to 1 in regions characterized by high  
157 thermal contrast with medium humidity, such as the mid-latitude continental Northern  
158 Hemisphere (N.H.) (Clerbaux et al., 2009). Lower DOFS values in the intertropical belt are  
159 explained by overlapping water vapor lines. In contrast, the DOFS for the MLST are globally  
160 almost constant and close to one, with only slightly lower values (0.9) over polar regions. The *a*  
161 *priori* contribution is anti-correlated with the sensitivity, as expected. It ranges between a few %  
162 to ~30% and does not exceed 20% on 20° zonal averages in the troposphere (see Supporting  
163 Information; Fig. S3, dashed lines), while the *a priori* contribution is smaller than ~12% in the  
164 middle stratosphere. These findings indicate that the IASI MLST time series should accurately  
165 represent stratospheric variations, while the time series in the troposphere may reflect to some  
166 extent variations from the upper layers in addition to the real variability in the troposphere. In  
167 order to quantify this effect, the contribution of the stratosphere into the tropospheric ozone as  
168 seen by IASI has been estimated with a global 3-D chemical transport model (MOZART-4).  
169 Details of the model-observation comparisons can be found in the Supplement (see Fig. S2 and  
170 S3). We interestingly show that the stratospheric contribution to the MLT columns measured by  
171 IASI varies between 30% and 60%, depending on latitude and season (Fig. S5). The limited  
172 vertical sensitivity of IASI contributes to this by a smaller part (~10%-20%) than the natural  
173 stratospheric influence (~20% to 45%) (See Fig. S4 and S5). In addition, we find that the  
174 contribution of the natural variability (from both the troposphere and the stratosphere) on the  
175 MLT O<sub>3</sub> columns is larger than 50% everywhere. In the 30N-50N band where the DOFS is the  
176 largest (See Fig.2 (b)), this contribution reaches ~85% from which ~25-35% originates from the  
177 stratosphere and ~55% from the troposphere (Fig.S6 (a) and (b)). Nevertheless, the

178 contamination of IASI MLT O<sub>3</sub> with variations in stratospheric O<sub>3</sub> has to be kept in mind when  
 179 analyzing IASI MLT O<sub>3</sub>.

### 180 **3 Fitting method**

#### 181 **3.1. Statistical model**

182 In order to characterize the changes in ozone measured by IASI and to allow a proper separation  
 183 of trend, we use a multiple linear regression model accounting for a linear trend and for inter-  
 184 annual, seasonal and non-seasonal variations related to physical processes that are known to  
 185 affect the ozone records. More specifically, the time series analysis is based on the fitting of  
 186 daily (or monthly) median partial columns in different latitude band following:

$$187 \quad O_3(t) = Cst + x_1 \cdot trend + \sum_{n=1,2} [a_n \cdot \cos(n\omega t) + b_n \cdot \sin(n\omega t)] + \sum_{j=2}^m x_j X_{norm,j}(t) + \varepsilon(t) \quad (1)$$

188 where  $t$  is the number of days (or months),  $x_1$  is the 6-year trend coefficient in the data,  $\omega =$   
 189  $2\pi/365.25$  for the daily model (or  $2\pi/12$  for the monthly model) and  $X_{norm,j}$  are independent  
 190 geophysical variables, the so-called “explanatory variables” or “proxies”, which are in this study  
 191 normalized over the period of IASI observation (2008-2013), as:

$$192 \quad X_{norm}(t) = 2[X(t) - X_{median}] / [X_{max} - X_{min}] \quad (2)$$

193  $\varepsilon(t)$  in Eq. (1) represents the residual variation which is not described by the model and which is  
 194 assumed to be autoregressive with time lag of 1 day (or 1 month). The constant term ( $Cst$ ) and  
 195 the coefficients  $a_n, b_n, x_j$  are estimated by least-squares method and their standard errors ( $\sigma_e$ )  
 196 are calculated from the covariance matrix of the coefficients and corrected to take into account  
 197 the uncertainty due to the autocorrelation of the noise residual as discussed in Santer et al. (2000)  
 198 and references therein:

$$199 \quad \sigma_e^2 = (Y^T Y)^{-1} \cdot \frac{\sum [O_3(t) - yY(t)]^2}{n - m} \cdot \frac{1 + \Phi}{1 - \Phi} \quad (3)$$

200 Where  $Y$  is the matrix with the covariates ( $trend, \cos(n\omega t), \sin(n\omega t), X_{norm,j}$ ) sorted by  
 201 column,  $y$  is the vector of the regression coefficients corresponding to the columns of  $Y$ ,  $n$  is the

202 number of daily (or monthly) data points in the time series,  $m$  is the number of the fitted  
203 parameters, and  $\Phi$  is the lag-1 autocorrelation of the residuals.

204

205 The median is used as a statistical average since it is more robust against the outliers than the  
206 normal mean (Kyrölä et al., 2006; 2010). Note that, similarly to Kyrölä et al. (2010), the model  
207 has been applied on  $O_3$  mixing ratios rather than on partial columns but without significant  
208 improvement on the fitting residuals and R values.

209

### 210 **3.2. Geophysical variables**

211 In Eq. (1), harmonic time series with a period of a year and a half year are used to account for the  
212 Brewer-Dobson circulation and the solar insolation ( $a_1$  and  $b_1$  coefficients), and for the  
213 meridional circulation ( $a_2$  and  $b_2$  coefficients), respectively (Kyrölä et al., 2010). While these  
214 effects are of a periodic nature, the geophysical variables ( $X_j$ ) are used here to parameterize the  
215 ozone variations on non-seasonal timescales. The chosen proxies are  $F_{10.7}$ ,  $QBO^{10}$ ,  $QBO^{30}$ ,  $ENSO$ ,  
216  $NAO/AAO$ , the first three being the most commonly used (“standard”) proxies to describe the  
217 natural ozone variability, i.e. the solar radio flux at 10.7 cm and the quasi-biennial oscillation  
218 ( $QBO$ ) which is represented by two orthogonal zonal components of the equatorial stratospheric  
219 wind measured at 10 hPa and 30 hPa, respectively (e.g. Randel and Wu, 2007). The three other  
220 proxies,  $ENSO$ ,  $NAO$  and  $AAO$ , are used to account for other important fluctuating dynamical  
221 features: the El Niño/Southern Oscillation, the North Atlantic Oscillation and the Antarctic  
222 Oscillation, respectively. Table 1 lists the selected proxies, their sources and their resolutions.  
223 The time series of these proxies normalized over the 2000-2013 period following Eq. 2 are  
224 shown in Fig.3 (a) and (b) and they are shortly described hereafter:

225 - *Solar flux*: the F10.7 cm solar radio flux is an excellent indicator of solar activity and is  
226 commonly used to represent the 11 year solar cycle. It is available from continuous routine  
227 consistent measurements at the Penticton Radio Observatory in British Columbia which are  
228 corrected for the variable Sun-Earth distance resulting from the eccentric orbit of the Earth  
229 around the Sun. Over the period 2008-2013, the radio solar flux increases from about 65 units in  
230 2008 to 180 units in 2013 and is characterized by a specific daily “fingerprint” (see Fig.3 (a)).  
231 Note that because the period of IASI observations does not cover a full 11 year solar cycle, it

232 could affect the determination of the trend in the regression procedure. The difficulty in  
233 discriminating the solar flux and linear trend terms is a known problem for such multivariate  
234 regression: it feeds into their uncertainties and it can lead to biases in the coefficients  
235 determination (e.g. Soukharev et al., 2006).

236 - *QBO terms*: The QBO of the equatorial winds is a main component of the dynamics of the  
237 tropical stratosphere (Chipperfield et al., 1994; 2003; Randel and Wu, 1996; 2007; Logan et al.,  
238 2003; Tian et al., 2006; Fadnavis and Beig, 2009; Hauchecorne et al., 2010). It strongly  
239 influences the distributions of stratospheric O<sub>3</sub> propagating alternatively westerly and easterly  
240 with a mean period of 28 to 29 months. Positive and negative vertical gradients alternate  
241 periodically. At the top of the vertical QBO domain, there is a predominance of easterlies, while,  
242 at the bottom, westerly winds are more frequent. In order to account for the out-of-phase  
243 relationship between the QBO periodic oscillations in the upper and in the lower stratosphere,  
244 orthogonal zonal winds measured at 10hPa (Fig.3a; orange) and 30hPa (Fig.3a; green) by the  
245 ground-station in Singapore have been considered here (Randel and Wu, 1996; Hood and  
246 Soukharev, 2006).

247 - *NAO, AAO and ENSO*: The El Niño/Southern Oscillation is represented by the 3-month  
248 running mean of Sea Surface Temperature (SST) anomalies (in degrees Celsius) in the Niño  
249 region 3.4 (region bounded by 120°W-170°W and 5°S- 5°N). Raw data are taken from marine  
250 ships and buoys observations. The North Atlantic and Antarctic Oscillations are described by the  
251 daily (or monthly) NAO and AAO indices which are constructed from the daily (or monthly)  
252 mean 500-hPa height anomalies in the 20°N-90°N region and 700-hPa height anomalies in the  
253 20°S-90°S region, respectively. Detailed information for these proxies can be found in  
254 <http://www.cpc.ncep.noaa.gov/>. These proxies describe important dynamical features which  
255 affect ozone distributions in both the troposphere and the lower stratosphere (e.g. Weiss et al.,  
256 2001, Frossard et al., 2013; Rieder et al., 2013; and references therein). The daily or 3-monthly  
257 average indexes used to parameterize these fluctuations are shown in Fig. 3 (b). The NAO and  
258 AAO indexes are used for the N.H. and the S.H. (Southern Hemisphere), respectively (both are  
259 used for the equatorial band). These proxies have been included in the statistical model for  
260 completeness even if they are expected to only have a weak apparent contribution to the IASI  
261 ozone time series due to their large spatial variability in a zonal band (e.g. Frossard et al., 2013;

262 Rieder et al., 2013). We have verified that including a typical time-lag relation between ozone  
263 and the ENSO variable from 0 to 4 months did not improve the regression model in terms of  
264 residuals and uncertainty of the fitted parameters. As a consequence, a time-lag has not been  
265 taken into account in our study.

266 - *Effective equivalent stratospheric chlorine (EESC)*: The EESC is a common proxy used for  
267 describing the influence of the ODS in O<sub>3</sub> variations. However, because the IASI time series  
268 starts several years after the turnaround for the ozone hole recovery in 1996/1997 (WMO, 2010),  
269 their influence is not represented by a dedicated proxy but is rather accounted for by the linear  
270 trend term.

271  
272 Even if some of the above proxies are only specific to processes occurring in the stratosphere, we  
273 adopt the same approach (geophysical variables, model and regression procedure) for adjusting  
274 the IASI O<sub>3</sub> time series in the troposphere. This proves useful in particular to account for the  
275 stratospheric contribution to the tropospheric layer (~30-60%; see Section 2 and Supporting  
276 Information, Fig. S5) due to stratosphere-troposphere exchanges (STE) and to the fact that this  
277 tropospheric layer is not perfectly decorrelated from the stratosphere. This has to be kept in mind  
278 when analyzing the time series in the troposphere in Section 4. Specific processes in the  
279 troposphere such as emissions of ozone precursors, long-range transport and in situ chemical  
280 processing are taken into account in the model in the harmonic and the linear trend terms of the  
281 Eq. 1 (e.g. Logan et al., 2012). Including harmonic terms having 4- and 3-month periods in the  
282 model has been tested to describe O<sub>3</sub> dependency on shorter scales (e.g. Gebhardt et al., 2014),  
283 but this did not improved the results in terms of residuals and uncertainty of correlation  
284 coefficients.

### 285 286 **3.3. Iterative backward variable selection**

287 Similarly to previous studies (e.g. Steinbrecht et al., 2004; Mäder et al. 2007, 2010; Knibbe et al.,  
288 2014), we perform an iterative stepwise backward elimination approach, based on p-values of the  
289 regression coefficients for the rejection, to select the most relevant combination of the above  
290 described regression variables (harmonic, linear and explanatory) to fit the observations. The  
291 minimum p-value for a regression term to be removed (exit tolerance) is set at 0.05, which

292 corresponds to a significance of 95%. The initial model which includes all regression variables is  
293 fitted first. Then, at each iteration, the variables characterized by p-values larger than 5% are  
294 rejected. At the end of the iterative process, the remaining terms are considered to have  
295 significant influence on the measured O<sub>3</sub> variability while the rejected variables are considered to  
296 be non-significant. The correction accounting for the autocorrelation in the noise residual is then  
297 applied to give more confidence in the coefficients determination.

298

#### 299 **4 Ozone variations observed by IASI**

300 In this section, we first examine the ozone variations in IASI time series during 2008-2013 in the  
301 four layers defined in the troposphere and the stratosphere to match the IASI sensitivity (Section  
302 2). The performance of the multiple linear model is evaluated in subsection 4.2 in terms of  
303 residuals errors, regression coefficients and associated uncertainties determined from the  
304 regression procedure (Section 3). Based on this, we characterize the principal physical processes  
305 that affect the IASI ozone records. Finally, the ability of IASI to derive apparent trends is  
306 examined in sub-section 4.3.

307

#### 308 **4.1 O<sub>3</sub> time series from IASI**

309 Figure 4 (a) shows the time development of daily O<sub>3</sub> number density over the entire altitude  
310 range of the retrieved profiles based on daily medians. The time series cover the six years of  
311 available IASI observations and are separated in three 20-degree latitude belts: 30°N-50°N (top  
312 panel), 10°N-10°S (middle panel) and 30°S-50°S (bottom panel). The figure shows the well-  
313 known seasonal cycle at mid-latitudes in the troposphere and the stratosphere with maxima  
314 observed in spring-summer and in winter-spring, respectively, and a strong stability of ozone  
315 layers with time in the equatorial belt. At high latitudes of both hemispheres, the high ozone  
316 concentrations and the large amplitude of the seasonal cycle observed in MLST and UTLS are  
317 mainly the consequence of the large-scale downward poleward Brewer-Dobson circulation  
318 which is prominent in late winter below 25 km.

319

320 Figure 4 (b) presents the estimated statistical uncertainty on the O<sub>3</sub> profiles retrieved from  
321 FORLI. This total error depends on the latitude and the season, reflecting, amongst other, the

322 influence of signal intensity, of interfering water lines and of thermal contrast under certain  
323 conditions (e.g. temperature inversion, high thermal contrast at the surface). It usually ranges  
324 between 10 and 30% in the troposphere and in the UTLS (Upper Troposphere-Lower  
325 Stratosphere), except in the equatorial belt due to the low O<sub>3</sub> amounts (see Fig.4 (a)) which leads  
326 to larger relative errors. The retrieval errors are usually less than 5% in the stratosphere.

327  
328 The relative variability (given as the standard deviation) of the daily median O<sub>3</sub> time series  
329 presented in Fig.4 (a) is shown in Fig.5, as a function of time and altitude. It is worth noting that,  
330 except in the UTLS over the equatorial band, the variability is larger than the estimated retrieval  
331 errors of the FORLI-O<sub>3</sub> data (~25% vs ~15% and ~10% vs ~5%, on average over the troposphere  
332 and the stratosphere, respectively), reflecting that the high natural temporal variability of O<sub>3</sub> in  
333 zonal bands is well captured with FORLI (Dufour et al., 2012; Hurtmans et al., 2012). The  
334 standard deviation is larger in the troposphere and in the stratosphere below 20 km where  
335 dynamic processes play an important role. The largest values (>70% principally in the northern  
336 latitudes during winter) are measured around 9-15km altitude. They highlight the influence of  
337 tropopause height variations and the STE processes. In the stratosphere, the variability is always  
338 lower than 20% and becomes negligible in the equatorial region. Interestingly, the lowest  
339 troposphere of the N.H. (below 700 hPa; <4km) is marked by an increase in both O<sub>3</sub>  
340 concentrations (Fig. 4a) and standard deviations (between ~30% and ~45%) in spring-summer,  
341 the latter being larger than the total retrieval error (less than 25%, see Fig. 4 (b)). The lower  
342 tropospheric column (e.g. ground-700 hPa) can generally not well be discriminated because of  
343 the weak sensitivity of IASI in the lowermost layers (Section 2). However, the measurements in  
344 northern mid-latitudes in spring-summer are characterized by a larger sensitivity. In the ground-  
345 700hPa columns, we find that the apriori contributions do not exceed 40% and they range  
346 between 10% and 20% over the continental regions. In addition, the stratosphere-troposphere  
347 exchanges are usually the weakest in summer. The stratospheric contributions into the IASI  
348 MLT columns are estimated to be the lowest in the summer mid-latitudes N.H. (e.g. ~35% in  
349 the 30°N-50°N band; See Fig. S5 (b) of the Supplement) and, as mentioned in Section 2, the  
350 real natural contribution originating from the troposphere reaches ~55% (cfr Fig.S6 (b) in  
351 Supplement). This certainly helps in detecting the real variability of O<sub>3</sub> in the N.H. troposphere,

352 and, the increase in the observed concentrations and in the variability may likely indicate a  
353 photochemical production of O<sub>3</sub> associated with anthropogenic precursor emissions (e.g. Logan  
354 et al., 1985; Fusco and Logan, 2003; Dufour et al., 2010; Cooper et al., 2010; Wilson, et al.,  
355 2012; Safieddine et al., 2013). Changes in biomass and biogenic emissions of NO<sub>x</sub>, CO and non-  
356 methane organic volatile compounds (NMVOC) may also play a role. However, they only  
357 represent a small part of the total emissions for NO<sub>x</sub> and CO (e.g. ~23% vs 72% for the  
358 anthropogenic NO<sub>x</sub> emissions and ~40% vs 60% for the anthropogenic CO emissions from the  
359 emissions dataset used in the Supplement), while the biogenic emissions of NMVOC represent  
360 the largest contribution to the total (~80%).

361  
362 The zonal representation of the O<sub>3</sub> variability seen by IASI is given in Fig. 6. It shows the daily  
363 number density at altitude levels corresponding to maximum of sensitivity in the four analyzed  
364 layers in most of the cases (600 hPa - ~6km; 240 hPa - ~10km; 80 hPa - ~20km; 6 hPa - ~35  
365 km) (Section 2). The top panel (~35 km) reflects well the photochemical O<sub>3</sub> production by  
366 sunlight with the highest values in the equatorial belt during the summer (~3x10<sup>12</sup>  
367 molecules/cm<sup>3</sup>). The middle panels (~20 km and ~10 km) shows the transport of ozone rich-air  
368 to high latitudes in late winter (up to ~6x10<sup>12</sup> molecules/cm<sup>3</sup> in the N.H.) which is induced by the  
369 Brewer-Dobson circulation. The fact that the patterns at ~10km are similar to those at ~20 km  
370 mainly reflects the low sensitivity of IASI to that level compared to the others. Finally, the lower  
371 panel (~6 km) presents high O<sub>3</sub> levels in spring at high latitudes (~1.4x10<sup>12</sup> molecules/cm<sup>3</sup> in the  
372 N.H.), which likely reflects both the STE processes and the contribution from the stratosphere  
373 due to the medium IASI sensitivity to that layer (see Section 2 and Supporting Information), and  
374 a shift from high to middle latitudes in summer which could be attributed to anthropogenic O<sub>3</sub>  
375 production. The MLT panel also reflects the seasonal oscillation of the Inter-Tropical  
376 Convergence Zone (ITCZ) around the Equator and the large fire activity in spring around 20°S-  
377 40°S.

#### 378 379 **4.2 Multivariate regression results: Seasonal and explanatory variables**

380 Figure 4(a) shows superimposed on the time series of the IASI ozone concentration profile, those  
381 of the partial columns (dots) for the 4 layers (color contours). The adjusted daily time series to

382 these columns with the regression model defined by Eq.1 is also overlaid and shown by colored  
383 lines. The model represents reasonably well the ozone variations in the four layers, with, as  
384 illustrated for three latitude bands, good coefficient correlations (e.g.  $R_{MLT}=0.94$ ;  $R_{UTLS}=0.91$ ;  
385  $R_{MLT}=0.90$  and  $R_{US}=0.91$  for the 30°N-50°N band) and low residuals (< 8%) in all cases. The  
386 regression model explains a large fraction of the variance in the daily IASI data over the  
387 troposphere (~85%-95%) and the stratosphere (~85%-95% in all cases, except for the UST with  
388 ~70-95%), as estimated from  $\frac{\sigma(O_3^{Fitted\_model}(t))}{\sigma(O_3(t))}$  where  $\sigma$  is the standard deviation relative to the  
389 fitted regression model and to the IASI O<sub>3</sub> time series.

390  
391 However, note that the fit fails to reproduce the highest ozone values ( $>5 \times 10^{12}$  molecules/cm<sup>3</sup>)  
392 above the seasonal maxima for 30°N-50°N latitude band, especially in the MLST during the  
393 springs 2009 and 2010. This could be associated with occasional downward transport of upper  
394 atmospheric NO<sub>x</sub>-rich air occurring in winter and spring at high latitudes (Brohede et al., 2008)  
395 following the strong subsidence within the intense Arctic vortex in 2009-2010 (Pitts et al., 2011)  
396 or with the missing time-lags in the regression model between the QBO and ENSO variables and  
397 the response of mid-latitude lower stratospheric ozone (Neu et al., 2014).

398  
399 Fig.7 displays the annual cycle averaged over the 6 years recorded by IASI (dots) for the studied  
400 layers and bands, as well as that from the fit of the daily O<sub>3</sub> columns (lines). The regression  
401 model follows perfectly the O<sub>3</sub> variations in terms of timing of O<sub>3</sub> maxima and of amplitude of  
402 the cycle. The fit is generally characterized by low residuals (<10%) and good correlation  
403 coefficients (0.70-0.95), which indicates that the regression model is suitable to describe the  
404 zonal variations. Exception is found over the Southern latitudes (residual up to 15% and R down  
405 to 0.61) probably because of the variation induced by the ozone hole formation which is not  
406 parameterized in the regression model, and because of the low temporal sampling of daytime  
407 IASI measurements in this region.

408  
409 From Figure 7, the following general patterns in the O<sub>3</sub> seasonal cycle can be isolated from the  
410 zonally averaged IASI datasets:

- 411 1- In UST (top left panel), the maxima is in the equatorial belt, around  $4.7 \times 10^{18}$  molecules/cm<sup>2</sup>  
412 throughout the year and the amplitudes are small compared to the averaged O<sub>3</sub> values. The  
413 largest amplitude in the annual cycle is found in the N.H. between 30N and 50N where O<sub>3</sub>  
414 peaks in July after the highest solar elevation (in June) following a progressive buildup  
415 during spring-summer. In agreement with FTIR observations (e.g. Steinbrecht et al., 2006;  
416 Vigouroux et al., 2008), a shift of the O<sub>3</sub> maximum from spring (March-April) to late  
417 summer (August-September) is found as one moves from high to low latitudes in the N.H. In  
418 the S.H., the general shape of the annual cycle which shows a peak in October-November  
419 before the highest solar elevation (in December), results from loss mechanisms depending on  
420 annual cycle of temperatures and other trace gases. Other effects such as changing Brewer-  
421 Dobson circulation, light absorption and tropical stratopause oscillations may also  
422 considerably impact the cycle in this layer (Brasseur and Solomon, 1984; Schneider et al.,  
423 2005).
- 424 2- In the lower stratosphere (MLST and UTLS, top right and bottom left panels), the  
425 pronounced amplitudes of the annual cycle is dominated by the influence of the Brewer  
426 Dobson circulation with the highest O<sub>3</sub> values observed over polar regions (reaching  $\sim 6 \times 10^{18}$   
427 molecules/cm<sup>2</sup> on average vs  $\sim 2 \times 10^{18}$  molecules/cm<sup>2</sup> on average in the equatorial belt). The  
428 maximum is shifted from late winter at high latitudes to spring at lower latitudes.
- 429 3- In MLT (bottom right panel), we clearly see a large hemispheric difference with the highest  
430 values over the N.H. (also in UTLS). Maxima are observed in spring, reflecting more  
431 effective STE processes. A particularly broad maximum from spring to late summer is  
432 observed in the 30N-50N band. It probably points to anthropogenic production of O<sub>3</sub>. This  
433 has been further investigated in the Supplement through MOZART4-IASI comparison by  
434 using constant anthropogenic emissions in the model settings (see Fig. S2). The results show  
435 clear differences between the modeled and the observed MLT seasonal cycles, which  
436 highlights the need for further investigation of the role of anthropogenically produced O<sub>3</sub> and  
437 the realism of anthropogenic emissions inventories.

438  
439 Figure 8 presents all the fitted regression parameters included in Eq. 1 (Section 3) in the four  
440 layers as a function of latitude. The uncertainty in the 95% confidence limits which accounts for

441 the autocorrelation in the noise residual is given by error bars. The constant term (Fig.8a) is  
 442 found to be statistically significant (uncertainty<10%) in all cases. It captures the two ozone  
 443 maxima in the stratosphere: one over the Northern Polar regions in the MLST and one at  
 444 equatorial latitudes in the UST ( $\sim 4.5 \times 10^{18}$  molecules/cm<sup>2</sup>), the important decrease of O<sub>3</sub> in the  
 445 lower stratospheric layers (UTLS and MLST) moving from high to equatorial latitudes, and the  
 446 weak negative and strong positive gradients in the Northern MLT and in the UST, respectively.  
 447 The sum of the constant terms of the four layers varies between  $7.50 \times 10^{18}$  (equatorial region) and  
 448  $9.50 \times 10^{18}$  molecules/cm<sup>2</sup> (polar regions) and is similar to the one of the fitted total column  
 449 (relative differences < 3.5%) (red line). Note that the constant terms in the UTLS region in the  
 450 mid-latitudes and in the tropics are certainly affected by the fact that the FORLI-O<sub>3</sub> profiles are  
 451 biased high by ~10-15% in this layer and latitude bands (Dufour et al., 2012; Gazeaux et al.,  
 452 2012). The representativeness of the 20-degree zonal averages in terms of spatial variability has  
 453 been examined by fitting the IASI time series for specific locations in the N.H. (results shown  
 454 with stars in Fig.8a): the constant terms are found to be consistent, within their uncertainties,  
 455 with those averaged per latitude bands in all cases. Over the polar region where O<sub>3</sub> shows a large  
 456 natural variability, the regression coefficient is characterized by a large uncertainty.

457

458 The regression coefficients for other variables (harmonic and proxy terms) which are retained in  
 459 the regression model by the stepwise elimination procedure are shown in Fig.8 (b). They are  
 460 scaled by the fitted constant term and the error bars represent the uncertainty in the 95%  
 461 confidence limits accounting for the autocorrelation in the noise residual. A positive (or  
 462 negative) sign of the coefficients indicates that the associated variables are correlated (or anti-  
 463 correlated) with the IASI O<sub>3</sub> time series. Note that if the uncertainty is larger than its associated  
 464 estimate (i.e. larger than 100%, corresponding to an error bar overlapping the zero line), it means  
 465 that the estimate becomes statistically non-significant when accounting for the autocorrelation in  
 466 the noise residuals at the end of the elimination process. This is summarized in Table S1 of the  
 467 Supplement. The contribution of the fitted variables into the IASI O<sub>3</sub> variations is estimated as

468 
$$\frac{\sigma([a_n; b_n; x_j] [\cos(n\omega t); \sin(n\omega t); X_{norm,j}])}{\sigma(O_3(t))}$$
 where  $\sigma$  is the standard deviation relative to the fitted

469 signal of harmonic or proxy terms and to the IASI O<sub>3</sub> time series. From Figure 8, we find that:

470 1- The annual harmonic term (upper left) is the main driver of the O<sub>3</sub> variability and largely  
471 dominates (scaled  $a_1+b_1$  around  $\pm 40\%$ ) over the semi-annual one (upper right; scaled  $a_2+b_2$   
472 around  $\pm 15\%$ ). In UTLS and MLST, its amplitude decreases from high to low latitudes  
473 likely following the cycle induced by the Brewer-Dobson circulation (*cfr.* Fig.6 and Fig.7)  
474 and the sign of the coefficient accounts for the winter-spring maxima in both hemispheres  
475 (negative values in the S.H. and positive ones in the N.H). The annual term contributes  
476 importantly around 45%-85% of the observed O<sub>3</sub> variations, except in the 10°N-30°N and  
477 equatorial bands (10%-30%), while the influence of the semi-annual variation on O<sub>3</sub> is  
478 smaller (10%-25%) and highly variable between the bands. In the UST, the amplitudes vary  
479 only slightly (around -5% to 5%) and account for the weak summer maximum. The  
480 contributions of the annual harmonic term are estimated between 5%-30%. As expected, the  
481 uncertainties associated with the annual terms are very weak and most of the harmonic terms  
482 (annual and seasonal) are statistically significant.

483 2- The QBO and solar flux proxies are generally minor (scaled coefficients <10% and  
484 contributions <15%) and they are often statistically non-significant contributors to O<sub>3</sub>  
485 variations after accounting for the autocorrelation in the noise residual (see Table S1 in the  
486 Supplement), except in equatorial region (scaled coefficients of 10-15% in UTLS and  
487 contributions up to 75% and 21% for QBO and SF, respectively) where they are important  
488 drivers of O<sub>3</sub> variations (e.g. Logan et al., 2003; Steinbrecht et al., 2006b; Soukharev and  
489 Hood, 2006; Fadnavis and Beig, 2009). Previous studies have indeed supported the solar  
490 influence on the lower stratospheric equatorial dynamics (e.g. Soukharev and Hood, 2006;  
491 McCormack et al., 2007). Note that the QBO<sup>30</sup> proxy (data not shown) has negative  
492 coefficients for the mid-latitudes, which is in line with Frossard et al. (2013).

493 3- The contributions described by the ENSO and NAO/AAO proxies are generally very weak  
494 (<10% and <5%, respectively), with scaled coefficients lower than 5%, and, in many cases  
495 for the NAO/AAO proxies, they are even not statistically significant when taking into  
496 account the correlation in the noise residuals (see Table S1 in Supplement). Despite of this, it  
497 is worth pointing out that their effects to the O<sub>3</sub> variations are comparable to the results  
498 published in the previous studies. The negative ENSO coefficient in the tropical UTLS is  
499 consistent with results from Neu et al. (2014). Rieder et al. (2013) and Frossard et al. (2013)

500 have also shown large regions of negative coefficients for NAO North of 40°N, and large  
501 regions of positive and negative coefficient estimates for ENSO, North of 30°N and South of  
502 30°S, respectively.

503  
504 We note that the non-representation of time-lags in the proxy time series may be underestimating  
505 the role of some geophysical variables on O<sub>3</sub> variations, in particular that of ENSO and QBO in  
506 zonal bands outside the regions where these geophysical quantities are measured (i.e. Niño  
507 region 3.4 for ENSO and Singapore for QBO). Finally, we see in Fig.8 (b), large uncertainties  
508 associated with the regression coefficients in UTLS in comparison with other layers, and in polar  
509 regions in comparisons with other bands. We interpret this as an effect from the high natural  
510 variability of O<sub>3</sub> measured by IASI in UTLS (see Fig.5) and from missing parameterizations and  
511 low temporal sampling of daytime IASI measurements over the poles, respectively.

512  
513 As a general feature, the results demonstrate the representativeness of the fitted models in each  
514 layer and latitude band. This good performance of the regression procedure allows examination  
515 of the adjusted linear trend term in Section 4.3 below.

516  
517 **4.3 Multivariate regression results: trend over 2008-2013**

518 An additional goal of the multivariate regression method applied to the IASI O<sub>3</sub> time series is to  
519 determine the linear trend term and its associated uncertainty. Despite the fact that more than 10  
520 years of observations, corresponding to the large scale of solar cycle, is usually required to  
521 perform such a trend analysis, we could argue that statistically relevant trends could possibly be  
522 derived from the first six years of IASI observations, owing to the high spatio-temporal  
523 frequency (daily) of IASI global observations, to the daily “fingerprint” in the solar flux (see  
524 Figure 3 (a)), possibly making it distinguishable from a linear trend, and to its weak contribution  
525 to O<sub>3</sub> variations (see section 4.2. and references therein). To verify the specific advantage of  
526 IASI in terms of frequency sampling, we compare, in the subsections below, the statistical  
527 relevance of the trends when retrieved from the monthly averaged IASI datasets *vs* the daily  
528 averages as above, in the 20° zonal bands for the 4 partial and the total columns.

529

### 530 **4.3.1. Regressions applied on daily vs monthly averages**

531 Figure 9 (top) provides, as an example, the 6-year time series of IASI O<sub>3</sub> daily averages (left  
532 panels) compared to the monthly averages (right panels) for the 30°S-50°S latitude band in the  
533 UST (dark blue), along with the results from the regression procedure (light blue). Note that  
534 either daily or monthly F10.7, NAO and AAO proxies (see Table 1) are used depending on the  
535 frequency of the IASI O<sub>3</sub> averages to be adjusted. The second row in Fig.9 provides the  
536 deseasonalised IASI and fitted time series, calculated by subtracting the model seasonal cycle  
537 from the time series, as well as the residuals (red curves). The averaged residuals relative to the  
538 deseasonalised IASI time series strongly vary with the layers and latitudinal bands and usually  
539 range between 30% and 60%. The fitted signal in DU of each proxy is shown on the bottom  
540 panels. The O<sub>3</sub> time series and the solar flux signal resulting from the adjustment without the  
541 linear term trend in the regression model are also represented (orange lines in 2<sup>d</sup> and bottom  
542 panels, respectively). When it is not included in the regression model, the linear trend term is  
543 only partly compensated by the solar flux term in the daily averages. This leads to an offset  
544 between the fitted O<sub>3</sub> time series resulting from the both regression models (with and without the  
545 linear term), which corresponds well to a trend over the IASI period, and, consequently, to larger  
546 residuals (e.g. 80% without vs 44% with the linear term for this example and 94% without vs  
547 58% with the linear term for the 30°S-50°S band in the MLST illustrated in Fig. S1 of the  
548 Supplement). This offset is observed for a lot of layers and latitudinal bands. On the contrary, the  
549 linear term can largely be compensated by the solar flux term in the monthly averages: the offset  
550 is weak and the relative difference between the both fitted models is smaller (averaged  
551 differences relative to the deseasonalised IASI time series of 10% in monthly data vs 17% in  
552 daily data for this example). In this example, the linear and solar flux terms are even not  
553 simultaneously retained in the iterative stepwise backward procedure when applied on the  
554 monthly averages while they are when applied on daily averages. This effective co-linearity of  
555 the linear and the monthly solar flux terms translates to larger model fit residuals (44% in daily  
556 averages vs 60% in monthly averages in UST, relative to the deseasonalised IASI time series), to  
557 smaller relative differences between both regression models (with and without the linear term)  
558 (17% in daily vs 10% in monthly data), and to larger uncertainty on the trend coefficients when  
559 using the monthly data in comparison with the daily data. This even leads, in this specific

560 example, to a not statistically significant linear term of  $1.21 \pm 1.30 \text{ DU/yr}$  when derived from  
561 monthly averages *vs* a significant trend of  $1.74 \pm 0.77 \text{ DU/yr}$  from daily averages.

562  
563 The same conclusions can be drawn from the fits in other layers and latitude bands, especially  
564 those where the solar cycle variation of ozone is large (MLST and UTLs) or where the ozone  
565 recovery occurs (UST). A larger trend uncertainty associated with monthly data *vs* daily data is  
566 found in all situations (see Table 2, Section 4.3.2 below).

567  
568 This brings us to the important conclusion that, thanks to the unprecedented sampling of IASI,  
569 apparent trends can be detected in FORLI-O<sub>3</sub> time series even on a short period of  
570 measurements. This supports the need for regular and high frequency measurements for  
571 observing ozone variations underlined in other studies (e.g. Saunio et al., 2012). The O<sub>3</sub> trends  
572 from the daily averages of IASI measurements are discussed and compared with results from the  
573 monthly averages in the subsection below.

574

#### 575 **4.3.2. O<sub>3</sub> trends from daily averages**

576 Table 2 summarizes the trends and their uncertainties in the 95% confidence limit, calculated for  
577 each 20° zonal band and for the 4 partial and the total columns. In the northern and southern  
578 polar regions, the polar night period is not covered because only IASI observations during  
579 sunlight (over Feb-Oct and Oct-Apr for N.H. and S.H., respectively) are used in this study (See  
580 Section 2). For the sake of comparison, the trends are reported for both the daily (top values) and  
581 the monthly (bottom values) averages, and their uncertainties account for the auto-correlation in  
582 the noise residuals considering a time lag of 1-day or 1-month, respectively. We show that the  
583 daily and monthly trends in all layers and all latitude bands fall within each other uncertainties,  
584 but that the use of daily median strongly helps in reducing everywhere the uncertainty associated  
585 with the trends for the reasons discussed above (Section 4.3.1). This is particularly observed in  
586 the UST where the ozone hole recovery would occur, but also in the MLST and the UTLs where  
587 the solar cycle variation of ozone is the largest (see Figure 8). As a consequence, the UST trends  
588 in monthly averages are shown to be mostly non-significant in comparison with those from daily

589 averages. Table 3 summarizes the trends in the daily averages for two 3-month periods: June-  
590 July-August (JJA) and December-January-February (DJF).

591  
592 From Tables 2 and 3, we observe very different trends according to the latitude and the altitude.  
593 From Table 2, we find for the total columns that the trends derived from the daily medians are  
594 only significant at high northern latitudes and that they are interestingly of the same order as  
595 those obtained from other satellites and assimilated satellite data (Weatherhead and Anderson,  
596 2006; Knibbe et al., 2014) or from ground-based measurements (Vigouroux et al., 2008)  
597 calculated over longer time periods. The non-significant trends calculated for the mid- and low  
598 latitudes of the N.H. are also comparable to the results published in the previous studies (Reinsel  
599 et al., 2005; Andersen et al., 2006a; Vigouroux et al., 2008). Regarding the individual layers, we  
600 find the following:

601 1- In the US, significant positive trends are observed in both hemispheres from the daily  
602 medians, particularly over the mid- and high latitudes of both hemispheres (e.g.  $1.74 \pm 0.77$   
603 DU/yr in the 30°S-50°S band, i.e., 12%/decade) where the changes in ozone trends before  
604 and after the turnaround in 1997 have been found to be the highest. Kyrola et al. (2013) and  
605 Laine et al. (2014) report for instance a change of up to 10%/decade in O<sub>3</sub> trends between  
606 1997-2011 vs 1984-1997. Positive trends in the UST are consistent with many previous  
607 observations if one considers the fact that the period covered by IASI is later than those  
608 reported in previous studies and that the recovery rate seems to increase since the beginning  
609 of the turnaround (Knibbe et al. (2014) reports a factor of two increase in the recovery rate  
610 between 1997-2010 with  $\sim 0.7$ DU/yr and 2001-2010 with  $\sim 1.4$ DU/yr in the S.H.). They could  
611 indicate a leveling off of the negative trends that were observed since the second half of the  
612 1990's mostly from satellites and ground-based monthly mean data (e.g. WMO 2006, 2011;  
613 Randel and Wu, 2007; Vigouroux et al., 2008; Steinbrecht et al., 2009; Jones et al., 2009;  
614 McLinden et al., 2009; Laine et al. 2014; Nair et al., 2014). The causes of this "turnaround"  
615 remain, however, uncertain. If the compensating impact of decreasing chlorine in recent  
616 years and maximum solar cycle (over 2011-2012 in the period studied here) is probably part  
617 of the answer (e.g. Steinbrecht et al., 2004), the effects of changing stratospheric  
618 temperatures and Brewer-Dobson circulation (Salby et al., 2002; Reinsel et al., 2005;

619 Dhomse et al., 2006; Manney et al., 2006) could also contribute and should be further  
620 investigated. The long-lasting cold winter/spring 2011 in the Arctic leading to unprecedented  
621 ozone loss (Manney et al., 2011), could explain the non-significant trend in the 70°N-90°N  
622 band. This is supported by the results in winter (Table 3). From Table 3, we generally find  
623 significant positive trends in summer N.H. and weaker positive or even non-significant trends  
624 in winter S.H. A non-significant trend is also calculated for the 70°S-90°S band in spring  
625 (data not shown). This could indicate the strong influence of changing stratospheric  
626 temperatures on ozone depletion from year to year (e.g. Dhomse et al., 2006), leading to  
627 larger uncertainties in our trends estimations and larger fitting residuals (see Section 4.2) due  
628 to the fact that the stratospheric temperature is not taken into account as an explanatory  
629 variable in the model.

630 2- In the MLST, one can see that, except in the high latitude bands, the trends are either non-  
631 significant or significantly negative. This is in agreement with the trend analysis of Jones et  
632 al. (2009) for the 20-25 km altitude range over the 1997-2008 period, as well as with other  
633 studies at N.H. latitudes, which investigated O<sub>3</sub> changes in the 18-25 km range between 1996  
634 and 2005 (Miller et al., 2006; Yang et al., 2006; Kivi et al., 2007). The results derived  
635 separately for summer and winter in Table 3 are also in line with those of Kivi et al. (2007)  
636 which reported contrasted trends in the Arctic MLST depending on season.

637 3- In the UTLS, negative trends are calculated in the tropics and significant positive trends are  
638 found in the mid- and high latitudes of N.H., these latter falling within the uncertainties of  
639 those reported by Kivi et al. (2007) for the tropopause-150 hPa layer between 1996 and 2003.  
640 The large positive trends calculated at Northern latitudes (e.g.  $1.28 \pm 0.82$  DU/year in the  
641 70°N-90°N band) contribute for ~ 30 % to the positive trend for the total column. This result  
642 is consistent with Yang et al. (2006) which reported that UTLS contributes 50% to positive  
643 trends for the total columns measured in the mid-latitudes of the N.H. from ozonesondes. In  
644 that study, these positive trends were linked to changes in atmospheric dynamics either  
645 related to natural variability induced by potential vorticity and tropopause height variations  
646 or related to anthropogenic climate change. Hence, the apparent increase in total ozone in the  
647 mid-latitudes of the N.H. seen by IASI would reflect the combined contribution of dynamical  
648 variability and declining ozone-depleting substances (e.g. Weatherhead and Andersen, 2006;

649 WMO, 2006; Harris et al., 2008, Nair et al., 2014). It is worth to keep in mind that these  
650 effects are not independently accounted for in the regression model. Previous studies  
651 reported, however, that dynamical and chemical processes are physically coupled in the  
652 atmosphere, making difficult to define unambiguously such drivers in a statistical model (e.g.  
653 Mäder et al., 2007; Harris et al., 2008). On a seasonal basis (see Table 3), the trends seen by  
654 IASI at Northern latitudes in summer are all significantly positive and increasing towards the  
655 pole. Note that the trends in upper layers may contribute to the ones calculated in UTLS due  
656 to the medium IASI sensitivity to that layer (*cf.* Section 2).

657 4- In the MLT, most of the trends are significantly negative (Tables 2 and 3). The non-  
658 significant trends in polar regions could be partly related to the lack of IASI sensitivity to  
659 tropospheric O<sub>3</sub> (see Section 2, Fig.2). On a seasonal basis, we see that the negative trends  
660 are more pronounced during the JJA period (around  $-0.25 \pm 0.10$  DU/yr) for all bands except  
661 between 30°N and 10°S. In the N.H., these results tend to confirm the leveling off of  
662 tropospheric ozone observed in recent years during the summer months (Logan et al., 2012).  
663 This trend, however, remains difficult to interpret because it could be linked to a variety of  
664 processes including most importantly: the decline of anthropogenic emissions of ozone  
665 precursors, the increase of UV-induced O<sub>3</sub> destruction in the troposphere and STE processes  
666 (Isaksen et al., 2005; Logan et al., 2012; Parrish et al., 2012; Hess and Zbinden, 2013). As a  
667 consequence, it is hard to reconcile the trends in tropospheric ozone with changes in  
668 emissions of ozone precursors. However, trends in emissions have already been able to  
669 qualitatively explain measured ozone trends over some regions but with inconsistent  
670 magnitude between observations and model simulations (e.g. Cooper et al., 2010; Logan et  
671 al., 2012; Wilson et al., 2012). It is also worth to keep in mind that due to medium sensitivity  
672 of IASI to the troposphere, the a priori contribution and ozone variations in stratospheric  
673 layers may largely influence the trends seen by IASI in the MLT layer (*cf.* Section 2 and  
674 Supporting Information).

675

#### 676 4.3.3. O<sub>3</sub> trends from IASI vs FTIR data

677 In order to validate the trends inferred from IASI in the UST and in the total columns, we  
678 compare them with those obtained from ground-based FTIR measurements at several NDACC

679 stations (Network for the Detection of Atmospheric Composition Change, available at  
680 [http://www.ndsc.ncep.noaa.gov/data/data\\_tbl/](http://www.ndsc.ncep.noaa.gov/data/data_tbl/)) by using the same fitting procedure and taking  
681 into account the autocorrelation in the noise residuals. A box of  $1^\circ \times 1^\circ$  centered on the stations  
682 has been used for the collocation criterion. The regression model is applied on the daily FTIR  
683 data for a series of time periods starting after the turnaround point (from 1998 for mid-latitude  
684 stations and from 2000 for polar stations), as well as for the same periods as recently studied in  
685 Vigouroux et al. (2014) for the sake of comparison. Note that because we are not interested here  
686 in validating the IASI columns which was achieved in previous papers (e.g. Dufour et al., 2014;  
687 Oetjen et al., 2014) but in validating the trends obtained from IASI, we did not correct biases  
688 between IASI and FTIR due to different vertical sensitivity and *a priori* information. The results  
689 are given in DU/year in Table 4. We see large significant positive total column trends from IASI  
690 at middle and polar stations (e.g.  $5.26 \pm 4.72$  DU/yr at Ny-Alesund), especially during spring.  
691 These values are consistent with those reported in Knibbe et al. (2014) for the 2001-2010 period  
692 in spring in the Antarctic (around 3-5DU/yr). This trend is not obtained from the FTIR data for  
693 which trends are found to be mostly non-significant (even not retained in the stepwise  
694 elimination procedure in some cases) as reported in Vigouroux et al. (2014), except at  
695 Jungfraujoch which shows a trend of  $5.28 \pm 4.82$  DU/yr over the 2008–2012 period. For the  
696 periods starting before 2000, we calculated from FTIR, in agreement with Vigouroux et al.  
697 (2014), a significantly negative trend at Ny-Alesund for the total column and significantly  
698 positive trends at polar stations for the US. In addition, we see from Table 4 a leveling off of  $O_3$   
699 at polar stations in the UST after 2003, as previously reported in Vigouroux et al. (2014), which  
700 was explained by a compensation effect between the decrease of solar cycle after its maximum in  
701 2001-2002 and a positive trend. These trends are, however, non-significant and inferred only  
702 from few FTIR measurements (see Number of days column, Table 4).

703

704 From IASI, it is worth to point out that, in all cases, positive trends are calculated in the UST  
705 (even if some are not significant) and that these trends are consistent with those calculated from  
706 FTIR data covering a ~11-year period and starting after the turnaround (e.g. at Thule;  
707  $1.24 \pm 1.09$  DU/yr from IASI for the period 2008-2013 vs  $1.42 \pm 0.78$  DU/yr from the FTIR over  
708 2001-2012). This is illustrated for three stations (Ny-Alesund, Thule and Kiruna) in Fig.10 which

709 compares the time series from IASI (2008-2013, in red) with those from FTIR covering periods  
710 starting after the turnaround (in blue). Their associated trends as well as the trend calculated from  
711 FTIR covering the IASI period (in green) are also indicated.

712  
713 In order to better characterize the effect of the temporal frequency on determining statistical  
714 trends, the IASI time series have been subsampled to match the temporal resolution of FTIR. The  
715 associated trend values are also indicated in Table 4 (2<sup>d</sup> row). In any cases, we observe that the  
716 fitted trends inferred from both IASI and FTIR with the same temporal samplings are within the  
717 uncertainties of each other and that those associated with the subsampled IASI datasets are  
718 significantly larger than those obtained with the daily ones, leading to statistically non-  
719 significant trends.

720  
721 Even if validating the IASI fitted trends with independent datasets is challenging due to the  
722 short-time period of available IASI measurements and the insufficient number of usable  
723 correlative measurements over such a short period, the results obtained for IASI vs FTIR tend to  
724 confirm the conclusion drawn in subsections 4.3.1 and 4.3.2, that the high temporal sampling of  
725 IASI provides good confidence in the determination of the trends even on periods shorter than  
726 those usually required from other observational means.

727  
728 **6 Summary and conclusions**

729 In this study, we have analyzed 6 years of IASI O<sub>3</sub> profile measurements as well as the total O<sub>3</sub>  
730 columns based on the profile. Four layers have been defined following the ability of IASI to  
731 provide reasonably independent information on the ozone partial columns: the mid-lower  
732 troposphere (MLT), the upper troposphere – lower stratosphere (UTLS), the mid-lower  
733 stratosphere (MLST) and the upper stratosphere (UST). Based on daily values of these four  
734 partial or of the total columns in 20-degree zonal averages, we have demonstrated the capability  
735 of IASI for capturing large scale ozone variability (seasonal cycles and trends) in these different  
736 layers. We have presented daytime vertical and latitudinal distributions for O<sub>3</sub> as well as their  
737 evolution with time and we have examined the underlying dynamical or chemical processes. The  
738 distributions were found to be controlled by photochemical production leading to a maximum in

739 summer at equatorial region in the UST, while they reflect the impact of the Brewer-Dobson  
740 circulation with maximum in winter-spring at mid- and high latitude in the MLST and in the  
741 troposphere. The effect of the photochemical production of O<sub>3</sub> from anthropogenic precursor  
742 emissions was also observed in the troposphere with a shift in the timing of the maximum from  
743 spring to summer in the mid-latitudes of the N.H.

744  
745 The dynamical and chemical contributions contained in the daily time development of IASI O<sub>3</sub>  
746 have been analyzed by fitting the time series in each layer and for the total column with a set of  
747 parameterized geophysical variables, a constant factor and a linear trend term. The model was  
748 shown to perform well in term of residuals (<10%), correlation coefficients (between 0.70 and  
749 0.99) and statistical uncertainties (<7%) for each fitted proxies. The annual harmonic terms  
750 (seasonal behavior) were found to be largely dominant in all layers but the US, with fitted  
751 amplitudes decreasing from high to low latitudes in agreement with the Brewer-Dobson  
752 circulation. The QBO and solar flux terms were calculated to be important only in the equatorial  
753 region, while other dynamical proxies accounted for in the regression (ENSO, NAO, AAO) were  
754 found negligible.

755  
756 Despite the short time period of available IASI dataset used in this study (2008-2013) and the  
757 potential ambiguity between the solar and the linear trend terms, statistically significant trends  
758 were derived from the six first years of daily O<sub>3</sub> partial columns measurements (on the contrary  
759 to monthly averages which lead to mostly non-significant trends). This result which was  
760 strengthened from comparisons with the regression applied on local FTIR measurements, is  
761 remarkable as it demonstrates the added value of IASI exceptional frequency sampling for  
762 monitoring medium to long-term changes in global ozone concentrations. We found two  
763 important apparent trends:

764 1) Significant positive trends in the upper stratosphere, especially at high latitudes in both  
765 hemispheres (e.g.  $1.74 \pm 0.77$  DU/yr in the 30°S-50°S band), which are consistent with a probable  
766 “turnaround” for upper stratospheric O<sub>3</sub> recovery (even if the causes of such a turnaround are  
767 still under investigations). In addition, the trends calculated for some local stations are in line  
768 with those calculated from FTIR measurements after the turnaround.

769 2) Negative trends in the troposphere at mid- and high Northern latitudes, especially during  
770 summer (e.g.  $-0.26 \pm 0.11$  DU/yr in the 30°N-50°N band) which are in line with the decline of  
771 ozone precursor emissions.

772

773 To confirm the above findings beyond the 6 first years of IASI measurements and to better  
774 disentangle the effects of dynamical changes, of the 11-year solar cycle and of the equivalent  
775 effective stratospheric chlorine (EESC) decline on the O<sub>3</sub> time series, further years of IASI  
776 observations will be required, and more complete fitting procedures (including, among others,  
777 proxies to account for the decadal trend in the EESC, for the ozone hole formation, for changes  
778 in the Brewer-Dobson circulation, as well as including time lags in ENSO and QBO proxies) will  
779 have to be explored. Further investigation on the regressors uncertainties and on the total error on  
780 ozone measurements should be performed as well to understand on the unexplained variations in  
781 IASI O<sub>3</sub> records.

782

783 This will be achievable with the long-term homogeneous records obtained by merging  
784 measurements from the three successive IASI instruments on MetOp-A (2006); -B (2012) and –  
785 C (2018), and by IASI successor on EPS-SG after 2021 (Clerbaux and Crevoisier, 2013;  
786 Crevoisier et al., 2014).

787

788

789

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802

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1158 **Table 1** List of the proxies used in this study and their sources

Proxy	Description ( <i>resolution</i> )	Sources
<b>F10.7</b>	The 10.7 cm solar radio flux ( <i>daily or monthly</i> )	NOAA National Weather Service Climate Prediction Center: <a href="ftp://ftp.ngdc.noaa.gov/STP/space-weather/solar-data/solar-features/solar-radio/noontime-flux/penticton/penticton_adjusted/listings/listing_drao_noontime-flux-adjusted_daily.txt">ftp://ftp.ngdc.noaa.gov/STP/space-weather/solar-data/solar-features/solar-radio/noontime-flux/penticton/penticton_adjusted/listings/listing_drao_noontime-flux-adjusted_daily.txt</a> or <a href="ftp://ftp.ngdc.noaa.gov/STP/space-weather/solar-data/solar-features/solar-radio/noontime-flux/penticton/penticton_adjusted/listings/listing_drao_noontime-flux-adjusted_monthly.txt">ftp://ftp.ngdc.noaa.gov/STP/space-weather/solar-data/solar-features/solar-radio/noontime-flux/penticton/penticton_adjusted/listings/listing_drao_noontime-flux-adjusted_monthly.txt</a>
<b>QBO<sup>10</sup></b> <b>QBO<sup>30</sup></b>	Quasi-Biennial Oscillation index at 10hPa and 30hPa ( <i>monthly</i> )	Free University of Berlin: <a href="http://www.geo.fu-berlin.de/en/met/ag/strat/produkte/qbo/">www.geo.fu-berlin.de/en/met/ag/strat/produkte/qbo/</a>
<b>ENSO</b>	El Niño /Southern Oscillation - Nino 3.4 Index ( <i>3-monthly averages</i> )	NOAA National Weather Service Climate Prediction Center: <a href="http://www.cpc.noaa.gov/data/indices/">http://www.cpc.noaa.gov/data/indices/</a>
<b>NAO</b>	North Atlantic Oscillation index ( <i>daily or monthly</i> )	<a href="ftp://ftp.cpc.ncep.noaa.gov/cwlinks/norm.daily.nao.index.b500101.current.ascii">ftp://ftp.cpc.ncep.noaa.gov/cwlinks/norm.daily.nao.index.b500101.current.ascii</a> or <a href="http://www.cpc.ncep.noaa.gov/products/precip/CWlink/pna/norm.nao.monthly.b5001.current.ascii">http://www.cpc.ncep.noaa.gov/products/precip/CWlink/pna/norm.nao.monthly.b5001.current.ascii</a>
<b>AAO</b>	Antarctic Oscillation index ( <i>daily or monthly</i> )	<a href="ftp://ftp.cpc.ncep.noaa.gov/cwlinks/norm.daily.aao.index.b790101.current.ascii">ftp://ftp.cpc.ncep.noaa.gov/cwlinks/norm.daily.aao.index.b790101.current.ascii</a> or <a href="http://www.cpc.ncep.noaa.gov/products/precip/CWlink/daily_aao_index/aao/monthly.aao.index.b79.current.ascii">http://www.cpc.ncep.noaa.gov/products/precip/CWlink/daily_ao_index/aao/monthly.aao.index.b79.current.ascii</a>

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1175 **Table 2** Ozone trends and associated uncertainties (95% confidence limits; accounting for the  
1176 autocorrelation in the noise residuals), given in DU/year, for 20-degree latitude bands, based on  
1177 daily (top values) and monthly (bottom values) medians over 6 years of IASI observations. Bold  
1178 (underlined) values refer to significant (positive) trends. Values marked with a star (\*) refer to  
1179 trends which are rejected by the iterative backward elimination procedure<sup>†</sup>.

<i>DU/yr</i>	# Days	Ground-300hPa (MLT)	300-150hPa (UTLS)	150-25hPa (MLST)	25-3hPa (UST)	Total columns
<b>70°N-90°N (Feb-Oct)</b>	1493	<b>-0.13±0.10</b>	<b><u>1.28±0.82</u></b>	<b><u>2.81±2.27</u></b>	-0.16±0.97*	<b><u>3.90±2.93</u></b>
		-0.03±0.29*	0.70±0.92	-0.04±2.60	-1.81±2.81*	1.37±3.62*
<b>50°N-70°N</b>	2103	-0.08±0.09	<b><u>0.73±0.51</u></b>	0.97±1.30	<b><u>0.55±0.36</u></b>	<b><u>1.93±1.71</u></b>
		0.17±0.35*	1.24±1.24	2.28±4.24*	0.66±0.76	4.72±5.58
<b>30°N-50°N</b>	2105	<b>-0.19±0.05</b>	<b><u>0.34±0.18</u></b>	-0.34±0.77	<b><u>0.89±0.41</u></b>	0.91±1.24
		<b>-0.15±0.13</b>	<u>0.75±0.75</u>	-0.37±1.65*	<b><u>0.87±0.52</u></b>	0.33±2.25*
<b>10°N-30°N</b>	2105	0.10±0.11	-0.03±0.10*	<b>-0.73±0.29</b>	<b><u>0.95±0.65</u></b>	0.21±0.30*
		0.12±0.15*	0.05±0.12*	-0.55±0.62*	<b><u>1.25±0.74</u></b>	0.82±1.01
<b>10°S-10°N</b>	2104	<b>-0.41±0.12</b>	<b>-0.25±0.07</b>	-0.11±0.26*	<b><u>0.44±0.19</u></b>	-0.16±0.34
		<b>-0.25±0.14</b>	-0.08±0.10	-0.11±0.64*	0.61±0.64	0.13±0.83*
<b>30°S-10°S</b>	2106	<b>-0.22±0.10</b>	<b>-0.08±0.04</b>	<b>-0.61±0.26</b>	<b><u>0.89±0.58</u></b>	-0.04±0.31*
		<b>-0.15±0.13</b>	<b>-0.09±0.07</b>	<b>-0.45±0.36</b>	0.80±1.23	-0.01±1.26*
<b>50°S-30°S</b>	2105	<b>-0.19±0.07</b>	<b>-0.22±0.08</b>	<b>-2.17±0.58</b>	<b><u>1.74±0.77</u></b>	-0.79±0.96
		<b>-0.18±0.09</b>	<b>-0.27±0.12</b>	<b>-2.36±1.80</b>	1.21±1.30	-0.64±1.45*
<b>70°S-50°S</b>	2105	<b>-0.13±0.05</b>	0.09±0.16	0.56±0.82	<b><u>0.54±0.29</u></b>	1.15±1.28
		-0.22±0.12	0.05±0.32*	0.02±1.15*	0.57±0.82	0.51±1.75*
<b>90°S-70°S (Oct-Apr)</b>	738	-0.15±0.21*	0.01±0.61*	0.00±2.36*	<b><u>1.04±0.57</u></b>	1.50±3.15*
		-0.17±0.40*	0.25±0.73*	2.59±3.80*	0.91±2.10	3.28±5.12*

1180 † The trend values result from the adjustment of the regression model where the linear term is  
1181 kept whatever its p-value calculated during the iterative process.

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1189 **Table 3** Same as Table 2 but for seasonal O<sub>3</sub> trends and associated uncertainties based on daily  
 1190 medians during JJA (top values) and DJF (bottom values) periods. Values marked with a star (\*)  
 1191 refer to trends which are rejected by the iterative backward elimination procedure<sup>†</sup>.

<i>DU/yr</i>	# Days	Ground-300hPa (MLT)	300-150hPa (UTLS)	150-25hPa (MLST)	25-3hPa (UST)	Total columns
<b>70°N-90°N</b> <b>(Feb-Oct)</b>	613 48	<b>-0.18±0.08</b> -	<b><u>1.13±0.65</u></b> -	-0.91±1.52 -	<b><u>1.72±0.51</u></b> -	<b><u>1.36±1.15</u></b> -
<b>50°N-70°N</b>	551 527	<b>-0.23±0.07</b> -0.09±0.12*	<b><u>1.03±0.37</u></b> <b><u>1.74±1.30</u></b>	0.62±1.64 0.73±1.73*	<b><u>1.67±0.48</u></b> -0.66±0.79	<b><u>3.01±1.64</u></b> 1.56±2.66*
<b>30°N-50°N</b>	551 529	<b>-0.30±0.10</b> <b>-0.24±0.09</b>	<b><u>0.42±0.30</u></b> 0.28±0.28	-0.30±0.65* -0.82±0.90	<b><u>0.84±0.25</u></b> <b><u>0.62±0.49</u></b>	1.17±1.35 -0.81±1.05
<b>10°N-30°N</b>	551 529	-0.05±0.16* <b><u>0.18±0.14</u></b>	<b><u>0.17±0.05</u></b> 0.01±0.09*	<b>-0.34±0.30</b> <b>-1.05±0.45</b>	<b><u>0.36±0.27</u></b> 0.49±0.54	-0.09±0.54* -1.14±0.44
<b>10°S-10°N</b>	551 529	-0.06±0.10 <b>-0.70±0.23</b>	0.04±0.05* <b>-0.32±0.10</b>	-0.84±0.86 1.64±1.77	0.32±0.42 0.53±0.59	-0.56±0.74* 0.34±0.93*
<b>30°S-10°S</b>	551 530	<b>-0.26±0.09</b> <b>-0.15±0.11</b>	-0.06±0.07 0.06±0.12*	<b>-0.56±0.40</b> -0.12±0.31*	<b><u>1.06±0.55</u></b> <b><u>1.48±0.53</u></b>	0.24±0.43 <b><u>1.56±0.92</u></b>
<b>50°S-30°S</b>	551 529	<b>-0.21±0.05</b> <b>-0.10±0.06</b>	<b>-0.16±0.09</b> <b>-0.14±0.06</b>	-0.52±0.54 <b>-2.83±0.64</b>	0.49±0.59 <b><u>3.40±0.85</u></b>	-0.44±0.83 0.47±0.52
<b>70°S-50°S</b>	551 529	<b>-0.25±0.06</b> <b>-0.10±0.04</b>	<b>1.03±0.60</b> 0.19±0.24*	<b><u>2.63±1.65</u></b> <b><u>0.52±0.48</u></b>	<b><u>0.98±0.62</u></b> <b><u>1.66±0.70</u></b>	<b><u>3.44±2.47</u></b> <b><u>1.72±0.74</u></b>
<b>90°S-70°S</b> <b>(Oct-Apr)</b>	- 523	- <b>-0.21±0.20</b>	- -0.46±0.80*	- 0.16±2.53*	- <b><u>1.18±0.67</u></b>	- 0.98±3.27*

1192 † The trend values result from the adjustment of the regression model where the linear term is  
 1193 kept whatever its p-value calculated during the iterative process.

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1202 **Table 4** Ozone trends and associated uncertainties (95% confidence limits), given in DU/year  
1203 over NDACC (Network for the Detection of Atmospheric Composition Change) stations in the  
1204 N.H. based on daily medians of IASI (within a grid box of 1°x1° centered on stations, two first  
1205 rows) and FTIR observations (successive rows for different time intervals). Italic values (2<sup>d</sup> row)  
1206 refer to trends inferred from subsampled IASI data and bold values refer to statistically  
1207 significant trends. Values marked with a star (\*) refer to trends which are rejected by the iterative  
1208 backward elimination procedure<sup>†</sup>.

<i>DU/yr</i>	<b>Data periods</b>	<b># days</b>	<b>25-3hPa (US)</b>	<b>Total columns</b>
Ny-Alesund (79°N) Mar-Sept	2008-2013	1239	0.56±0.73	<b>5.26±4.72</b>
	<i>Subsamp.</i> 2008-2012	82	-0.29±4.58	6.26±18.11
	2008-2012	84	-3.58±4.58	2.24±20.78*
	2003-2012	168	-0.17±0.70*	<b>-4.84±3.01</b>
	2000-2012	288	<b>0.64±0.60</b>	-1.02±2.40*
	1999-2012	320	<b>0.62±0.55</b>	<b>-2.35±1.40</b>
	1995-2012	383	<b>1.03±0.66</b>	1.31±2.39*
	1995-2003	167	<b>1.25±1.05</b>	3.33±3.41
Thule (77°N) Mar-Sept	2008-2013	1094	<b>1.24±1.09</b>	<b>4.97±4.72</b>
	<i>Subsamp.</i> 2008-2012	231	1.31±2.69	0.10±7.36
	2008-2012	340	-2.10±2.89	0.39±11.59*
	2003-2012	697	0.86±0.89	-2.77±2.99
	2000-2012	776	<b>1.33±0.86</b>	-1.29±1.73
	1999-2012	779	<b>1.69±0.88</b>	-1.25±1.74
	1999-2003	138	<b>3.73±2.90</b>	4.86±10.13*
	Kiruna (68°N) Mar-Sept	2008-2013	1236	0.21±1.42
<i>Subsamp.</i> 2008-2012		226	0.97±4.05	3.78±6.03
2008-2012		254	-1.97±6.04*	-3.75±6.64*
2003-2012		678	0.15±0.67*	2.26±3.68
2000-2012		913	<b>1.60±1.29</b>	3.69±4.20
1999-2012		984	<b>1.10±0.98</b>	-0.43±1.64*
1996-2012		1183	<b>1.11±0.54</b>	<b>1.82±1.77</b>
1996-2003		596	<b>1.26±1.21</b>	1.12±3.77*
Jungfraujoch (47°N)	2008-2013	1580	<b>2.95±0.61</b>	<b>5.64±3.15</b>
	<i>Subsamp.</i> 2008-2012	524	<b>3.72±1.14</b>	<b>5.61±5.11</b>
	2008-2012	565	1.60±1.80	<b>5.28±4.82</b>
	1998-2012	1582	0.10±0.35	-0.28±0.86*
	1995-2012	1771	0.02±0.33*	<b>0.85±0.79</b>

Zugspitze (47°N)	2008-2013	1729	<b>3.17±0.56</b>	<b>5.53±2.92</b>
	<i>Subsamp.</i>			
	2008-2012	538	<b>3.56±1.63</b>	<b>5.99±4.49</b>
	2008-2012	597	0.71±1.22	3.46±3.79
	1998-2012	1472	0.08±0.32*	0.81±0.98
	1995-2012	1525	0.23±0.32	<b>1.36±1.01</b>
Izana (28°N)	2008-2013	1803	0.56±0.65	<b>1.28±0.77</b>
	<i>Subsamp.</i>			
	2008-2012	380	0.32±1.28	0.11±1.95
	2008-2012	443	0.24±0.80*	0.91±2.44*
	1999-2012	1257	<b>0.46±0.25</b>	0.20±0.33*

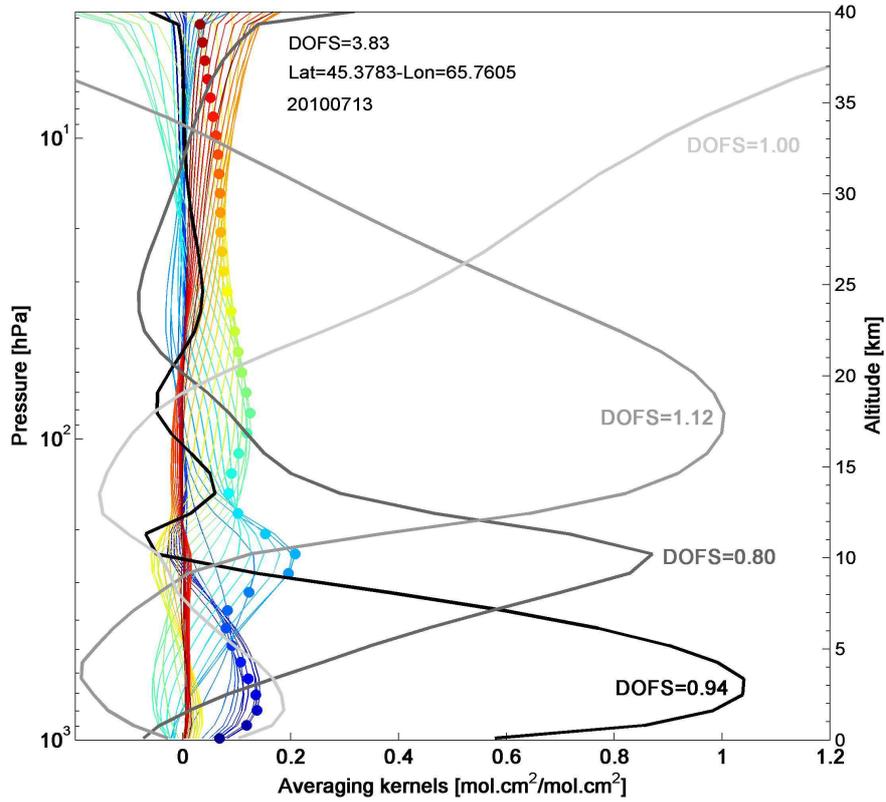
1209 † The trend values result from the adjustment of the regression model where the linear term is  
1210 kept whatever its p-value calculated during the iterative process.

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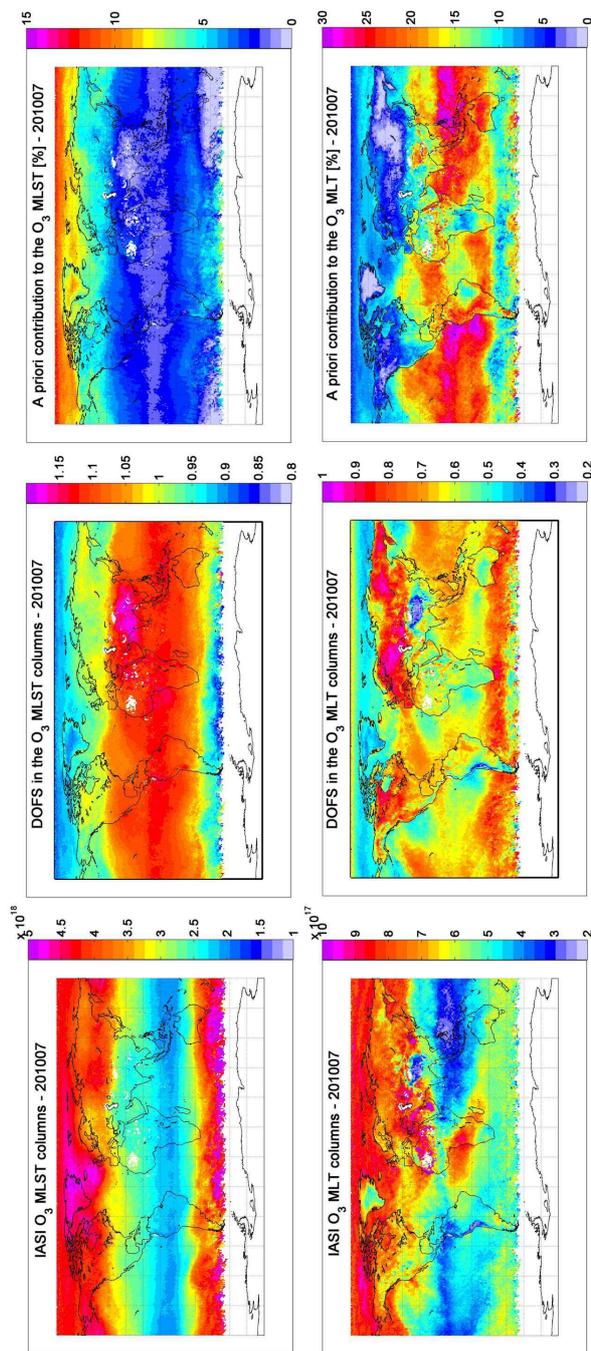
1233 **Figure captions**



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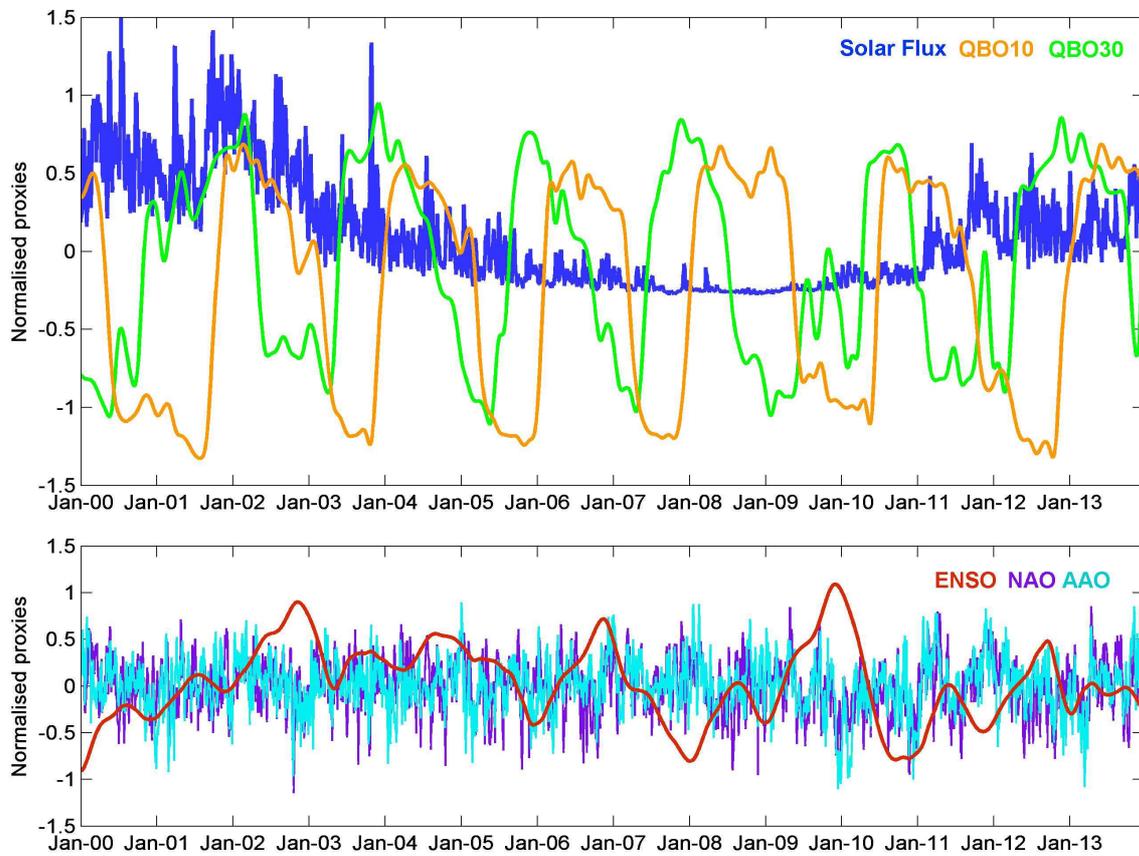
1235 **Figure 1.** Typical IASI FORLI-O<sub>3</sub> averaging kernels, in partial column units, corresponding to  
1236 one mid-latitude observation in July (45°N/66°E) for each 1 km retrieved layers from ground to  
1237 40 km altitude (color scale) and for 4 merged layers: ground-300 hPa; 300-150 hPa; 150-25 hPa;  
1238 25-3 hPa (grey lines). The total DOFS and the DOFS for each merged layers are also indicated.

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1241 **Figure 2.** Distributions of (a) O<sub>3</sub> columns, (b) DOFS and (c) *a priori* contribution (given as a %)  
 1242 in the ground-300hPa (MLT) and 150-25hPa (MLST) layers for IASI O<sub>3</sub>, averaged over July  
 1243 2010 daytime data. Note that the scales are different.



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1245 **Figure 3.** Normalized proxies as a function of time for the period 2000-2013 for the solar F10.7  
 1246 cm radio flux (blue) and the equatorial winds at 10 (green) and 30 hPa (orange), respectively (top  
 1247 panel), and for the El Niño (red), north Atlantic oscillation (purple) and Antarctic oscillation  
 1248 (light blue) indexes (bottom panel).

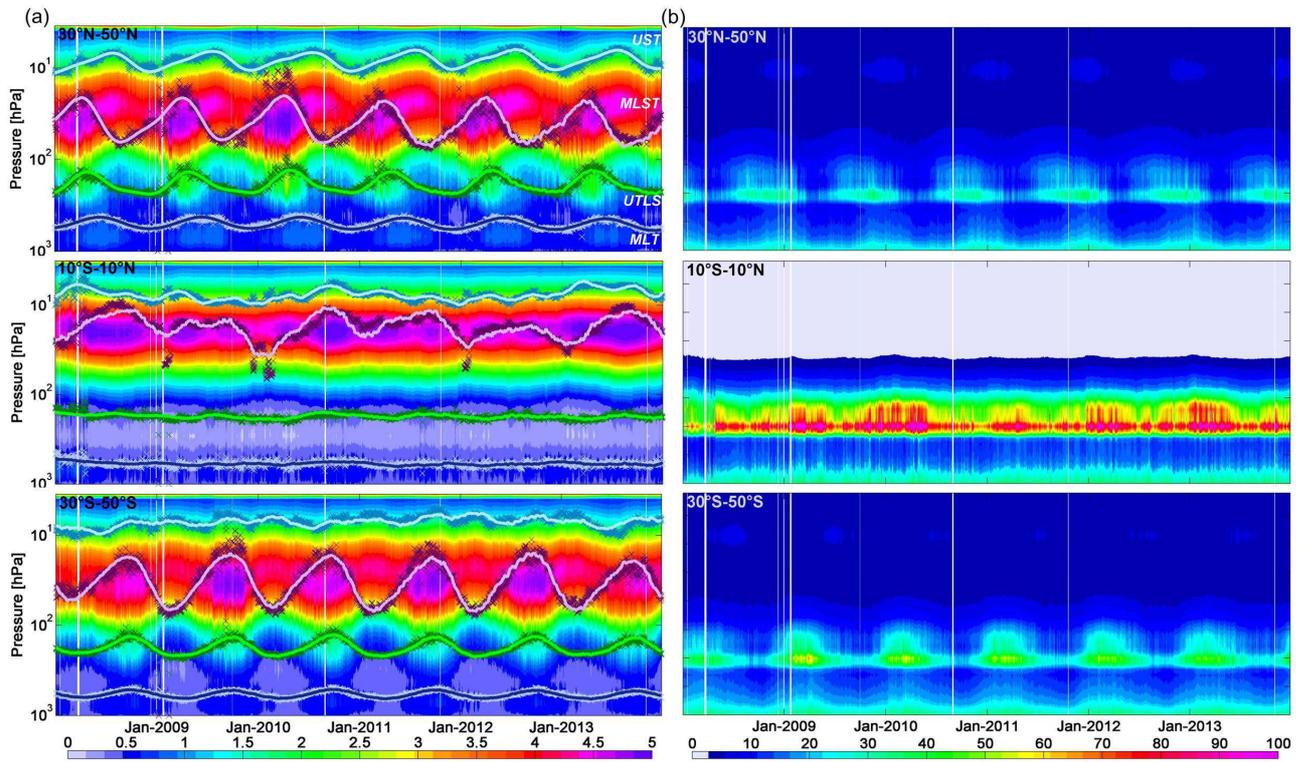
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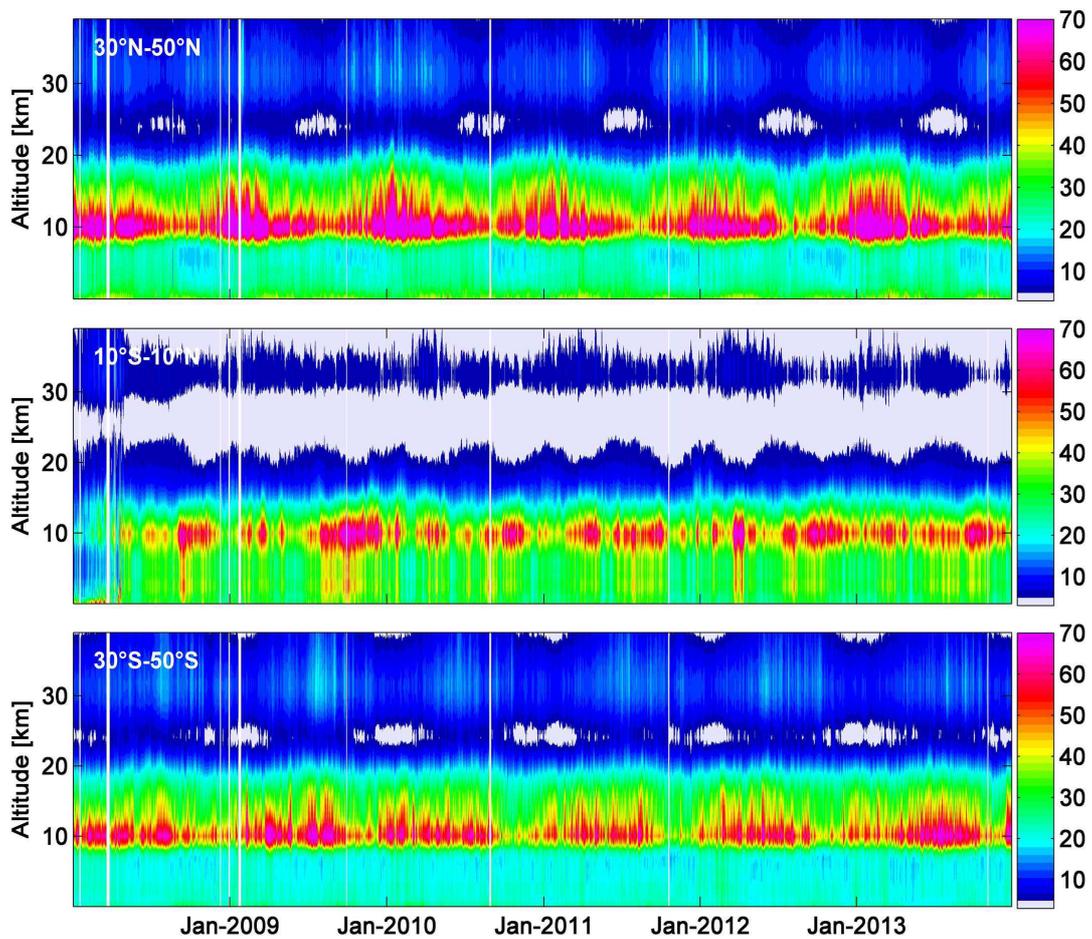
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1255 **Figure 4.** (a) Daily IASI O<sub>3</sub> profiles ( $1 \times 10^{12}$  molecules/cm<sup>3</sup>) for the period 2008-2013 and over  
 1256 the range of the retrieved profiles as a function of time and altitude, in three latitude bands:  
 1257 30°N-50°N (top), 10°S-10°N (middle), 30°S-50°S (bottom). Superimposed daily IASI O<sub>3</sub> partial  
 1258 columns (scatters) and the associated fits (solid lines) from the multivariate regressions for the  
 1259 MLT (ground-300hPa), UTLS (300-150hPa), MLST (150-25hPa) and UST (above 25hPa)  
 1260 layers. The IASI measurements and the fits have been scaled for clarity. (b) Estimated total  
 1261 retrieval errors (%) associated with daily IASI O<sub>3</sub> profiles.

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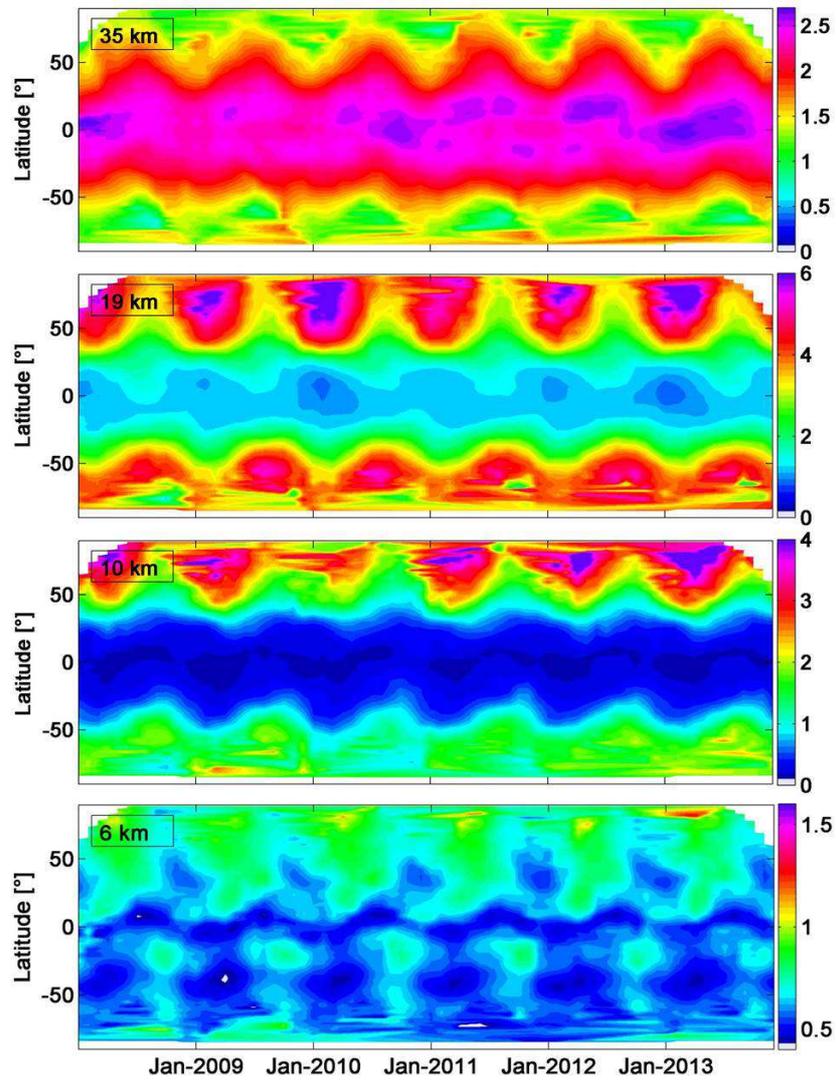
1266 **Figure 5.** Daily IASI O<sub>3</sub> variability (%), expressed as  $[\sigma(O_3(t))/O_3(t)]100\%$ , where  $\sigma$  is the  
 1267 standard deviation, as a function of time and altitude in three latitude bands: 30°N-50°N (top),  
 1268 10°S-10°N (middle), 30°S-50°S (bottom).

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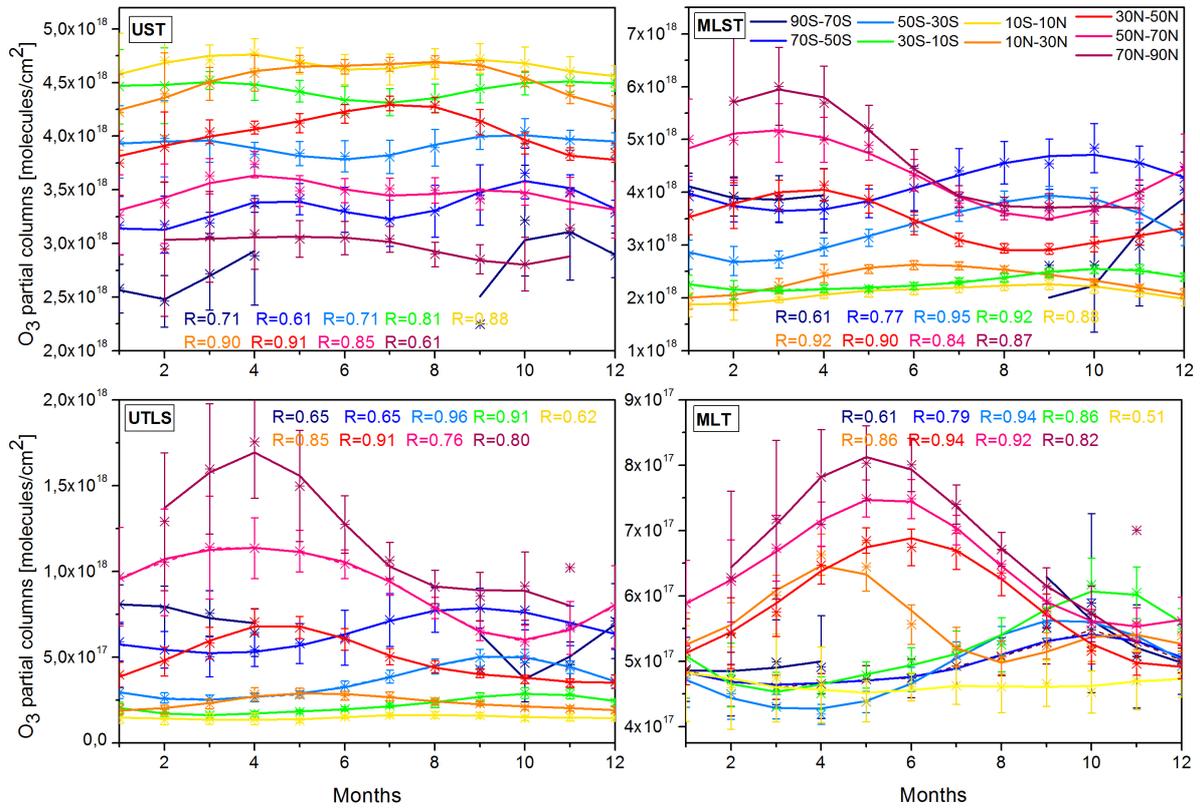
1274 **Figure 6.** Daily IASI O<sub>3</sub> number density ( $1 \times 10^{12}$  molecules/cm<sup>3</sup>) at 35 km (top row), 19 km  
 1275 (second row), 10 km (third row) and 6 km (bottom row) as a function of time and latitude. Note  
 1276 that the color scales are different.

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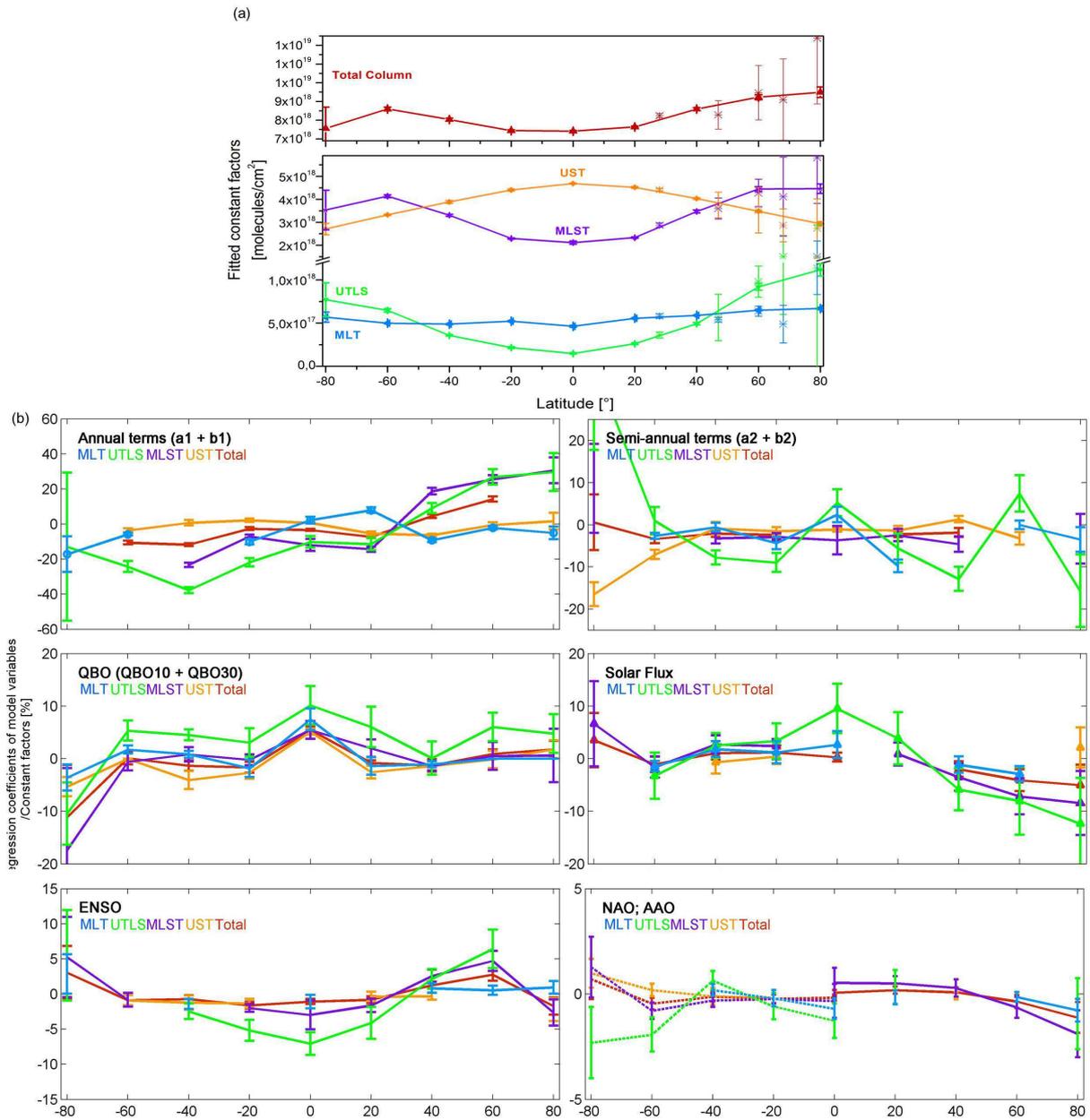
1282 **Figure 7.** Monthly medians of measured (scatters) and of fitted (line) IASI O<sub>3</sub> columns averaged  
 1283 over the period 2008-2013, for the UST, MLST, UTLS and MLT layers and for each 20-degrees  
 1284 latitude bands (color scale in the top-right panel). The fit is based on daily medians. Error bars  
 1285 give the 1 $\sigma$  standard deviation relative to the monthly median values. Correlation coefficient (R)  
 1286 between the daily median observations and the fit are also indicated. Note that the scales are  
 1287 different.

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**Figure 8.** (a) Fitted constant factors (Cst, see Eq.1, Section 3) from the 6-years IASI daily O<sub>3</sub> time series for the 20-degree latitude belts, separately given for the 4 layers and for the total column. The stars correspond to the constant factors fitted above ground-based measurement stations: Ny-Ålesund (79°N), Kiruna (68°N), Harestua (60°N), Jungfraujoch (47°N), Izana (28°N). (b) Regression coefficients of the variables retained by the stepwise procedure, given in % as [(regression\_coefficients)]/fitted\_Cst]x100%. Identification for the variables: Annual (top

1299 left) and Semi-Annual variations (top right) terms, QBO at 10 and 30 hPa (bottom left), solar  
1300 flux (bottom right). Note that the scales are different. The associated fitting uncertainties (95%  
1301 confidence limits) are also represented (error bars).

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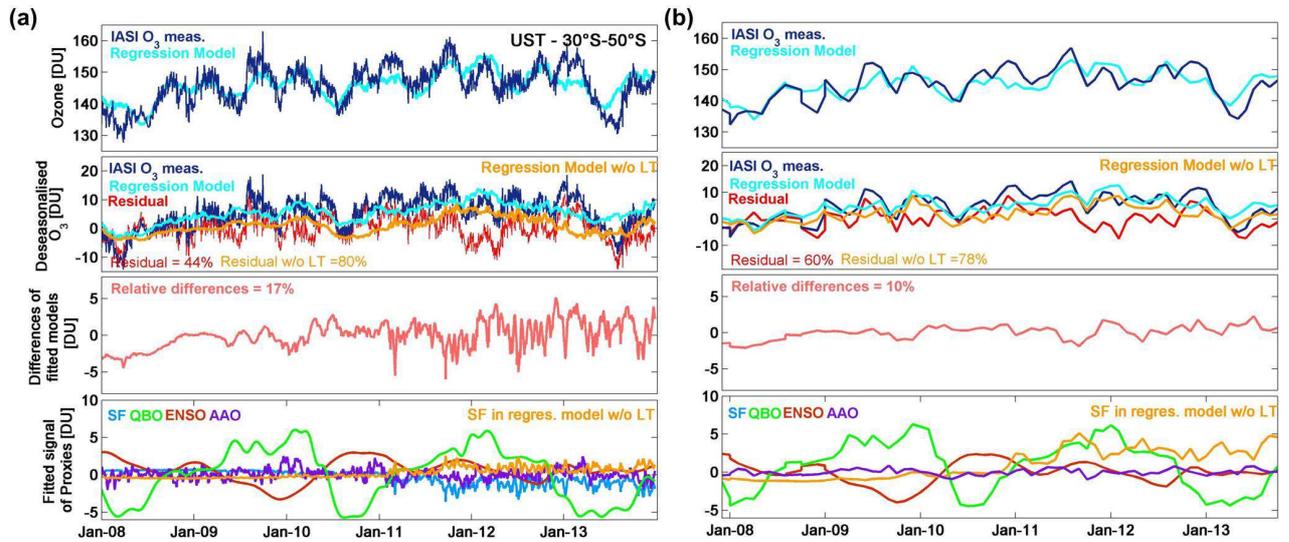
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1330 **Figure 9.** Daily (a) and monthly (b) time series of O<sub>3</sub> measurements and of the fitted regression  
 1331 model in the UST for the 30°S-50°S latitude band (top row), of the deseasonalised O<sub>3</sub> (2<sup>d</sup> row),  
 1332 of the difference of the fitted models with and without the linear term (3<sup>d</sup> row), and of the fitted  
 1333 signal of proxies ([regression coefficients\*Proxy]): SF (blue), QBO (QBO<sup>10</sup> + QBO<sup>30</sup>; green),  
 1334 ENSO (red) and AAO (purple) (bottom) (given in DU). The averaged residuals relative to the  
 1335 deseasonalised IASI time series are also indicated (%).

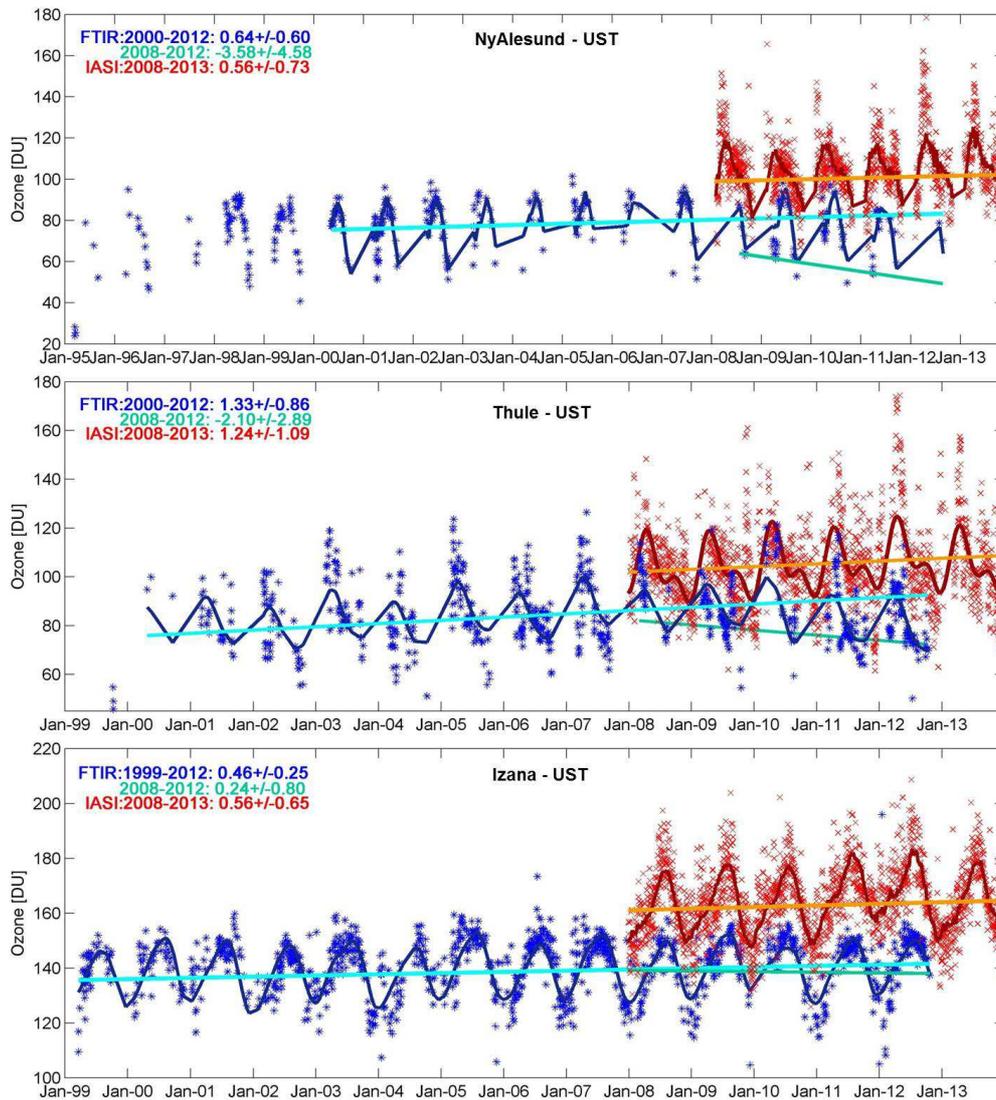
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 1342 **Figure 10.** Daily time series of O<sub>3</sub> FTIR (blue symbols) and IASI (red symbols) measurements in  
 1343 the UST at Ny-Alesund (top), Thule (middle) and Izana (bottom), covering the 1995-2012 and  
 1344 the 1999-2012 periods, respectively (given in DU). The fitted regression models (dark blue and  
 1345 dark red lines, for FTIR and IASI, respectively) and the linear trends calculated for periods  
 1346 starting after the turnaround over 1999/2000-2012 and over 2008-2012 for FTIR (light blue and  
 1347 green lines), and the 2008-2013 period for IASI (orange line) are also represented (DU/yr). The  
 1348 trend values given in DU/year are indicated.  
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