Determination and climatology of the planetary boundary layer height by in-situ and remote sensing methods as well as the COSMO model above the Swiss plateau

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Abstract

The planetary boundary layer (PBL) height is a key parameter in air quality control and pollutant dispersion. The PBL height can however not be directly measured and its estimation relies on the analysis of the vertical profiles of the temperature, the turbulence or the atmospheric composition. An operational PBL height detection including several remote sensing instruments (windprofiler, Raman lidar, microwave radiometer) and several algorithms (Parcel and bulk Richardson number methods, surface-based temperature inversion, aerosol or humidity gradient analysis) were developed and the first year of application allowed validating these various detection methods against radio sounding measurements. The microwave radiometer provides convective boundary layer heights in good agreement with the radio sounding (median bias < 25 m, $R^2 > 0.70$) and allows to fully analyzing the PBL height diurnal cycle due to its smaller time granularity. The Raman lidar also leads to good results whereas the windprofiler yields some more dispersed results. Comparisons with the numerical weather prediction model COSMO-2 were also established and point out a general overestimation by the model. Finally the seasonal cycles of the daytime and nighttime PBL heights are discussed for each instrument and each detection algorithm for two stations on the Swiss plateau.

1 Introduction

The height of the planetary boundary layer (PBL), also called atmospheric boundary layer is a key parameter for air quality analysis, pollutants dispersion and quantification of pollutant emissions and sources. It controls the interactions of the atmosphere with the oceans and the land and determines the air volume available for the dispersion of all atmospheric constituents (including anthropogenic pollution and water vapor) emitted at the earth surface, and hence contributes to the assessment of the pollutant concentration near the surface. The PBL height is therefore a key parameter of all air pollution
models that is however not directly measured but has to be estimated by upper-air instruments.

The Cost Action 710 defined the daytime PBL height as “the height of the layer adjacent to the ground over which pollutants or any constituents emitted within this layer or entrained into it become vertically dispersed by convection or mechanical turbulence within a time scale of about an hour” (Cost Action 710 – Final report, 1998). The PBL height can consequently be estimated by the measurement of mechanical turbulence, of the temperature enabling convection or of the concentration of any PBL constituent. These detection methods are based on various atmospheric parameters, various measuring instruments and different analyzing algorithms, leading to several PBL height estimations that are not always straightforward comparable. The intense development of remote sensing instruments offers nowadays a wide field of vertical profiles up to several kilometers allowing PBL height detection with high temporal resolution.

The PBL presents usually a marked diurnal cycle that depends on the synoptic and local weather conditions. In case of fair-weather days, the PBL height has a well-defined structure and diurnal cycle (Fig. 1) leading to the development of a Convective Boundary Layer (CBL), also called mixing layer, during the day and of a Stable Boundary Layer (SBL) surmounted by a Residual Layer (RL) during the night (Stull, 1988). In case of cloudy or rainy conditions as well as in case of advective weather situations, free convections are no more primarily driven by solar heating, but by ground thermal inertia, cold air advection, forced mechanical convection or cloud top radiative cooling. In those overcast cases the CBL development remains weaker than in case of clear sky conditions, with slower growth and lower height maxima. The boundary layer is said to be neutral if the buoyancy is near zero; these neutral cases are found for overcast conditions with strong winds but little temperature difference between the air and the ground. Neutral conditions are frequently met in the RL but rarely near the ground. For clarity purpose, the PBL development under clear sky conditions (i.e. more than 50% of solar radiation during the CBL development) that leads to strong convections driven
principally by solar heating will be called CBL in this paper. For all “no clear-sky” cases with partial or total cloud coverage but without precipitation, it will be called cloudy-CBL.

While the definition and the measurement of the CBL, the neutral boundary layer (NBL) and the cloudy-CBL are well assessed, the nocturnal SBL presents a more complicated internal structure. It comprises a stable layer caused by radiative cooling of the ground, which gradually merges into a neutral layer called RL (Stull, 1988; Salmond and McKendry, 2005; Mahrt et al., 1998). The stable layer can be characterized by the surface-based temperature inversion (SBI), and its top can be estimated by the height at which the gradient of the potential temperature ($\theta$) vanishes. Local and short-term turbulence can occur within this stable layer. The RL height is the top of the neutral layer and the begin of the stable free troposphere. The pollutants emitted during the night are trapped into the SBI whereas the pollutants of the previous days tend to stay in the RL.

Contrary to RS launched usually only twice a day, the continuous remote sensing measurements allow to fully analyze the diurnal cycles of the different layers constituting the PBL. The use of remote sensing instrumentation to detect the PBL height was recently overviewed by Emeis (2009). Recent studies compared several detection methods or retrieval techniques (Bianco and Wilczac, 2002; Seidel et al., 2010; Beyrich and Leps, 2012; Haefelin et al., 2012; Summa et al., 2013), remote sensing with RS measurements (Baars et al., 2008; Liu and Liang, 2010; Granados-Muñoz et al., 2012; Milroy et al., 2012; Sawyer and Li, 2013; Cimini et al., 2013) and/or several remote sensing instruments (Wang et al., 2012; Zahng et al., 2012). In most of the cases, good correlations are found in case of convective weather situations with differences of 100–300 m between the various instruments and/or methods. The differences become greater for less well-defined weather situations and for methods/instruments detecting various types of PBL height, even if these studies are scarce. If temperature profiles are measured, bulk Richardson number (bR) or Parcel (PM) methods are usually considered as the most relevant methods for daytime PBL height detection regarding the reliability and the uncertainties. Some studies also compared measurements with
models predictions (Baars et al., 2008; Seidel et al., 2012; Ketterer et al., 2013), the results depending on both the model and the measurement types.

Some of these studies were done on long enough time series (between 1 and 25 years) to analyze the PBL height climatology at some stations in Europe and US (Baars et al., 2008; Schmid and Niyogy, 2012; Beyrich and Leps, 2012; Granados-Muñoz et al., 2012; Sawyer and Li, 2013) or over continents (Seidel et al., 2010, 2012). For continental stations, a clear CBL seasonal cycle is usually found with summer maxima reaching 1000 to 2000 m above ground level (a.g.l.) and winter minima between 500 and 1200 m a.g.l. The seasonal cycle of the nocturnal SBL was to our knowledge only addressed on the basis of temperature (T) profiles from RS measurements (Seidel et al., 2010; Beyrich and Leps, 2012; Seidel et al., 2012). Both authors found a minimum in summer and a maximum in winter explained by greater wind speeds and consequently stronger mechanical turbulences during winter. Few of the PBL height detections run operationally, meaning that they run as fully automatic processes delivering continuous PBL estimation in real time. Some authors specify that a visual inspection is necessary to increase the results reliability.

In this study an operational system for PBL height detection has been developed based on the analysis of vertical atmospheric profiles of T, wind turbulence and atmospheric constituents, which are measured by different remote sensing techniques like radar windprofiler (WP), microwave radiometer (MWR) and lidar as well as by radiosounding (RS). One year (2012) of measurements was used to compare all these PBL height determination methods against a reference method being the PM applied either on RS or on MWR measurements. The PBL heights computed by the COSMO-2 model (numerical weather prediction model of the Consortium for Small Scale Modelling; see www.cosmo-model.org) were also compared to the instrumental PBL height determination. A two-year climatology of the CBL, the cloudy-CBL and the different layers constituting the nocturnal SBL was moreover computed for Payerne (PAY) and Schaffhausen (SHA) situated both on the Swiss plateau.
The paper comprises the description of the instruments and the PBL detection methods, some examples of PBL operational estimations, the inter-comparison and validation of the experimental methods, a comparison with COSMO-2 model and a two-years climatology. Recommendations about the most comprehensive set of instruments for an operational detection of the PBL diurnal cycle are given in the conclusion. Abbreviations for the sites, instruments and methods as well as for the different PBL layers are summarized in Table 1. Along the paper the various PBL detections are named by the measuring instrument and the applied method, RS/PM being for example given for Parcel method (PM) applied on radio-sounding (RS) measurements. If not specified, elevations or heights are given as above ground level (a.g.l.) and time as LT (= UTC + 1).

2 Experimental

2.1 Site and instrumentation

For this study, a two-year (2012–2013) dataset from the two upper-air remote sensing sites Payerne (491 m a.s.l., 46.799° N, 6.932° E) and Schaffhausen (437 m a.s.l., 47.672° N, 8.604° E) of the CN-Met (Centrales Nucléaires et METéorologie) measurement network (Calpini et al., 2011) were used. Both sites are located on the Swiss plateau in rural areas in proximity of a small city. Both stations include a WP, a MWR and a SwissMetNet surface station. A ceilometer and a Raman lidar were moreover available at PAY, as well as RS measurements twice a day.

The WPs are Degreane PCL1300 (Degreane Horizon, 2006) with five antennas operating at 1290 MHz (λ = 23.3 cm) in two modes. For this study only the low mode has been used with a vertical resolution of 150 m and the first level at 105 m. The windprofiler measures the clear air radar echo generated by inhomogeneities in the refractive index due to atmospheric turbulent structures that are assumed to travel with the background wind. The vertical profile of horizontal and vertical wind is derived from the radial velocities along each direction. The effective time resolution is 40 min for one
vertical profile; profiles are generated every 10 min using a gliding average. The signal to noise ratio (SNR) is calculated from the five SNR values corresponding to the five antenna directions by taking the minimum SNR. This procedure minimized high SNR values generated by hard, non-atmospheric targets. Both wind and SNR data undergo a quality check and their availability depends on the atmospheric conditions.

The MWR is a passive remote sensing instrument that measures electromagnetic radiation emitted from the atmosphere in the microwave band. From the measured radiation spectrum, the atmospheric $T$ profile between 0 and 5 km is retrieved. The used MWRs are TEMPRO from Radiometer Physics GMbH (RPG) (Radiometer Physics GMbH, 2011) with 7 channels between 51 and 58 GHz for $T$ profiling. The radiometer alternates between elevation scanning (6 elevations between 5 and 90°) and zenith observations. The $T$ profile is combined from the profiles retrieved from the elevation scan and the zenith observations. The vertical resolution decreases with altitude, from 50 m for $z < 1200$ m to 200 m at 3000 m according to the manufacturer. The time resolution is set to one profile every 10 min. The elevated $T$ inversions were not measured by the MWR.

The PAY aerological station is equipped with a fully automated and operational Raman lidar (Dinoev et al., 2013). The laser source is a Nd:YAG-laser emitting UV pulses (300 mJ per pulse, 30 Hz repetition rate) at $\lambda = 355$ nm. The receiver consists of four telescopes with 0.3 m diameter each, which are fiber coupled to the water vapor and the aerosol polychromator isolating the nitrogen (387 nm) and the water vapor (407 nm) signals as well as two portions of the pure rotational Raman spectrum and the elastic signals. The aerosol scattering ratio is derived from the sum of the rotational Raman signals and the elastic signal (Dinoev et al., 2010). The maximum range varies from 4000 m during the day up to 8000 m during the night for the water vapor measurements and from 700 m (day) to 12 000 m (night) for aerosol backscatter ratio measurements. The first range level is located at 110 m. The vertical resolution is dynamically adapted to the measurement conditions, varying from 30 m to a maximum of 300 m. However, the signal to noise ratio is generally very high in the boundary layer and the vertical
resolution remains constant (30 m). The effective time resolution of profiles is 30 min. No measurements are possible during precipitation and in presence of low clouds, i.e. the lidar powers down if the clouds are below 900 m and powers up as soon as the cloud base rises above 2000 m.

A ceilometer (CBME80 from Eliasson) measuring at $\lambda = 905$ nm with a time resolution of a few seconds is interfaced to the lidar system to provide independent cloud information. This model was not configured to record backscatter profiles but only to provide the height of the cloud bases detected by a strong gradient in the backscattered signal.

In addition to the remote sensing instruments, the station of Payerne performs RS providing pressure ($p$), $T$, humidity and wind speed and direction profiles up to 35 km. Meteolabor SRS 400 C34 radiosondes are launched twice a day at 00:00 and 12:00 LT. The horizontal displacement of the sonde can reach up to 200 km. However, only the first vertical 3500 m corresponding to approximately 12 min of rise are used to determine the PBL height allowing neglecting the RS horizontal displacement. RS has a constant height resolution of 5–6 m corresponding to a one second time resolution.

The SwissMetNet meteorological surface network provides surface $T$, humidity, $p$, wind direction and speed as well as sunshine duration and precipitations every 10 min. The wind components are measured at 10 m and all the other parameters at 2 m. In addition, the PAY station is equipped with a sonic anemometer on a 10 m mast measuring several parameters related to turbulence, including sensible heat flux that characterizes the thermal energy exchanged and is used to estimate the intensity of the convective forces.

The COSMO-2 model (http://www.cosmo-model.org) was used in assimilation mode. It has a horizontal grid spacing of 2.2 km and a total of 60 vertical levels, of which 15 lie within the first 500 m. The time step is 20 s and data are written out every 1 h. The bulk Richardson number method is used to estimate the boundary layer height in the model (see Sect. 2.2.1).
The cloud cover is detected by Automatic Partial Cloud Amount Detection Algorithm (APCADA) that estimates in real-time the sky cloud cover from surface based measurements of long-wave downward radiation, $T$ and humidity (Dürr and Philipona, 2004). APCADA does not take into account the cirrus clouds.

Measurements of both the MWR and the lidar are necessary to calculate the virtual potential temperature ($\theta_v$), and they are combined with WP data to calculate the bulk Richardson number (see Sect. 2.2.1). These three instruments have however different vertical levels and time constants. For these cases, a vertical scale (35 levels of 100 m between 0 and 3500 m) is set and the mean of the parameters in each level are used. Despite the rather long integration times in the case of the windprofiler and the lidar all observational data have been assumed to be instantaneous. These different time granularities are sometimes visible by a time shift of the CBL growth measured by MWR/PM and WP/SNR or lidar/ASR.

2.2 Methods to determine PBL height

2.2.1 Methods based on $T$ profiles

The Parcel method (Holzworth, 1964; Fisher et al., 1998) defines the PBL height as the elevation to which an air parcel with ambient surface $T$ can rise adiabatically from the ground by convection. As depicted in Fig. 2, the PBL height is set to the elevation $z$ where the $T$ profile crosses the dry adiabatic, or where the potential temperature $\theta$ is equal to the surface $\theta$. The PM needs only the $T$ profile and a precise surface $T$ measurement. To apply the PM, the condition $\theta (z_1) < \theta (z)$, with $z_1 > z$, corresponding to unstable $\theta$ vertical profile, has to be fulfilled. No excess $T$ has been added to the surface $T$. The PM was applied to RS and MWR $\theta$ profiles to detect daytime PBL in case of strong or weak convective conditions (CBL and cloudy-CBL).

The bulk Richardson number ($Ri_b$) is a dimensionless quantity combining the potential energy and the vertical wind shear. It corresponds to the ratio of convective and wind shear produced turbulences and is widely used in turbulence characterization. In order
to be consistent with the $Ri_b$ used in the COSMO-2 model (Szintai, 2010), the following formulation was applied:

$$Ri_b = \frac{gz(\theta(z) - \theta(z_0))}{\bar{\theta}(U^2(z) + V^2(z))}$$  \hspace{1cm} (1)

where $z$ is the height ($z > z_0$), $U$ and $V$ the two horizontal wind velocity components, $g$ the Earth gravitational constant and $\bar{\theta}$ the mean $\theta$ between $z_0$ and $z$. The PBL height corresponds to the first elevation $z$ with $Ri_b$ greater than a critical threshold taken as 0.22 or 0.33 in case of unstable (day) or stable (night) conditions, respectively (Fisher et al., 1998; Jericevic and Grisogono, 2006; Szintai, 2010). It has to be noted that, in most cases, the exact threshold value has only very little impact on the PBL height due to the great $Ri_b$ slope in this interval (see for ex. Figure 2b). PBL height detected by bR is by definition higher than PBL height detected by PM, because both methods are similar if the threshold value is set to 0 involving $\theta(z) = \theta(z_0)$. The WP wind velocities were used to calculate the $Ri_b$ from the MWR $T$ profile. The bR method was applied on daily RS and MWR $\theta$ and COSMO-2 $\theta_v$ profiles for CBL, cloudy-CBL and SBL detection.

For both PM and bR methods the surface $T$ has a large impact on the determined PBL height and hence it is crucial to take a representative measurement that is not biased by micrometeorological effects. The surface $T$ was therefore taken from the meteorological surface network at 2 m. If needed, a linear interpolation between two measured $\theta$ is applied to determine the PBL height. Uncertainties in PBL height for both methods were calculated by varying the surface $T$ by $\pm 0.5^\circ$ and were found to be in the order of $\pm 50–150$ m around the PBL maximal height reached in the early afternoon. Far larger uncertainties were found for the PBL height decrease in the late afternoon. For RS, $\theta$ and $\theta_v$ were calculated using the $p$ and RH provided by RS measurements, whereas for MWR RH was provided by the lidar and $p$ was calculated from the MWR
\[ p(z,T) = p_0 \cdot \exp \left(- \int_{z_0}^{z} \frac{M_a g}{R T(z')} \, dz' \right) \] (2)

where \( M_a \) is the mass of the air, \( R \) the specific gas constant and \( p \) is measured at 2 m from the meteorological surface network.

The nocturnal SBL can only be detected by the \( T \) profiles measured by RS and MWR, since wind turbulence, aerosol and relative humidity (RH) profiles retrieve the RL height during the night. The SBI is defined as the height of the surface-based \( T \) inversion, where \( T \) first decreases with elevation (\( dT/dz = 0 \)) as depicted in Fig. 3a (Bradley et al., 1993; Stull, 1988). A surface-based \( T \) inversion is a clear indicator of a stable boundary layer that can be defined as a SBL height (Seidel et al., 2010). The SBL top can also be defined as the transition between the stable surface layer and the neutral residual layer (Stull, 1988). This height is detected by a vanishing \( \theta \) gradient (\( d\theta/dz = 0 \)), which will be called SBLp\( T \) (Fig. 3b and c). SBLp\( T \) is per definition higher than SBI since the \( \theta \) gradient is still positive at the height of the surface-based \( T \) inversion and does not correspond to the top of the stable layer.

### 2.2.2 Method based on wind turbulence profiles

The radar echo measured by the WP is generated by inhomogeneities in the refractive index, which are characterized by the structure constant \( C_n^2 \). It can be shown that the range corrected SNR is proportional to \( C_n^2 \), which has a maximum at the top of the capping inversion, which marks the PBL top (White et al., 1991; Angevine et al., 1994, and references therein). Therefore a peak in the SNR profile can be associated to the PBL height under convective conditions. However, turbulence as well as humidity and \( T \) gradients associated with clouds and other dynamical processes can generate high SNR values, which do not correspond to the PBL height, leading to an attribution problem. To get rid of part of the false PBL height attribution, a time continuity algorithm was
applied: each SNR peak with local maximum greater than 75% of the absolute maximum was weighted by a Gaussian function with mean equals to the PBL height of the former time step and a standard deviation $\sigma$ depending on the hour of the day. The PBL height is then attributed to the maximum of the weighted SNR peak. The uncertainty of this method is considered equal to the full width at half maximum (FWHM) of the selected SNR peak after subtraction of the noise floor and is in the order of 100–500 m. CBL starting height at sunrise was set to ground height with a large $\sigma$ attributed to the first hours after sunrise. Similar algorithms taking into account the SNR slope and curvature were tested but have shown a lower consistency with respect to the other PBL height detection methods and a higher rate of false detections. This WP/SNR method was used to detect the CBL during the day and the RL during the night, but cannot be used in case of precipitation.

### 2.2.3 Method based on concentration profiles

The aerosol scattering ratio (ASR) is the ratio between the total and the molecular backscatter coefficients. Since the PBL top is characterized by a sharp decrease in concentration of all pollutants, the absolute minima in the vertical gradient of the lidar/ASR and of the RS/RH profiles can be associated to the CBL height during day and to the RL during night. A continuity algorithm similar to the WP/SNR method (see Sect. 2.2.2) was applied, with the modified condition that the local minimum has to be lower than 10% of the absolute minimum. According to the WP/SNR method the uncertainty is considered equal to the FWHM of the selected in the lidar/ASR gradient profile and is in the order of 100–250 m.
3 Results and discussion

3.1 Comparison of PBL height determined from potential and virtual potential temperature.

Both the PM and bR methods can be applied not only to $\theta$ but also to $\theta_v$ that also corrects for air moisture, water vapor being lighter than dry air. The humidity profile of the lidar was taken to calculate $\theta_v$ from MWR $T$ profile, restricting the data availability to cloud and precipitation free cases. A comparison of PBL height detected by either $\theta$ or $\theta_v$ was achieved for 35 convective days (12:00 to 15:00) taken between February and September 2012 (Table 2). The PBL heights computed from $\theta_v$ are slightly greater (3–8 %) than those computed from $\theta$, but the agreement is however good resulting in coefficient of determination of at least 0.95 and median bias smaller than 120 m in all cases. Considering these small differences and the greater $\theta$ availability, the RS and MWR potential temperatures $\theta$ were used for this study.

3.2 PBL height operational measurement

The operational procedure calculates the PBL heights each hour. Examples of the resulting plots are presented in Figs 4 and 5. All PBL heights from the various instruments and methods are plotted on the lidar/ASR (Fig. 4) or on the WP/SNR (Fig. 5) signal in the upper panel, whereas the vertical heat flux, the sunshine duration and the temporal gradient of the surface $T$ are plotted in the lower panel.

The first example (Fig. 4) is a clear CBL height diurnal cycle measured at PAY during a nice clear-sky convective day on the 23 July 2012, where all principal PBL features of Fig. 1 were clearly measured:

– The layered structure of the nocturnal PBL between midnight and the sunrise: (1) the SBI is detected by both RS (dark blue triangles) and MWR (reversed dark blue triangles) at about 100 m, (2) the SBL detected by MWR/bR (white squares) and the top of the stable layer detected by the MWR/SBLpT (magenta triangles)
peaks both at the same altitude of 500 m until 03:00 and decreases to about 200 m thereafter, (3) the SBL detected by the COSMO-2 bR (orange diamonds) stays constant at 250 m until sunrise, (4) the RL is detected by both the WP/SNR (light blue circles) and the lidar/ASR (green circles) at 1500 m, the WP catching another turbulent layer at 700–800 m between 03:00 and 09:00 corresponding to a jet of north-east wind (15 m s\(^{-1}\), not shown). These two layers measured by WP/SNR and lidar/ASR before sunrise are finally merged into the developing CBL at 07:00 and 09:00, respectively.

– The CBL development from sunrise to mid afternoon: (1) one hour after sunrise, the CBL height increase is very well caught by all the methods based on \(T\) profiles, MWR/bR and COSMO-2/bR showing a quicker CBL increase and a higher CBL height between sunrise and 09:00 than MWR/PM. This difference between the PM and bR methods is due to the horizontal wind component that is taken into account in the bR method. In this case, the air moisture seems to have minor influence since the bR method leads to similar results if applied on the \(\theta_v\) (COSMO-2) or on \(\theta\) profiles (MWR). (2) The CBL remains then constant from 12:00 to 15:00, when the temporal gradient of the ground \(T\) vanishes before becoming negative (see the lower panel of Fig. 4). This CBL height maximum is consistently measured by all methods. (3) The CBL decrease after 15:00 is also well depicted by the methods based on \(T\) profiles (MWR/bR and MWR/PM), whereas the RL is thereafter measured by the WP/SNR and the lidar/ASR. The PM method, which is devised for CBL detection, becomes non applicable as soon as the vertical sensible heat flux becomes negative (see the red curve in the lower panel of Fig. 4), generating a positive or vanishing gradient of \(\theta\).

– The nocturnal SBL development: after 18:00, the bR method continues to follow the CBL decrease whereas the development of the nocturnal SBL can be detected by the MWR/SBI and MWR/SBLpT methods.
The second example (Fig. 5) is a winter day with a stable cloud cover at 800–1200 m detected by the ceilometer. In that cloudy-CBL case, only the PBL height detection methods based respectively on $T$ profiles remain robust and provide reasonable results. Due to the presence of low clouds, the lidar is powered down and the WP/SNR detects actually the cloud top that decreases from about 1800 m at midnight to 1000 m 24 h later. The cloud thickness diminishes then gradually from about 1000 m to some 100 m before vanishing at the end of the day. Both the MWR/PM and the COMSMO-2/bR catch very well the cloudy-CBL that peaks at 500 m during the afternoon and decreases in height when the vertical heat flux becomes negative. During night, a MWR/SBI is only detected in the evening when the cloud coverage decreases allowing radiative cooling of the ground. MWR/bR is most of the time not available due to a vanishing $\theta$ gradient involving an already positive RI$_0$ at the first level and to missing wind velocity data from WP at some levels; the available MWR/bR heights are greatly influenced by northerly wind at about 500 m. Similarly, the SBL detected by COSMOS-2/bR is most of the time found at the first COSMO-2 level that is attributed to cases with vanishing $\theta$ gradient. The presence of a neutral layer measured by MWR/SBLpT over the ground is detected at a constant height of 200 m between 00:00 and 08:00.

### 3.3 Inter-comparison and validation for the CBL

The inter-comparison and validation process was performed at PAY on a set of 119 clear-sky convective days, representing $1/3$ of the total measured days in 2012. RS/PM at 12:00 was chosen as the reference method for the validation due to the availability and reliability of RS $T$ profiles and the stability of PM method. Table 3 and Fig. 6 allow drawing the following conclusions:

- Due to the use of the same RS data having a very great vertical resolution, the RS/bR and RS/RH gradient methods are the closest to RS/PM with regression slopes near 1, coefficients of determination ($R^2$) with the fit or with the 1 : 1 line greater than or equal to 0.85, and a very small median bias. As expected by the
bR definition, its PBL heights are higher than the ones computed by PM, the median bias remaining however very low at about 20 m.

- The MWR results are somewhat more scattered, but with very small median bias (< 25 m) and interquartile ranges (100 m). The MWR/PM has the smallest interquartile ranges and whiskers size due to the same applied detection method that, contrary to bR, do not use the WP wind velocity.

- The WP/SNR method has the lowest correlation coefficients (0.47), the largest median bias (−63 m) and the largest interquartile range (−560 to 460) of all the experimental methods. It also contains several large positive outliers that may be explained by the detection of elevated cloud layers falsely attributed to PBL height.

- The comparison with lidar/ASR can be only done on a reduced dataset (61 cases) due to its lower data availability. Taking into account the very different detection methods based on $T$ and aerosol profiles, the comparison with RS/PM is very good with a slope of 1.00, correlation coefficients of 0.81 and a median bias of −5 m.

Since the CBL may not always be at its maxima at 12:00, an inter-comparison on the same set of 119 convective days was performed with MWR/PM as reference for the 12:00–15:00 time interval corresponding to CBL height maxima for all seasons (Fig. 7). Similarly to the 12:00 case, the difference between PM and bR is rather small with interquartile ranges of 5 and 71 m and whiskers far below 200 m. The lidar/ASR also shows a very good agreement with a median bias of 20 m and an interquartile range of about ±150 m. Finally the great number of false detections of the WP due to either cloud, high humidity layers or turbulences are visible in the WP/SNR larger median bias (71 m) and interquartile range of about 200 m.

Each of the considered method has its own uncertainties in the PBL height detection as explained in Sect. 2.2. The uncertainty minimum is usually obtained for fully developed CBL that is the easiest case to detect. These uncertainties provide a similar
picture as the inter-comparison, with a greatest precision for methods based on $T$ profile and the lowest one for WP/SNR.

Finally, considering all these statistical differences found between the various instruments and methods as well as their related uncertainties, one has however to remain conscious that the measured parameter (PBL height) is in reality not a fixed point but rather a transition layer between 2 states of the atmosphere, which thickness reaches probably several tens of meters, and that the remote sensing instruments measure an air volume and not a precise point. The results of the inter-comparison and the instrumental uncertainties are of course greater than the thickness of this transition layer, but they stay however in the same order of magnitude.

### 3.4 Comparison between PBL height measured and computed by COSMO-2

Table 3 and Figs. 6 and 7 show that the PBL height given by COSMO-2 model has a positive bias compared to experimentally determined PBL heights. The median biases are of 275 m and 299 m when compared to the RS/PM (12:00) and to the MWR/PM (12:00 to 15:00), respectively. The interquartile ranges reach 200 to 350 m and the maximal whiskers are higher than 1000 m. A detailed analysis of the individual plots (see Fig. 8 for example) reveals that COSMO-2 often overestimates the PBL height during the whole day and tends to show a too rapid PBL growth in the morning. This behavior is not limited to clear-sky convective days and is observed throughout the year. This significant positive bias compared to all experimental methods and the asymmetry of the distribution, which is obvious on the histograms of Figs 6 and 7, may be explained by several reasons:

- Contrary to all the experimental methods, COSMO-2 determines the PBL height from $\theta_v$ profile, leading to a physically meaningful systematic positive bias. This bias of 3–8 % (see Sect. 3.1) cannot however explain the large discrepancy with the experimental methods.
– The use of the bR method also induces a positive bias compared to the PM method, but the difference does not exceed some tens of meters as demonstrated by the RS and MWR results.

– The bR method is very sensitive to the surface $T$ and an overestimation of this parameter can induce a systematic positive bias of PBL height. Errors and uncertainties in both $T$ and RH profiles of COSMO-2 could also explain the large observed bias.

– The occurrence of clouds, which may be missing in the model, can reduce for a while the surface heating and the convection of air masses leading to a lower measured PBL height. This phenomenon is clearly visible in some cases (not shown).

Further studies are necessary to assess the impact of these various parameters and determine the main causes of the PBL height overestimation by COSMO-2.

### 3.5 PBL height two-year climatology at PAY and SHA

#### 3.5.1 CBL climatology

The two years climatology of CBL heights calculated from all instruments and COSMO-2 is presented in Fig. 9 for PAY (256 days) and SHA (289 days). It has to be noted that the same subset of days was taken for the MWR, the WP and COSMO-2, whereas the lower availability of lidar/ASR data leads to a smaller dataset that still allows the comparison with the CBL heights estimated from the other instruments. The CBL heights have a clear annual cycle with a minima at 300–700 m in winter and a maxima at 1200–1500 m during the Mai–August period. It has to be noted that the CBL extremes occur at the solstices and not at the $T$ extremes (January–February and July–August), allowing to conclude that the solar radiation would be a better climatic variable to predict CBL cycle than the $T$. 
The systematic overestimation of the COSMO-2 model observed at both stations presents a clear annual cycle with a winter minimum and a summer maximum that can reach 500–700 m. At PAY and to a lesser extend at SHA, the WP/SNR and lidar/ASR detect a higher CBL (300–500 m) than the MWR/PM in winter. This can probably be related to meteorological conditions with high-altitude $T$ inversion leading to a stable and sometimes decoupled aerosol layer at altitudes higher than the CBL top. The lidar/ASR measurement of this aerosol layer top should be better associated to a RL than a CBL height and the WP/SNR measures the turbulences due to wind shear at the $T$ inversion altitude.

The CBL maxima measured over the Swiss plateau are similar to the PBL heights maxima measured over Europe by RS (Seidel et al., 2012; Beyrich and Leps, 2012), but lower than the lidar measured PBL height over Leipzig (Baars et al., 2008) and the PBL height detected by several methods (RS, MWR and lidar) over Granada (Granados-Muños et al., 2012). The higher PBL height over both regions can be explained for Leipzig by its lower altitude (135 m a.s.l.), its northerly latitude leading to longer summer days and similar annual $T$ cycle (0°C in winter and 18°C in summer (www.dwd.de)), and for Granada by the far greater mean $T$ (6°C in winter and 25°C in summer (www.aemet.es)) even if the city lies at higher altitude (730 m a.s.l.) than PAY.

### 3.5.2 Cloudy-CBL climatology

Cloudy-CBL cases have been selected as non CBL days without rain between 6:00 and 15:00 and correspond to various meteorological situations (high altitude clouds, fog, advections, mixed situations, . . . ). As expected by its more heterogeneous atmospheric structure, the cloudy-CBL climatology (271 days at PAY and 223 at SHA) presents more scattered results with larger quartiles (Fig. 10) than the CBL one. The cloudy-CBL annual cycle based on the $T$ profile (MWR/PM and COSMO-2) is similar to the CBL cycle but with lower PBL heights. The difference between CBL and cloudy-CBL heights is greater at SHA (500–700 m) than at PAY (< 300 m). COSMO-2 cloudy-CBL height has a positive bias compared to MWR/PM at both stations that is somewhat
higher than for the CBL case. The WP/SNR cloudy-CBL heights are most of the time more than 500 m higher than MWR/PM ones and measures probably the cloud top in a number of cases (see for example Fig. 5). Despite the low amount of available data, the lidar/ASR results are very similar to MWR/PM PBL heights during summer and somewhat higher in winter, similarly to the CBL case.

### 3.5.3 SBL climatology

The SBL climatology was divided into clear-sky (Fig. 11) and cloudy nights (Fig. 12) in order to differentiate cases with large and small radiative cooling. Clear-sky (186 at PAY and 163 at SHA) and cloudy nights (126 at PAY and 151 at SHA) were selected as days without precipitation between 00:00 and 5:00 and with 0–2 and 7–8 octa of the sky covered by clouds estimated by APCADA, respectively. While some features of the SBL annual behavior can be deduced, the low number of cases for some months, particularly for cloudy conditions, does not allow us to draw strict conclusions on the effective seasonal cycle of the different layers forming the SBL. The following points can however be observed:

- During clear-sky nights, the complete SBL structure can be clearly observed at PAY with SBI heights being between 100 and 500 m during the whole year, SBLpT being lower than 500 m in winter and rising up to 800 m during the other seasons. The RL measured by the lidar/ASR has a seasonal cycle completely similar to the CBL one (see Fig. 9), so that the pollutant emitted during the preceding days remain at the altitude of the CBL maxima during the night.

- During clear-sky nights, the WP/SNR method, which is more frequently subjected to false attribution than the other methods, leads to much more scattered results and large quartiles. The WP/SNR results are however comparable to the RL heights measured by the lidar/ASR.
During cloudy nights, the ground \( T \) remains higher due to lower radiative cooling and a different SBL structure is observed. First, the SBI is found at a lower altitudes (100–200 m) than during clear-sky nights, second the SBLpT also reaches lower altitudes remaining usually under 500 m, third the cloud base is found between 500 and 2000 m, finally the cloud top is measured by the WP/SNR between 1000 and 2000 m. A mean cloud depth between 200 and 1000 m is therefore measured over PAY. One can say that the various SBL heights measured by \( T \) profiles are all compressed under 500 m by the cloud base.

The COSMO-2 bR method frequently computes SBL height lower than 50 m that can hardly represent a real physical PBL height. These false estimations are due to a stable \( \theta \) profile near ground leading to an already positive bR number at the first levels and occur more frequently during calm and clear-sky nights with large ground radiative cooling than during cloudy nights with higher ground \( T \) and less turbulence. This phenomenon is clearly visible in Fig. 11: in case of clear-sky nights, COSMO-2 SBL heights are always lower or equal to 50 m whereas MWP/bR measures a higher valid SBL height but in much fewer cases (see lowest panels). During cloudy nights (Fig. 12), COSMO-2 produces more reliable results with SBL heights in the same order of magnitude than the MWR/SBI, MWR/bR and MWR/SBLpT methods.

The MWR/bR method gives results usually similar to SBI in case of clear-sky but clearly higher in case of cloudy nights. This difference is probably due to the direct dependence of SBI height on the ground radiative cooling, whereas the bR method is more affected by wind turbulences and katabatic jets that are not discriminated by the cloud amount.

Few SBL climatologies have been yet published probably due to the greater complexity of PBL heights detection during night than during day. Cimini et al. (2013) found MWR/SBL height lower than 500 m near Paris during the March–August period that are comparable to our climatology over the Swiss plateau. Martucci et al. (2007) found...
nighttime RL heights detected by lidar/ASR between 500 and 1500 m in Neuchâtel (Switzerland) similarly to our results. Finally, Beyrich and Leps (2012) and Seidel et al. (2010) studied the 10 year climatology of PBL height detected by RS measurements (twice a day). The SBL seasonal cycles over Europe were found to depend on the method applied to the RS profiles: the PM method leads to almost constant SBL during the whole year, whereas SBI has a seasonal minima in summer and a maxima in winter. Unfortunately our two-year dataset restricted by the cloud coverage is not large enough to compare our SBL seasonal cycles with these results. Finally, similarly to our results, the gradient method applied to the RH or specific humidity profiles is maximal during summer and minimal during winter. As expected, they also found that SBI yields the smallest heights, followed by the PM method, while the humidity and the ASR profiles similarly lead to much greater heights, corresponding to RL top.

4 Conclusion: strengths and limitations of an operational mode

A system for automatic real time detection of the PBL height based on several methods applied to various remote sensing observations was implemented and operated for two years (2012–2013) for two upper air stations on the Swiss plateau. The numerical weather prediction model COSMO-2 PBL height was also compared to instrumental results. All the remote sensing and model results were validated on a subset of 119 convective days, the RS/PM at 12:00 or the MWR/PM between 12:00 and 15:00 being taken as references. A two years climatology for daytime and nighttime PBL heights were calculated for convective days and clear-sky nights, as well as for non-convective days and for cloudy nights without precipitation. The system for automatic detection of the PBL height is now implemented in an operational environment and the data are visualized and provided to end users in real time.

The difficulty of the PBL height detection comes first from the complexity of the troposphere itself, which can be composed of several layers with different thermal structure, wind regimes and concentrations of atmospheric constituents. Secondly, each
detection method has good performances only for defined PBL structures and under specific meteorological conditions. Only the combination of several methods and instruments allows to follow the complete diurnal cycle of the complex PBL layered structure. The advantages and limitations of each detection/measurement method as an operational mode are summarized in Table 4 and in the next paragraphs.

The greatest advantage of PBL detection by the various profiles measured by RS is its very good measurement precision and vertical resolution. Its temporal resolution (2 measurement per day) does however not provide the PBL diurnal cycle.

The MWR provides $T$ profiles under all non-precipitating conditions with a lower vertical but a higher temporal resolution than RS allow to analyze the whole PBL diurnal cycle. The four PBL height detection methods applied on MWR data allow to follow (1) the PBL growth after sunrise, its maximal elevation at the beginning of the afternoon and its decrease as soon as the vertical heat flux vanishes up to some hours after sunset, (2) the SBI development and maximal height from sunset to sunrise that corresponds to the layer in which the pollutants emitted during the night are trapped, (3) the top of the nocturnal stable layer (measured by SBLpT). MWR is therefore able to detect the daytime and nighttime layers in which ground emitted atmospheric constituents are trapped, but not the RL corresponding to the air volume keeping the atmospheric constituents emitted some hours/days before.

The Raman lidar has a higher vertical resolution than MWR and its data availability is restricted by fog, low cloud coverage and precipitation. The profiles of the aerosol or the humidity concentrations allow to measure the dynamics of atmospheric constituents and are consequently a direct measure of the pollutant dispersion in the PBL. The comparison with RS/PM and MWR/PM proves that the lidar/ASR is able to detect the CBL maxima during the afternoon with a good precision and also sometimes part of the CBL formation. During night this method provides the RL height and can therefore be considered as complementary to the MWR methods.

The comparison of WP/SNR with RS/PM and MWR/PM shows that, in most cases, the CBL maximum is well detected by WP, but with a lower precision and a greater
amount of outliers. De facto WP/SNR maxima can be generated by turbulences at the PBL top, but also at cloud top or at wind shears. An operational PBL height measurement by WP is therefore much more difficult to implement without a human visual control to attribute the SNR maxima to the real atmospheric phenomena. In case of cloudy condition, the WP/SNR tends to measure the cloud top instead of the PBL height, which could be exploited for other applications. It has to be noted that the WP and the Raman lidar have been used in their operational configuration. However, it would technically be possible for both systems to go to higher temporal and vertical resolutions optimized for PBL height detection which could slightly improve their performance.

The forecast model COSMO-2 uses the bR method applied to the $\theta_v$ profile and relies therefore on bR qualities (day and night detection, detection of CBL growth, maxima and decrease) and weaknesses (often false detection during night particularly in case of clear-sky conditions). COSMO-2 is found to often overestimate both CBL and cloudy-CBL by 500–1000 m. The most probable causes for this discrepancy are systematic differences in terms of surface $T$, $T$ or RH profiles. This issue will be addressed in future work. The SBL detection during night is attributed to the lowest level in case of stable $\theta_v$ gradient, which could lead to misinterpretation of this value that does not really correspond to a PBL height. To avoid such misunderstanding, a missing value or a flagged value should be introduced instead of the lowest level for these cases.

We conclude that the MWR/PM is the most robust among the experimental methods under consideration and best suited for automatic real time detection of the PBL height. It provides good results under a wide range of meteorological conditions. Moreover, the MWR/SBI and SBLpT allow to characterize the nocturnal SBL. It has however to be associated with a ceilometer or a lidar to monitor the RL height. A potential for improvement is the combination of the experimental methods and to compute the best guess of PBL height based on all available methods.

Taking advantages of all available upper-air measurements, the principal features of the PBL are well depicted by the two years climatology. The annual cycle of the CBL height with its maxima at 1500 m during the May–August period is caught by all
instruments and seems to follow the solar radiation cycle rather than the $T$ cycle. In case of partial or total cloudy conditions, a similar annual cycle but with lower PBL heights is measured, the WP results being however largely influenced by wind turbulence at the cloud top. The nocturnal PBL structure can be clearly observed under clear-sky conditions, with the SBI height remaining rather constant during the year at 200–300 m, the top of the stable layer lying at 800 m for most of the non-winter months and finally the RL nocturnal seasonal cycle following the CBL diurnal maximal. In case of total cloud coverage, the SBI height is lower than in case of clear-sky, and the SBL layers seems to be compressed and not well-structured under the cloud base. Further meteorological phenomena such as fog, neutral boundary layer height, main pollutant advections or nocturnal jets will be further addressed either as case studies or statistically after a longer measurement period.

References


Table 1. List of abbreviations.

<table>
<thead>
<tr>
<th>Atmospheric layers</th>
<th></th>
</tr>
</thead>
<tbody>
<tr>
<td>CBL</td>
<td>Convective Boundary Layer</td>
</tr>
<tr>
<td>cloudy-CBL</td>
<td>CBL for overcast conditions but without precipitations</td>
</tr>
<tr>
<td>NBL</td>
<td>Neutral Boundary Layer</td>
</tr>
<tr>
<td>PBL</td>
<td>Planetary Boundary Layer</td>
</tr>
<tr>
<td>RL</td>
<td>Residual Layer</td>
</tr>
<tr>
<td>SBL</td>
<td>Stable Boundary Layer</td>
</tr>
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</table>

<table>
<thead>
<tr>
<th>Instruments</th>
<th></th>
</tr>
</thead>
<tbody>
<tr>
<td>lidar</td>
<td>Raman lidar</td>
</tr>
<tr>
<td>MWR</td>
<td>Microwave radiometer</td>
</tr>
<tr>
<td>RS</td>
<td>Radio sounding</td>
</tr>
<tr>
<td>WP</td>
<td>Windprofiler</td>
</tr>
</tbody>
</table>

<table>
<thead>
<tr>
<th>Methods</th>
<th></th>
</tr>
</thead>
<tbody>
<tr>
<td>APCADA</td>
<td>Automatic Partial Cloud Amount Detection Algorithm</td>
</tr>
<tr>
<td>ASR</td>
<td>Aerosol Scattering Ratio</td>
</tr>
<tr>
<td>bR</td>
<td>Bulk Richardson number method</td>
</tr>
<tr>
<td>COSMO-2</td>
<td>COnsortium for Small-scale MOdeling</td>
</tr>
<tr>
<td>PM</td>
<td>Parcel Method</td>
</tr>
<tr>
<td>SBI</td>
<td>Surface-based Temperature Inversion</td>
</tr>
<tr>
<td>SBLpT</td>
<td>Stable Boundary Layer detected by potential Temperature</td>
</tr>
<tr>
<td>SNR</td>
<td>Signal to Noise Ratio</td>
</tr>
</tbody>
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<table>
<thead>
<tr>
<th>Measuring sites</th>
<th></th>
</tr>
</thead>
<tbody>
<tr>
<td>PAY</td>
<td>Payerne</td>
</tr>
<tr>
<td>SHA</td>
<td>Schaffhausen</td>
</tr>
</tbody>
</table>
Table 2. Linear regression of PBL height detected by $\theta_v$ as a function of the PBL height detection by $\theta$: slope, intercept, coefficient of determination between the data and the fit ($R^2$), coefficient of determination between the data and the 1:1 line ($R_{th}^2$), root mean square error of the $x-y$ difference (RMS), median of the difference between $x$ and $y$ coordinates (median bias) and the number of considered data ($N$). The results are given for PM and bR methods applied on RS and MWR data.

<table>
<thead>
<tr>
<th>Instrument/method</th>
<th>Slope</th>
<th>Intercept</th>
<th>$R^2$</th>
<th>$R_{th}^2$</th>
<th>RMS [m]</th>
<th>Median bias [m]</th>
<th>$N$</th>
</tr>
</thead>
<tbody>
<tr>
<td>RS/PM</td>
<td>1.08</td>
<td>-12</td>
<td>0.95</td>
<td>0.92</td>
<td>110</td>
<td>18</td>
<td>35</td>
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<tr>
<td>RS/bR</td>
<td>1.05</td>
<td>5</td>
<td>0.97</td>
<td>0.94</td>
<td>99</td>
<td>26</td>
<td>35</td>
</tr>
<tr>
<td>MWR/PM</td>
<td>1.03</td>
<td>104</td>
<td>0.95</td>
<td>0.84</td>
<td>161</td>
<td>117</td>
<td>437</td>
</tr>
<tr>
<td>MWR/bR</td>
<td>1.05</td>
<td>67</td>
<td>0.95</td>
<td>0.88</td>
<td>154</td>
<td>93</td>
<td>420</td>
</tr>
</tbody>
</table>
**Table 3.** Linear regression of PBL height computed with various methods and instruments as a function of RS/PM. See Table 2 for parameters description.

<table>
<thead>
<tr>
<th>Instrument/method</th>
<th>Slope</th>
<th>Intercept</th>
<th>$R^2$</th>
<th>$R_{th}^2$</th>
<th>RMS [m]</th>
<th>Median bias [m]</th>
<th>N</th>
</tr>
</thead>
<tbody>
<tr>
<td>RS/bR</td>
<td>1.02</td>
<td>46</td>
<td>0.95</td>
<td>0.93</td>
<td>122</td>
<td>21.5</td>
<td>118</td>
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<tr>
<td>RS/RH</td>
<td>1.01</td>
<td>3.64</td>
<td>0.86</td>
<td>0.90</td>
<td>154</td>
<td>0</td>
<td>118</td>
</tr>
<tr>
<td>MWR/PM</td>
<td>0.89</td>
<td>73</td>
<td>0.75</td>
<td>0.74</td>
<td>228</td>
<td>−25.5</td>
<td>100</td>
</tr>
<tr>
<td>MWR/bR</td>
<td>0.84</td>
<td>173</td>
<td>0.72</td>
<td>0.69</td>
<td>239</td>
<td>2.33</td>
<td>85</td>
</tr>
<tr>
<td>WP/SNR</td>
<td>0.73</td>
<td>210</td>
<td>0.49</td>
<td>0.41</td>
<td>351</td>
<td>−64</td>
<td>105</td>
</tr>
<tr>
<td>Lidar/ASR</td>
<td>1.00</td>
<td>−3</td>
<td>0.81</td>
<td>0.81</td>
<td>211</td>
<td>−50</td>
<td>61</td>
</tr>
<tr>
<td>COSMO-2</td>
<td>1.20</td>
<td>141</td>
<td>0.72</td>
<td>0.43</td>
<td>472</td>
<td>275</td>
<td>114</td>
</tr>
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</table>
Table 4. Advantages and limits of detection methods and instruments to estimate the PBL height.

<table>
<thead>
<tr>
<th>Method</th>
<th>Profiles</th>
<th>PBL height detected</th>
<th>Advantages</th>
<th>Limits</th>
</tr>
</thead>
<tbody>
<tr>
<td>PM</td>
<td>$\theta$ or $\theta_v$</td>
<td>CBL, cloudy-CBL</td>
<td>– also efficient under weak convective condition</td>
<td>– requires negative gradient in $\theta$ at the ground</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– early growth after sunrise until decrease when temporal gradient of surface $T$ and vertical heat flux become negative</td>
<td>– not available during night</td>
</tr>
<tr>
<td>bR</td>
<td>$\theta$ or $\theta_v$ + wind</td>
<td>CBL, cloudy-CBL, SBL</td>
<td>– nighttime and daytime detection</td>
<td>– requires wind profiles from WP or RS</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– transition between SBL and CBL at sunrise</td>
<td>– often false SBL detection in case of constant $\theta$ profile</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– CBL decrease also after the vertical heat flux and temporal $T$ gradient become negative</td>
<td></td>
</tr>
<tr>
<td>SBI</td>
<td>$T$</td>
<td>SBL</td>
<td>– SBI formation after sunset</td>
<td></td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– describe the layer where the pollutants emitted during night are trapped</td>
<td></td>
</tr>
<tr>
<td>SBLpT</td>
<td>$\theta$ or $\theta_v$</td>
<td>SBL</td>
<td>– Formation and top of the stable nocturnal layer</td>
<td>– Not well defined limit of the SBL layered structure</td>
</tr>
<tr>
<td>Aerosol/humidity gradient</td>
<td>ASR, RH</td>
<td>CBL, RL</td>
<td>– measures the dynamics of aerosol dispersion</td>
<td>– no measure of the SBL</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– a real measure of the pollutants ML</td>
<td></td>
</tr>
<tr>
<td>SNR maxima</td>
<td>wind</td>
<td>CBL, (RL)</td>
<td>– sometimes retrieves PBL height early growth after sunrise</td>
<td>– large number of outliers due to false attributions</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– only method based on the vertical structure of turbulence.</td>
<td>– can also retrieve the cloud top</td>
</tr>
</tbody>
</table>

<table>
<thead>
<tr>
<th>Instrument</th>
<th>Profiles</th>
<th>PBL height detected</th>
<th>Advantages</th>
<th>Limits</th>
</tr>
</thead>
<tbody>
<tr>
<td>Microwave radiometer</td>
<td>$T$ , RH, wind</td>
<td>CBL, cloudy-CBL, SBL</td>
<td>– captures diurnal cycle</td>
<td>– low vertical resolution</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– good data availability</td>
<td></td>
</tr>
<tr>
<td>Windprofiler</td>
<td>Wind, SNR ratio</td>
<td>CBL, RL</td>
<td>– daily cycle</td>
<td>– no PBL detection in case of precipitation</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– can also retrieve the cloud top based on the vertical structure of turbulence.</td>
<td>– low data availability at low altitude</td>
</tr>
<tr>
<td>Lidar</td>
<td>ASR, RH</td>
<td>CBL, RL</td>
<td>– daily cycle</td>
<td>– no data in case of fog, low clouds</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– direct measurement of atmospheric composition</td>
<td>– needs maintenance</td>
</tr>
<tr>
<td>Radio Sounding</td>
<td>$T$ , $\rho$, RH, wind</td>
<td>CBL, cloudy-CBL, SBL</td>
<td>– most accurate and precise data</td>
<td>– only twice a day at 00:00 and 12:00</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>– best vertical resolution</td>
<td></td>
</tr>
</tbody>
</table>

Determination and climatology of the planetary boundary layer height

M. Collaud Coen et al.
Figure 1. Diurnal cycle of the PBL height over land for a clear convective day (adapted from Stull, 1988).
Figure 2. PBL detection methods based on $T$ profiles: (a) Parcel method applied on MWR $T$ and $\theta$ profiles and (b) bulk Richardson number method applied on RS $\theta$ profile. Both profiles were measured at about 12:00 on the 16 July 2012 in convective conditions.
Figure 3. Detection of the SBL from RS T-profile of 17 July 2012: (a) the surface-based temperature inversion (SBI) determined by the first $T$ decrease as a function of altitude, (b) the top of the stable layer (SBLpT) detected by the stability of $\theta$ profile or by the vanishing $T$ gradient.
Figure 4. Upper panel: automatic detection of PBL height from all remote sensing instruments, RS and COSMO-2 model for a convective day in summer 2012 (23 July 2012) at PAY; the background signal corresponds to the lidar/ASR. Lower panel: sunshine duration, vertical heat flux and temporal gradient of surface $T$. Vertical heat flux greater than 10 or lower than $-10$ W m$^{-2}$ are limited to $\pm 10$ with a dashed line.
Figure 5. Example of cloudy-CBL detection under cloudy conditions in winter (14 February 2013) plotted on WR/SNR signal as background. For symbol description see Fig. 4.
Figure 6. Boxplots of PBL height differences $\Delta H$ between RS/PM and other methods/instrumentation computed at 12:00. The central box line is the median, the edges of the box are the 25th and 75th percentiles ($q_1$ and $q_3$), the whiskers enclose all data points not considered outliers, and the red crosses are the outliers. Data are considered as outliers if they are larger than $q_3 + 1.5 \cdot (q_3 - q_1)$ or smaller than $q_1 - 1.5 \cdot (q_3 - q_1)$, which means that whiskers cover 99% of data assuming a normal distribution. The $\Delta H$ statistical distribution and the number of data points $N$ for each boxplot are given in the sub plotted histograms, data points greater than 1000 m being displayed in the last column of the corresponding histogram.
Figure 7. Same as Fig. 6 but between 12:00 and 15:00 UT and with MWR/PM taken as the reference.
Figure 8. Example of CBL overestimation by COSMO-2, the background signal corresponds to the lidar/ASR. For a description of the symbols, see Fig. 4.
Figure 9. Upper panel: CBL height two-years climatology at PAY (left) and SHA (right). The dots are the monthly median of the daily medians of the CBL height taken between 12:00 and 15:00; the error bars are the 25th and 75th percentiles. Lower panel: the number of convective days are given in black for MWR/PM, WP and COSMO-2 and in green for lidar/ASR.
Figure 10. Cloudy-CBL top height climatology at PAY (left) and SHA (right). Symbols as in Fig. 9.
Figure 11. SBL and RL heights for clear-sky conditions at PAY (left) and SHA (right). Symbols as in Fig. 9.
Figure 12. SBL and RL heights for cloudy conditions at PAY (left) and SHA (right). Symbols as in Fig. 9.