Brown carbon in tar balls from smoldering biomass combustion

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Abstract

We report the direct observation of large-scale production of spherical, carbonaceous particles – “tar balls” – from smoldering combustion of two commonly occurring dry mid-latitude fuels. Real-time measurements indicate that brown carbon is an important component of tar balls. The spectrum of the imaginary parts of their complex refractive indices can be described with a Lorentzian-like model with an effective resonance wavelength in the ultraviolet (UV) spectral region. Sensitivity calculations for aerosols containing traditional organic carbon (no absorption at visible and UV wavelengths) and brown carbon suggest that accounting for UV absorption by brown carbon leads to a significant increase in aerosol radiative forcing efficiency and increased atmospheric warming. Since particles from smoldering combustion account for nearly three-fourths of the total carbonaceous aerosol mass emitted globally, inclusion of the optical properties of tar balls into radiative forcing models has significance for the Earth’s radiation budget, optical remote sensing, and understanding of anomalous UV absorption in the troposphere.

1 Introduction

A recent report by the Intergovernmental Panel on Climate Change (IPCC) (Forster et al., 2007) highlights large uncertainty in current estimates of direct radiative forcing (RF) due to organic carbon (OC) aerosols from biomass and fossil fuel burning. Most RF and optical remote sensing retrieval models use traditional optical properties of OC particles (i.e., OC absorption can be neglected in the visible and near visible spectral region) (Forster et al., 2007; Chung and Seinfeld, 2002). Recently, a class of OC compounds – optically defined as brown carbon (Andreae and Gelencser, 2006; Moosmüller et al., 2009) – has been shown to strongly absorb solar radiation in the blue and near ultraviolet (UV), thereby modifying RF (Forster et al., 2007). Global inventories of black (BC) and organic carbon (OC) emitted from combustion (Bond et
al., 2004) show that approximately 88% of total carbonaceous aerosol mass is emitted from biomass combustion, with approximately 80% of this emitted from the smoldering combustion phase (Einfeld et al., 1991).

Duff combustion has been shown to be one of the largest contributors to smoldering smoke production (Hille and Stephens, 2005; Neary et al., 1999). Duffs are often the dominant fuel in low intensity wildland fires and prescribed burns, and may smolder for an extended time period (days to months). Consumption of duff in wildland fires is a direct function of fuel moisture and duff with a moisture content below 14% gets totally consumed during a forest fire (Hille and Stephens, 2005). While flaming combustion produces mostly fractal-like black carbon (BC) particles, there is limited knowledge about the properties of particles from smoldering duff combustion.

In this article, we report the direct observation of near-spherical particles (identified as “tar balls”; Chakrabarty et al., 2006) produced from the smoldering combustion of two commonly occurring dry duffs: (a) Ponderosa Pine Duff (PPDuff), the partly decayed litter in a ponderosa pine (Pinus ponderosa) forest including a portion of the uppermost soil layer collected from a rural location near Missoula, Montana, USA; and (b) Alaskan Duff (AKDuff), the uppermost layer of soil with live and dead feathermoss (Pleurozium schreberi) collected from a spruce-feathermoss forest near Tok, Alaska, USA. Ponderosa pine is commonly found in mountainous terrain of southwestern Canada and the Cascades, Sierra Nevada, and Rocky Mountains of the United States; spruce-feathermoss forests dominate the southern boreal forest zone, which includes a large portion of Alaska. We discuss the unique optical properties of tar balls, and highlight the need to include them into RF and optical remote sensing models due to their impact on climate change and aerosol retrieval.

2 Experiment and analysis

Experiments were conducted during 2003-06 in the combustion facility of the United States Department of Agriculture (USDA) Forest Service Fire Sciences Laboratory
(FSL) as part of a comprehensive study to investigate the physical, optical, and chemical properties of smoke particles emitted from combustion of common mid-latitude wildland fuels (McMeeking et al., 2009). Smoldering combustion of approximately 200 g of PPDuff and AKDuff on a flat fuel bed (Fig. 1) produced white smoke that filled a large (∼3300 m³) chamber from which aerosol was sampled through a PM_{2.5} inlet and distributed through a manifold (Fig. 1) to a suite of instruments, namely a particle sampling unit for scanning electron microscopy (SEM), one dual wavelength (405 nm and 532 nm) and a single wavelength (780 nm) integrated photoacoustic-nephelometers (IPNs) (Arnott et al., 1999; Abu-Rahmah et al., 2006; Lewis et al., 2008), and an electrical low pressure impactor (ELPI) (Marjamäki et al., 2000). Conductive tubing was used to transport the particles to the various instruments in order to minimize particle losses. PPDuff had a fuel moisture content (as fraction of dry mass) of 13.9% and burned with a modified combustion efficiency (Ward et al., 1996) of 0.915, while AKDuff had a fuel moisture content (as a fraction of dry mass) of 9.2% and burned with a modified combustion efficiency of 0.902. Smoke sampling for each smoldering burn lasted for approximately two hours, with only minimal flaming combustion during ignition.

Simultaneous measurements, with one-second time resolution, were performed to determine the scattering and absorption coefficients (IPN), and the particle size distribution and number concentration (ELPI). The IPN consists of a photoacoustic spectrometer and a reciprocal integrating nephelometer in one instrument using the same sample volume and laser beam. The photoacoustic spectrometer measures particle light absorption with a power-modulated laser beam inducing particle temperature changes and subsequent pressure changes of the surrounding air at the modulation frequency (Arnott et al., 1999). The reciprocal integrating nephelometer adds a cosine-weighted optical detector to the instrument that measures the integrated (over ∼4π) scattering from the sample volume yielding the scattering coefficient $B_{\text{sca}}$ (Abu-Rahmah et al., 2006). The two IPNs used in this study were operated with a modulation frequency of 1500 Hz to measure the scattering ($B_{\text{sca}}$) and absorption ($B_{\text{abs}}$).
coefficients of the particles. For the SEM analysis, particles were impacted at a flow rate of 2 L/min onto 10-µm thick nuclepore clear polycarbonate 13-mm diameter filters (Whatman Inc, Chicago, IL) mounted on Costar Pop-Top Membrane holders. The filter exposure times for each of these burns were chosen, with the help of theoretical calculations involving factors such as flow rate and particle number concentration, to ensure moderate filter loading to help perform single particle analysis under a SEM. After sampling, the filter samples were kept in refrigerated storage and later prepared for SEM analysis by coating them with a 1-nm thick layer of platinum to prevent aerosol charging during SEM analysis. The coated filters were analyzed using a Hitachi Scanning Electron Microscope (Model S-4700).

The Electrical Low Pressure Impactor (ELPI) incorporates aerodynamic size classification with the electrical detection of particles in the range of several nm to 10 µm (Marjamäki et al., 2000; Marjamaki and Keskinen, 2004; Virtanen et al., 2001). In ELPI, aerosol undergoes unipolar corona charging before passing through a low-pressure cascade impactor to separate particles of different aerodynamic diameters ($D_a$). Electrometers record the current deposited on each impactor stage can be up to 1 Hz. Due to such a high time resolution, ELPI has been deployed to sample transient-mode particle emission from automotive and biomass burning sources (Maricq et al., 2006; Latva-Somppi et al., 1998; Hays et al., 2004). The size resolution of ELPI is limited by 13 impactor stages (including a filter stage in this study). Primary particle collection efficiency of each stage is characterized by a size-cut $D_{a,50}$ and a steepness coefficient $s$. Beside impaction, secondary collection mechanisms include diffusion, image charge and space charge effects, which are dependents of particle mobility diameter ($D_p$) (Virtanen et al., 2001). An ELPI response function developed by (Marjamäki et al., 2005) was followed in this study for retrieval of particle size-number concentration distribution. Kernel functions that transfer particle size-number concentration to measured current on stages by taking into account all primary and secondary collection mechanisms were constructed. $D_a$ is related to $D_p$ using:

$$D_a^2 c_c(D_a, \kappa_i) \rho_0 = D_p^2 c_c(D_p, \kappa_i) \rho_p$$  \hspace{1cm} (1)
where \( C_c \): Cunningham slip correction factor, which depends on particle size and mean free path

\[ k_i : \text{Particle mean free path at stage } i, i = 1, 2, \ldots, 13 \]

\[ \rho_0 : \text{Reference density (1 g cm}^{-3}\text{)} \]

\[ \rho_p : \text{Effective particle density} \]

Equation (1) assumes nonporous spherical particles (Kelly and McMurry, 1992), which can apply to the smoldering-combustion particles observed in this study. An iterative, least-square-based inversion algorithm similar to (Dong et al., 2004) was then used to retrieve particle size-number concentration from the kernel functions and ELPI current measurements. The size-number concentration was constrained to bimodal log-normal distribution, though it was found later that best fits usually consist of only single mode. To verify the robustness of the solution, multiple runs with random initial distribution were performed. They were compared with mobility number size distribution measurements from a Scanning Mobility Particle Sizer (SMPS) operated by a collaborating group during this experiment (McMeeking et al., 2009). The ELPI modes generally agreed with those from a Scanning Mobility Particle Sizer (SMPS). However, due to a coarse time resolution of \( \sim 10 \) min and size coverage of \( \sim 0.04–0.6 \) µm, the SMPS data were not suitable for examining the optical closure of smoke particles. Further, the SMPS under-stated the total number concentrations due to the saturation limit of the condensation particle counter (CPC) detectors (TSI inc., 2003). Uncertainty of the ELPI retrieval results partly from the unknown particle density. From the literature, organic matter from biomass burning shows \( \rho_p \) ranging from 0.79–1.53 g cm\(^{-3}\) (Reid et al., 2005). In this study the mean value of 1 g cm\(^{-3}\) was used (Radke et al., 1988; Radke et al., 1991), which gives equal aerodynamic and mobility diameter.
3 Results and discussions

SEM images (Fig. 2) reveal that typical smoke particles in this study are spherical and homogeneous tar balls (Pósfai et al., 2004; Hand et al., 2005; Alexander et al., 2008; Chakrabarty et al., 2006). Tar balls are a recently recognized type of near-spherical, atmospheric aerosol particle consisting of amorphous carbonaceous material, and have been found to exist in abundance in polluted continental air masses (Pósfai et al., 2004; Hand et al., 2005; Alexander et al., 2008). Electron dispersive spectroscopy (EDX) of tar balls shows that these particles consist primarily of carbon and oxygen with an average molar ratio of about six (Pósfai et al., 2004; Hand et al., 2005; Alexander et al., 2008).

Figure 3 shows four-minute, time-averaged number size distributions of smoke particles determined by the ELPI. These size distributions are narrow with a single peak at \( \sim 70 \text{ nm} \) diameter, similar to the size distributions of tar balls obtained from the SAFARI and K-pustza samples (Pósfai et al., 2004). Previous chemical characterization of the particles, performed using scanning transmission x-ray microscopy coupled with near-edge x-ray absorption fine structure spectroscopy (STXM/NEXAF), indicated a high proportion of oxygen-containing functional groups and very low \( \text{sp}^2 \) hybridization (Hopkins et al., 2007). These findings clearly differentiate tar ball composition from that of BC particles, which are fractal-like aggregates of spherules that consist of concentrically wrapped graphitic layers with a low proportion of oxygen-containing functional groups and relatively high \( \text{sp}^2 \) hybridization (Hopkins et al., 2007). The BC content of our aerosol samples was previously determined with a thermal-optical analysis method (Chow et al., 2004) and found to be below the detection limit.

Table 1 summarizes the experimentally determined optical properties of tar balls emitted from smoldering combustion of the two fuels described above. The complex refractive indices of the particles were retrieved from Mie “closure” calculations based on particle size, scattering, and absorption measurements. Use of Mie theory is justified by the near-spherical shape of the particles, and a Mie code similar to that of Guyon...
et al., 2003 and Hoffer et al., 2006 was used. Using the method of quadrature sum of
the individual errors, the measurement errors in the real part of the index of refraction
were calculated to be around 4.1% and that in the imaginary part were approximately
13.6%. The absorption Ångström coefficient (AAC) for a pair of wavelengths $\lambda_1$ and $\lambda_2$
is defined as

$$AAC(\lambda_1, \lambda_2) = -\ln\left[\frac{\beta_{\text{abs}}(\lambda_1)}{\beta_{\text{abs}}(\lambda_2)}\right]/\ln\left[\frac{\lambda_1}{\lambda_2}\right],$$

where $\beta_{\text{abs}}$ is the absorption coefficient. The AAC characterizes the wavelength
dependence of light absorption with $AAC = 1$ typical for BC (Reid et al., 2005; Moosmüller
et al., 2009; Lewis et al., 2008).

The spectrally varying AAC of the tar balls indicates that they can be categorized
as brown carbon. Recent studies on humic-like substances (HULIS, related to brown
carbon) and water-soluble OC (WSOC) (Hoffer et al., 2006; Kirchstetter et al., 2004;
Moosmüller et al., 2009) have reported observation of AAC ranging between 6 and 7
for wavelengths between 300 nm and 532 nm. The very high AAC of 6.4, observed
between wavelengths 405 nm and 532 nm for tar balls from AKDuff, suggest the pos-
sibility of HULIS being a primary component. This inference is further strengthened
by the similarities of the complex refractive index of tar balls and HULIS (Hoffer et al.,
2006). In the visible and near-visible range (i.e., 405 nm, 532 nm, and 780 nm), the
single scattering albedo (SSA) (Lewis et al., 2008) of tar balls emitted from smoldering
combustion of PPDuff and AKDuff was fairly high (>0.93), increasing with wavelength –
contrary to the decreasing SSA with increasing wavelength, typically observed for
biomass burning aerosols (Chen et al., 2008). A very recent study also observed con-
sistently higher SSA at 870 nm than at 405 nm for particles emitted from wildfires in
Northern California (Gyawali et al., 2009). SSA is the ratio of scattering to extinction
coefficients, with the extinction coefficient being the sum of scattering and absorption
coefficients (Bohren and Huffman, 1983). For ultra fine and fine (diameter < 1000 nm)
aerosol particles containing mixtures of BC and OC, the wavelength dependence of
their scattering and absorption coefficients are characterized by a scattering Ångström
coefficient (SAC) of $3 < \text{SAC} < 4$ (He et al., 2009) and AAC of $1 < \text{AAC} < 4$ (Bergstrom et

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al., 2007). Therefore, the scattering efficiency generally drops faster with increasing wavelength than the absorption efficiency, explaining why aerosols containing mixtures of BC and OC have been typically observed to have decreasing SSA with increasing wavelength, corresponding to a positive SSA Ångström coefficient (defined similarly to AAC in Eq. 2). For the particles studied here however, their optical properties are dominated by brown carbon yielding an AAC as high as 6.4, significantly larger than their SAC. Therefore, for short wavelengths, the absorption coefficient drops faster with increasing wavelength than the scattering coefficient yielding negative SSA Ångström coefficients as shown in Table 1. For carbonaceous particles, negative SSA Ångström coefficients may become a useful criteria for the definition of brown carbon.

The derived refractive indices of the observed combustion particles are fuel and time dependent, and agree with recently reported results (Guyon et al., 2003; Hoffer et al., 2006; Hungershoefler et al., 2008). The imaginary part of refractive index corresponding to the fuel types in Table 1 can be fitted with a simple function of wavelength \( \lambda \) as:

\[
k(\lambda) = k_{BC} + k_{BrC} = k_{BC} + \frac{a}{(\lambda - \lambda_0)^2},
\]

where \( \lambda_0 \) is an effective resonance wavelength and \( a \) is a constant (see Fig. 4a). The imaginary refractive index data can be described as the sum of a wavelength-independent component \( k_{BC} \) due to small BC volume fraction and a wavelength-dependent component \( k_{BrC} \) due to brown carbon volume fraction. The wavelength-dependent component is modeled with a Lorentzian-like \( 1/(\lambda - \lambda_0)^2 \) function. For wavelengths closer to the effective resonance wavelength, a line width term would need to be added to the denominator in Eq. (3), resulting in a Lorentzian function. In Fig. 4a, note that the assumption of identically wavelength dependent \( k_{BrC} \) for the two Ponderosa Pine Duff data sets still yields excellent fits (fit errors \( \approx 1\% \)) of the data points. Calculating \( k_{BrC} \) by subtracting \( k_{BC} \) from our data (Fig. 4b) makes it obvious that the brown carbon component of the imaginary part of the refractive index is virtually identical for smoke particles from PPDuff1 and PPDuff2 (Fig. 4b). However, the \( k_{BC} \) component of
PPDu2 is less than 2/3 of that of PPDu1. This is probably due to the later sampling time for PPDu2 when additional smoldering emissions and/or additional condensation of semi-volatile organic compounds added more brown carbon to the aerosol, while BC content stayed constant as minor flaming combustion during the ignition of the fire (the source of BC) had already terminated. The effective resonance wavelength of smoke particles from PPDu was found to be further in the UV (i.e., 289 nm) than that of AKDu (i.e., 348 nm).

The knowledge of $k_{BrC}$ allows to estimate the radiative forcing efficiency (RFE; the radiative forcing per unit optical depth and bandwidth) at 405 nm, 532 nm, and 780 nm by brown carbon emitted from PPDu2 and AKDu combustion (Table 2). In order to assess the contribution of absorption by brown carbon, the estimated RFE for brown carbon is compared with that of traditional OC (i.e., complex refractive index = 1.55 + 0i) (Horvath, 1993). Most calculations of RF for traditional OC containing aerosols have assumed that OC does not absorb in the visible and near-visible part of the spectrum. The aerosol RFE was estimated in a similar fashion to Dinar et al. (2008) and Randles et al. (2004). The shortwave aerosol RFE at the top of the atmosphere caused by a uniform, optically thin aerosol layer in the lower troposphere was calculated using

$$\frac{\Delta F}{\tau} = SD(1 - A_{cl}) T_{atm}^2 (1 - R_{sfc})^2 \left[ 2 R_{sfc} \left( 1 - (SSA) \right) \left( 1 - R_{sfc} \right)^2 - \beta (SSA) \right],$$

(4)

where $\Delta F$ is the change in net solar flux at the top of the atmosphere due to the presence of the aerosols, $\tau$ is the aerosol optical depth, $S$ is the solar constant, set to 1370 Wm$^{-2}$, $D$ is the fractional day length, set to 0.5, $A_{cl}$ is the fractional cloud cover, set to 0.6, $T_{atm}$ is the solar atmospheric transmittance, set to 0.76, $R_{sfc}$ is the surface albedo, set to 0.15 (appropriate for an urban area), and SSA is the aerosol single scattering albedo. The parameter $\beta$ is the fraction of scattered sunlight that is scattered into the upward hemisphere, which in turn is a function of hemispheric backscatter fraction, $b$, defined as the ratio of backscattering coefficient to total scattering coefficient. The
parameter $b$ was determined from $g$, the asymmetric parameter, using

$$b = \frac{1 - g^2}{2g} \left[ \frac{1}{\sqrt{1 + g^2}} - \frac{1}{1 + g} \right]. \quad (5)$$

From $b$, the parameter $\beta$ was calculated using the following functional relationship derived from the Henyey–Greenstein phase function (Henyey and Greenstein, 1941) as

$$\beta = 0.082 + 1.85b - 2.97b^2. \quad (6)$$

In all of the calculations, the relative humidity is assumed to be 0%.

In Table 2, positive values of RFE indicate heating and negative values cooling of the atmosphere by aerosols. As would be expected, the RFE of brown carbon is significantly less negative than the RFE of OC at 405 nm, emphasizing the increasing importance of the absorbance by brown carbon near the UV compared to the visible.

4 Conclusion

Inclusion of the properties of tar balls into current RF aerosol models and optical remote sensing retrievals is an important step towards improving model accuracy. Given their abundance in the atmosphere (Alexander et al., 2008; Lukacs et al., 2007), these brown carbonaceous tar balls may account for a significant increase in the RFE (resulting in additional atmospheric absorption and warming) in the near-UV and UV region. Additionally, their presence may help explain anomalously large near–UV and UV solar light absorption observed in the troposphere (Jacobson, 1999).

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Table 1. Summary of the experimentally determined optical properties of tar balls emitted from the combustion of two fuels: Ponderosa Pine Duff (PPDuff, at two sampling times) and Alaskan Duff (AKDuff).

<table>
<thead>
<tr>
<th>Fuel Name</th>
<th>405 nm</th>
<th>532 nm</th>
<th>780 nm</th>
<th>405 nm</th>
<th>532 nm</th>
<th>780 nm</th>
<th>405–532 nm</th>
<th>532–780 nm</th>
<th>405–532 nm</th>
<th>532–780 nm</th>
</tr>
</thead>
<tbody>
<tr>
<td>PPDuff1a</td>
<td>N/A</td>
<td>1543 &amp; 60.4</td>
<td>731 &amp; 21</td>
<td>N/A</td>
<td>1.72 + 0.008i</td>
<td>1.71 + 0.006i</td>
<td>N/A</td>
<td>2.8</td>
<td>N/A</td>
<td>0.027</td>
</tr>
<tr>
<td>PPDuff2b</td>
<td>1807 &amp; 136</td>
<td>1527 &amp; 42.8</td>
<td>792 &amp; 18</td>
<td>1.78 + 0.015i</td>
<td>1.87 + 0.006i</td>
<td>1.86 + 0.004i</td>
<td>4.2</td>
<td>2.3</td>
<td>−0.15</td>
<td>−0.027</td>
</tr>
<tr>
<td>AKDuffb</td>
<td>1150 &amp; 56</td>
<td>907 &amp; 9.7</td>
<td>467 &amp; 3.9</td>
<td>1.83 + 0.0076i</td>
<td>1.75 + 0.0020i</td>
<td>1.60 + 0.0014i</td>
<td>6.4</td>
<td>2.4</td>
<td>−0.15</td>
<td>−0.008</td>
</tr>
</tbody>
</table>

a Data set collected four minutes after ignition of fuel. During sampling, the IPN at 405 nm was not operating.
b Data set collected 30 min after ignition of the fuel.
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Table 2. Comparison of radiative forcing efficiency (i.e., radiative forcing per unit optical depth and bandwidth) at 405 nm, 532 nm, and 780 nm for brown carbon emitted from Ponderosa Pine Duff and Alaskan Duff combustion with that of traditional organic carbon.

<table>
<thead>
<tr>
<th>Fuel Name</th>
<th>$\lambda = 405$ nm</th>
<th>$\lambda = 532$ nm</th>
<th>$\lambda = 780$ nm</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>Refractive Index</td>
<td>RFE (W/nm m$^2$)</td>
<td>Difference (W/nm m$^2$)</td>
</tr>
<tr>
<td>PPDuff2</td>
<td>1.78 + 0.015i</td>
<td>-22.7</td>
<td>+3.3</td>
</tr>
<tr>
<td>OC</td>
<td>1.55 + 0.0i</td>
<td>-26.0</td>
<td></td>
</tr>
<tr>
<td>AKDuff</td>
<td>1.83 + 0.006i</td>
<td>-22.0</td>
<td>+2</td>
</tr>
<tr>
<td>OC</td>
<td>1.55 + 0.0i</td>
<td>-24.0</td>
<td></td>
</tr>
</tbody>
</table>
Fig. 1. Schematic diagram of the experimental setup used for smoldering combustion and characterization of emitted particles.
**Fig. 2.** SEM image of a typical tar ball emitted from smoldering biomass combustion. Electron dispersive spectroscopy (EDX) of tar balls shows that these particles consist primarily of carbon and oxygen with an average molar ratio of about six.
Fig. 3. Number size distribution of tar balls from the smoldering combustion of Ponderosa Pine Duff (PPDuff1 and PPDuff2) and Alaskan Tundra Duff (AKDuff) obtained from ELPI measurements. Each curve represents a four-minute average of second-by-second ELPI retrievals based on spherical, uniform particles and fitted with bi-modal, lognormal size-distribution functions.
Fig. 4. (a) Data (individual points) and models (solid lines) of the imaginary part of the refractive index $k$ as function of wavelength for Ponderosa Pine Smoke 1 and 2 and Alaskan Duff Smoke. The wavelength independent components $k_{BC}$ of $k$ are shown as dotted lines for the three individual smoke samples. (b) Data (individual points) and models (solid lines) of the imaginary part of the refractive index $k$ as function of wavelength for the brown carbon component of Ponderosa Pine Smoke 1 and 2 and Alaskan Duff Smoke. The fact that the data points for Ponderosa Pine Smoke 1 and 2 are nearly identical indicates that differences in the imaginary parts of their refractive indices are predominantly due to different black carbon content.