



1 **Depolarization Ratios Retrieved by AERONET Sun/Sky Radiometer Data**
2 **and Comparison to Depolarization Ratios Measured With Lidar**

3 Youngmin Noh¹, Detlef Müller^{2*}, Kyunghwa Lee³, Kwanchul Kim³, Kwonho Lee⁴

4 ¹ International Environmental Research Center, Gwangju Institute of Science and Technology
5 (GIST), Korea

6 ² University of Hertfordshire, United Kingdom

7 ³ Gwangju Institute of Science and Technology (GIST), Korea

8 ⁴ Gangneung-Wonju National University, Korea

9

10

11

12

13

14

15

16

17

18

19 *Corresponding Author:

20 School of Physics, Astronomy and Mathematics, University of Hertfordshire, Hertfordshire,

21 UK

22 Tel: +44-1707-28-4365

23 E-mail address: d.mueller@herts.ac.uk (Detlef Müller)



24

Abstract

25 The linear particle depolarization ratios at 440, 675, 870, and 1020 nm were derived using
26 data taken with AERONET sun/sky radiometer at Seoul (37.45° N, 126.95° E), Kongju
27 (36.47° N, 127.14° E), Gosan (33.29° N, 126.16° E), and Osaka (34.65° N, 135.59° E). The
28 results are compared to the linear particle depolarization ratio measured by lidar at 532 nm.
29 The correlation coefficient R^2 between the linear particle depolarization ratio derived by
30 AERONET data at 1020 nm and the linear particle depolarization ratio measured with lidar at
31 532 nm is 0.90, 0.92, 0.79, and 0.89 at Seoul, Kongju, Gosan, and Osaka, respectively. A
32 good correlation between the lidar-measured depolarization ratio at 532 nm and the one
33 retrieved by AERONET at 870 nm. We find correlation coefficients R^2 of 0.89, 0.92, 0.76,
34 and 0.88 at Seoul, Kongju, Gosan, and Osaka, respectively. The correlation coefficient for the
35 data at 675 nm is lower than the correlation coefficient at 870 and 1020 nm. We find
36 correlation values of 0.81, 0.90, 0.64, and 0.81 at Seoul, Kongju, Gosan, and Osaka,
37 respectively. The lowest correlation values are found for the AERONET-derived linear
38 particle depolarization ratio at 440 nm. We find values of 0.38, 0.62, 0.26, and 0.28 at Seoul,
39 Kongju, Gosan, and Osaka, respectively. The linear particle depolarization ratio can be used
40 as a parameter to obtain insight into the variation of optical and microphysical properties of
41 dust when it mixed with anthropogenic pollution particles. The single-scattering albedo
42 decreases with increasing measurement wavelength for low linear particle depolarization
43 ratios. In contrast, single-scattering albedo increases with decreasing wavelength for high
44 linear particle depolarization ratios. The retrieved volume particle size distributions are
45 dominated by the fine-mode fraction if linear particle depolarization ratios are less than 0.15
46 at 532 nm. The fine-mode fraction of the size distributions decreases and the coarse-mode



47 fraction of the size distribution increases for increasing the linear particle depolarization ratio
48 at 1020 nm. The dust ratio based on using the linear particle depolarization ratio derived from
49 AERONET data is 0.12 to 0.17 lower than the coarse-mode fraction derived from the volume
50 concentrations of particle size distributions in which case we can compute the coarse-mode
51 fractions of dust.

52

53 **Key words:** linear particle depolarization ratio, lidar, AERONET sun/sky radiometer, dust,
54 single-scattering albedo, size distribution

55

56 **1. Introduction**

57 There are various aerosol types of natural (primarily desert dust and sea salt) and
58 anthropogenic (primarily combustion of biomass and fossil fuels) origin. A precise
59 understanding of the radiative forcing of these aerosol types is the key to quantifying the
60 aerosol impact on regional and global climate change (IPCC, 2013). In order to better
61 estimate the aerosol effect (direct and indirect radiative forcing) on global climate change
62 many studies have been performed to classify aerosol types (Burton et al., 2013; Eck et al.,
63 2010; 1999; Lee et al., 2010a; Dubovik et al, 2002). However, those studies do not separate
64 aerosol types according to their contribution in a plume of mixed aerosol (respectively
65 mixtures of different aerosol types), but merely classify dominant aerosol types based on the
66 optical properties of aerosols.

67 Dust is one of the major aerosol components in the global atmosphere. Dust affects Earth's
68 climate by interacting with solar as well as thermal infrared radiation. Dust also affects
69 atmospheric dynamics, atmospheric chemistry, air quality, and ocean biogeochemistry over a



70 wide range of spatial and temporal scales, e.g., Haywood et al. (2005), Jickells et al. (2005)
71 and Husar et al. (2001). On a global average, dust contributes to about one quarter of aerosol
72 optical depth (τ) in the mid-visible wavelength range (Kinne et al., 2006). Dust is also light-
73 absorbing (Lafon et al., 2006 and 2004; Alfaro et al., 2004; Sokolik and Toon, 1999). It is
74 estimated that more than half of aerosol absorption optical depth at 550 nm may come from
75 dust (Chin et al., 2009).

76 The size distribution and absorption properties of desert dust and other anthropogenic
77 aerosols show properties that can be clearly distinguished (Russel et al., 2010; Dubovik et al.,
78 2002). Desert dust predominately consists of coarse mode particles (typically radius $> \sim 1 \mu\text{m}$).
79 In contrast, combustion-produced particles are predominately found in the fine- mode
80 fraction of particle size distributions (typically radius $< \sim 1 \mu\text{m}$). Aerosols in which fractions
81 of fine-mode and coarse-mode particles are mixed are among the most challenging aerosol
82 types to characterize. If we can separate desert dust from other aerosols in mixed dust plumes,
83 we improve our understanding of the effect of those mixed aerosol plumes on climate change.

84 The linear particle depolarization ratio (δ_p) strongly depends on particle shape. Since dust
85 particles have non-spherical shape, the linear particle depolarization ratio can be used to
86 identify the presence of dust particles in the atmosphere. In that regard lidar is a particularly
87 powerful measurement technique (Tesche et al., 2009; Noh et al., 2008; 2007; Iwasaka et al.,
88 2003; Cairo et al., 1999). The δ_p has also been used to identify biogenic aerosols. Noh et al.
89 (2013a,b) and Sassen et al. (2008) identified the vertical distribution of pollen in the
90 atmosphere using the δ_p measured by lidar. The possibility that dust particles are mixed
91 with other, man-made pollution and/or biomass burning particle is very high except in source
92 regions of dust emissions where population density and thus emissions caused by human



93 activities are comparably low (Sun et al., 2010; 2005; Yu et al., 2006). Mixtures of dust
94 particles and anthropogenic particles cause changes in the δ_p of dust plumes. Shin et al.
95 (2015) reported on lidar measurements and show that the δ_p is decreased when dust is
96 mixed with pollution particles. The δ_p is also a very useful parameter that allows us to
97 separate dust from non-dust particles (Anthropogenic, smoke, and sulphate particles) in
98 mixed dust plumes by retrieving the dust ratio (R_D). Shimizu et al. (2004) estimated the
99 contributions of dust and pollution particles in a mixed-dust plume with the assumption that
100 both aerosol types are externally mixed. The optical data of the mixed dust plumes were
101 separated by the R_D into pure dust content and the anthropogenic particles (Noh et al., 2016a;
102 Bravo-Aranda et al., 2015; Noh 2014; Noh et al., 2012b;). Tesche et al. (2011) separated the
103 optical properties of desert dust and biomass burning particles in mixed dust and smoke
104 plumes over the tropical North Atlantic west of the African continent using multi-wavelength
105 aerosol Raman lidar in combination with polarization lidar. Burton et al. (2014) provides a
106 generalized version of the separation methodology between two aerosol types, urban
107 pollution plus dust, marine plus dust, and smoke plus marine by modifying the methodology
108 suggested by Shimizu et al. (2004) and Tesche et al. (2011). Noh (2014) and Tesche et al.
109 (2011; 2009) used the δ_p to retrieve vertically-resolved single-scattering albedo of mixed
110 dust plumes by separating the contribution of dust and non-dust particles. Ansmann et al.
111 (2011) and Navas-Guzmán et al. (2013) separated the contribution of volcanic ash and
112 sulphate particles to total backscatter and extinction coefficient by the δ_p .



113 However, the numbers of lidar measurement sites are limited and provide only spatially and
114 temporally sparse information on the global scale. Thus we need other methods that could
115 allow us to measure δ_p .

116 AERONET (Aerosol Robotic NETwork) is an automated, robotic Sun-and-sky-scanning
117 measurement network that currently encompasses more than 797 measurement sites
118 worldwide (<http://aeronet.gsfc.nasa.gov/>) which span everything from temporally limited
119 observations at sites at which field campaigns were carried out to sites that carry out long-
120 term observations since AERONET started with its observations. AERONET sun/sky
121 radiometers provide globally distributed observations of spectrally-resolved aerosol optical
122 depth (τ) and data inversion products such as particle size distributions and complex
123 refractive indices of different aerosol types (Holben et al., 1998). Dubovik et al. (2006)
124 suggested to use AERONET sun/sky radiometer data to retrieve the δ_p . Müller et al. (2012;
125 2010) calculated the δ_p of Saharan dust using AERONET sun/sky radiometer data. Noh et
126 al. (2016b) and Lee et al. (2010b) used AERONET sun/sky radiometer data to retrieve the
127 δ_p of Asian dust. However, only cases of nearly pure desert-dust particle were analyzed in
128 these studies. There exist no studies in which the δ_p for various mixtures (mixing ratio)
129 between desert dust and anthropogenic pollution particles using AERONET data has been
130 determined.

131 In this contribution we are trying to verify the reliability of AERONET-derived δ_p by
132 comparing these values to values of δ_p measured by lidar. Section 2 presents the methods
133 used in this study. Section 3 presents our results. We discuss our results and summarize our
134 findings in section 4.



135

136

137 **2. Methodology**

138 **2.1 Study sites**

139 The AERONET sites considered in this study are all located along the pathway of
140 storms (regime of prevailing westerly winds) that serve as major transport routes of Asian
141 dust carried from the arid regions of China and Mongolia. Figure 1 shows the locations of the
142 AERONET sun/sky radiometers and lidar measurement sites used in this study. Seoul (37.45°
143 N, 126.95° E) and Kongju (36.47° N, 127.14° E) are located inland (continental influenced),
144 whereas Gosan (33.29°N, 126.16° E) and Osaka (34.65° N, 135.59° E) are coastal sites.
145 The Gosan site faces the Yellow Sea and is considered an ideal location for monitoring re
146 gional background aerosols in East Asia because there are few local industrial sources
147 in that region. The other three sites are located inside large cities. We also use data from the
148 AERONET site at Dunhuang (40.49° N, 94.95° E) in our study, and we analyzed the
149 depolarization ratios and optical properties of pure Asian dust at this source region.
150 Lidar data are obtained from the lidar network of the National Institute of Environmental
151 Research (NIES), Japan. The lidars operated in this network are two wavelength (1064
152 nm, 532nm) Mie-scattering lidars that measure the linear particle depolarization ratio at 532
153 nm. The details of these lidar systems are explained by Shimizu et al. (2004) and Sugimoto
154 et al. (2008). The locations of the lidar systems used for our research work are the same as the
155 locations of the AERONET systems for the sites in Seoul, Gosan, and Osaka. The lidar used



156 for the Kongju site is located approximately 32 km away from the AERONET site.

157

158 **2.2 Depolarization ratios derived from data taken with AERONET sun/sky radiometer**

159 Dubovik et al. (2006) introduced kernel look-up tables that describe mixtures of spheroid
160 particles. These kernel look-up tables were used to infer the δ_p of mineral dust observed
161 with Sun/sky radiometer. The details of the AERONET inversion algorithm that processes
162 data of mineral dust are given by Dubovik et al. (2006).

163 Briefly, the retrieval of the depolarization ratios works as follows. The elements
164 $F_{11}(\lambda)$ and $F_{22}(\lambda)$ of the Müller scattering matrices (Bohren and Huffman, 1983) are
165 computed from the retrieved complex refractive index and particle size distributions. For
166 unpolarized incident light, $F_{11}(\lambda)$ is proportional to the flux of the scattered light (Volten et
167 al., 2001). The $F_{22}(\lambda)$ in turn follows from the angular and spectral distribution of the
168 radiative intensity which is measured with the AERONET instrument (Dubovik et al., 2006).
169 Another input parameter that is needed for the retrieval of the δ_p is the aspect ratio
170 distribution. The aspect ratio indicates the ratio of a particle's longest axis to its shortest axis.
171 In the case of prolate particles its polar diameter is greater than the equatorial diameter, in
172 contrast to oblate particles where this ratio is vice versa. The aspect ratio distribution is kept
173 to a fixed distribution in the AERONET model since scattering elements are nearly
174 equivalent for all mixtures of spheroid particles (Dubovik et al., 2006).

175 From the ratio of the elements $F_{11}(\lambda)$ and $F_{22}(\lambda)$ at the scattering angle 180°
176 the $\delta_p(\lambda)$ can be computed as



$$\delta_p(\lambda) = \frac{1 - F_{22}(\lambda, 180^\circ) / F_{11}(\lambda, 180^\circ)}{1 + F_{22}(\lambda, 180^\circ) / F_{11}(\lambda, 180^\circ)} \times 100 (\%) \quad (1)$$

177 The δ_p derived from the sun/sky radiometer data is written as δ_p^S in order to
178 distinguish it from the lidar-derived δ_p (δ_p^L).

180 The contributions of dust and anthropogenic pollution particles to the total backscatter
181 coefficients of mixed aerosol plumes were estimated from the δ_p under the assumption that
182 both types of aerosol particles are externally mixed. The dust ratio (R_D) of the dust-related
183 backscatter coefficient to the total backscatter coefficient was calculated using Eq. (1), based
184 on the method suggested by Shimizu et al. (2004):

$$R_D = \frac{(\delta_p^S - \delta_2)(1 + \delta_1)}{(\delta_1 - \delta_2)(1 + \delta_p^S)} \quad (2)$$

186 where δ_1 and δ_2 denote the δ_p^S of pure dust and non-dust particles (i.e. the total aerosol
187 plume without the contribution by dust), respectively, in the external mixture of aerosol
188 particles. The values δ_1 and δ_2 can be empirically determined. In the present study, we used
189 the value 0.34 for δ_1 , which was derived from adding 0.01 to the maximum value observed at
190 the Dunhuang site (Asian dust source region). The value of 0.02 was used for δ_2 , which is the
191 minimum value used in this study. When δ_p^S was higher than δ_1 or lower than δ_2 , R_D was set
192 to 1 or 0, respectively.

193 Two kinds of coarse-mode fraction (CMF) were calculated. The coarse-mode fraction of the
194 aerosol optical depth (τ) (CMF τ) is calculated from the ratio of the coarse-mode τ to the total
195 (coarse + fine mode) τ at the same wavelength at which the δ_p^S is available. The coarse-
196 mode fraction is also calculated on the basis of the volume concentration (CMF $_{vc}$).

197



198

199 **2.3 Column-integrated depolarization ratio measured by LIDAR**

200 The lidar systems used in our study measure the linear volume depolarization ratio
201 (aerosols + molecules; δ^L) from the linearly and perpendicularly polarized components of
202 the Mie/Rayleigh backscatter signals at 532 nm wavelength (Sakai et al., 2000). The value of
203 δ^L is defined as

$$204 \quad \delta^L(z) = \frac{P_{\perp}(z)}{P_{\parallel}(z) + P_{\perp}(z)} \times 100 (\%) = \frac{\beta_{p,\perp}(z) + \beta_{m,\perp}(z)}{\beta_{p,\parallel}(z) + \beta_{m,\parallel}(z) + \beta_{p,\perp}(z) + \beta_{m,\perp}(z)} \times 100 (\%) \quad (3)$$

205 where $P(z)$ is the backscatter signal with respect to height z ; β_p and β_m are the volume
206 backscatter coefficients of aerosol particles and air molecules. The symbols \parallel and \perp
207 denote the linearly and perpendicularly polarized components with respect to the plane of
208 polarization of the emitted light, respectively.

209 The δ_p^L differs from δ^L as it depends on the concentration of particles without
210 taking account of the contribution (concentration) of air molecules. In this contribution, δ_p^L
211 can be calculated according to the definition by Sakai et al. (2000)

$$212 \quad \delta_p^L(z) = \frac{\beta_{p,\perp}(z)}{\beta_{p,\parallel}(z) + \beta_{p,\perp}(z)} \times 100 (\%) = \frac{\delta^L(z)R(z) - \delta_m^L}{R(z) - 1} (\%). \quad (4)$$

213 The backscatter ratio R is the ratio of the sum of the aerosol backscatter coefficient
214 ($\beta_p + \beta_m$) to the pure molecular backscatter coefficient (β_m), which, according to Whiteman
215 et al. (1992) can be expressed by



216
$$R(z) = \frac{\beta_P(z) + \beta_M(z)}{\beta_M(z)} \quad (5)$$

217 The molecular depolarization ratio (δ_M) is assumed to be 0.0044 (Behrendt and Nakamura,
218 2002).

219 The parameter δ_P^L can be derived by lidar measurements in terms of vertical profiles. In
220 contrast, δ_P^S describes a column-integrated value. For that reason, δ_P^L had to be changed to
221 column-integrated values in our study, in order to allow for a direct comparison with δ_P^S .

222 The column integrated weighted δ_P^L (δ_P^{CL}) can be calculated with Eq. (6).

223
$$\delta_P^{CL} = \int_0^z \delta_P^L(z) W(z) dz \quad (6)$$

224 where the term $W(z)$ is a weight factor that is calculated on the basis of the measured aerosol
225 backscatter coefficient (β_P) according to the following Eq. (7):

226
$$W(z) = \frac{\beta_P(z)}{\int_0^z \beta_P(z) dz} \quad (7)$$

227 Figure 2 shows a retrieval example of δ_P^{CL} . Three cases, corresponding to lidar
228 measurements carried out at 23:00 UTC (start time of measurement) on 13 March 2010 (a), at
229 06:00 UTC on 22 March 2010 (b), and at 23:15 on 3 May 2010 (c) are shown. The
230 measurement on 13 March 2010 describes an aerosol plume that has a high value of δ_P^L and
231 a high β_P (case 1, Fig. 2 (a)). The measurement on 22 March 2010 describes an aerosol plume
232 with high β_P below and above the planetary boundary layer (PBL), but a high value of δ_P^L
233 detected only above the PBL (case 2, Fig. 2 (b)). The measurement on 2 May 2010 describes
234 an aerosol plume with low backscatter coefficient and a high value of δ_P^L (case 3, Fig. 2 (c)).



235 The averaged values of δ_p^L (δ_{ave}^L) which do not consider the backscatter intensity but just the
236 averaged δ_p^L from bottom to the top height of the profiles of δ_p^L are also listed in Figure 2.
237 The values of δ_p^S at 1020 nm are 0.250, 0.140, and 0.164 for the cases 1, 2, and 3,
238 respectively. The values of δ_p^{CL} are similar to the values of δ_p^S which are 0.243, 0.129, and
239 0.157 for the cases 1, 2, and 3, respectively. However, the values of δ_{ave}^L in case 1 and 3 are
240 different compared to δ_p^S . Since δ_p^L is not directly related to aerosol concentration but only
241 to the non-sphericity of (an ensemble of) aerosol particles (in a given volume of air) large
242 values of δ_p^L can occur for high as well as for low aerosol backscatter coefficients. High
243 values of δ_p^L for the situation of high aerosol backscatter coefficients of a thin aerosol layer
244 (Figure 2 (a)) show lower values of δ_{ave}^L compared to δ_p^S . We find that δ_{ave}^L are higher than
245 δ_p^S for the situation of a high value δ_p^L in combination with a low aerosol backscatter
246 coefficient (Figure 2 (3)). Those examples in Figure 2 explain that δ_p^{CL} has to be compared
247 with δ_p^S .

248

249 3. Results and Discussion

250 3.1. Comparison with δ_p^{CL}

251 Figure 3 shows the temporal variation of τ at 500 nm and the values of δ_p^S at 1020
252 nm at the four AERONET sites. The combined data of the four sites are shown in the same
253 figure. The number of measurement cases for the four sites are listed in Table 1. The total
254 number of retrieved values of δ_p^S is 163, 44, 139, and 234 at Seoul, Kongju, Gosan, and
255 Osaka, respectively. Since the measurement cases are limited it is hard to analyze seasonal



256 trends. However, we find high values of τ and δ_p^S in spring (March to May). We assume that
257 these high values are caused by transport of dust from East Asia to the Pacific Ocean.

258 Figure 4 presents scatterplots of δ_p^{CL} and δ_p^S at the four AERONET sites. We compare the
259 values of the δ_p^S at these four wavelengths (440, 675, 870, and 1020 nm) to the δ_p^{CL} at 532
260 nm. The correlation coefficients R^2 at 1020 nm are high. We find 0.90, 0.92, 0.79, and 0.89 at
261 Seoul, Kongju, Gosan, and Osaka, respectively.

262 We find similarly high correlation between δ_p^{CL} and δ_p^S at 870 nm, i.e. numbers are 0.89,
263 0.92, 0.76, and 0.88 at Seoul, Kongju, Gosan, and Osaka, respectively. The correlation at 675
264 nm is lower compared to the values we find at 870 and 1020 nm. Values are 0.81, 0.90, 0.64,
265 and 0.81 at Seoul, Kongju, Gosan, and Osaka, respectively. The correlation is comparably
266 low at 440 nm. Values are 0.38, 0.62, 0.26, and 0.28 at Seoul, Kongju, Gosan, and Osaka,
267 respectively. The correlation at 440 nm at Kongju is much higher than at the other sites. This
268 higher correlation may be caused by the limited number of observational data and/or
269 observation time. Only 44 cases were taken during a short period of two months, from April
270 to May 2012 at Kongju.

271 Figure 4 shows that the differences between δ_p^{CL} and δ_p^S are high when the δ_p^{CL} is less
272 than 0.10 at Seoul, Gosan and Osaka. However, the number of cases of low δ_p^{CL} (< 0.10) is
273 comparably low at Kongju compared to the other sites. The number of cases with high δ_p^{CL}
274 (< 0.25) is comparably high (with respect to all cases) compared to what we find for the
275 other sites (see Table 1).

276 We classified the observational data into 6 groups based on the values of δ_p^S at 1020 nm.
277 Group 1 contains values of less than 0.05 of δ_p^S at 1020 nm. Groups 2, 3, 4, and 5 include



278 values between 0.05-0.1, 0.1-0.15, 0.15-0.2, and 0.2-0.25 at 1020 nm, respectively. Values
279 are above 0.25 in group 6.

280 Table 1 shows the number of data sets (observation cases) for each of the 6 data groups at the
281 four observation sites. The averaged τ at 500 nm, the optical-depth-related Ångström
282 exponent (\AA) between 440 to 870 nm and the light-absorption-related Ångström exponent (\AA_A ,
283 440 – 870 nm) of each group are also listed in Table 1. The values of τ are similar in all six
284 groups. Reason for that is because τ is insensitive to the shape and size of the particles. The
285 values of \AA_A increase but values of \AA decrease with increasing value of δ_p^S .

286 Figure 5 shows the values δ_p^S of the six groups at the four measurement wavelengths of the
287 AERONET sun/sky radiometers. There is a rather clear increase of δ_p^S with respect to
288 increasing measurement wavelength in group 6. We see a similar pattern in groups 4 and 5,
289 respectively. In contrast, groups 1 - 3 show the highest values of δ_p^S at 440 nm whereas the
290 values of δ_p^S are similar at the other three measurement wavelengths.

291 Values of δ_p^L of pure mineral dust plumes were measured at three wavelengths (355, 532,
292 and 1064 nm) with lidar (Freudenthaler et al., 2009) during the Saharan Mineral Dust
293 Experiment (SAMUM) in 2006. Freudenthaler et al. (2009) found values of 0.31 of δ_p^L at
294 532 nm. Müller et al. (2010; 2012) compared those data with data derived from collocated
295 AERONET sun/sky radiometer observations, see Figure 3 in Müller et al. (2010) and Figure
296 7 in Müller et al. (2012). Values of δ_p from both instruments agree at 1064-nm wavelength.

297 If the sun/sky radiometer results are extrapolated to the lidar wavelength of 355 nm, the δ_p
298 from the sun/sky radiometer is 20 % lower than the value obtained from the lidar
299 observations.



300 Müller et al. (2010) find that the values of δ_p inferred from the sun/sky radiometer
301 observations tend to be lower than the values measured with lidar for the case of pure mineral
302 dust. Only group 6 (Kongju and Gosan) shows 20% - 30 % lower values of δ_p^S compared to
303 δ_p^{CL} in the visible wavelength range, which is a similar trend reported by Müller et al. (2010).
304 This feature, i.e. that δ_p^{CL} is higher than δ_p^S is also found in group 5. However, the
305 differences of the numbers are less compared to the differences we find for group 6. The
306 values of δ_p^S increase from group 1 to group 6 and the values of δ_p^S are becoming more
307 and more similar to the values of δ_p^{CL} from group 1 to group 6.
308 Figure 5 shows the δ_p^{CL} at 532 nm of each group. We find 0.27 ± 0.02 and 0.27 ± 0.03 at
309 Kongju and Gosan, respectively, for group 6. The values at the Osaka site are lower. We find
310 0.21 ± 0.04 . The highest values of the δ_p^{CL} is 0.29 at the Kongju and Gosan sites, see Figure
311 4. The highest value of δ_p^{CL} for the Osaka site is 0.27.
312 The differences of the δ_p at the observation sites likely are caused by the appearance of dust.
313 The transport distance to the four observation sites may have influence on the values of δ_p .
314 Kongju and Gosan have similar transport distances from the source regions of East Asian
315 dust. Osaka is located in a distance of 1 or 2 days of transport time from Kongju and Gosan.
316 It means that more dust particles can be removed by gravity sedimentation during transport
317 (Maring et al., 2003; Gong et al., 2003). Another reason may be that more anthropogenic
318 pollution particles are mixed into these Asian dust plume because of the longer transport time
319 (Kanayama et al., 2002; Noh et al., 2014; Shin et al., 2015).
320 Figure 6 shows the averaged values of the vertically resolved β_p , the values of δ_p^L , and the



321 weighted δ_p^L ($\delta_p^L W$) measured by lidar for each group. The backscatter coefficients generally
322 decrease with height. However, high values are found in the upper atmosphere in those cases
323 in which δ_p^L is high. Values of δ_p^L are as high as 0.1 in group 1 and increase with
324 increasing group number (from group 1 to group 6). The increase of the δ_p^L is more obvious
325 above 2-km height than below 2-km height above ground. The higher values above 2-km
326 above ground may be caused by the fact that Asian dust has a relatively lower chance of
327 being mixed with other pollutants if it is transported in the upper parts of the atmosphere
328 (Shin et al., 2015). Especially, group 5 and group 6 show high values of δ_p^L , i.e. larger than
329 0.3 above 2 km height.

330 We see that the β_p shows different trends in these two groups. Group 5 has low values of β_p
331 and high values of δ_p^L . In contrast group 6 shows high values of β_p and high values of δ_p^L .
332 This different behavior in these two groups is clearly visible in the values of $\delta_p^L W$. Values of
333 $\delta_p^L W$ in group 5 are less than 2 throughout the whole altitude range. Values of $\delta_p^L W$ are lar
334 ger than 2 in group 6.

335 We find that δ_p^S , especially at 1020 nm, is rather similar to δ_p^{CL} and the values of δ_p^S at
336 1020 nm are large for high dust concentrations and small for low dust concentrations, see
337 Figures 4, 5, and 6. Thus we think that δ_p^S can be a reliable information for identifying the
338 presence of Asian dust particles in East Asian pollution plumes. This means that δ_p^S can be
339 used to retrieve the dust ratio in mixed dust plumes even if vertically-resolved information on
340 the linear particle depolarization ratio is not available.

341



342 **3. 2. Correlation between δ_p^S and single-scattering albedo**

343 One main purpose of our study is to estimate the mixing ratio of dust particles with other
344 pollutants in the atmosphere by analyzing δ_p^S . Another purpose of our study is to estimate on
345 the basis of δ_p^S the variation of the optical and microphysical properties of dust when it
346 mixes with anthropogenic pollution particles. Correlations between δ_p^S and other optical
347 parameters allow us to gain insight into these variations.

348 Variations of aerosol absorption properties can be described by the single-scattering albedo
349 (SSA). The knowledge of the variability of light-absorption of aerosol mixtures discussed in
350 this contribution allows us to assess the direct forcing of mixed-dust plumes. We can also
351 investigate the semi-direct forcing that may occur from atmospheric heating by absorbing
352 aerosol layers (Noh et al., 2012b; 2016b; Noh, 2014). We also investigate how SSA varies
353 with the volume particle size distribution that is corresponding to SSA. For that purpose we
354 use the values of δ_p^S .

355 Figure 7 depicts SSA and the volume particle size distribution for each of the 6 groups. The
356 SSA spectra vary with changing δ_p^S in clearly distinguishable patterns. The SSA spectra of
357 group 1 (low δ_p^S) show decreasing SSA with increasing wavelength. We find that SSA
358 decreases with increasing measurement wavelength for particle plumes that are dominated by
359 urban-industrial and biomass-burning particles (Dubovik et al., 2002; Giles et al., 2012).
360 Black carbon particles have the strongest light-absorption capacity in the near-infrared
361 wavelength region.

362 In contrast, the SSA spectra of group 6 (high δ_p^S) show an increase of SSA with increasing
363 wavelength. The wavelength dependence (i.e. increasing, decreasing, or constant with



364 wavelength) of SSA is an important property that is used in aerosol type classification
365 because the spectral absorption characteristics depend on aerosol type (Giles et al., 2012;
366 Russell et al., 2010; Eck et al., 2010; 2005; Dubovik et al., 2002).

367 The increase of SSA with increasing wavelength is a characteristic optical feature of desert
368 dust particles (Giles et al., 2012). Dust particles are aggregates of combinations of clay,
369 quartz, and hematite in variable concentration. Dust exhibits strong light-absorption in the
370 UV and at short visible wavelengths (e.g., 440 nm) and lower light-absorption from mid-
371 visible to near infrared wavelengths (Sokolik and Toon, 1999).

372 Kim et al. (2011) define particles with $\alpha < 0.2$ as “pure dust” based on observations of dust
373 particles over North Africa and the Arabian Peninsula. The average value of SSA of the “pure
374 dust” part of the aerosol plumes observed at the four observation sites is 0.91, 0.97, 0.97, and
375 0.97 at 440, 675, 870, and 1020 nm wavelength, respectively.

376 The SSA spectra of Group 6 at Kongju, Gosan and Osaka resemble the SSA spectra of “pure
377 dust” described by Kim et al. (2011). The SSA of group 6 at Kongju and Gosan show similar
378 values reported by Kim et al. (2011), i.e. 0.94, 0.98, 0.98, and 0.99 at 440, 675, 870, and 1020
379 nm wavelength, respectively. Lower values of SSA are observed at Osaka. We find 0.88, 0.95,
380 0.96, and 0.95 at 440, 675, 870, and 1020 nm wavelength, respectively. The differences of
381 SSA at Osaka may be caused by the mixing of pollution particles with dust.

382 Except for the SSA at 440 nm, the SSA at 675, 870, and 1020 nm show higher values for
383 high δ_p^S at each wavelength. This increase of SSA with increasing δ_p^S results from the
384 mixing of fine-mode pollution particles and coarse-mode Asian dust.

385 Mixtures of desert dust and pollution aerosols contain two primary particulate light-absorbing
386 species, black carbon in fine-mode particles (Bond and Bergstrom, 2006) and iron oxides in



387 coarse-mode dust particles (Sokolik and Toon, 1999). Iron oxides cause strong light-
388 absorption in the UV and in the short-wavelength range of the visible spectral range of light
389 (Derimian et al., 2008). In pollution particles the principal absorber is soot or black carbon
390 which exhibits light-absorption throughout the entire solar spectrum due to an imaginary part
391 of the complex refractive index that is spectrally relatively constant (Bergstrom et al., 2002).
392 The SSA as a function of δ_p^S shows very little variation at 440 nm compared to the other
393 wavelengths considered in our study (675, 870, and 1020 nm) at the Seoul and Gosan sites.
394 Eck et al. (2010) suggest that this kind of restricted SSA-values at 440 nm is induced when
395 both, coarse-mode-aerosol dominated mixtures (desert dust) and fine-mode-dominated
396 aerosol mixtures (pollution) have relatively similar magnitudes of light-absorption with
397 regard to the light-scattering at that wavelength.

398 The lower (440 nm) and similar SSAs (675, 870, and 1020 nm) in group 6 compared to the
399 SSAs in group 5 are shown for the Osaka site. High concentrations of fine-mode particles
400 with strong light-absorbing property can cause a decreasing SSA at the four wavelengths 440,
401 675, 870, and 1020 nm.

402 The variations of the fine-mode and the coarse-mode part of the size distributions are clearly
403 shown in Figure 7. Group 1 contains fine-mode dominated particle size distributions. The
404 fine-mode part of the particle size distributions decreases and the coarse-mode part of the size
405 distributions increases for increasing δ_p^S . The variations of C_{v_c}/C_{v_f} for each group in Table 2
406 show these tendencies more clearly.

407 The values of C_{v_c}/C_{v_f} increase as we move from group 1 to 6. The values of C_{v_c}/C_{v_f} in group
408 1 are similar at Seoul, Kongju, Gosan, and Osaka. We find 0.29, 0.32, 0.35, and 0.28,
409 respectively. The values of C_{v_c}/C_{v_f} are similar for all three observation sites in each of the



410 groups 1 - 5.

411 However, in the case of group 6, C_{vc}/C_{vf} is 7.6, 11.3, and 8.6 at Kongju, Gosan, and Osaka,
412 respectively. The values of C_{vc}/C_{vf} in group 6 are smaller than the average value at
413 Dunhuang. The average value of C_{vc}/C_{vf} at Dunhuang is 15.0 ± 2.6 , see Table 4. C_{vc}/C_{vf}
414 decreases if fine-mode pollution particles are mixed into a dust plume and/or if coarse-mode
415 dust particles are removed from the plume during long-range transport.

416

417 3. 3. δ_p^S , SSA, and particle size distribution at the dust source region

418 We analyzed the AERONET sun/sky radiometer data taken in the source region of Asian dust
419 and evaluated the optical properties and δ_p^S of “pure Asian dust”. Figure 8 shows δ_p^S , SSA
420 and volume particle size distributions observed on 5 days in one of the source regions of
421 Asian dust, i.e. Dunhuang in 2012. Values of δ_p^S at 440 nm are characteristic of pure dust
422 particles, i.e. we find values larger than 0.25. SSA increases with increasing measurement
423 wavelength which is also characteristic of pure dust. We find this behavior on all days except
424 for the data representing 8 April 2012. We find that the SSA at each wavelength (at the
425 Dunhuang site) is higher than the SSA retrieved at the corresponding wavelengths for sites in
426 North Africa and the Arabian Peninsula (Kim et al., 2011; Müller et al., 2010).

427 Table 3 lists τ , δ_p^S at 1020 nm, and \hat{a} , \hat{a}_A , CMF_{vc} and R_D . The values of \hat{a} observed at
428 Dunhuang on all measurement dates used in this study are such that they can be considered as
429 representing pure dust, as suggested by Kim et al. (2011). The exception is the measurement
430 on 8 April. On that day we find $\delta_p^S = 0.25$ at 1020 nm and $\hat{a}_A = 1.76$, which are comparably
431 lower values than the values retrieved for the other observation days. In addition, volume
432 particle size distributions retrieved for 8 April (see Figure 8) show a higher peak modal



433 volume radius for the coarse-mode size distribution compared to what is typically found for
434 size distributions of desert dust (Dubovik et al. 2002; Müller et al., 2010; 2012). We conclude
435 that particles observed on 8 April describe a mixed-dust plume rather than a pure dust plume.
436 We compare the average values of δ_p^S , SSA, and the volume particle size distributions
437 observed at Dunhuang with the respective values of group 6 at Kongju, Gosan, and Osaka.
438 We exclude the data taken on 8 April at Dunhuang for the calculation of the average values
439 because that data likely do not represent pure Asian dust. Figure 9 shows our comparison
440 results.

441 The highest values of δ_p^S at Dunhuang are 0.26, 0.28, 0.30, and 0.33 at 440, 675, 870, and
442 1020 nm wavelength, respectively. The spectral behavior of δ_p^S (at the four wavelengths) at
443 Kongju, Gosan, and Osaka is similar to the spectral behavior of δ_p^S retrieved for the
444 Dunhuang site. However, the values of δ_p^S at the four measurement wavelengths at Kongju,
445 Gosan, and Osaka are 0.04 - 0.05 lower than the respective values at the Dunhuang site. This
446 difference between Dunhuang and the other three sites may be caused by gravitational
447 settling of coarse mode dust particles during transport and/or a higher share of anthropogenic
448 pollution particles that may enter the dust plume during long-range transport from the source
449 region to the other three sites.

450 The volume particle size distributions shown in Figure 9 (c) corroborate our assumption. The
451 volume concentration of the coarse mode particles is as low as 0.36, 0.44, and 0.33 ($\mu\text{m}^3/\mu\text{m}^2$)
452 in the far-field sites of Kongju, Gosan, and Osaka, respectively. We find that among all days
453 during which we observed pure Asian dust at Dunhuang, the minimum value of 0.49
454 ($\mu\text{m}^3/\mu\text{m}^2$) was found on 9 April 2012, see Figure 9 (c) and Table 4.

455 Figure 9 (b) shows a comparison of the spectral SSA between Dunhuang and the other three



456 sites. The average values of SSA at Dunhuang are 0.94, 0.98, 0.98, and 0.98 at 440, 675, 870,
457 and 1020 nm, respectively. The value of 0.94 for SSA at 440 nm at the Dunhuang site is
458 higher than the SSA of “pure dust” observed over North Africa and the Arabian Peninsula
459 (Kim et al., 2011; Müller et al., 2011; Müller et al., 2010).

460 As noticed previously, iron oxides cause the strongest light-absorption in the ultraviolet and
461 at visible wavelengths (Derimian et al., 2008). We assume that the differences of SSA at 440
462 nm between the Dunhuang site and observation sites in North Africa and the Arabian
463 Peninsula are caused by differences of the chemical composition of dust particles, as for
464 example the concentration of iron oxides in the dust particles at these different sites.

465 Our results show that nearly-pure dust may be transported to Kongju and Gosan from long
466 distances, and that the coarse mode fraction of the particle size distribution may not
467 necessarily increase in that case. The spectral behavior of SSA and its values at each
468 wavelength at the Kongju and Gosan sites which are long-range transport sites match with the
469 values of SSA at Dunhuang. This match suggests that there may be a similar chemical
470 composition of the dust, and perhaps similar concentrations of iron oxide in the dust observed
471 at Dunhuang, Kongju, and Gosan.

472 SSA values at Osaka are lower than those at Kongju and Gosan. We find values of 0.88, 0.95,
473 0.96, and 0.95 at 440, 675, 870, and 1020 nm, respectively. These low SSAs can be caused
474 by the mixing of pollution particles. We first investigate the vertical distribution of the dust
475 plumes of group 6 at Osaka. The vertical distribution of particles in the dust plumes of group
476 6 at Osaka can be clearly distinguished according to their observation date.

477 Figure 10 shows the separated values of β_p , δ_p^L , and $\delta_p^L W$ (dust and non-dust contribution)
478 in terms of case 1 and 2 for group 6 at Osaka according to the observation date. The data



479 taken on 14 and 15 March 2010 (case 1) show that the dust plumes are distributed up to 3.5
480 km height above ground. The main part of the dust plumes is located between 1 and 2 km
481 height on both days. The values of δ_p^L are above 0.25 in that height range. The β_p -values of
482 the dust plume are lower below 1 km height compared to what we find for β above 1 km
483 above ground. The value of δ_p^L varies between 0.1 and 0.12 and thus is lower compared to
484 what we find between 1 and 2 km height.

485 The dust plumes extend to 2 km height above ground on 2 May 2011 (measurement case 2).
486 We find high values of β_p near the surface. Values of 0.3 – 0.4 for δ_p^L are higher than
487 depolarization ratios found for case 1.

488 Case 1 is quite different from case 2 if we look at $\delta_p^L W$. The values of $\delta_p^L W$ of case 1 are
489 mainly affected by Asian dust that is present above 1 km height above ground. Values for
490 case 2 are mainly influenced by Asian dust near the surface.

491 These differences of the height above ground of the main portions of the dust plumes may be
492 one reason why the light-absorption capacity of the two cases differs. There is a higher
493 possibility that pollutants can be mixed during long-range transport in case 2.

494 Figure 11 corroborates our assumption. Figure 11 shows values of δ_p^S , SSA, and volume
495 particle size distributions for two cases. The values of δ_p^S and SSA in case 1 show a similar
496 values compared to the case of pure dust observed at Dunhuang. Values of these parameters
497 are different from those at Dunhuang for case 2. Particularly, the values of SSA of case 2 are
498 significantly lower than SSA-values at Dunhuang. Moreover, SSAs at 1020 nm are lower
499 than those at 870 nm for case 2. This difference of absolute values and the differences of the
500 spectral behavior may be caused by the mixing of dust with light-absorbing pollutants, such



501 as black carbon, when we take account of the vertical distribution of the dust plume; see
502 Figure 10.

503 The differences of \hat{a}_A between case 1 and case 2 (Table 5) support this observation. \hat{a}_A in case
504 1 is 2.12 which points to pure dust. In contrast, \hat{a}_A in case 2 is 1.58 which is more likely
505 representative of a mixture of dust with pollutants (Russell et al., 2010).

506 Figure 12 shows backward trajectories. The transport pattern is different for case 1 and case 2.
507 The main portions of the dust plumes of case 1 passed over source regions of major pollution
508 emissions in China and Korea at heights above 1.5 – 2 km, i.e., above the planetary boundary
509 layer, before the plumes arrived over Osaka. Since most of the pollution resides in the PBL
510 the possibility is low that the main dust layer mixed with pollutants. In contrast to case 1,
511 case 2 indicates that pollutants were mixed into the Asian dust layers while they were
512 transported near the surface and within the PBL over industrialized areas and before they
513 arrived over Osaka. Shin et al. (2015) reported that more pollution particles can be mixed into
514 dust plumes if these plumes are transported at low altitude above ground.

515 Table 5 shows that C_{v_c}/C_{v_f} for case 2 is 11.8 which is higher than the value of 6.8 for case 1.

516 The increase of the fine-mode particle concentration may not be the only reason that can
517 cause this decrease. Other reasons that may contribute to a decrease of C_{v_c}/C_{v_f} are for
518 example the coating of coarse-mode dust particles by absorbing fine-mode pollution particles.

519 There is considerable evidence of the coating of dust particles by absorbing fine-mode
520 pollution particles in the East Asian region. Respective observations were made during the
521 ACE-Asia campaign in spring 2001 (Huebert et al., 2003; Kim et al., 2004). Arimoto et al.
522 (2006) also present scanning electron microscopy images showing black carbon particles
523 adhering onto the surface of coarse-mode dust, with typically 15% - 30% of the dust surface



524 coated by black carbon, which likely increases the absorption.

525

526 **3. 4. Retrieval of Dust ratio using sun/sky radiometer derived Depolarization ratio**

527

528 The concept of separating lidar backscatter signals caused by the contribution of dust
529 particles in mixed dust plumes to the total backscatter signals using δ_p has already been
530 applied to lidar measurements (Shimizu et al., 2004; Noh, 2014; Tesche et al., 2009). With
531 regard to AERONET sun/sky radiometer data R_D is calculated from Eq. (2) (see section 2)
532 and the use of δ_p^S at 1020 nm.

533 The retrieved R_D is compared with CMF_{vc} . The comparison between R_D retrieved from
534 AERONET data and CMF_{vc} allows us to distinguish between non-dust coarse-mode particles
535 and dust.

536 Figure 13 shows the correlation between R_D and CMF_{vc} in terms of R^2 . We find a comparably
537 high correlation between R_D and CMF_{vc} . Values are 0.72, 0.95, 0.77, and 0.93 at Seoul,
538 Kongju, Gosan, and Osaka, respectively.

539 R_D describes the ratio of dust particles to other types of non-spherical aerosols in the
540 atmosphere. Unlike R_D , CMF_{vc} considers the size of particles and is uncorrelated to the shape
541 of the particles. Since most of the dust particles belong to the coarse-mode fraction, the value
542 of CMF_{vc} increases alongside with the dust ratio. However, the coarse-mode of a particle size
543 distribution does not include dust only but also contains large particles that are generated by
544 physical and chemical reactions, e.g. coagulation, condensation processes, and hygroscopic
545 growth.

546 Figure 13 shows that CMF_{vc} is on average 0.12-0.17 higher than R_D at the four observation



547 sites, which implies that non-dust particles are present, i.e., even though dust particles for the
548 most part belong to the coarse-mode fraction of particle size distributions, not all coarse-
549 mode particles are dust particles. The average values of CMF_{vc} and the differences between
550 R_D and CMF_{vc} for each of the 6 groups are listed in Table 1.
551 R_D of groups 1 – 5 is 0.11 - 0.19 lower than CMF_{vc} at the four sites. The differences between
552 R_D and CMF_{vc} are lower for group 6. We find 0.4, 0.4, and 0.3 at Kongju, Gosan, and Osaka,
553 respectively.
554 We find a similar difference for the Dunhuang site, see Table 3. The average values of R_D and
555 CMF_{vc} are 0.97 ± 0.02 and 0.94 ± 0.01 , respectively, except for the case of 8 April. This
556 means that most of the coarse-mode particles are composed of dust particles. However,
557 pollution and/or biomass burning particles can be injected into the dust plume during
558 transport from the source region. These particles may contribute to the coarse mode of the
559 volume particle size distribution. If the dust plume is transported at low altitude above ground,
560 there is an increased possibility that dust mixes with other aerosols (Shin et al., 2015).
561 This increased possibility is corroborated by the results for the case of Osaka in group 6.
562 Table 5 shows for this case that δ_p^S , SSA, and the particle size distribution clearly depend on
563 the altitude of the dust plume. R_D and CMF_{vc} show very similar values with respect to case 1.
564 The average value of the difference between R_D and CMF_{vc} is 0.004 ± 0.008 for the case that
565 the main part of the dust plume is transported above the planetary boundary layer (case 1).
566 This difference between R_D and CMF_{vc} increases to 0.11 ± 0.04 when the dust plume is
567 transported near the surface, as can be seen from $\delta_p^L W$ of case 2. From these results we can
568 confirm that the coarse mode is mostly composed of pure dust particles without that mixing
569 with other types of particle has occurred during transport above the PBL. However, the ratio



570 of dust particles to non-dust coarse-mode particles decreases as the result of mixing processes
571 during transport. These phenomena can be explained by the variation of the volume median
572 radius (R_v).

573 The volume median radius of the coarse mode (R_{v_c}) of 3.21-3.27 μm for urban-industrial
574 aerosols generated by fossil fuel combustion and biomass burning aerosol produced by forest
575 and grassland fires is higher than R_{v_c} of 2.62-3.05 μm of desert dust (Dubovik et al., 2002;
576 Eck et al., 2010). The average R_{v_c} decreases as δ_p^S increases. In this study, R_{v_c} shows low
577 values when the observed particles are nearly pure dust. The average R_{v_c} at Dunhuang is 1.88
578 μm (except 8 April 2012). This value is quite similar to 2 μm of R_{v_c} for dust that originated
579 from China and was measured over Japan (Tanaka et al., 1989). The average R_{v_c} at Osaka
580 decreases from 2.85 to 2.46 to 2.20 to 1.94 to 2.08 to 1.85 μm between group 1 and group 6,
581 respectively. The other sites show the same pattern of decreasing values between groups 1 to
582 6. The average R_{v_c} of 2.15 at Kongju and 1.77 μm at Gosan (group 6) also shows similar
583 values to the values found in the dust source region. The average R_{v_c} of the other groups is
584 higher than R_{v_c} of dust in the source region. We find the highest values in group 1, i.e. 3.05,
585 2.72, 2.58, and 2.85 μm at Seoul, Kongju, Gosan, and Osaka, respectively.

586

587 **4. Summary and Conclusion**

588 In this study we attempt to verify the reliability of the AERONET sun/sky radiometer derived
589 linear particle depolarization ratio (δ_p^S) that can be used for detecting dust particles by
590 comparing this parameter to the linear particle depolarization ratios measured by lidar (δ_p^L).
591 We considered low (cases dominated by pollution particles) to high linear particle
592 depolarization ratios (cases dominated by Asian dust) at four downwind regions (Seoul,



593 Kongju, Gosan and Osaka) of Asian dust. We calculated the column-integrated weighted δ_p^L
594 (δ_p^{CL}) and compared these values with δ_p^S . The strongest correlation exists between δ_p^S at
595 1020 nm and δ_p^{CL} . Values are 0.90, 0.92, 0.79, and 0.89 for the sites at Seoul, Kongju, Gosan,
596 and Osaka, respectively. A good correlation was also observed at 870 nm. We find values of
597 0.89, 0.92, 0.76, and 0.88 at Seoul, Kongju, Gosan, and Osaka, respectively. Although the
598 correlation at 675 nm is weaker than the one at 870 and 1020 nm, we still find a comparably
599 high correlation of 0.81, 0.90, 0.64, and 0.81 at Seoul, Kongju, Gosan, and Osaka,
600 respectively. The correlation coefficient at 440 nm is comparably low. We find values of 0.38,
601 0.62, 0.26, and 0.28 at Seoul, Kongju, Gosan, and Osaka, respectively.

602 We are of the opinion that δ_p^S can be used as a parameter to estimate the variation of optical
603 and microphysical properties of dust when it is mixed with anthropogenic pollution particles.
604 There is a clear pattern of variation of the fine and coarse modes of the volume particle size
605 distributions with regard to changes of the value of δ_p^S . Fine-mode dominant volume particle
606 size distributions are present for low values of δ_p^S . The fine-mode fraction of the volume
607 particle size distribution decreases and the coarse-mode fraction of the particle size
608 distribution increases when the δ_p^S at 1020 nm increases.

609 The SSA spectra show clearly distinguishable patterns according to the variation of δ_p^S at
610 1020 nm. The SSA decreases with increasing measurement wavelength for low values of δ_p^S .
611 In contrast, the SSA increases with decreasing wavelength for high values of δ_p^S .

612 The dust ratio (R_D) can be derived from the δ_p^S . The R_D is approximately 0.12 to 0.17 lower
613 than what we find from the coarse-mode fraction of the volume concentration (CMF_{vc}).



614 However, if δ_p^S is larger than 0.25, then R_D is similar to CMF_{vc} , and the difference between
615 these two parameters is less than 0.04. We conclude that the δ_p^S at 1020 nm can be used to
616 estimate the dust ratio. The dust ratio estimated in that way can then be used to calculate the
617 mass concentration of Asian dust and air pollutants as column-integrated value.

618 We can confirm that δ_p^S , especially at 1020 nm, is in good agreement with δ_p^{CL} and
619 provides comparably reliable information that allows us to distinguish the presence of Asian
620 dust particles in mixed aerosol plumes. The consistency indicates that the values of δ_p^S at
621 1020 nm are high for high dust concentrations and small for low dust concentrations. This
622 means that δ_p^S can be used to retrieve the dust ratio in mixed dust plumes. However, we
623 need to keep in mind that we cannot identify the vertical distribution of dust particles on the
624 basis of δ_p^S because δ_p^S is a column-integrated value.

625

626

627 **Acknowledgement**

628 This work was funded by the Korea Meteorological Administration Research and
629 Development Program under Grant KMIPA 2015-6150. This work was supported by a
630 National Research Foundation of Korea (NRF) grant funded by the Korean government
631 (MEST) (NRF-2015R1D1A1A09058269). This research was also supported by the
632 International Environmental Research Center (IERC).

633

634

635 **References**

636 Alfaro, S. C., Lafon, S., Rajot, J. L., Formenti, P., Gaudichet, A. and Maillé, M.: Iron oxides
637 and light absorption by pure desert dust: An experimental study, *J. Geophys. Res.*, 109(D8),
638 D08208, doi:10.1029/2003JD004374, 2004.



- 639 Ansmann, A., et al.: Ash and fine-mode particle mass profiles from EARLINET-AERONET
640 observations over central Europe after the eruptions of the Eyjafjallajökull volcano in 2010, *J.*
641 *Geophys. Res.*, 116, D00U02, doi:10.1029/2010JD015567., 2011.
- 642 Arimoto, R., Kim, Y. J., Kim, Y. P., Quinn, P. K., Bates, T. S., Anderson, T. L., Gong, S.,
643 Uno, I., Chin, M., Huebert, B. J., Clarke, A. D., Shinozuka, Y., Weber, R. J., Anderson, J. R.,
644 Guazzotti, S. A., Sullivan, R. C., Sodeman, D. A., Prather, K. A. and Sokolik, I. N.:
645 Characterization of Asian Dust during ACE-Asia, *Global and Planetary Change*, 52(1–4), 23–
646 56, doi:10.1016/j.gloplacha.2006.02.013, 2006.
- 647 Behrendt, A. and Nakamura, T.: Calculation of the calibration constant of polarization lidar
648 and its dependency on atmospheric temperature, *Opt. Express*, OE, 10(16), 805–817,
649 doi:10.1364/OE.10.000805, 2002.
- 650 Bergstrom, R. W., Russell, P. B., and Hignett, P.: Wavelength dependence of the absorption
651 of black carbon particles: Predictions and results from the TARFOX experiment and
652 implications for the aerosol single scattering albedo, *J. Atmos. Sci.*, 59, 567–577, 2002.
- 653 Bohren, C. F. and Huffman, D. R.: Absorption and Scattering by a Sphere, in *Absorption and*
654 *Scattering of Light by Small Particles*, pp. 82–129, Wiley-VCH Verlag GmbH. [online]
655 Available from: <http://onlinelibrary.wiley.com/doi/10.1002/9783527618156.ch4/summary>
656 (Accessed 27 December 2016), 1983.
- 657 Bond, T. C. and Bergstrom, R. W.: Light Absorption by Carbonaceous Particles: An
658 Investigative Review, *Aerosol Science and Technology*, 40(1), 27–67,
659 doi:10.1080/02786820500421521, 2006.
- 660 Bravo-Aranda, J. A., Titos, G., Granados-Munoz, M. J., GuerreroRascado, J. L., Navas-
661 Guzman, F., Valenzuela, A., Lyamani, H., Olmo, F. J., Andrey, J., and Alados-Arboledas, L.:
662 Study of mineral dust entrainment in the planetary boundary layer by lidar depolarisation
663 technique, *Tellus B*, 67, 26180, doi:10.3402/tellusb.v67.26180, 2015.
- 664 Burton, S. P., Ferrare, R. A., Vaughan, M. A., Omar, A. H., Rogers, R. R., Hostetler, C. A.
665 and Hair, J. W.: Aerosol classification from airborne HSRL and comparisons with the
666 CALIPSO vertical feature mask, *Atmos. Meas. Tech.*, 6(5), 1397–1412, doi:10.5194/amt-6-
667 1397-2013, 2013.
- 668 Burton, S.P., Vaughan, M.A., Ferrare, R.A., Hostetler, C.A.: Separating mixtures of aerosol
669 types in airborne High Spectral Resolution Lidar data. *Atmos. Meas. Tech.* 7, 419-436. 2014.
- 670 Cairo, F., Donfrancesco, G. D., Adriani, A., Pulvirenti, L. and Fierli, F.: Comparison of
671 various linear depolarization parameters measured by lidar, *Appl. Opt.*, AO, 38(21), 4425–
672 4432, doi:10.1364/AO.38.004425, 1999.
- 673 Chin, M., Diehl, T., Dubovik, O., Eck, T. F., Holben, B. N., Sinyuk, A. and Streets, D. G.:
674 Light absorption by pollution, dust, and biomass burning aerosols: a global model study and
675 evaluation with AERONET measurements, *Ann. Geophys.*, 27(9), 3439–3464,
676 doi:10.5194/angeo-27-3439-2009, 2009.



- 677 Derimian, Y., Karnieli, A., Kaufman, Y. J., Andreae, M. O., Andreae, T. W., Dubovik, O.,
678 Maenhaut, W. and Koren, I.: The role of iron and black carbon in aerosol light absorption,
679 Atmos. Chem. Phys., 8(13), 3623–3637, doi:10.5194/acp-8-3623-2008, 2008.
- 680 Dubovik, O., Holben, B., Eck, T. F., Smirnov, A., Kaufman, Y. J., King, M. D., Tanré, D.
681 and Slutsker, I.: Variability of Absorption and Optical Properties of Key Aerosol Types
682 Observed in Worldwide Locations, J. Atmos. Sci., 59(3), 590–608, doi:10.1175/1520-
683 0469(2002)059<0590:VOAAOP>2.0.CO;2, 2002.
- 684 Dubovik, O., Sinyuk, A., Lapyonok, T., Holben, B. N., Mishchenko, M., Yang, P., Eck, T. F.,
685 Volten, H., Muñoz, O., Veihelmann, B., van der Zande, W. J., Leon, J.-F., Sorokin, M. and
686 Slutsker, I.: Application of spheroid models to account for aerosol particle nonsphericity in
687 remote sensing of desert dust, J. Geophys. Res., 111(D11), D11208,
688 doi:10.1029/2005JD006619, 2006.
- 689 Eck, T. F., Holben, B. N., Reid, J. S., Dubovik, O., Smirnov, A., O'Neill, N. T., Slutsker, I.
690 and Kinne, S.: Wavelength dependence of the optical depth of biomass burning, urban, and
691 desert dust aerosols, J. Geophys. Res., 104(D24), 31333–31349, doi:10.1029/1999JD900923,
692 1999.
- 693 Eck, T. F., Holben, B. N., Dubovik, O., Smirnov, A., Goloub, P., Chen, H. B., Chatenet, B.,
694 Gomes, L., Zhang, X.-Y., Tsay, S.-C., Ji, Q., Giles, D. and Slutsker, I.: Columnar aerosol
695 optical properties at AERONET sites in central eastern Asia and aerosol transport to the
696 tropical mid-Pacific, J. Geophys. Res., 110(D6), D06202, doi:10.1029/2004JD005274, 2005.
- 697 Eck, T. F., Holben, B. N., Sinyuk, A., Pinker, R. T., Goloub, P., Chen, H., Chatenet, B., Li,
698 Z., Singh, R. P., Tripathi, S. N., Reid, J. S., Giles, D. M., Dubovik, O., O'Neill, N. T.,
699 Smirnov, A., Wang, P. and Xia, X.: Climatological aspects of the optical properties of
700 fine/coarse mode aerosol mixtures, J. Geophys. Res., 115(D19), D19205,
701 doi:10.1029/2010JD014002, 2010.
- 702 Freudenthaler, V., Esselborn, M., Wiegner, M., Heese, B., Tesche, M., Ansmann, A., Müller,
703 D., Althausen, D., Wirth, M., Fix, A., Ehret, G., Knippertz, P., Toledano, C., Gasteiger, J.,
704 Garhammer, M. and Seefeldner, M.: Depolarization ratio profiling at several wavelengths in
705 pure Saharan dust during SAMUM 2006, Tellus B, 61(1), 165–179, doi:10.1111/j.1600-
706 0889.2008.00396.x, 2009.
- 707 Giles, D. M., Holben, B. N., Eck, T. F., Sinyuk, A., Smirnov, A., Slutsker, I., Dickerson, R.
708 R., Thompson, A. M. and Schafer, J. S.: An analysis of AERONET aerosol absorption
709 properties and classifications representative of aerosol source regions, J. Geophys. Res.,
710 117(D17), D17203, doi:10.1029/2012JD018127, 2012.
- 711 Gong, S. L., Zhang, X. Y., Zhao, T. L., McKendry, I. G., Jaffe, D. A. and Lu, N. M.:
712 Characterization of soil dust aerosol in China and its transport and distribution during 2001
713 ACE-Asia: 2. Model simulation and validation, J. Geophys. Res., 108(D9), 4262,
714 doi:10.1029/2002JD002633, 2003.



- 715 Haywood, J. M., Allan, R. P., Culverwell, I., Slingo, T., Milton, S., Edwards, J. and Clerbaux,
716 N.: Can desert dust explain the outgoing longwave radiation anomaly over the Sahara during
717 July 2003?, *J. Geophys. Res.*, 110(D5), D05105, doi:10.1029/2004JD005232, 2005.
- 718 Holben, B. N., Eck, T. F., Slutsker, I., Tanré, D., Buis, J. P., Setzer, A., Vermote, E., Reagan,
719 J. A., Kaufman, Y. J., Nakajima, T., Lavenu, F., Jankowiak, I. and Smirnov, A.:
720 AERONET—A Federated Instrument Network and Data Archive for Aerosol
721 Characterization, *Remote Sensing of Environment*, 66(1), 1–16, doi:10.1016/S0034-
722 4257(98)00031-5, 1998.
- 723 Huebert, B. J., Bates, T., Russell, P. B., Shi, G., Kim, Y. J., Kawamura, K., Carmichael, G.
724 and Nakajima, T.: An overview of ACE-Asia: Strategies for quantifying the relationships
725 between Asian aerosols and their climatic impacts, *J. Geophys. Res.*, 108(D23), 8633,
726 doi:10.1029/2003JD003550, 2003.
- 727 Husar, R. B., Tratt, D. M., Schichtel, B. A., Falke, S. R., Li, F., Jaffe, D., Gassó, S., Gill, T.,
728 Laulainen, N. S., Lu, F., Reheis, M. C., Chun, Y., Westphal, D., Holben, B. N., Gueymard, C.,
729 McKendry, I., Kuring, N., Feldman, G. C., McClain, C., Frouin, R. J., Merrill, J., DuBois, D.,
730 Vignola, F., Murayama, T., Nickovic, S., Wilson, W. E., Sassen, K., Sugimoto, N. and Malm,
731 W. C.: Asian dust events of April 1998, *J. Geophys. Res.*, 106(D16), 18317–18330,
732 doi:10.1029/2000JD900788, 2001.
- 733 IPCC: Climate Change 2013 - The Physical Science Basis: Working Group I Contribution to
734 the Fifth Assessment Report of the Intergovernmental Panel on Climate Change [Stocker,
735 T.F., D. Qin, G.-K. Plattner, M. Tignor, S.K. Allen, J. Boschung, A. Nauels, Y. Xia, V. Bex
736 and P.M. Midgley (eds.)], Cambridge University Press Cambridge, UK, and New York, NY,
737 USA., 2013.
- 738 Iwasaka, Y., Shibata, T., Nagatani, T., Shi, G.-Y., Kim, Y. S., Matsuki, A., Trochne, D.,
739 Zhang, D., Yamada, M., Nagatani, M., Nakata, H., Shen, Z., Li, G., Chen, B. and Kawahira,
740 K.: Large depolarization ratio of free tropospheric aerosols over the Taklamakan Desert
741 revealed by lidar measurements: Possible diffusion and transport of dust particles, *J. Geophys.*
742 *Res.*, 108(D23), 8652, doi:10.1029/2002JD003267, 2003.
- 743 Jickells, T. D., An, Z. S., Andersen, K. K., Baker, A. R., Bergametti, G., Brooks, N., Cao, J.
744 J., Boyd, P. W., Duce, R. A., Hunter, K. A., Kawahata, H., Kubilay, N., laRoche, J., Liss, P.
745 S., Mahowald, N., Prospero, J. M., Ridgwell, A. J., Tegen, I. and Torres, R.: Global Iron
746 Connections Between Desert Dust, Ocean Biogeochemistry, and Climate, *Science*, 308(5718),
747 67–71, doi:10.1126/science.1105959, 2005.
- 748 Kanayama, S., Yabuki, S., Yanagisawa, F. and Motoyama, R.: The chemical and strontium
749 isotope composition of atmospheric aerosols over Japan: the contribution of long-range-
750 transported Asian dust (Kosa), *Atmospheric Environment*, 36(33), 5159–5175,
751 doi:10.1016/S1352-2310(02)00587-3, 2002.



- 752 Kim, D., Chin, M., Yu, H., Eck, T. F., Sinyuk, A., Smirnov, A. and Holben, B. N.: Dust
753 optical properties over North Africa and Arabian Peninsula derived from the AERONET
754 dataset, *Atmos. Chem. Phys.*, 11(20), 10733–10741, doi:10.5194/acp-11-10733-2011, 2011.
- 755 Kim, K. W., He, Z. and Kim, Y. J.: Physicochemical characteristics and radiative properties
756 of Asian dust particles observed at Kwangju, Korea, during the 2001 ACE-Asia intensive
757 observation period, *J. Geophys. Res.*, 109(D19), D19S02, doi:10.1029/2003JD003693, 2004.
- 758 Kinne, S., Schulz, M., Textor, C., Guibert, S., Balkanski, Y., Bauer, S. E., Bernsten, T.,
759 Berglen, T. F., Boucher, O., Chin, M., Collins, W., Dentener, F., Diehl, T., Easter, R.,
760 Feichter, J., Fillmore, D., Ghan, S., Ginoux, P., Gong, S., Grini, A., Hendricks, J., Herzog, M.,
761 Horowitz, L., Isaksen, I., Iversen, T., Kirkevåg, A., Kloster, S., Koch, D., Kristjansson, J. E.,
762 Krol, M., Lauer, A., Lamarque, J. F., Lesins, G., Liu, X., Lohmann, U., Montanaro, V.,
763 Myhre, G., Penner, J., Pitari, G., Reddy, S., Seland, O., Stier, P., Takemura, T. and Tie, X.:
764 An AeroCom initial assessment – optical properties in aerosol component modules of global
765 models, *Atmos. Chem. Phys.*, 6(7), 1815–1834, doi:10.5194/acp-6-1815-2006, 2006.
- 766 Lafon, S., Rajot, J.-L., Alfaro, S. C. and Gaudichet, A.: Quantification of iron oxides in desert
767 aerosol, *Atmospheric Environment*, 38(8), 1211–1218, doi:10.1016/j.atmosenv.2003.11.006,
768 2004.
- 769 Lafon, S., Sokolik, I. N., Rajot, J. L., Caquineau, S. and Gaudichet, A.: Characterization of
770 iron oxides in mineral dust aerosols: Implications for light absorption, *J. Geophys. Res.*,
771 111(D21), D21207, doi:10.1029/2005JD007016, 2006.
- 772 Lee, J., Kim, J., Song, C. H., Ryu, J.-H., Ahn, Y.-H. and Song, C. K.: Algorithm for retrieval
773 of aerosol optical properties over the ocean from the Geostationary Ocean Color Imager,
774 *Remote Sensing of Environment*, 114(5), 1077–1088, doi:10.1016/j.rse.2009.12.021, 2010a.
- 775 Lee, K.-H., Muller, D., Noh, Y.-M., Shin, S.-K. and Shin, D.-H.: Depolarization Ratio
776 Retrievals Using AERONET Sun Photometer Data, *J. Opt. Soc. Korea*, JOSK, 14(3), 178–
777 184, 2010b.
- 778 Maring, H., Savoie, D. L., Izaguirre, M. A., Custals, L. and Reid, J. S.: Mineral dust aerosol
779 size distribution change during atmospheric transport, *J. Geophys. Res.*, 108(D19), 8592,
780 doi:10.1029/2002JD002536, 2003.
- 781 Mikami, M., Shi, G. Y., Uno, I., Yabuki, S., Iwasaka, Y., Yasui, M., Aoki, T., Tanaka, T. Y.,
782 Kurosaki, Y., Masuda, K., Uchiyama, A., Matsuki, A., Sakai, T., Takemi, T., Nakawo, M.,
783 Seino, N., Ishizuka, M., Satake, S., Fujita, K., Hara, Y., Kai, K., Kanayama, S., Hayashi, M.,
784 Du, M., Kanai, Y., Yamada, Y., Zhang, X. Y., Shen, Z., Zhou, H., Abe, O., Nagai, T.,
785 Tsutsumi, Y., Chiba, M. and Suzuki, J.: Aeolian dust experiment on climate impact: An
786 overview of Japan–China joint project ADEC, *Global and Planetary Change*, 52(1–4), 142–
787 172, doi:10.1016/j.gloplacha.2006.03.001, 2006.
- 788 Müller, D., Ansmann, A., Freudenthaler, V., Kandler, K., Toledano, C., Hiebsch, A.,
789 Gasteiger, J., Esselborn, M., Tesche, M., Heese, B., Althausen, D., Weinzierl, B., Petzold, A.
790 and von Hoyningen-Huene, W.: Mineral dust observed with AERONET Sun photometer,



- 791 Raman lidar, and in situ instruments during SAMUM 2006: Shape-dependent particle
792 properties, *J. Geophys. Res.*, 115(D11), D11207, doi:10.1029/2009JD012523, 2010.
- 793 Müller, D., Lee, K.-H., Gasteiger, J., Tesche, M., Weinzierl, B., Kandler, K., Müller, T.,
794 Toledano, C., Otto, S., Althausen, D. and Ansmann, A.: Comparison of optical and
795 microphysical properties of pure Saharan mineral dust observed with AERONET Sun
796 photometer, Raman lidar, and in situ instruments during SAMUM 2006, *J. Geophys. Res.*,
797 117(D7), D07211, doi:10.1029/2011JD016825, 2012.
- 798 Müller, T., Schladitz, A., Kandler, K. and Wiedensohler, A.: Spectral particle absorption
799 coefficients, single scattering albedos and imaginary parts of refractive indices from ground
800 based in situ measurements at Cape Verde Island during SAMUM-2, *Tellus B*, 63(4),
801 doi:10.3402/tellusb.v63i4.16348, 2011.
- 802 Navas-Guzma, J. A., Bravo-Aranda, J. L., Guerrero-Rascado, M. J.,
803 Granados-Munoz, D., PeDrez-Ramirez, F. J., Olmo, and L. Alados-Arboledas.: Eruption of
804 the Eyjafjallajökull Volcano in spring 2010: Multiwavelength Raman lidar measurements
805 of sulphate particles in the lower troposphere, *J. Geophys. Res. Atmos.*, 118, 1804–1813,
806 doi:10.1002/jgrd.50116, 2013.
- 807 Noh, Y. M.: Single-scattering albedo profiling of mixed Asian dust plumes with
808 multiwavelength Raman lidar, *Atmospheric Environment*, 95, 305–317,
809 doi:10.1016/j.atmosenv.2014.06.028, 2014.
- 810 Noh, Y. M., Kim, Y. J., Choi, B. C. and Murayama, T.: Aerosol lidar ratio characteristics
811 measured by a multi-wavelength Raman lidar system at Anmyeon Island, Korea,
812 *Atmospheric Research*, 86(1), 76–87, doi:10.1016/j.atmosres.2007.03.006, 2007.
- 813 Noh, Y. M., Kim, Y. J. and Müller, D.: Seasonal characteristics of lidar ratios measured with
814 a Raman lidar at Gwangju, Korea in spring and autumn, *Atmospheric Environment*, 42(9),
815 2208–2224, doi:10.1016/j.atmosenv.2007.11.045, 2008.
- 816 Noh, Y. M., Müller, D., Lee, H., Lee, K. and Kim, Y. J.: Columnar aerosol optical and
817 radiative properties according to season and air mass transport pattern over East Asia,
818 *Environ Monit Assess*, 184(8), 4763–4775, doi:10.1007/s10661-011-2300-7, 2012a.
- 819 Noh, Y. M., Müller, D., Lee, H., Lee, K., Kim, K., Shin, S. and Kim, Y. J.: Estimation of
820 radiative forcing by the dust and non-dust content in mixed East Asian pollution plumes on
821 the basis of depolarization ratios measured with lidar, *Atmospheric Environment*, 61, 221–
822 231, doi:10.1016/j.atmosenv.2012.07.034, 2012b.
- 823 Noh, Y. M., Müller, D., Lee, H. and Choi, T. J.: Influence of biogenic pollen on optical
824 properties of atmospheric aerosols observed by lidar over Gwangju, South Korea,
825 *Atmospheric Environment*, 69, 139–147, doi:10.1016/j.atmosenv.2012.12.018, 2013a.
- 826 Noh, Y. M., Lee, H., Mueller, D., Lee, K., Shin, D., Shin, S., Choi, T. J., Choi, Y. J. and Kim,
827 K. R.: Investigation of the diurnal pattern of the vertical distribution of pollen in the lower



- 828 troposphere using LIDAR, Atmos. Chem. Phys., 13(15), 7619–7629, doi:10.5194/acp-13-
829 7619-2013, 2013b.
- 830 Noh, Y. M., Shin, S.-K., Lee, K., Müller, D. and Kim, K.: Utilization of the depolarization
831 ratio derived by AERONET Sun/sky radiometer data for type confirmation of a mixed
832 aerosol plume over East Asia, International Journal of Remote Sensing, 37(10), 2180–2197,
833 doi:10.1080/01431161.2016.1176274, 2016a.
- 834 Noh, Y. M., Lee, K., Kim, K., Shin, S.-K., Müller, D. and Shin, D. H.: Influence of the
835 vertical absorption profile of mixed Asian dust plumes on aerosol direct radiative forcing
836 over East Asia, Atmospheric Environment, 138, 191–204,
837 doi:10.1016/j.atmosenv.2016.04.044, 2016b.
- 838 Russell, P. B., Bergstrom, R. W., Shinozuka, Y., Clarke, A. D., DeCarlo, P. F., Jimenez, J. L.,
839 Livingston, J. M., Redemann, J., Dubovik, O. and Strawa, A.: Absorption Angstrom
840 Exponent in AERONET and related data as an indicator of aerosol composition, Atmos.
841 Chem. Phys., 10(3), 1155–1169, doi:10.5194/acp-10-1155-2010, 2010.
- 842 Sakai, T., Nagai, T., Nakazato, M., Mano, Y. and Matsumura, T.: Ice clouds and Asian dust
843 studied with lidar measurements of particle extinction-to-backscatter ratio, particle
844 depolarization, and water-vapor mixing ratio over Tsukuba, Appl Opt, 42(36), 7103–7116,
845 2003.
- 846 Sassen, K., Wang, Z. and Liu, D.: Global distribution of cirrus clouds from CloudSat/Cloud-
847 Aerosol Lidar and Infrared Pathfinder Satellite Observations (CALIPSO) measurements, J.
848 Geophys. Res., 113(D8), D00A12, doi:10.1029/2008JD009972, 2008.
- 849 Shimizu, A., Sugimoto, N., Matsui, I., Arao, K., Uno, I., Murayama, T., Kagawa, N., Aoki,
850 K., Uchiyama, A. and Yamazaki, A.: Continuous observations of Asian dust and other
851 aerosols by polarization lidars in China and Japan during ACE-Asia, J. Geophys. Res.,
852 109(D19), D19S17, doi:10.1029/2002JD003253, 2004.
- 853 Shin, S.-K., Müller, D., Lee, C., Lee, K. H., Shin, D., Kim, Y. J. and Noh, Y. M.: Vertical
854 variation of optical properties of mixed Asian dust/pollution plumes according to pathway of
855 air mass transport over East Asia, Atmos. Chem. Phys., 15(12), 6707–6720, doi:10.5194/acp-
856 15-6707-2015, 2015.
- 857 Sokolik, I. N. and Toon, O. B.: Incorporation of mineralogical composition into models of the
858 radiative properties of mineral aerosol from UV to IR wavelengths, J. Geophys. Res.,
859 104(D8), 9423–9444, doi:10.1029/1998JD200048, 1999.
- 860 Sugimoto, N., Matsui, I., Shimizu, A., Nishizawa, T., Hara, Y., Xie, C., Uno, I., Yumimoto,
861 K., Wang, Z. and Yoon, S.-C.: Lidar network observations of tropospheric aerosols, vol. 7153,
862 p. 71530A–71530A–13., 2008.
- 863 Sun, Y., Zhuang, G., Wang, Y., Zhao, X., Li, J., Wang, Z. and An, Z.: Chemical composition
864 of dust storms in Beijing and implications for the mixing of mineral aerosol with pollution



865 aerosol on the pathway, *J. Geophys. Res.*, 110(D24), D24209, doi:10.1029/2005JD006054,
866 2005.

867 Sun, Y., Zhuang, G., Huang, K., Li, J., Wang, Q., Wang, Y., Lin, Y., Fu, J. S., Zhang, W.,
868 Tang, A. and Zhao, X.: Asian dust over northern China and its impact on the downstream
869 aerosol chemistry in 2004, *J. Geophys. Res.*, 115(D7), D00K09, doi:10.1029/2009JD012757,
870 2010.

871 Tanaka, M., Nakajima, T., Shiobara, M., Yamano, M., and Arao, K.: Aerosol optical
872 characteristics in the yellow sand events observed in May, 1982 at Nagasaki—Part 1.
873 Observations. *J. Meteor. Soc. Japan*, 67, 267–278, 1989.

874 Tesche, M., Ansmann, A., Müller, D., Althausen, D., Engelmann, R., Freudenthaler, V. and
875 Groß, S.: Vertically resolved separation of dust and smoke over Cape Verde using
876 multiwavelength Raman and polarization lidars during Saharan Mineral Dust Experiment
877 2008, *J. Geophys. Res.*, 114(D13), D13202, doi:10.1029/2009JD011862, 2009.

878 Tesche, M., Müller, D., Gross, S., Ansmann, A., Althausen, D., Freudenthaler, V., Weinzierl,
879 B., Veira, A. and Petzold, A.: Optical and microphysical properties of smoke over Cape
880 Verde inferred from multiwavelength lidar measurements, *Tellus B*, 63(4), 677–694,
881 doi:10.1111/j.1600-0889.2011.00549.x, 2011.

882 Volten, H., Muñoz, O., Rol, E., de Haan, J. F., Vassen, W., Hovenier, J. W., Muinonen, K.
883 and Nousiainen, T.: Scattering matrices of mineral aerosol particles at 441.6 nm and 632.8
884 nm, *J. Geophys. Res.*, 106(D15), 17375–17401, doi:10.1029/2001JD900068, 2001.

885 Whiteman, D. N., Melfi, S. H. and Ferrare, R. A.: Raman lidar system for the measurement of
886 water vapor and aerosols in the Earth's atmosphere, *Appl Opt*, 31(16), 3068–3082, 1992.

887 Yu, X., Cheng, T., Chen, J. and Liu, Y.: A comparison of dust properties between China
888 continent and Korea, Japan in East Asia, *Atmospheric Environment*, 40(30), 5787–5797,
889 doi:10.1016/j.atmosenv.2006.05.013, 2006.

890

891 **Figure captions**

892 Figure 1. Map of the observation sites. Measurements with AERONET sun/sky radiometer
893 and lidar were performed at Seoul, Kongju, Gosan, and Osaka. AERONET sun/sky
894 radiometer measurements were made at Dunhuang.

895



896 Figure 2. Lidar derived aerosol backscatter coefficient (β_p ; black line), the linear particle
897 depolarization ratio (δ_p^L ; green), and the weighted linear particle depolarization ratio ($\delta_p^L W$;
898 gray) at 532 nm observed (a) from 23:00 – 23:15 UTC on 13 March 2010, (b) from 06:00 –
899 06:15 UTC on 22 March 2010, and (c) from 23:15 - 23:30 on 3 May 2010. The wavelength of
900 δ_p^S is 1020 nm. The height is expressed as a.g.l.

901

902 Figure 3. Aerosol optical depth (τ) measured with sun/sky radiometer at 500 nm (black
903 squares) and linear particle depolarization ratio derived at 1020 nm from sun/sky radiometer
904 data (δ_p^S ; blue open circles). (a) data of the four sites taken from 2010 to 2014, (b) data taken
905 at Seoul for two years (2012 and 2013), (c) data taken at Kongju in spring 2012, and (d) data
906 taken at Gosan during four years (2011 – 2014), mostly during spring, and (e) data taken
907 from 2010 – 2014 at Osaka.

908

909 Figure 4. The correlation coefficients R^2 (the coefficient of determination) between δ_p^{CL} at
910 532 and δ_p^S at 440 (black squares), 675 (red circles), 870 (blue open triangles), and 1020 nm
911 (orange diamonds) at (a) Seoul, (b) Kongju, (c) Gosan, and (d) Osaka.

912

913 Figure 5. The average value of the δ_p^S at 440, 675, 870, and 1020 nm for each group. Each
914 group is distinguished by color: black (group 1), red (group 2), blue (group 3), pink (group 4),
915 gray (group 5), and orange (group 6). The average values of the δ_p^{CL} at 532 nm are shown as
916 open circles and the same color as the δ_p^S .

917



918 Figure 6. Average values of the vertical profiles of (1) the particle backscatter coefficient, (2)
919 the linear particle depolarization ratios (δ_p^L), and (3) the weighted linear particle
920 depolarization ratios ($\delta_p^L w$) for group 1 (a), group 2 (b), group 3 (c), group 4 (d), group 5 (e),
921 and group 6 (f). The sites are Seoul (red), Kongju (orange), Gosan (blue), and Osaka (black).
922

923 Figure 7. Average value of the SSA and the volume particle size distributions of each of the 6
924 groups considered in this study: group 1 (black), group 2 (red), group 3 (blue), group 4
925 (pink), group 5 (gray), and group 6 (orange).
926

927 Figure 8. (a) Linear particle depolarization ratios, (b) single-scattering albedos, and (c)
928 volume particle size distributions derived from sun/sky radiometer observations at Dunhuang.
929

930 Figure 9. Comparison between (a) δ_p^S , (b) SSA, and volume particle size distributions
931 representing the dust source region (Dunhuang, black) and group-6-data (Kongju, blue),
932 (Gosan, gray), and (Osaka, red).
933

934 Figure 10. Separation of data of group 6 at Osaka according to case 1 (black) and case 2 (red).
935 Shown are (a) backscatter coefficients, (b) linear particle depolarization ratios (δ_p^L), and (c)
936 weighted linear particle depolarization ratios ($\delta_p^L W$).
937



938 Figure 11. The (a) δ_p^S at 440, 675, 870, and 1020 nm, (b) SSA at 440, 675, 870, and 1020
939 nm, and (c) volume particle size distributions. Shown are the results for case 1 (black) and
940 case 2 (red). C_{V_c}/C_{V_f} is inserted in (c). The observation site is Osaka.

941

942 Figure 12. HYSPLIT 5-days backward trajectories of dust plumes for case 1 (a) and case (2).

943 The start height for case 1 is 1200 m (blue), 1500 m (red), and 1800 m (yellow). For case 2 it
944 is 500 m (blue), 1000 m (red), and 1500 m (yellow). The start time is 0:00 UTC in each case.

945

946 Figure 13. Correlation plots of the dust ratio at 1020 nm versus the volume concentration in
947 the coarse-mode fraction.

948

949

950

951

952

953

954

955

956

957

958

959

960



Table 1. Average value of aerosol optical depth (τ) at 500 nm, optical-depth-related Ångström exponent (\hat{a} , 440 – 870 nm), absorption-related Ångström exponent (\hat{a}_A), dust ratio (R_D) derived by δ_p^2 at 1020 nm, coarse-mode fraction of the τ at 1020 nm, coarse-mode fraction of the volume concentration (CMF_{vc}), and difference between CMF_{vc} and R_D .

DPR (1020)		Group 1	Group 2	Group 3	Group 4	Group 5	Group 6	Total
		0 - 0.05	0.05 - 0.1	0.1 - 0.15	0.15 - 0.2	0.2 - 0.25	0.25 >	
Seoul	#	25	74	38	21	5		163
	τ (500 nm)	0.62 ± 0.26	0.74 ± 0.31	0.77 ± 0.29	0.62 ± 0.25	0.89 ± 0.29		0.72 ± 0.29
	\hat{a}	1.42 ± 0.21	1.38 ± 0.16	1.22 ± 0.16	1.05 ± 0.12	0.8 ± 0.15		1.29 ± 0.22
	\hat{a}_A	1.32 ± 0.37	1.56 ± 0.43	1.78 ± 0.40	1.8 ± 0.29	1.88 ± 0.36		1.62 ± 0.43
	R_D	0.08 ± 0.04	0.22 ± 0.05	0.37 ± 0.05	0.51 ± 0.05	0.63 ± 0.05		0.29 ± 0.15
	CMF_{τ^1}	0.13 ± 0.06	0.29 ± 0.14	0.46 ± 0.12	0.69 ± 0.06	0.78 ± 0.02		0.37 ± 0.21
	CMF_{vc}	0.22 ± 0.06	0.38 ± 0.11	0.56 ± 0.09	0.70 ± 0.05	0.79 ± 0.06		0.45 ± 0.18
	$CMF_{vc} - R_D$	0.14	0.16	0.19	0.19	0.16		0.16
Kongju	#	1	22	8	3		10	44
	τ (500 nm)	0.67	0.68 ± 0.29	0.69 ± 0.24	0.41 ± 0.01		0.61 ± 0.21	0.65 ± 0.25
	\hat{a}	1.74	1.55 ± 0.12	1.27 ± 0.16	1.09 ± 0.07		0.4 ± 0.10	1.21 ± 0.47
	\hat{a}_A	1.25	1.32 ± 0.21	1.54 ± 0.47	1.43 ± 0.72		2.28 ± 0.29	1.58 ± 0.50
	R_D	0.08	0.21 ± 0.04	0.40 ± 0.06	0.54 ± 0.03		0.85 ± 0.06	0.41 ± 0.26
	CMF_{τ^1}	0.23	0.29 ± 0.09	0.47 ± 0.16	0.66 ± 0.03		0.90 ± 0.02	0.49 ± 0.26
	CMF_{vc}	0.24	0.36 ± 0.07	0.54 ± 0.11	0.65 ± 0.01		0.88 ± 0.02	0.53 ± 0.22
	$CMF_{vc} - R_D$	0.16	0.15	0.14	0.11		0.04	0.12
Gosan	#	18	47	47	19	5	3	139
	τ (500 nm)	0.6 ± 0.24	0.54 ± 0.23	0.54 ± 0.17	0.53 ± 0.14	0.41 ± 0.02	0.85 ± 0.17	0.55 ± 0.20
	\hat{a}	1.55 ± 0.14	1.34 ± 0.20	1.27 ± 0.18	0.92 ± 0.18	0.77 ± 0.11	0.24 ± 0.24	0.124 ± 0.30
	\hat{a}_A	0.75 ± 0.36	0.91 ± 0.48	0.94 ± 0.43	0.95 ± 0.59	1.01 ± 0.11	2.39 ± 0.70	0.94 ± 0.51
	R_D	0.08 ± 0.02	0.20 ± 0.05	0.37 ± 0.04	0.53 ± 0.04	0.71 ± 0.02	0.85 ± 0.05	0.32 ± 0.18
	CMF_{τ^1}	0.24 ± 0.07	0.33 ± 0.15	0.53 ± 0.11	0.69 ± 0.10	0.81 ± 0.02	0.90 ± 0.02	0.47 ± 0.21
	CMF_{vc}	0.26 ± 0.06	0.37 ± 0.10	0.56 ± 0.09	0.68 ± 0.08	0.75 ± 0.05	0.89 ± 0.08	0.49 ± 0.18
	$CMF_{vc} - R_D$	0.18	0.17	0.19	0.15	0.04	0.04	0.17
Osaka	#	54	80	32	38	20	10	234
	τ (500 nm)	0.54 ± 0.21	0.47 ± 0.18	0.47 ± 0.12	0.54 ± 0.15	0.61 ± 0.17	0.58 ± 0.15	0.51 ± 0.18
	\hat{a}	1.64 ± 0.13	1.52 ± 0.16	0.13 ± 0.17	1.01 ± 0.14	0.70 ± 0.18	0.26 ± 0.08	1.31 ± 0.39
	\hat{a}_A	1.18 ± 0.25	1.23 ± 0.23	1.37 ± 0.43	1.69 ± 0.39	1.57 ± 0.43	1.98 ± 0.41	1.37 ± 0.39
	R_D	0.07 ± 0.04	0.20 ± 0.06	0.37 ± 0.06	0.55 ± 0.05	0.68 ± 0.04	0.86 ± 0.04	0.32 ± 0.23
	CMF_{τ^1}	0.17 ± 0.09	0.30 ± 0.12	0.47 ± 0.14	0.69 ± 0.07	0.81 ± 0.04	0.91 ± 0.01	0.43 ± 0.26
	CMF_{vc}	0.21 ± 0.06	0.35 ± 0.07	0.52 ± 0.08	0.66 ± 0.05	0.79 ± 0.06	0.89 ± 0.03	0.45 ± 0.22
	$CMF_{vc} - R_D$	0.14	0.15	0.15	0.11	0.11	0.03	0.13

¹1020 nm

961

962

963

964

965

966

967

968

969

970

971

972

973

974



Table 2. Averaged volume concentration of the fine (CV_f) and the coarse mode (CV_c), and the ratio (CV_c/CV_f) for the 6 groups.

		Group 1	Group 2	Group 3	Group 4	Group 5	Group 6	Total
		0 - 0.05	0.05 - 0.1	0.1 - 0.15	0.15 - 0.2	0.2 - 0.25	0.25 >	
Seoul	CV_f	0.112 ± 0.051	0.117 ± 0.052	0.103 ± 0.042	0.067 ± 0.032	0.082 ± 0.021		0.11 ± 0.05
	CV_c	0.031 ± 0.013	0.069 ± 0.032	0.132 ± 0.052	0.162 ± 0.071	0.352 ± 0.172		0.098 ± 0.081
	CV_c/CV_f	0.29 ± 0.10	0.66 ± 0.29	1.39 ± 0.54	2.45 ± 0.65	4.26 ± 1.60		1.11 ± 0.99
Kongju	CV_f	0.136	0.109 ± 0.044	0.092 ± 0.038	0.046 ± 0.004		0.048 ± 0.018	0.088 ± 0.044
	CV_c	0.044	0.062 ± 0.027	0.105 ± 0.029	0.085 ± 0.012		0.361 ± 0.148	0.138 ± 0.140
	CV_c/CV_f	0.32	0.59 ± 0.19	1.29 ± 0.63	1.84 ± 0.12		7.61 ± 1.44	2.37 ± 2.95
Gosan	CV_f	0.127 ± 0.043	0.104 ± 0.044	0.081 ± 0.036	0.074 ± 0.038	0.042 ± 0.002	0.054 ± 0.039	0.091 ± 0.043
	CV_c	0.043 ± 0.016	0.059 ± 0.027	0.105 ± 0.042	0.159 ± 0.068	0.136 ± 0.059	0.444 ± 0.125	0.098 ± 0.077
	CV_c/CV_f	0.35 ± 0.12	0.63 ± 0.31	1.40 ± 0.57	2.36 ± 0.89	3.18 ± 1.21	11.26 ± 7.07	1.41 ± 1.92
Osaka	CV_f	0.103 ± 0.037	0.077 ± 0.030	0.065 ± 0.019	0.066 ± 0.018	0.059 ± 0.016	0.039 ± 0.008	0.076 ± 0.031
	CV_c	0.027 ± 0.011	0.040 ± 0.015	0.072 ± 0.028	0.135 ± 0.047	0.253 ± 0.147	0.326 ± 0.094	0.087 ± 0.097
	CV_c/CV_f	0.28 ± 0.11	0.56 ± 0.18	1.13 ± 0.35	2.04 ± 0.43	4.49 ± 2.58	8.49 ± 2.55	1.49 ± 2.10

975

976

977

978

979

980

981

982

983

Table 3. Aerosol optical depth (τ) at 500 nm, linear particle depolarization ratio (δ_p^s) derived from the sun/sky radiometer data, optical-depth-related Ångström exponent (\hat{a} , 440-870 nm), absorption-related Ångström exponent (\hat{a}_A), coarse-mode fraction in terms of the volume concentration (CMF_{vc}), and dust ratio (R_D) at 1020 nm. The observation site is Dunhuang.

Date	τ (500 nm)	δ_p^s (1020 nm)	\hat{a} (440-870 nm)	\hat{a}_A	CMF_{vc}	R_D
8 Apr.	0.71	0.25	0.13	1.76	0.94	0.77
9 Apr.	0.92	0.31	0.12	2.17	0.93	0.94
26 Apr.	1.17	0.32	0.14	2.14	0.94	0.96
27 Apr.	1.00	0.34	0.15	2.46	0.95	0.99
28 Apr.	1.16	0.32	0.17	2.19	0.93	0.96
Ave. ¹	0.97 ± 0.17	0.31 ± 0.03	0.14 ± 0.02	2.14 ± 0.25	0.94 ± 0.01	0.93 ± 0.09
Ave. ²	1.04 ± 0.11	0.33 ± 0.01	0.14 ± 0.02	2.24 ± 0.15	0.94 ± 0.01	0.97 ± 0.02

¹Averaged for all data ²Averaged except 8 Apr. data

984



Table 4. Volume concentration of the fine (C_{Vf}) and the coarse mode (C_{Vc}), and the ratio (C_{Vc}/C_{Vf}) at Dunhuang

Date	8 Apr.	9 Apr.	26 Apr.	27 Apr.	28 Apr.	Ave. ¹	Ave. ²
C_{Vf}	0.03	0.037	0.048	0.038	0.051	0.041 ± 0.01	0.044 ± 0.01
C_{Vc}	0.44	0.49	0.77	0.7	0.64	0.61 ± 0.14	0.65 ± 0.12
C_{Vc}/C_{Vf}	14.8	13.2	16	18.3	12.6	14.9 ± 2.3	15.0 ± 2.6

¹Averaged for all data ²Averaged except 8 Apr. data

Table 5. Parameters for cases 1 and 2 for group 6 at Osaka, linear particle depolarization ratio (δ_p^s) at 440, 675, 870, and 1020 nm derived from the Sun/sky radiometer data, absorption-related Ångström exponent (\hat{a}_A), ratio of the volume concentration (C_{Vc}/C_{Vf}), coarse-mode fraction on the basis of the volume concentration (CMF_{vc}), and dust ratio (R_D) at 1020 nm.

	δ_p^s				\hat{a}_A	C_{Vc}/C_{Vf}	CMF_{vc}	R_D
	440 nm	675 nm	870 nm	1020 nm				
Case 1	0.19 ± 0.02	0.24 ± 0.01	0.27 ± 0.01	0.29 ± 0.01	2.12 ± 0.38	6.8 ± 0.7	0.87 ± 0.01	0.88 ± 0.02
Case 2	0.15 ± 0.02	0.24 ± 0.01	0.26 ± 0.01	0.27 ± 0.01	1.58 ± 0.09	11.8 ± 1.8	0.92 ± 0.01	0.81 ± 0.03

985

986

987

988

989

990

991

992

993

994

995

996

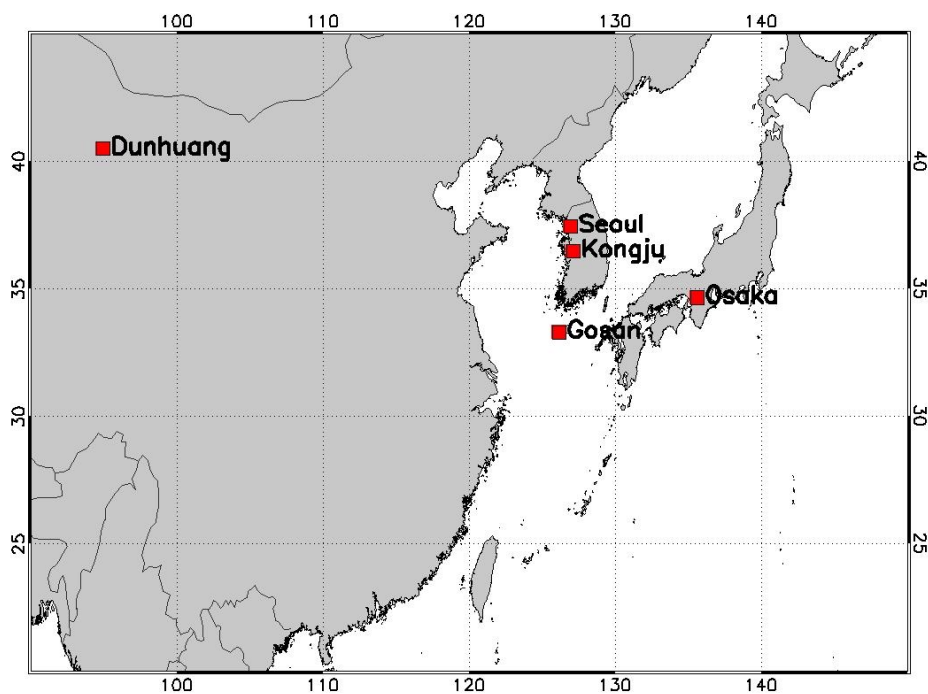
997

998

999



1000 Figure 1



1001

1002

1003

1004

1005

1006

1007

1008

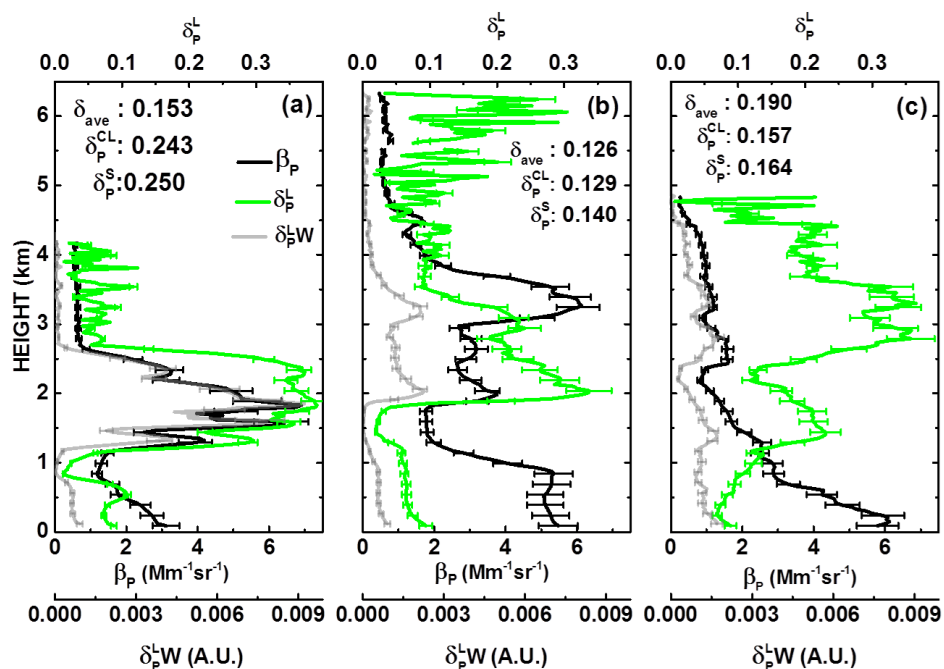
1009

1010

1011



1012 Figure 2



1013

1014

1015

1016

1017

1018

1019

1020

1021

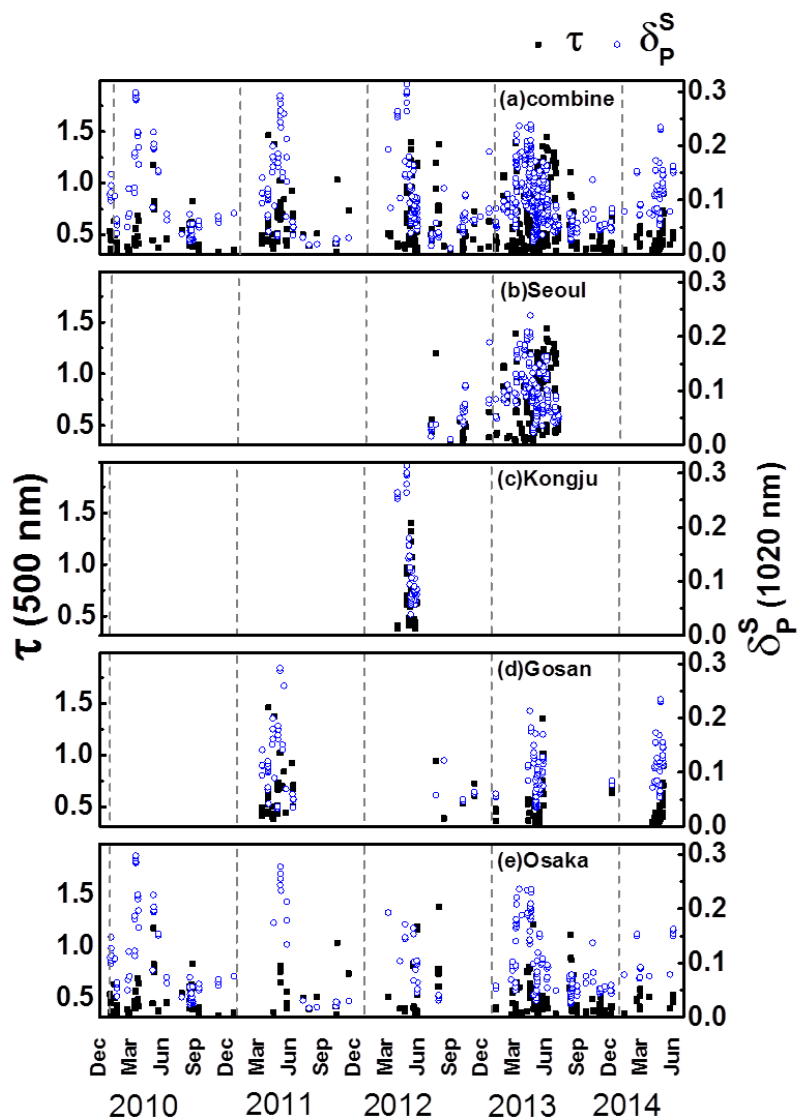
1022

1023

1024



1025 Figure 3



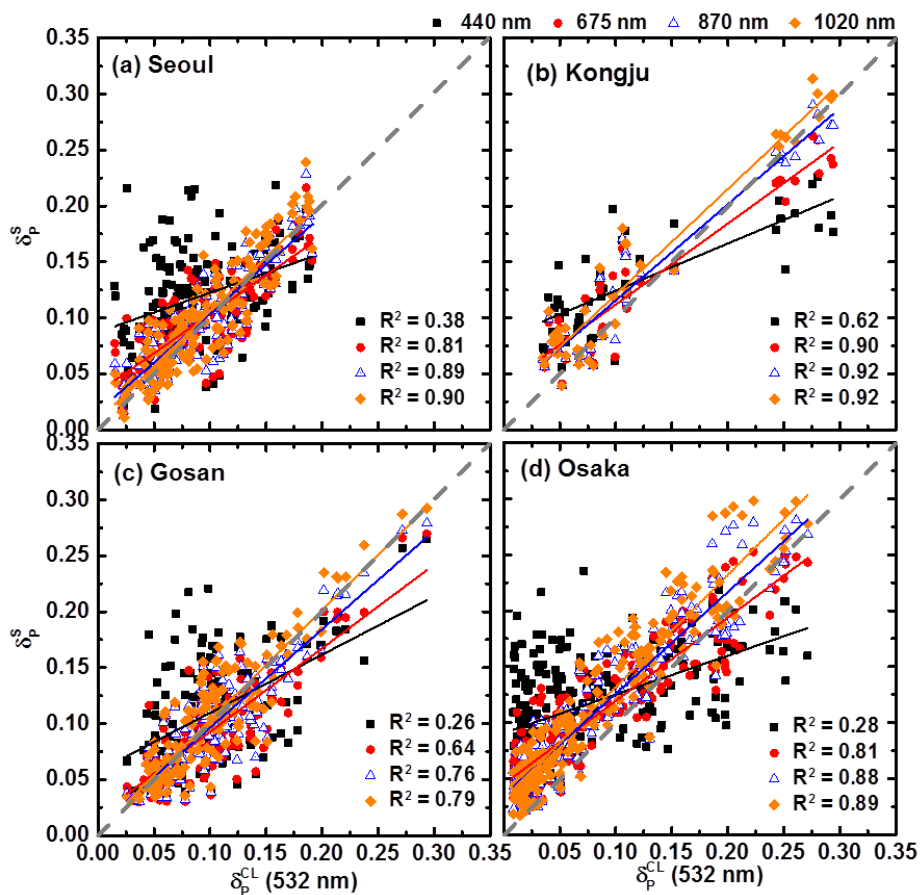
1026

1027

1028



1029 Figure 4



1030

1031

1032

1033

1034

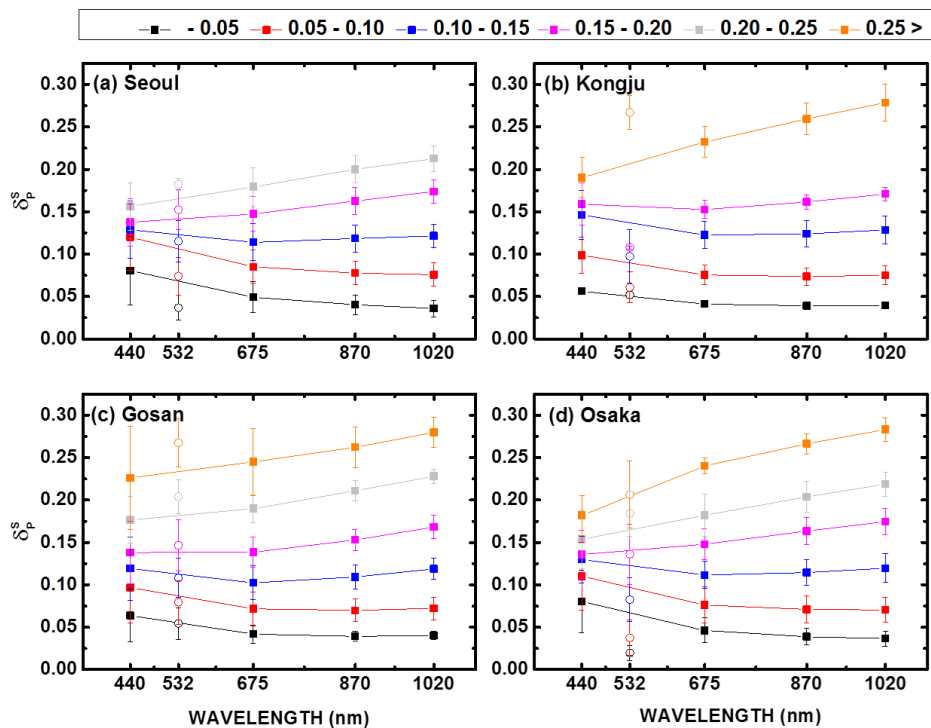
1035

1036

1037



1038 Figure 5



1039

1040

1041

1042

1043

1044

1045

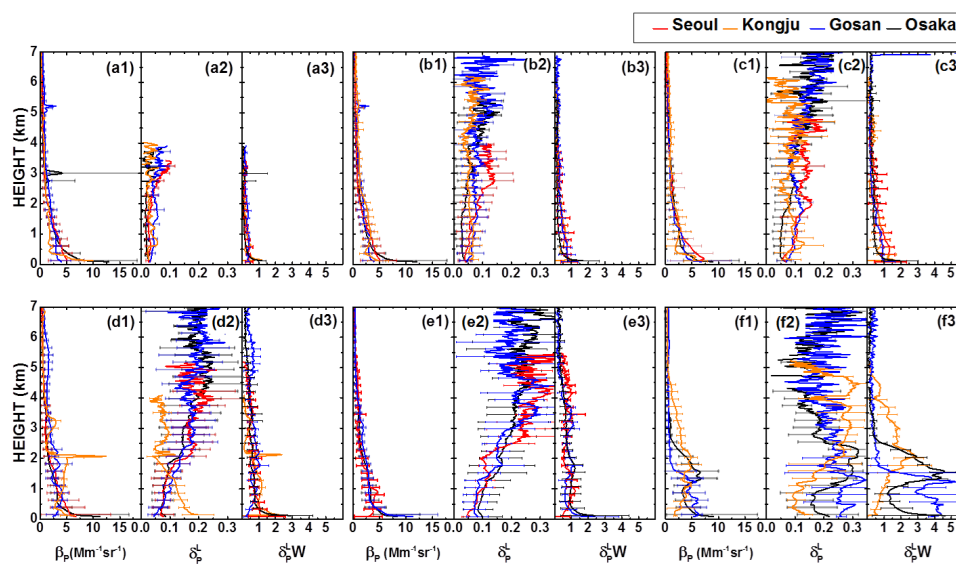
1046

1047

1048



1049 Figure 6



1050

1051

1052

1053

1054

1055

1056

1057

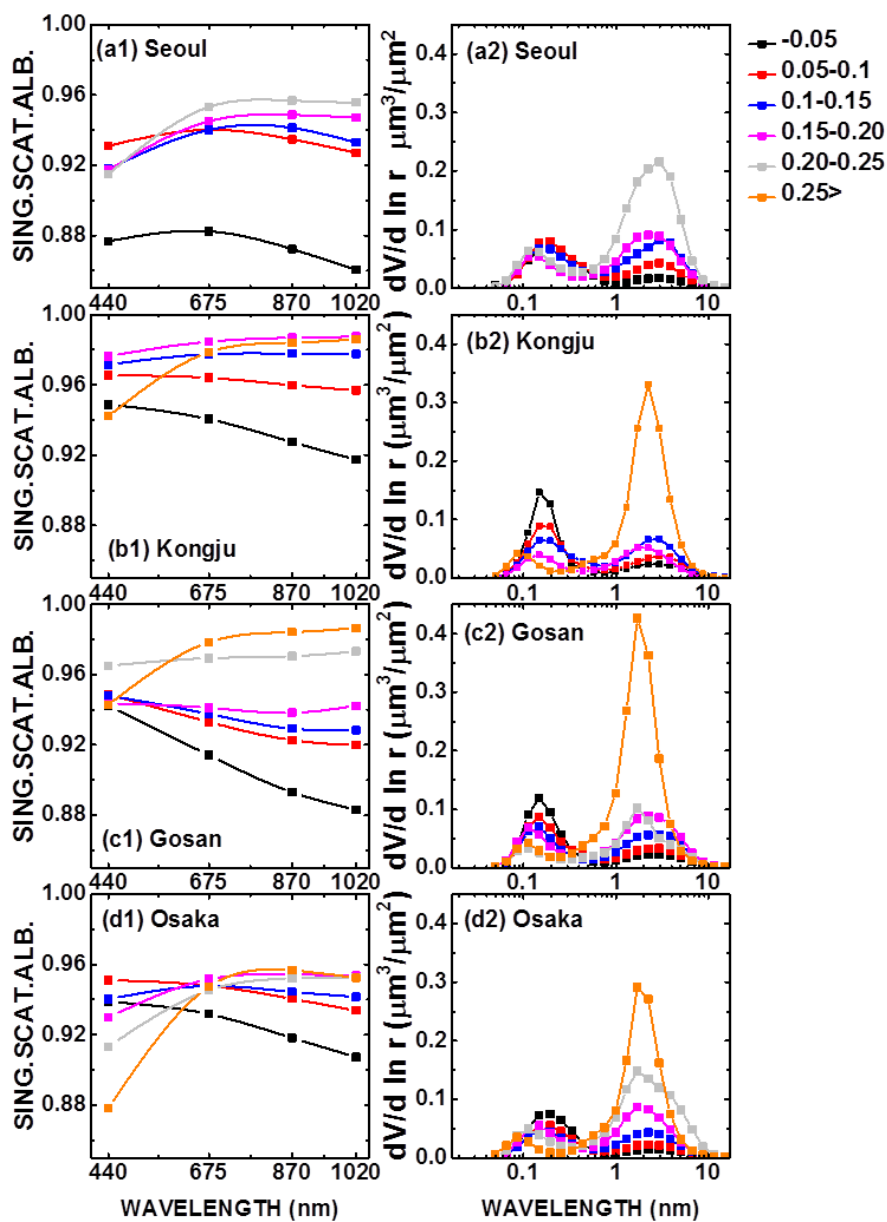
1058

1059

1060



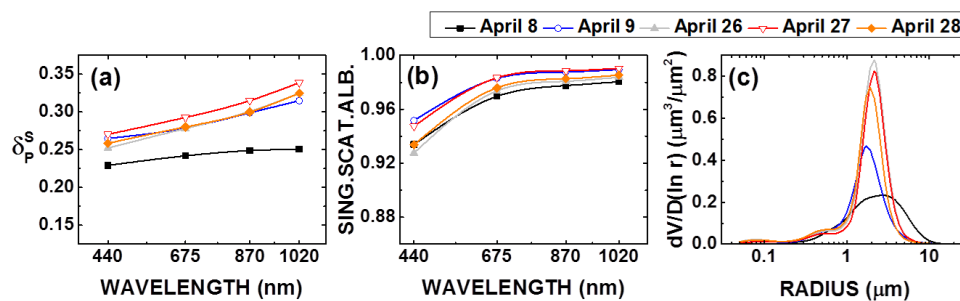
1061 Figure 7



1062



1063 Figure 8



1064

1065

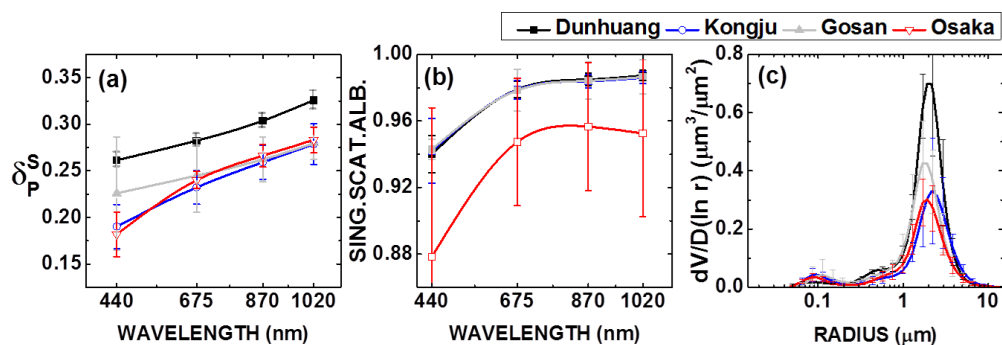
1066

1067

1068

1069

1070 Figure 9



1071

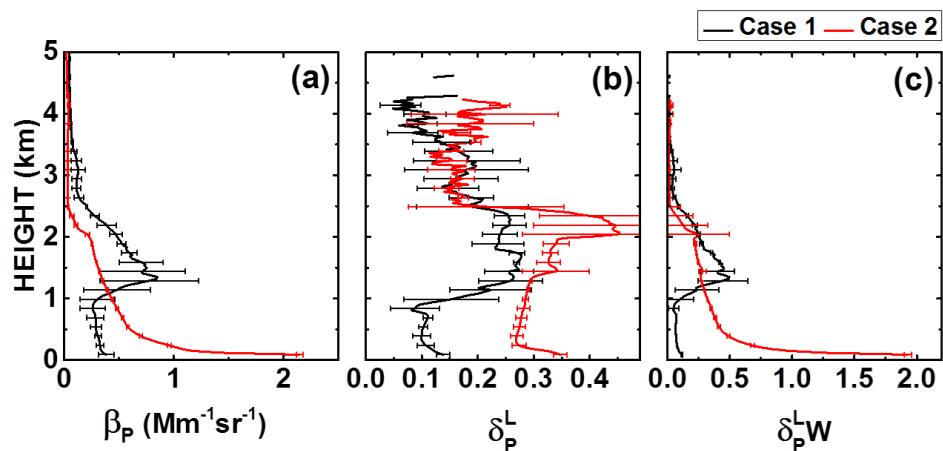
1072

1073

1074



1075 Figure 10



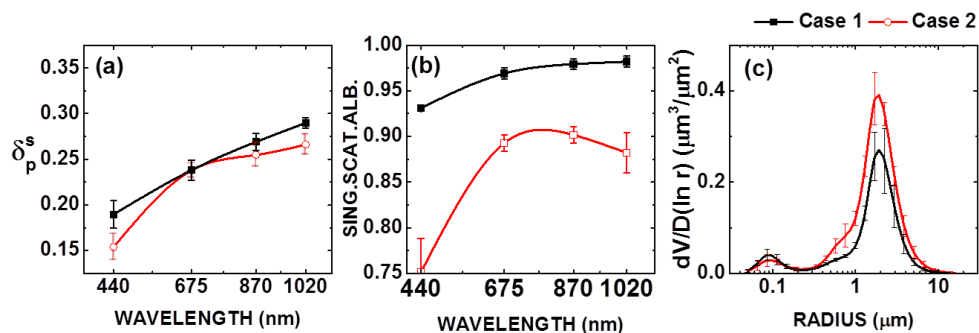
1076

1077

1078

1079

1080 Figure 11



1081

1082

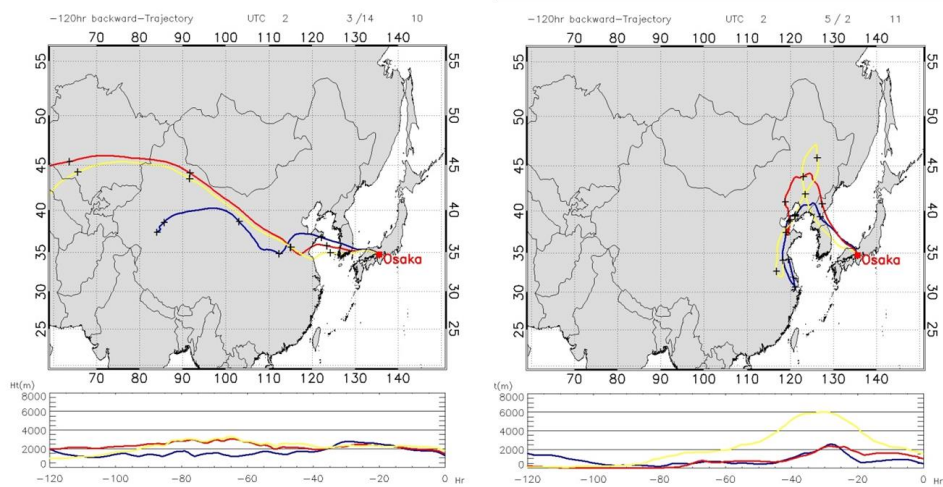
1083

1084

1085



1086 Figure 12



1087

1088

1089

1090

1091

1092

1093

1094

1095

1096

1097

1098

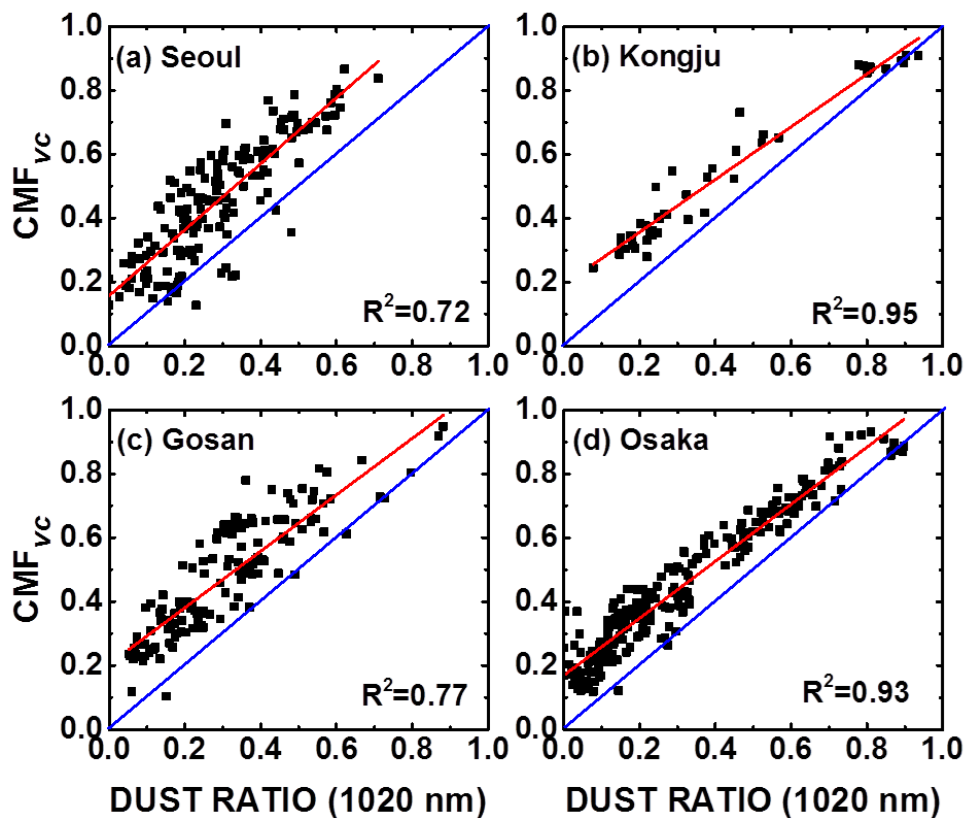
1099

1100

1101



1102 Figure 13



1103