



**Unusual  
stratospheric ozone  
anomalies observed  
in 22 years**

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**Unusual stratospheric ozone anomalies  
observed in 22 years of measurements  
from Lauder, New Zealand**

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## Abstract

The Microwave Ozone Profiling Instrument (MOP11) has provided ozone ( $O_3$ ) profiles for the Network for the Detection of Atmospheric Composition Change (NDACC) at Lauder, New Zealand, since 1992. We present the entire 22 year dataset and compare with satellite  $O_3$  observations. We will study in detail two particularly interesting variations in  $O_3$ . The first is a large positive  $O_3$  anomaly which occurs in the mid-stratosphere at  $\sim 10$ – $30$  hPa in June 2001, and which is caused by an anticyclonic circulation that persists for several weeks over Lauder. We find that this  $O_3$  anomaly is associated with air with the highest June average tracer equivalent latitude (TrEL) over the 35 year period (1980–2014). A second, and longer-lived feature, is a positive  $O_3$  anomaly in the mid-stratosphere ( $\sim 10$  hPa) from mid-2009 until mid-2013. Coincident measurements from the Aura Microwave Limb Sounder (MLS) show that these high  $O_3$  mixing ratios are well correlated with high nitrous oxide ( $N_2O$ ) mixing ratios. This correlation suggests that the high  $O_3$  over this 4 year period is driven by unusual dynamics. The beginning of the high  $O_3$  and high  $N_2O$  period at Lauder (and throughout this latitude band) occurs nearly simultaneously with a sharp decrease in  $O_3$  and  $N_2O$  at the equator, and the period ends nearly simultaneously with a sharp increase in  $O_3$  and  $N_2O$  at the equator.

## 1 Introduction

Observations of total column  $O_3$  show that, over most of the globe,  $O_3$  loss has leveled off since  $\sim 2000$ , and has even begun to increase. The large decline observed from the 1960s to the late 1990s has ended (WMO, 2014). While global  $O_3$  may be recovering, the magnitude and sign of stratospheric  $O_3$  trends over multi-decadal timescales in the mid-stratosphere may strongly depend on geographical location. It is important to understand the causes of this geographical variation.

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and a 4 year anomaly from 2009–2013. This 4 year anomaly, when analyzed from the beginning of the Aura MLS time series, results in a positive linear O<sub>3</sub> trend in the mid-stratosphere over Lauder, in the opposite sense to the trend over this time period in the tropics. An improved understanding of how dynamical variations affect mid-stratospheric O<sub>3</sub> variations at this southern mid-latitude site is important for interpreting measurements from mid-latitude sites in terms of long-term global O<sub>3</sub> change.

## 2 Measurements

The MOPI1 instrument has been making measurements of stratospheric O<sub>3</sub> from the Network for the Detection of Atmospheric Composition Change (NDACC) station at Lauder, New Zealand (45.0° S, 169.7° E) since 1992. During these 22 years the MOPI1 instrument has been used as a ground-based reference for a number of satellite instruments. Satellite measurements can provide a global perspective for the MOPI measurements, and we will use measurements from Aura MLS to provide such a global perspective for ozone variations since 2004. Here we present a brief description of both the ground-based microwave and satellite measurement techniques.

### 2.1 Ground-based microwave measurements

Each MOPI instrument has a heterodyne receiver coupled to a 120 channel filter spectrometer, which are used to measure the emission spectrum of a line produced by a thermally excited, purely rotational ozone transition at 110.836 GHz (2.7 mm wavelength). The spectral intensities and measurements of the tropospheric thermal emission are calibrated with black body sources at ambient and liquid nitrogen temperatures. The tropospheric opacity is calculated from hourly emission measurements. The experimental technique was described in Parrish et al. (1992), and technical details on the instrument used for this work are given in Parrish (1994). MOPI measurements

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have been employed in several validation and trend studies (e.g. Boyd et al., 2007; Steinbrecht et al., 2009).

MOPI observations are made continuously, with selected spectra chosen based on the absolute tropospheric opacity and its variability, where limits for these values are set empirically at each site. This allows measurements in weather ranging from clear to some overcast conditions. The standard MOPI retrieval product, which will be used here, provides retrievals four times per day, using 6 h data periods starting at local midnight. The diurnal variations in the  $O_3$  measurements from the MOPI2 instrument at Mauna Loa (using a 1 h retrieval product) have been validated and compared to models (Parrish et al., 2014), and measurements and model values generally agree to better than 1.5 % of the midnight value. For this study we make use of three of the four daily retrievals, and do not include the daytime afternoon measurements (i.e. 12:00–18:00 LT). This selection has been made because, at Lauder, these measurements show a slightly anomalous vertical profile in the mid-stratosphere, with values at 10 hPa ~ 3 % lower than at other times of the day. We believe that these variations may be caused by the strong thermal cycles in the building, especially in the afternoon.

In Fig. 1 we show a typical spectrum and  $O_3$  profile retrieval from the MOPI1 instrument. As described in Parrish et al. (1992), the measurement shown is obtained using a switching technique (Parrish et al., 1988) so that the spectrum used in the retrievals is the difference between measurements being made at a low elevation angle (in this case  $22.6^\circ$  above the horizon), and measurements made near the zenith through an attenuating sheet of plexiglass. The measured temperature in the two positions is approximately balanced, and any remaining slope or offset in this difference spectrum is removed before retrieving the  $O_3$  profile. The  $O_3$  mixing ratio profiles are retrieved from the spectra using an adaptation of the optimal estimation method of Rodgers (1976), discussed in Parrish et al. (1992) and Connor et al. (1995). Error analysis techniques are discussed in the latter paper. The independent variable in the retrieval system is pressure.

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In Fig. 2 we show typical averaging kernels for the MOPI1 version 6 retrievals. The vertical resolution of the MOPI1 measurements is  $\sim 7\text{--}8$  km (FWHM) at 10 hPa, slightly coarser than at Mauna Loa, where the MOPI2 resolution at this pressure level is  $\sim 6$  km. The MOPI retrievals have a measurement contribution of near 100 % at 10 hPa, as defined by the technique of Connor et al. (1995). Unless otherwise indicated the satellite measurements which will be shown have not been convolved with the MOPI averaging kernels.

### 2.2 Satellite measurements near Lauder

We compare MOPI  $\text{O}_3$  measurements with observations from three satellites that provide coincident measurements. Two of these are solar occultation instruments, which make  $\sim 15$  sunrise and  $\sim 15$  sunset high vertical resolution ( $\sim 1$  km) profile measurements each day, generally at different latitudes, with the latitude bands varying differently over the course of the season depending on the satellite orbit. The majority of MOPI-satellite comparisons come from Aura MLS, which provides measurements over all latitudes between  $82^\circ$  S and  $82^\circ$  N on a daily basis.

The SAGE II instrument was launched in October 1984 aboard the Earth Radiation Budget Satellite, and continued making measurements through August 2005. It consisted of a seven channel solar photometer using ultraviolet and visible channels between  $0.38$  and  $1.0\ \mu\text{m}$  in solar occultation mode to retrieve atmospheric profiles of ozone, water vapor, nitrogen dioxide and aerosol extinction. Measurements were made over a latitude range from  $80^\circ$  S to  $80^\circ$  N. The measurements are retrieved as ozone number density as a function of altitude, but are also provided as ozone mixing ratio as a function of pressure. The version 6.1 data is described in Wang et al. (2002).

The v7.00 dataset (Damadeo et al., 2013), released in December 2012, is used in this analysis. This latest processing version implements an algorithm that is consistent across all SAGE missions. The most significant change in the new version is that, whereas the previous SAGE retrievals made use of the meteorological profiles from the Climate Prediction Center (CPC) NCEP analysis, the new retrievals make use of the

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Modern Era Retrospective-Analysis for Research and Applications (MERRA) reanalysis. Retrievals using the meteorological profiles from the MERRA reanalysis do show significantly different  $O_3$  mixing ratios as a function of pressure in the upper stratosphere and mesosphere (pressures below  $\sim 4$  hPa), however for the pressure levels of interest for this study the differences between the v6.1 and v7.00 SAGE  $O_3$  retrievals are insignificant.

HALOE solar occultation measurements of  $O_3$  are available from 1991–2005. The latitude bands drifted daily so that near global latitudinal coverage was provided in both sunrise and sunset modes five times over the course of a year. The trends in the HALOE  $O_3$  measurements have been compared against SAGE II (Nazaryan et al., 2005) and differences have been found to be on the order of less than  $0.3\% \text{ year}^{-1}$  in a majority of latitude bands at 25, 35, 45, and 55 km.

Aura MLS measurements of  $O_3$  and  $N_2O$  are available since 2004. The stratospheric  $O_3$  product has been validated by Froidevaux et al. (2008). The vertical resolution of the MLS  $O_3$  measurements at 10 hPa is  $\sim 3$  km. The  $N_2O$  measurements have been validated by Lambert et al. (2007) and have a vertical resolution of  $\sim 4$  km. UARS MLS measurements of  $O_3$  are available from 1991–1999, and were validated by Froidevaux et al. (1996) for the v2.2 retrievals and by Livesey et al. (2003) for the v5 retrievals.

### 3 The MOPI $O_3$ timeseries

In Fig. 3 we show the monthly anomalies at selected pressure levels for the entire MOPI1 data record. The anomalies are calculated by first fitting the data with a sinusoidal seasonal cycle (including annual and semi-annual terms) and then subtracting this seasonal cycle from the data. There are two interesting features that particularly stand out and which we will further discuss. One feature is the very large positive anomaly that occurs in June 2001 at 31.6, 17.7, and 10.0 hPa (green boxes). The second interesting, and much longer-term, feature is the positive anomaly at 10.0 and 5.6 hPa from August 2009 through July 2013 (red boxes). During this period the mean

monthly  $O_3$  anomaly at 10 hPa is 0.32 ppmv, only 7 of the 47 measurement months show a negative anomaly, and the 3 month smoothing never shows a negative anomaly. This period ends with a sharp drop in  $O_3$  in August 2013.

### 3.1 Unusually high mid-stratospheric $O_3$ in June 2001

5 Since there are no MLS measurements available to document the global variation in  $O_3$  during June 2001, we use Tracer Equivalent Latitude (TrEL) simulations in order to better understand the June 2001 anomaly over Lauder. TrEL is determined from isentropic passive tracer advection calculations on the sphere as described by Allen and Nakamura (2003). The tracer mixing ratio is converted to an equivalent latitude  
10 by matching the area enclosed by tracer contours to that enclosed by an equivalent latitude line. Specific details of the 35 year TrEL calculation used for this paper, based on MERRA winds, are provided by Allen et al. (2012). The average TrEL in June 2001 over Lauder on potential temperatures surfaces from 550 to 850 K ( $\sim 35$  to 10 hPa) was the highest (i.e., most equatorward TrEL) June average observed throughout the  
15 35 year period from 1980–2014. At 650 K the mean TrEL value at Lauder in June 2001 was  $\sim 31^\circ$  S, indicating unusually tropical air relative to the latitude of the site.

From  $\sim 30$  to  $\sim 3$  hPa,  $O_3$  generally increases from pole to equator throughout the year, hence the unusually high (equatorward) TrEL is associated with high  $O_3$ . We calculated the monthly  $O_3$  latitudinal gradient from  $45$  to  $35^\circ$  S from the MLS measurements, and found that from 20–10 hPa, this gradient peaks during the months  
20 of March–June. Hence  $O_3$  mixing ratios measured over Lauder are particularly sensitive to changes in TrEL during these months. Thus the unusual high  $O_3$  anomaly in June 2001 seems to be the result of an unusual amount of equatorward air over Lauder at a time when  $O_3$  variations are particularly sensitive to such transport.

25 To better explain the dynamics that caused this unusually high TrEL (and hence  $O_3$ ) over Lauder in June 2001 we show, in Fig. 4, the TrEL at 650 K ( $\sim 20$  hPa) for the entire Southern Hemisphere during the second half of June 2001. Low TrEL occurs throughout the polar vortex, also identified by streamlines (white contours) circling the

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pole. A strong anticyclone, identified by closed streamlines and elevated TrEL (marked with red “H”), moves eastward from 16–20 June, before remaining relatively stationary for the next 10 days. This is an unusually strong “blocking” type pattern that kept high TrEL/high O<sub>3</sub> air over Lauder. While this quasi-stationary stratospheric “Australian high” signature is known to occur in the SH spring (Harvey et al., 2002), this is by far the largest and most persistent episode observed in June in the 35 year TrEL simulation.

### 3.2 Unusually high mid-stratospheric O<sub>3</sub> from August 2009 through July 2013

The 4 years of elevated O<sub>3</sub>, from the increase in 10 hPa O<sub>3</sub> in 2009 to the sharp drop in 2013, occurred during the period when coincident Aura MLS measurements are available. Aura MLS overpasses occur near 01:15 and 14:30 LST at this latitude, but since we are not using the MOPI1 measurements from 12:00–18:00 only the 01:15 overpass coincidences are shown in Fig. 5. There are two daily 01:15 MLS overpasses that fulfill the ±6° longitude criterion. The Aura MLS measurements show very good agreement with the MOPI measurements over the entire period, and both instruments show the large O<sub>3</sub> increase in mid-2009 and the large decrease in mid-2013. The increase in mid-2009 shows that mixing ratios increased from values near the lowest observed during the Lauder winter over this 10 year period, to values at, or near, the highest observed in a Lauder summer. There was a month-long gap in the MOPI1 measurements in July 2013, but the observed drop in the coincident MLS measurements is very similar to that of the MOPI1 measurements between June and August 2013. Both instruments show that, in August 2013, the O<sub>3</sub> values were lower than they had been since 2009.

The unusual nature of the 2009–2013 period is even more clearly emphasized in Fig. 6, which shows annual average anomalies from MOPI and from four satellite instruments which measured O<sub>3</sub> over extended periods since the early 1990’s. Note that, with the exception of the MOPI measurements, the O<sub>3</sub> anomalies shown in Fig. 6 are zonally averaged. Since Fig. 6 shows annual averages it helps to emphasize the Quasi-Biennial Oscillation (QBO). The annual-average MOPI measurements show local minima in 1997, 1999, 2002, 2004, late 2006/early 2007, and late 2008/early 2009.

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Following the minimum in late 2008/early 2009 the annual  $O_3$  rises and remains well above the long-term average until 2013. All of the measurements shown in Fig. 6 are provided on their native grid. For the SAGE II measurements the native grid is altitude, and results are shown at 30 km. For HALOE, UARS and Aura MLS, and MOPI we show results at a 10 hPa.

### 4 $O_3$ and $N_2O$ at Lauder and at the equator

In order to better understand the global implications of the observed  $O_3$  variations over Lauder, we investigated how the variations in  $O_3$  observed over Lauder compare with global changes in  $O_3$ . First, we calculated monthly average MLS profiles for  $2^\circ$  latitude bins. We then calculated a climatological average for each month at each of these latitudes based on the MLS measurements. Finally, using this climatology, we calculated an anomaly for each month and latitude.

In Fig. 7 we show the correlation (based upon the months where both Aura MLS and MOPI measurements were available) between monthly  $O_3$  anomalies measured at 10 hPa by MOPI and the monthly  $O_3$  anomalies measured at different pressures and latitudes by MLS. Figure 7 shows an anti-correlation ( $r < -0.5$ ) between the 10 hPa  $O_3$  variations measured by MOPI at Lauder, and the 10 hPa  $O_3$  variations measured by MLS at the equator. The correlation between the 10 hPa  $O_3$  MLS measurements at  $45^\circ S$  and the 10 hPa  $O_3$  MLS measurements at the equator is similar. Because of measurement differences resulting from errors, differences in vertical resolution, and because the MOPI1 measurement is for local conditions near Lauder and not a zonal average, the strongest correlation between the 10 hPa MOPI measurements at Lauder and the Aura MLS  $O_3$  measurements falls at a slightly different geophysical location, in this case slightly equatorward and at a slightly higher pressure level.

Of course the temporal correlations shown in Fig. 7 give no indication of the time period over which the correlation is taking place (by using anomalies we have eliminated only the seasonal cycle), and could, e.g., represent QBO-like variations, solar



lag relative to the 30 hPa QBO wind anomaly, but the generally positive correlation between this 30 hPa wind anomaly and  $O_3$  and  $N_2O$  mixing ratios suggests that an anomalously fast ascent rate near 10 hPa is associated with westerly (positive) winds at 30 hPa.

If we look at Fig. 8 in more detail, we see that in 2006, 2008, 2010, and 2013 there is a sharp increase in  $O_3$  and  $N_2O$  from  $5^\circ$  S to  $5^\circ$  N near the middle of the year, while in 2007, 2009, and 2011 there is a sharp decrease. Following these sharp changes the equatorial anomaly often remains high (or low) until the next June/July period. Thus the variation is often nearly biennial except for the absence of a change in sign for the  $O_3$  and  $N_2O$  anomalies from  $5^\circ$  S to  $5^\circ$  N in June/July 2012.

While the variation shown in Fig. 8 seems to be primarily nearly biennial, the period from 2009–2013 shows, as is apparent in the annual averages shown in Fig. 6, lower average equatorial  $O_3$  and  $N_2O$  mixing ratios than are observed from 2004–2008. Figure 8 shows that the  $5^\circ$  S to  $5^\circ$  N  $O_3$  and  $N_2O$  mixing ratios have both lower maxima and lower minima at a similar phase of the QBO. While these equatorial  $N_2O$  and  $O_3$  anomalies are correlated with the phase of the QBO wind anomalies, it is not clear whether or not the unusually low  $O_3$  and  $N_2O$  mixing ratios in 2009–2013 are associated with unusual QBO wind anomalies.

There are some peculiarities in the QBO winds during the 2009–2013 period. For instance, the westerly wind anomalies in 2010 do not reach quite the same maximum strength as in the other four cycles during this period ( $16.0 \text{ ms}^{-1}$  in August 2010 is the lowest maximum since 1992). The easterly 30 hPa wind anomalies in 2009/10 are, unusually, stronger than  $20 \text{ ms}^{-1}$  for only 3 months before an unusually fast transition back to westerly winds, while the  $21.4 \text{ m s}^{-1}$  maximum easterly wind anomaly in 2012 is the weakest over the four cycles shown. The 30 hPa wind anomalies during the 2008–2013 period persist for slightly longer than usual. The winds switched from easterly to westerly in March 2008, August 2010, and March 2013, producing QBOs of length 29 months and 31 months respectively.

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The 10 hPa  $O_3$  and  $N_2O$  anomalies at 40 to 50° S are not as strongly correlated as at the equator (see Fig. 4, Nedoluha et al., 2015), but nonetheless there is clearly a positive correlation between the anomalies of these two species. Not unexpectedly, given the correlations shown in Fig. 7, these southern mid-latitude anomalies show variations that are usually opposite to those seen at the equator. Most clearly the sharp changes in June/July are anti-correlated with those near the equator. Figure 8 shows that, like the  $O_3$  values that have been shown previously, the  $N_2O$  values over latitudes near Lauder are elevated from 2009–2013.

The anomalies in  $O_3$  and  $N_2O$  at 40 to 50° S are likely to be caused by the rate at which  $N_2O$  moves from the tropics to the southern mid-latitudes. Increased exchange between these latitudes will bring higher  $N_2O$  air from the tropics to the southern mid-latitudes. The same tropical 30 hPa westerly winds which are associated with the increased ascent rate in the tropics seem to be correlated with a decrease in transport from the tropics into the Southern Hemisphere, resulting in an anti-correlation between  $N_2O$  (and hence  $O_3$ ) anomalies at the tropics and the southern mid-latitudes.

In order to provide a global perspective on the 2009–2013 anomalies, we have fit the MLS data globally from August 2004 through May 2013 with an 8 parameter fit including annual and semi-annual sine-waves, the 30 and 50 hPa QBO winds, and a linear trend term. The results of these fits are shown in Fig. 9. Since there is no clear correlation between the amplitude of the QBO wind variation and the depth of the  $N_2O$  and  $O_3$  changes in 2009–2013, these changes are fit by the linear trend term. The  $O_3$  linear trend fit plot (Fig. 9a) has been shown previously in Nedoluha et al. (2015), where it was shown that the decrease observed at 10 hPa near the equator has been occurring for more than 20 years. While a linear trend is clearly a very coarse representation of the MLS data from 2004–2013, it does allow us to show the strong global correlations between  $N_2O$  and  $O_3$  increases (and decreases) during this time period.

## 5 Discussion

We have investigated two unusual  $O_3$  variations which occurred in the mid-stratosphere over Lauder, New Zealand during the 22 years of ground-based microwave measurements from the site. One feature is the very large positive  $O_3$  anomaly that was observed by the MOPI instrument in June 2001. We associated  $O_3$  anomaly with an unusual blocking pattern which keeps air from more equatorial latitudes (with high ozone) over Lauder for much of this month. The very unusual nature of this event was emphasized by comparing the average TrEL in June 2001 over Lauder on potential temperatures surfaces from 550 to 850 K ( $\sim 35$  to 10 hPa) with values found in other years. It was found that the TrEL in June 2001 was higher (i.e., more equatorward TrEL) than in any other June throughout the 35 year period 1980–2014.

The second interesting, and much longer-term, feature is the positive  $O_3$  anomaly near  $\sim 10$  hPa which persists over southern mid-latitudes from 2009–2013. During this period  $N_2O$  in this region is also unusually high, and the same chemical-dynamical relationship that causes the very strong  $N_2O$ - $O_3$  correlation in the tropics is likely the cause of the high  $O_3$ . Briefly,  $N_2O$  decreases rapidly both as a function of increasing altitude and increasing distance from the tropics due to photodissociation and oxidation. Thus the high  $N_2O$  at southern mid-latitudes from 2009–2013 suggests that air was transported into this region from the tropical lower stratosphere more quickly during this period, thus decreasing the amount of photodissociation and oxidation of  $N_2O$ . At the same time, air was being transported more slowly into the tropical 10 hPa region. The mid-2013 decrease in mid-latitude  $N_2O$  suggests that air is now again being transported more quickly upwards in the tropics as opposed to being shifted towards southern mid-latitudes, but it remains to be seen whether this is a brief interruption, a halt, or a reversal of a decadal scale trend.

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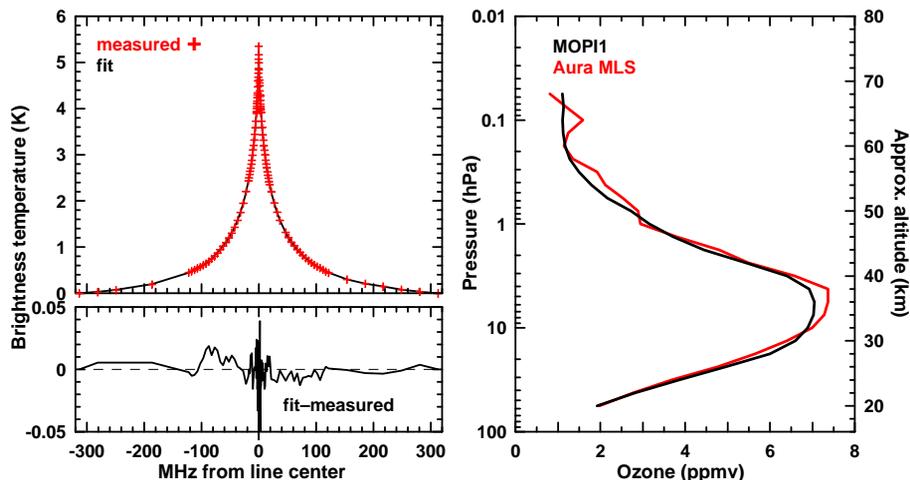
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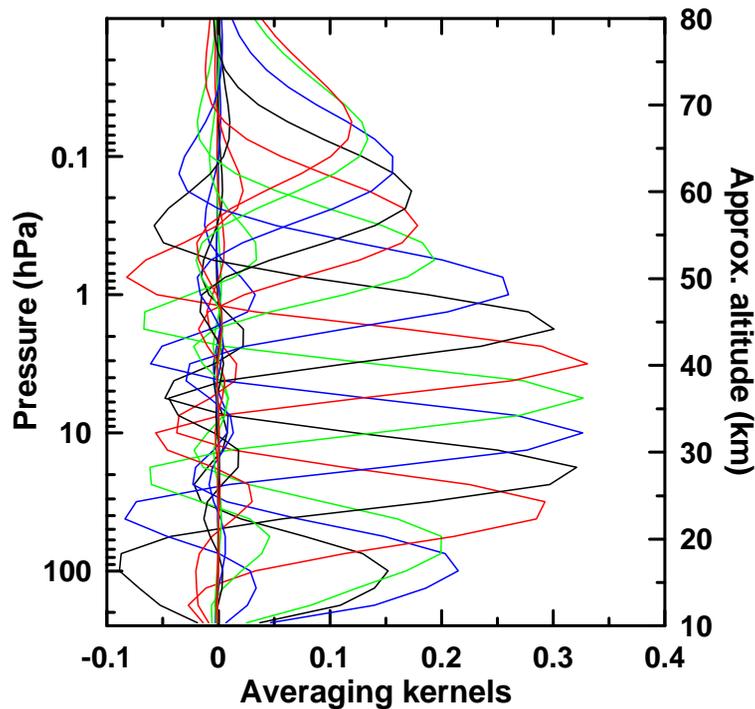
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**Figure 1.** Top left: the spectrum centered at 110.836 GHz as measured by MOP11 from Lauder over 3 h on 11 March 2014 (red crosses), and the model fit to this spectrum (black line). Bottom left: the residual difference between the measured and modeled spectrum. Right: the retrieved O<sub>3</sub> profile from MOP11 (black) and from a coincident Aura MLS measurement (red).

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**Figure 2.** Typical averaging kernels for the MOPI1 instrument based on 6 h of spectral integration. Averaging kernels are shown for every second level, with the averaging kernels in blue shown at 100, 10, 1, and 0.1 hPa.

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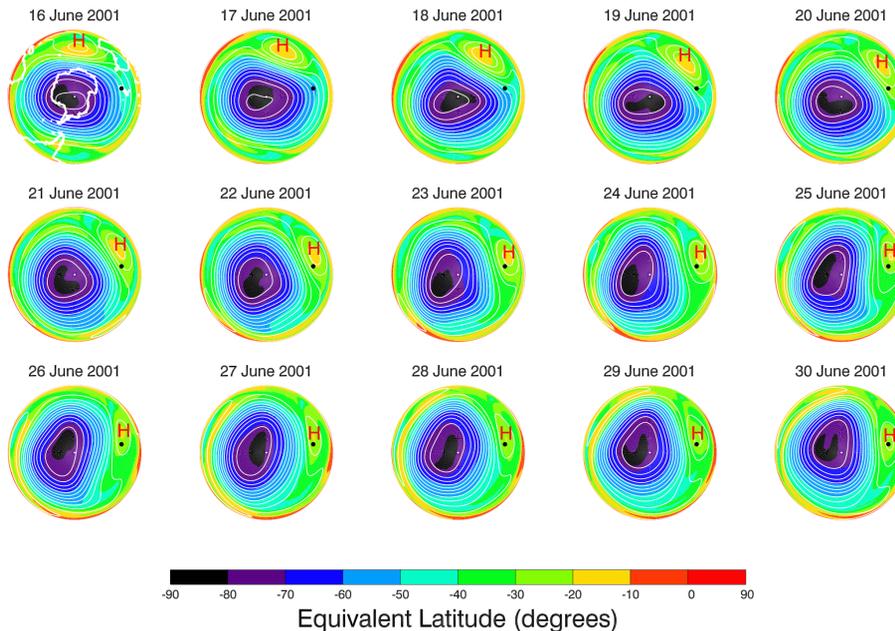
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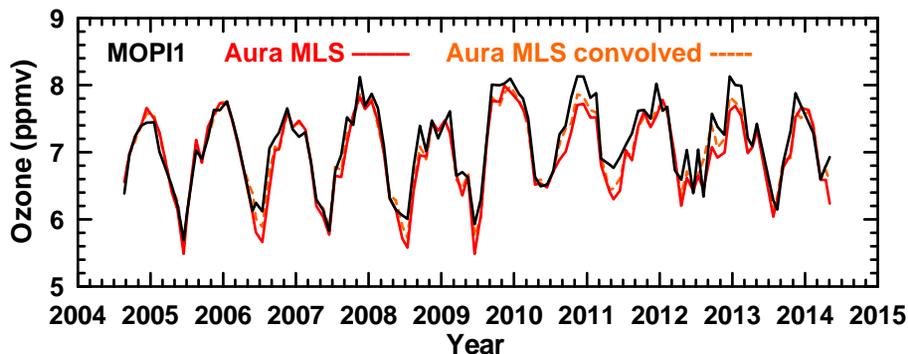
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**Figure 4.** The tracer equivalent latitude (see text) for the Southern Hemisphere at 650 K. The location of Lauder ( $45^{\circ}$  S,  $169.7^{\circ}$  E) is indicated by a black dot. White contours are 650 K streamlines at constant intervals. The red “H” indicates the location of strong anticyclonic circulation.

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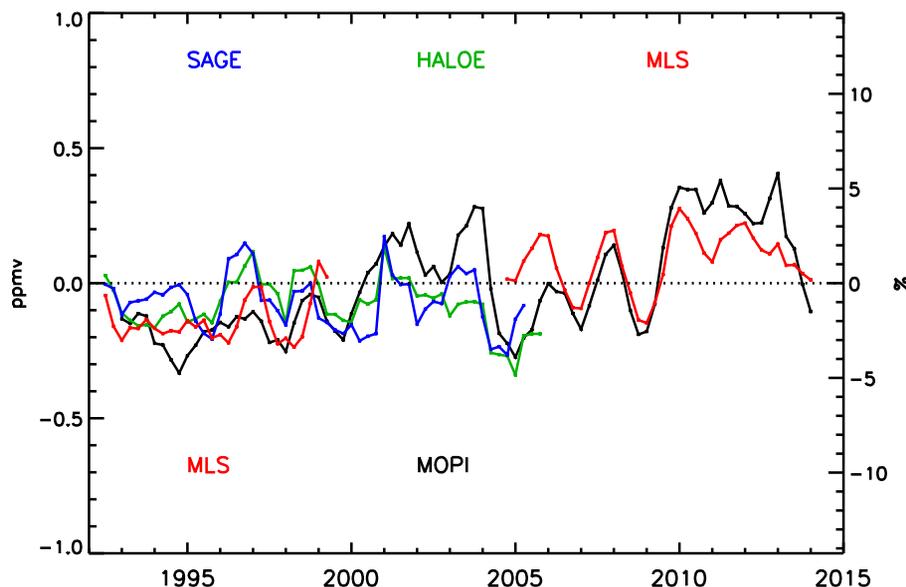


**Figure 5.** Monthly ozone averages for MOPI1 (black), Aura MLS (red), and Aura MLS convolved with the MOPI1 averaging kernels (dashed orange) measurement pairs at 10 hPa. Measurements are shown when there is an MLS measurement taken within  $\pm 1^\circ$  latitude and  $\pm 6^\circ$  longitude within 6 h of a MOPI1 measurement.

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**Figure 6.** Annual average ozone anomalies at 10 hPa (30 km for SAGE II) shown 4-times annually (January–December, April–March, July–June, and October–September). Results are shown for SAGE II (blue), HALOE (green), UARS and Aura MLS (both red), and MOPI1 (black). Satellite measurements (latitudinal averages from 40–50° S) have been offset so that the average ozone matches that of MOPI1 during the period of coincidence.

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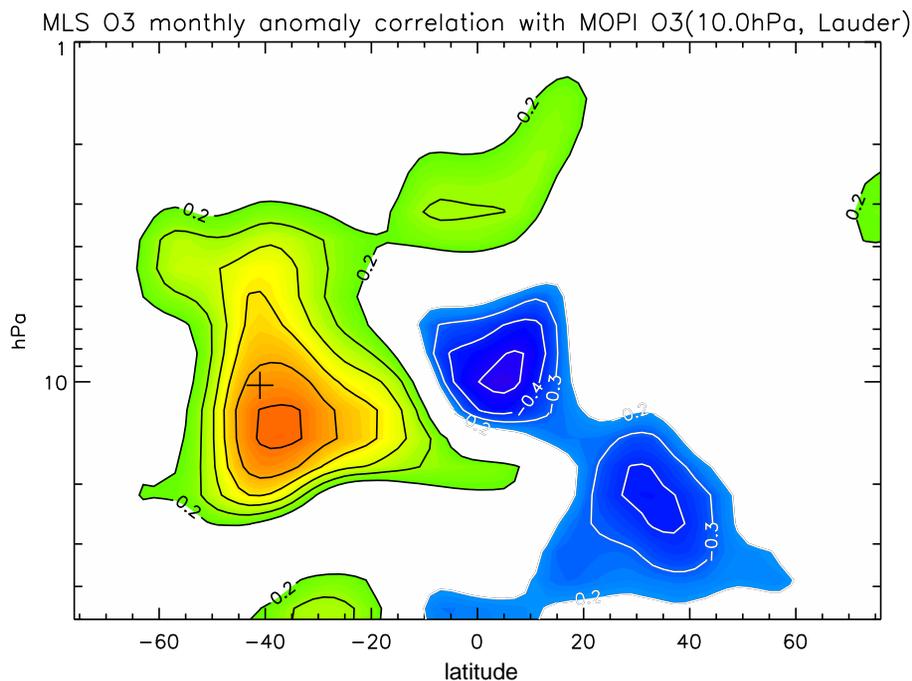
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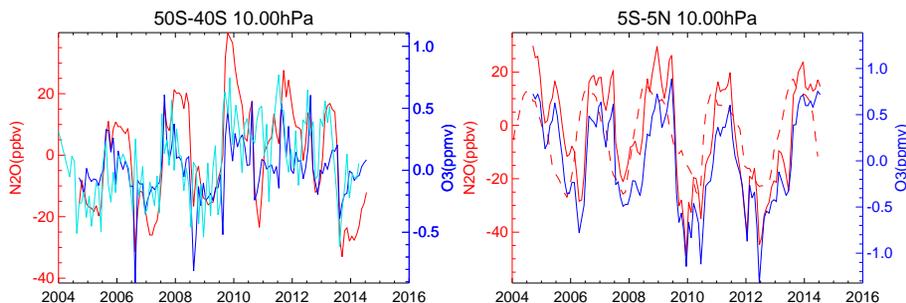
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**Figure 7.** The correlation coefficient of the monthly MLS O<sub>3</sub> anomalies with the monthly anomalies of the MOPI1 O<sub>3</sub> measurements at 10 hPa. Contours are shown at intervals of 0.1 for correlation coefficients  $> 0.2$  or  $< -0.2$ . The cross represents the latitude of Lauder at 10 hPa.

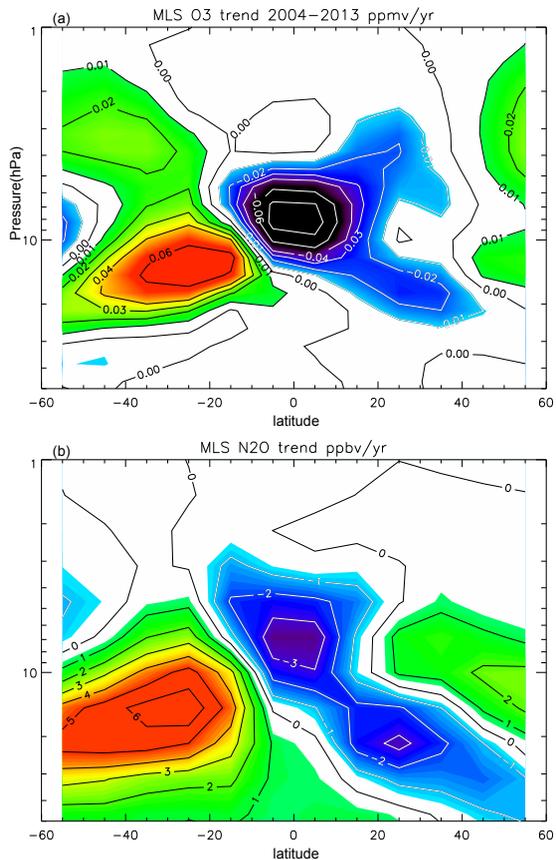
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**Figure 8.** Monthly average anomalies for  $\text{N}_2\text{O}$  (red) and  $\text{O}_3$  (blue) as measured by MLS at 10 hPa within  $5^\circ$  of the Lauder latitude ( $45^\circ\text{S}$ ) (left) and within  $5^\circ$  of the equator (right). The left hand plot also shows the monthly average  $\text{O}_3$  anomalies (based on the 2004–2014 averages) for MOPI (cyan). The right hand plot also shows (dashed red line) the 30 hPa QBO index in  $\text{m s}^{-1}$ , using the same numerical scale as the  $\text{N}_2\text{O}$  in ppbv.

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**Figure 9.** Linear trend fits to MLS O<sub>3</sub> (a) and N<sub>2</sub>O (b) measurements from August 2004 through May 2013. Contour lines for O<sub>3</sub> are shown at  $\pm 0.01$ , 0.02, 0.03, 0.04, 0.06, 0.08 ppmv yr<sup>-1</sup>. Contour lines for N<sub>2</sub>O are shown at intervals of 1 ppbv yr<sup>-1</sup> for N<sub>2</sub>O. The O<sub>3</sub> figure is from Nedoluha et al. (2015).