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**Exceptional  
emissions of NH<sub>3</sub> and  
HCOOH in the 2010  
Russian wildfires**

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# Exceptional emissions of NH<sub>3</sub> and HCOOH in the 2010 Russian wildfires

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## Abstract

In July 2010, several hundred forest and peat fires broke out across Central Russia during its hottest summer on record. Here, we analyze these wildfires using observations of the Infrared Atmospheric Sounding Interferometer (IASI). Carbon monoxide (CO), ammonia (NH<sub>3</sub>) and formic acid (HCOOH) total columns are presented for the year 2010. Maximum total columns have been observed reaching over 40 (for CO and HCOOH) and 200 (for NH<sub>3</sub>) times higher than typical background values. The temporal evolution of NH<sub>3</sub> and HCOOH enhancement ratios relative to CO are presented. Strong evidence of secondary formation of HCOOH is found, with enhancement ratios exceeding 10 times reported emission ratios in fresh plumes. We estimate the total emitted masses for the period July–August 2010 over the center of Western Russia; they are 19–33 Tg (CO), 0.7–2.6 Tg (NH<sub>3</sub>) and 0.9–3.9 Tg (HCOOH). For NH<sub>3</sub> and HCOOH, these quantities are comparable to what is emitted in the course of a whole year by all extratropical forest fires.

## 1 Introduction

Biomass burning is a major source of atmospheric trace gases and aerosols (Crutzen and Andreae, 1990). Boreal forest fires contribute significantly to these emissions and are also important as they can severely affect air quality in populated regions of the Northern Hemisphere (Langmann et al. (2009) and references therein). Many laboratory, field and aircraft studies have been conducted to characterize the composition of fire plumes (Andreae and Merlet, 2001; Akagi et al., 2011), but uncertainties remain due to the lack of measurements over larger areas or periods; this is particularly true for the total emission of reactive species.

The polar orbiting Infrared Atmospheric Sounding Interferometer (IASI), which measures globally the Earth's outgoing infrared radiation at a high spectral resolution and low instrumental noise (Clerbaux et al., 2009), is an excellent tool for observing and

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analyzing trace gas emissions from strong fire events (Coheur et al., 2009; Turquety et al., 2009; Clarisse et al., 2011). Its potential to monitor CO from vegetation burning has frequently been used. Measurements of other fire emission products, such as NH<sub>3</sub> or volatile organic compounds, have in contrast not been exploited much. Such measurements can directly improve our knowledge of fire emissions and fluxes. These currently rely on the remote sensing of fire counts (e.g. Kasischke et al., 2003) and fire radiation power (e.g. Wooster et al., 2005), with application of tabulated average values of emission factors. Here we study simultaneous CO, NH<sub>3</sub> and HCOOH fire emissions in the 2010 Russian fires. It is the first time that such a large spatial and temporal scale case study is undertaken for NH<sub>3</sub> and HCOOH. IASI CO observations from these fires have been reported in two recent studies (Yurganov et al., 2011; Krol et al., 2012). Here we include it as well for completeness, as an opportunity to test our methodology and for the calculation of enhancement ratios. We use total columns of NH<sub>3</sub> and CO from the FORLI (Fast-Optimal Estimation Retrievals on Layers for IASI) near-real-time retrieval software (Hurtmans et al., 2012) and a retrieval scheme based on brightness temperature differences (BTDs) for HCOOH (Razavi et al., 2011). Note that only the daytime data, which are more sensitive to surface concentrations, have been used. In addition, NH<sub>3</sub> and HCOOH are retrieved only when enough spectral information is available (see Sect. 5).

Formic acid is an important volatile organic compound in the atmosphere; but uncertainties in its global budget remain large (Stavrakou et al., 2011). The majority is photochemically produced from non-methane hydrocarbons. It is also released directly in the atmosphere from terrestrial vegetation and biomass burning, for which the total (primary) emissions account for 1–4 Tgyr<sup>-1</sup> (Paulot et al., 2011; Stavrakou et al., 2011). NH<sub>3</sub> is the most important base in the atmosphere and is released primarily by agriculture (76 %), natural sources (19 %) and fossil fuel burning (5 %). Natural fires only account for 4 % of the total emissions totaling 5–15 Tgyr<sup>-1</sup> (European Commission, Joint Research Centre (JRC)/Netherlands Environmental Assessment Agency (PBL), 2011).

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During the summer of 2010, Central European Russia endured the driest and hottest July ever recorded. Witte et al. (2011) reported anomalously high temperatures (35–41 °C, compared to average temperatures of 15–20 °C usually observed in the summer season) and low relative humidity (9–25 %) from mid–June to mid–August 2010, from radiosonde measurements in Western Russia. The intense heat and drought wave initiated hundreds of wildfires in European Russia and burned forests, grasslands, croplands, fields and urbanized areas across a massive region around Moscow (Shvidenko et al., 2011), severely affecting the air quality (Elansky et al., 2011a,b; Fokeeva et al., 2011; Konovalov et al., 2011; ZvyaginsteV et al., 2011; Golitsyn et al., 2012). The unusual magnitude of the event was revealed for instance by high values of the fire radiative power, which reached 19 000 MW close to Moscow (Witte et al., 2011); this is 400 times more than what is observed on average in Russia (Wooster and Zhang, 2004). Along this, large emissions of aerosols and trace gases such as CO, tropospheric nitrogen oxides (NO<sub>2</sub>), ozone (O<sub>3</sub>) and formaldehyde (HCHO) have been reported, using measurements from a variety of satellite sounders (Konovalov et al., 2011; Yurganov et al., 2011; Chubarova et al., 2012; Huijnen et al., 2012; Krol et al., 2012).

In the next section we present the CO, NH<sub>3</sub> and HCOOH total column maximums observed within a 2 month period (1 July to 1 September) in the summer of 2010, and the mean total columns for these species for the year 2010 over the area 40°–75° N and 30°–150° E. Section 3 presents the calculation and temporal evolution of the enhancement ratios of NH<sub>3</sub> and HCOOH relative to CO. In Sect. 4, we calculate the total masses and the extra burdens due to fires, the fluxes in Tgday<sup>-1</sup>, as well as the total emitted masses over the entire burning period. These issues of surface sensitivity, the related measurement uncertainties and how they affect the presented results are briefly discussed in Sect. 5. In Sect. 6 we present our conclusions.

## 2 Distributions and time series of total columns

Average total columns of CO, NH<sub>3</sub> and HCOOH are shown in Fig. 1 over Europe and Asia in grid cells of 0.5° × 0.5°, from 27 July (corresponding to the first IASI observation of the fire plumes) to 27 August (a few days after the end of the fires) of 2010. There is some difference between the spatial extents of CO, NH<sub>3</sub> and HCOOH. These can be explained by differences in lifetime (see Sect. 4), but also by the fact that we have a more limited sensitivity above 55° N for NH<sub>3</sub> and HCOOH related to thermal contrast (see Sect. 5). Note also that the red box in Fig. 1, extending from 40°–75° N to 30°–150° E, is the same area as the one considered in the study of Yurganov et al. (2011). This choice was made intentionally to facilitate the comparison of our results to that earlier study.

Very high total columns were observed during the event for the three species. Figure 2a shows the temporal evolution of the retrieved total column maximums during a two months period which includes the fires (1 July to 1 September 2010) in the area corresponding to the red box in Fig. 1; Fig. 2b shows the mean total columns for the year 2010 covering the same area. From Fig. 2a one sees a significant increase of emissions at the end of July which corresponds to the beginning of the fires (more precisely on July 27) reaching a maximum emission on 2, 10 and 15 August, respectively for CO, NH<sub>3</sub> and HCOOH which represent 40 times for CO and HCOOH and 200 times for NH<sub>3</sub> higher than typical background values. These maximum emissions are followed by a rapid decrease in the concentrations of these species and return to typical background values after 20 August. This decrease is especially pronounced for NH<sub>3</sub> which is the shortest-lived species of the three. The mean CO total columns, presented here (Fig. 2b), are in good agreement with recent studies of Yurganov et al. (2011), Witte et al. (2011) and Mielonen et al. (2012) which reported mean values between 2 and 4 10<sup>18</sup> molec cm<sup>-2</sup> in different areas around Moscow. The higher than average CO total columns observed above Central Russia between January and June 2010 (Fig. 2b) are due to the usual seasonality of that species, maybe intensified in springtime by the

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transport of fire plumes from outside the studied region (Parrington et al., 2012). Similarly, the progressive increase in  $\text{NH}_3$  and  $\text{HCOOH}$  from March to the start of the fires can be explained with the increase in spring of agricultural activities (for  $\text{NH}_3$ ) and plant growth (for  $\text{HCOOH}$ ). As explained above, the lack of retrieved  $\text{HCOOH}$  columns in autumn and winter is due to too low thermal contrasts, not exceeding the 5 K threshold for the conditional retrievals.

### 3 Enhancement ratios

Fires emit trace gases in amounts that depend on the specific fire conditions and source material. A useful parameter to quantify trace gas emissions from fires is the emission factor, which is the amount of emitted species per amount of dry fuel (in  $\text{g kg}^{-1}$ ) (Reid et al., 2009). Emission factors are measured both in situ (e.g. (Delmas et al., 1995; Yokelson et al., 1999) and in laboratories (Andreae et al., 1988; Koppmann et al., 2005). Large compilations of emission factors, grouped in different source types, can be found in Andreae and Merlet (2001) and Akagi et al. (2011). A related parameter is the emission ratio of a trace gas  $X$ , which is the ratio of emitted molecules of  $X$  over the number of emitted molecules of a target species such as  $\text{CO}$ , used hereafter. Emission ratios can be calculated from emission factors by multiplying with the ratio of the molar masses  $M_{\text{CO}}/M_X$ . In fresh plumes, the emission ratio can be estimated from the ratio of concentrations, the so-called enhancement ratio (Goode and Yokelson, 2000):

$$\frac{\Delta X}{\Delta \text{CO}} = \frac{[X]_{\text{smoke}} - [X]_{\text{ambient}}}{[\text{CO}]_{\text{smoke}} - [\text{CO}]_{\text{ambient}}} \quad (1)$$

As the plume ages, the enhancement ratios will typically decrease for species with a lifetime shorter than  $\text{CO}$ . The time evolution of the enhancement ratio therefore gives insight into the chemical loss processes within the plume (Xiao et al., 2007). When a lot of observations are available, the enhancement ratios can be calculated for each

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measurement pair ( $[X],[CO]$ ), while an average enhancement ratio can be estimated from the slope of the linear regression  $\Delta X$  vs  $\Delta CO$  (Coheur et al., 2009).

Enhancement ratios for the Russian fires are shown in Fig. 3 over the area defined above. Figure 3a shows an example of the calculation of the enhancement ratio, the scatter plots of total columns of  $NH_3$  and  $HCOOH$  vs  $CO$  on 5 August. The high correlations (correlation coefficients around 0.75) are indicative of the common emission source for the three species. Note that the series of low  $NH_3$  columns ( $< 10^{16}$  molec $cm^{-2}$ ) in Fig. 3a, which are associated with rather high columns of  $CO$  ( $> 4 \cdot 10^{18}$  molec $cm^{-2}$ ), likely correspond to aged air masses. For this day of 5 August, the calculated enhancement ratios are 0.032 for  $NH_3$  and 0.021 for  $HCOOH$ . Enhancement ratios are calculated similarly for each day and the corresponding time evolution from 27 July to 24 August is shown in Fig. 3b. For  $NH_3$  enhancement ratios start around 0.015 on the 27 July and reach a maximum of 0.055 on the 10 August when  $NH_3$  total columns are largest (Fig. 2). After that, the enhancement ratio decreases gradually, pointing to a decrease in fire activity and a return to background  $CO$  columns. Less pronounced variations of the enhancement ratio are found for  $HCOOH$ .

As the 2010 Russian fires are associated with the burning of both peat lands (accounting for 30 % of emitted  $CO$  (Konovalov et al., 2011) and other types of vegetation, it is difficult to compare the results to a single emission ratio. We find, however, that for  $NH_3$  the range of values measured from IASI (0.010–0.055) is consistent with the values reported in the literature, which average 0.033 for extratropical forest fires, 0.035 for boreal forest fires and 0.097 for peat lands (Akagi et al., 2011). In contrast, the calculated values vary between 0.010 and 0.030 for  $HCOOH$ , and are one order of magnitude larger than emission ratios values in Akagi et al. (2011). A likely explanation for this discrepancy is rapid secondary production of formic acid within the fire plume (e.g. Yokelson et al., 2009). For instance; Goode and Yokelson (2000) and Yokelson et al. (2007) describe this process and report enhancement ratios in downwind plumes between 0.023 and 0.029, which are in much better agreement with the enhancement ratios found here. A more general interpretation of the enhancement ratios measured

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by IASI is, however, hampered by the continuous emission, production and complex transport patterns: the buildup of CO during the entire burning period could in particular mean that the real emission ratios are still higher than our reported enhancement ratios.

#### 4 Total mass

Daily observed total masses  $TM_X$  of species  $X$  have been calculated here as follows:

$$TM_X = \frac{M_X C_X S}{N_a} \quad (2)$$

where  $M_X$  is the molar mass of  $X$  (in  $\text{g mol}^{-1}$ ),  $C_X$  is the mean column retrieved from IASI daytime observations (in  $\text{molec cm}^{-2}$ , see Fig. 2b),  $N_a$  is the Avogadro number and  $S$  represents a defined surface area. For CO and HCOOH the latter corresponds to the red rectangle identified in Fig. 1 and is estimated at  $2.75 \times 10^{17} \text{ cm}^2$ . For  $\text{NH}_3$   $S$  is a smaller area ( $40^\circ\text{--}75^\circ \text{ N}$  and  $30^\circ\text{--}90^\circ \text{ E}$ ) and this is justified by the fact that FORLI- $\text{NH}_3$  is a conditional retrieval method, as discussed previously, and that the short lifetime of this species prevents its long-range transport. Figure 4 shows, on the left panel, the daily evolution of the total masses for the year 2010 (in black daily values and in red smoothed). The total masses reach in mid August maximums of 36.1 Tg for CO, 0.06 Tg for  $\text{NH}_3$  and 0.70 Tg for HCOOH.

After having calculated the daily total masses over the area of interest it is possible to estimate the total emissions by making some extra assumptions. Here we follow an approach similar to that of Yurganov et al. (2011). In a first step we estimate the daily total mass due to the fires (the so-called burdens) by subtracting total mass background values. The background values (Fig. 4 left panel in blue) were obtained by smoothing the minimum of the total masses observed in 10 day intervals. The resulting burdens are plotted in the middle panel in Fig. 4. Maximum burdens are calculated at around

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10 Tg for CO, 0.04 Tg for NH<sub>3</sub> and 0.35 Tg for HCOOH. Note that a 10 Tg of CO burden represents more than 85 % of total annual anthropogenic emissions in that region (Konovalov et al., 2011) and that the NH<sub>3</sub> and HCOOH maximum values reported here account alone for 5 and 20 % of the total annual emissions from extratropical forests fires (Andreae and Merlet, 2001). It should be noted though that for HCOOH the values in Andreae and Merlet (2001) represent primary emissions only.

From the burdens we can calculate emission fluxes assuming a simple box model and first order loss terms (Jacob, 1999):

$$E_{i+1}(X) = \frac{B_{i+1} - B_i e^{-t/\tau_{\text{eff}}}}{\tau_{\text{eff}}(1 - e^{-t/\tau_{\text{eff}}})} \quad (3)$$

Here  $E_i$  and  $B_i$  are respectively the flux and the burden on day  $i$ ,  $t$  is time between two observations (here one day) and  $\tau_{\text{eff}}$  is the effective lifetime of the species  $X$ . This lifetime term includes chemical losses, but also losses due to transport outside the considered area. Here, we assume that all burdens are due to the Russian fires and neglect potential inflow of excess columns. The calculation thus depends exclusively on the effective lifetime  $\tau_{\text{eff}}$  of the species. Following estimates of the CO effective lifetime given in Yurganov et al. (2011), the daily fluxes for CO have been estimated using two values of  $\tau_{\text{eff}}$  namely 7 and 15 days. As we do not have accurate knowledge of the effective lifetime of HCOOH and NH<sub>3</sub>, especially not in such an extraordinary event, we calculate also for these species the fluxes for a wide range of different lifetimes. This approach allows to conservatively estimate a range of total emissions compatible with the IASI measurements. For HCOOH, we use 2, 4 and 10 days following its global estimated lifetime of 3–4 days (see Stavrou et al. 2011 and references therein). For NH<sub>3</sub>, we use lifetimes of 6, 12 and 24 h as its lifetime ranges typically from a couple of hours to days (Dentener and Crutzen, 1994; Asman et al., 1998; Aneja et al., 2001). The right column of Fig. 4 shows the time evolution of the (smoothed) fluxes in Tg day<sup>-1</sup> of these species for these various lifetimes, from 29 June to 27 September 2010. The maximum values calculated this way are 1.41–1.87 Tg day<sup>-1</sup>, 0.04–0.07–0.15 Tg day<sup>-1</sup>

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and 0.07–0.11–0.19 Tg day<sup>-1</sup> for CO, NH<sub>3</sub> and HCOOH respectively, with the larger values logically for the smaller values of  $\tau_{\text{eff}}$ . Finally, total emissions were estimated from these fluxes between 15 July and 31 August 2010 and are 19–33 Tg (CO), 0.7–2.6 Tg (NH<sub>3</sub>) and 0.9–3.9 Tg (HCOOH). For NH<sub>3</sub> and HCOOH, this is comparable to what is normally emitted in the course of a whole year by all extratropical forest fires (Galloway et al., 2004; Stavrakou et al., 2011).

Several estimates of total emitted CO have been reported in the literature for the 2010 Russian fires, and these have been summarized in Table 1. These include on the one hand estimates derived from CO measurements (from the infrared sounders IASI, AIRS and MOPITT) possibly coupled with an atmospheric transport model, and on the other hand estimates derived from MODIS fire radiative power measurements. In the studies of Yurganov et al. (2011) and Fokeeva et al. (2011) who use direct CO measurements, allowance is made for the limited sensitivity of infrared instruments to boundary layer concentrations by adding an offset estimated from the difference between satellite and ground based measurements. For these studies the labels “standard” and “adjusted” indicate whether adjustments were made.

There is a large scatter in the data, both for the maximum daily emissions (1.2–2.4 Tg) and the total emitted CO (10–40 Tg). The standard estimates from CO measurements found in Yurganov et al. (2011) (26.2 Tg) and Fokeeva et al. (2011) (19–26 Tg) agree well with our estimates (19–33 Tg). This is not surprising as all three studies use similar approaches. Their adjusted values, with enhanced boundary layer concentrations are about 40 % higher. The total emissions reported here also agree well with the values 20–25 Tg given in Krol et al. (2012) and obtained using inverse modeling of IASI CO data. As they take into account IASI’s sensitivity to the boundary layer using averaging kernels this is somehow surprising but points to a possible overcorrection in the papers of Yurganov et al. (2011) and Fokeeva et al. (2011). Differences and potential sources of errors are discussed in detail in Krol et al. (2012). The wide scatter in the data is evidence of the difficulty in estimating total emitted masses from satellite data. For CO our approach gives a rather large but realistic range of

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values, and gives confidence to our methodology. Likewise, for  $\text{NH}_3$  and  $\text{HCOOH}$ , we have taken the range of possible effective lifetimes quite large, as to obtain a range of reasonable values for the total emissions.

We can calculate effective emission ratios using the median estimates of the total emissions. Assuming a total CO emission of 26 Tg we obtain for  $\text{NH}_3$  an emission ratio of 0.082. This is naturally larger than the observed enhancement ratios 0.010–0.055. For  $\text{HCOOH}$  an average emission ratio of 0.047 is found, also higher than the observed enhancement ratios (0.020–0.030 for the most active period).

## 5 Discussion

The sensitivity of IASI to lower tropospheric concentrations of CO,  $\text{NH}_3$  and  $\text{HCOOH}$  depends heavily on the surface-atmospheric temperature difference, the so-called thermal contrast (Clarisse et al., 2010; Razavi et al., 2011). Daytime observations typically have a larger thermal contrast than night time observations and are the preferred overpass to study these species. FORLI CO retrievals have on average a retrieval error below 10–15 % (Hurtmans et al., 2012). However, the total error can be much larger such as in the case of strong fires (George et al., 2009), especially when there is limited sensitivity to the boundary layer. As we have mentioned in Sect. 4, Yurganov et al. (2011) and Fokeeva et al. (2011) found large underestimates in CO total column retrievals as compared to ground-based measurements close to Moscow. However, as our total mass estimate is in excellent agreement with a more sophisticated inverse modelling approach of Krol et al. (2012), we do not expect a consistent low bias of the retrieved CO columns.

Retrieval errors for ammonia are larger than for CO, because of smaller signal to noise ratios and larger dependence on thermal contrast. Here,  $\text{NH}_3$  was conditionally retrieved from IASI spectra, on the basis of a firm spectral signature (which depends on both the total column of  $\text{NH}_3$  and the thermal contrast). While we obtain less measurements in this way, the measurements which are retained have limited dependence on

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the a priori column (low smoothing error). An estimate of the error for  $\text{NH}_3$  in the case of extreme conditions as those reported here is outside the scope of this paper, as collocated in situ measurement are the only way to assess it. For agricultural pollution in the boundary layer a retrieval error of 30–50 % for the total column has been estimated (Clarisse et al., 2010). Note that the selection criteria might lead to an underestimate of the total burden and mass of  $\text{NH}_3$  (see Sect. 4). Finally, the retrieval of HCOOH relies on conversion from BTDs and is also dependent on the thermal contrast. As in Razavi et al. (2011) we have restricted retrieval to observations with a thermal contrast above 5 degrees. Razavi et al. (2011) estimate average uncertainties to be about 60 % for such conditions. The algorithm does however not provide errors on individual observations, but simulate to  $\text{NH}_3$  an accurate error could only be obtained via validation. Note that uncertainties related to limited sensitivity of the boundary layer partially cancel out in the calculation of enhancement ratios.

## 6 Conclusions

The fires which occurred in Russia during the summer 2010 emitted important quantities of trace gases and aerosols for almost a month. In this work, we focused on three trace gases namely CO,  $\text{NH}_3$  and HCOOH. We presented total columns (maximum and mean) from IASI observations for 2010. We have calculated enhancement ratios of HCOOH and  $\text{NH}_3$  relative to CO for each day and reported their time evolution during the fire period. For  $\text{NH}_3$ , the enhancement ratios were shown to be in good agreement with tabulated values of emission ratios, while for HCOOH our reported values are one order of magnitude larger. This is suspected to be due to the rapid secondary formation of HCOOH in the plumes. For both species the maximum enhancement ratios were found around mid-August where the total columns are largest. In addition, we have calculated total masses and the burdens from fires for each day, as well as fluxes using assumptions on the species lifetimes. For CO, the results obtained are comparable to previous studies. For  $\text{NH}_3$  and HCOOH, these are the first values reported. They

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are of the order of  $0.1 \text{ Tg day}^{-1}$  depending on the choice of reference lifetime. When integrated over the entire fire period (July–August of 2010), the emission fluxes translate to a total emitted mass of 19–33 Tg for CO, 0.7–2.6 Tg for  $\text{NH}_3$  and 0.9–3.9 Tg for HCOOH. For all these species, there is a significant contribution to the yearly global emission. For instance this is for  $\text{NH}_3$  and HCOOH comparable to what is emitted in a year by all extratropical forest fires.

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**Table 1.** Comparative analysis of total emissions of CO, NH<sub>3</sub> and HCOOH from current and previous studies of the 2010 Russian wildfires.

	CO	NH <sub>3</sub>	HCOOH	Comment
	<b>Max. daily emission (Tg day<sup>-1</sup>)</b>			
Fokeeva et al. (2011)	1.2–1.4	–	–	AIRS-MOPITT, standard
Fokeeva et al. (2011)	2.3–2.4	–	–	AIRS-MOPITT, adjusted
Yurganov et al. (2011)	1.5	–	–	IASI-OE, standard
Yurganov et al. (2011)	2.1–2.2	–	–	Estimate from 3 different sounders (IASI, AIRS and MOPITT), adjusted
<b>This study</b>	<b>1.41–1.87</b>	<b>0.04–0.15</b>	<b>0.07–0.19</b>	IASI-OE standard
	<b>Total emitted (Tg)</b>			
Fokeeva et al. (2011)	19–26	–	–	AIRS-MOPITT, standard
Fokeeva et al. (2011)	36–42	–	–	AIRS-MOPITT, adjusted
Fokeeva et al. (2011)	12–14	–	–	MODIS, inventory method
Huijnen et al. (2012)	12.2	–	–	MODIS, inventory method (GFASv1.0)
Konovalov et al. (2011)	9.7	–	–	MODIS, inverse modeling
Krol et al. (2012)	20–25	–	–	IASI, inverse modeling
Yurganov et al. (2011)	26.2	–	–	IASI-OE, standard
Yurganov et al. (2011)	34–40	–	–	Estimate from 3 different sounders (IASI, AIRS and MOPITT), adjusted
<b>This study</b>	<b>19–33</b>	<b>0.7–2.6</b>	<b>0.9–3.9</b>	IASI-OE standard

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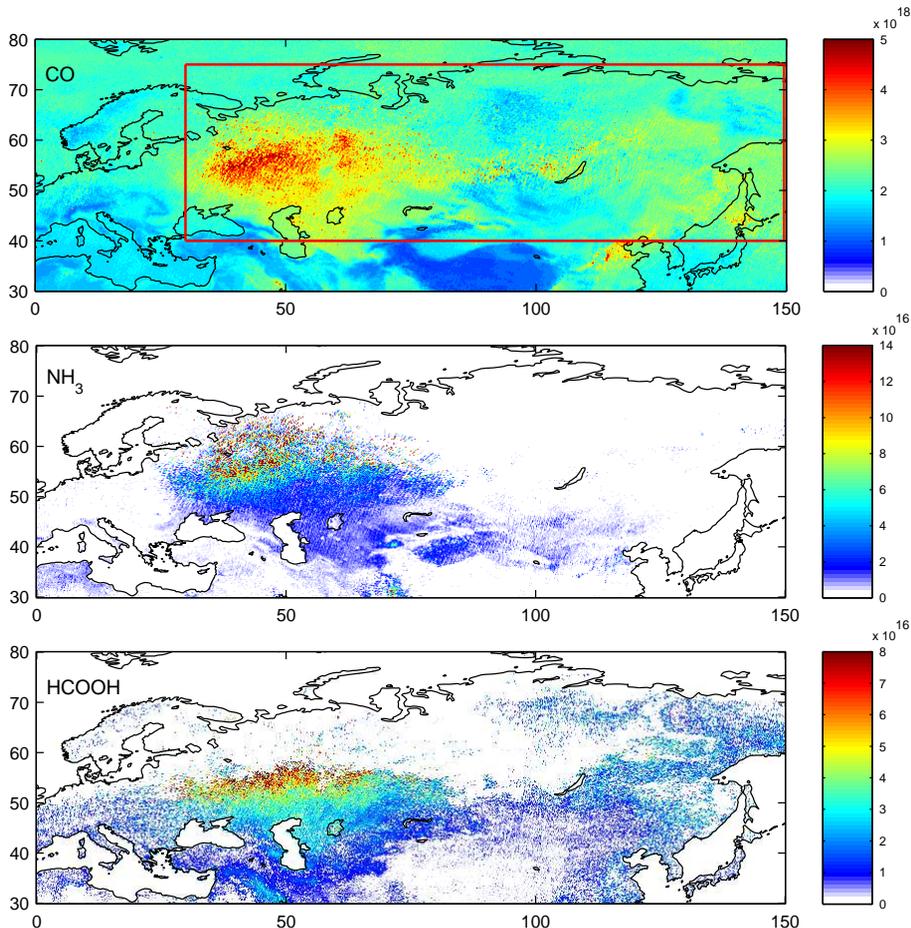
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**Fig. 1.** IASI morning observations of the 2010 Russian wildfires plumes. The 3 panels show the mean CO, NH<sub>3</sub> and HCOOH total columns in molec cm<sup>-2</sup>, in grid cells of 0.5° × 0.5° for one month, from 27 July to 27 August 2010.

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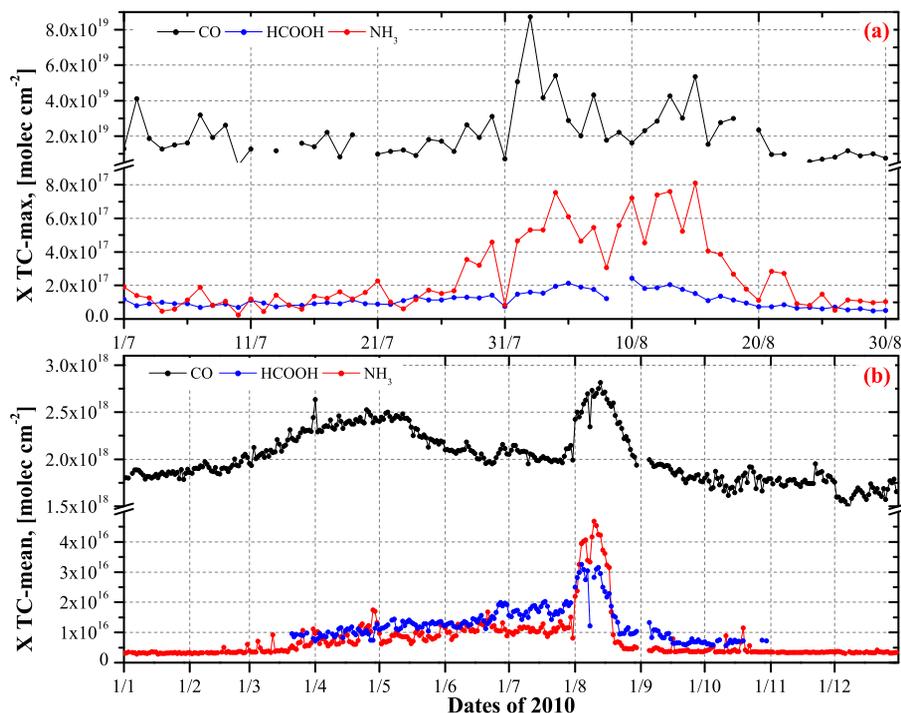
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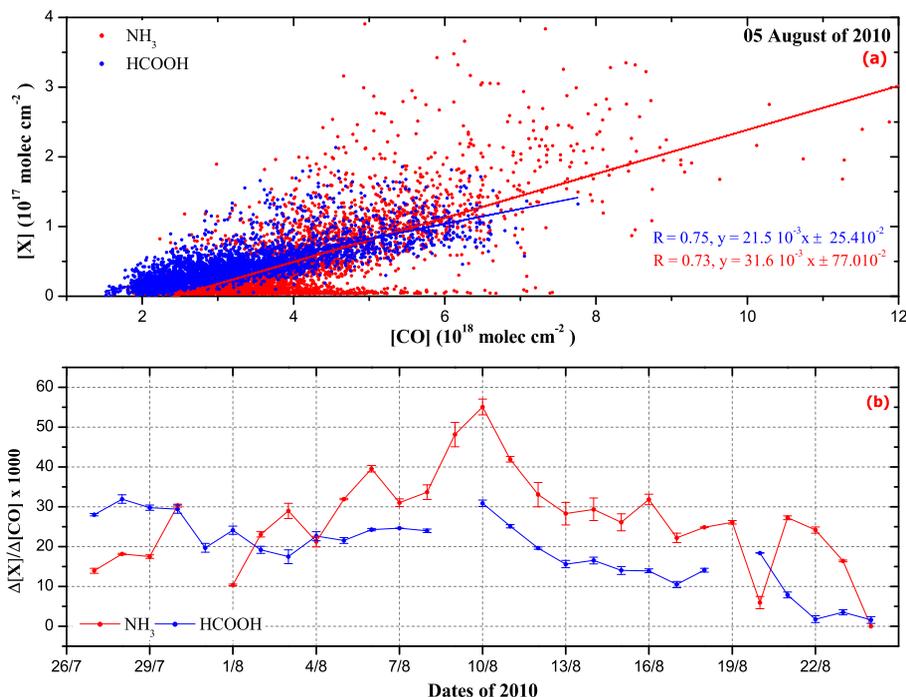


**Fig. 2.** (a) Temporal evolution of CO (black), NH<sub>3</sub> (red) and HCOOH (blue) total column maxima between 1 July and 1 September 2010. Only retrievals with spectral residual root-mean-square less than  $3 \cdot 10^{-8}$  and  $3 \cdot 10^{-6}$  ( $W\text{cm}^{-2}\text{srcm}^{-1}$ ) for CO and NH<sub>3</sub> respectively were considered. (b) Mean total column over the area 40°–75° N, 30°–150° E for CO, NH<sub>3</sub> and HCOOH, for the year 2010.

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**Fig. 3.** (a) Scatter plot of NH<sub>3</sub> and HCOOH total columns vs. CO total columns for 5 August 2010. Linear regressions are plotted for both species as red and blue lines, respectively. The values for the resulting slopes and their errors are given in the figure, along with the correlation coefficient  $R$ . (b) Temporal evolution of the enhancement ratios of NH<sub>3</sub> and HCOOH relative to CO, calculated as the slopes of the linear regressions from 27 July to 24 August, in the region 40°–75° N and 30°–90° E. The errors bars are the corresponding slopes errors.

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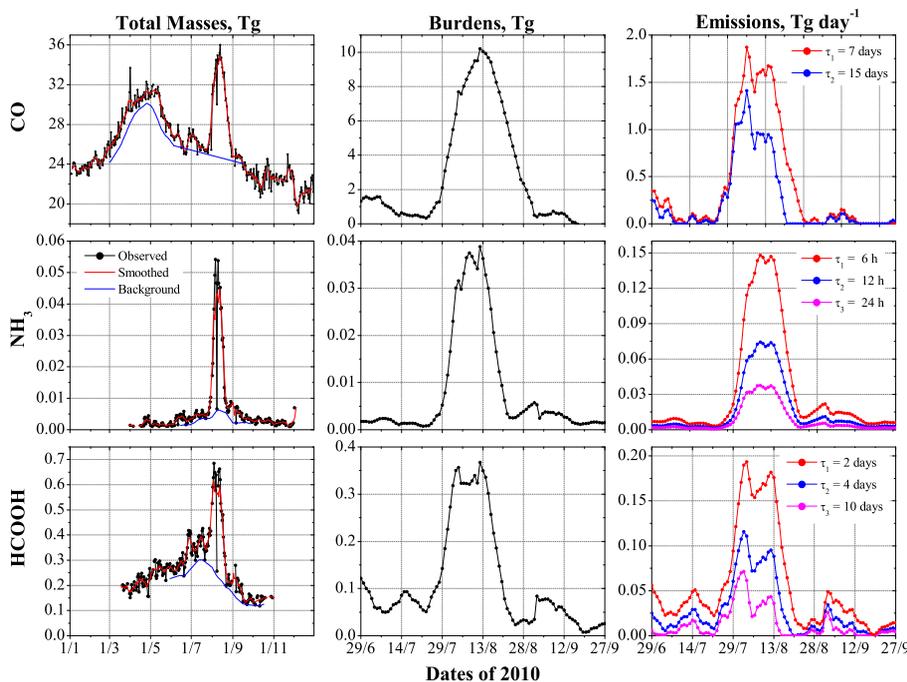
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**Fig. 4.** Left: total CO, NH<sub>3</sub> and HCOOH masses in Tg (black dots and line) over the area 40°–75° N, 30°–150° E for CO and HCOOH and 40°–75° N, 30°–90° E for NH<sub>3</sub>. The red lines are the smoothed total masses in 10-days interval; the blue lines are total masses representative of background concentrations. Middle: extra CO, NH<sub>3</sub> and HCOOH smoothed burdens in Tg from fires obtained by subtracting background values from the total masses. Right: fluxes in Tg day<sup>-1</sup> respectively of NH<sub>3</sub> and HCOOH calculated by Eq. (3).

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